

# Colour Documentation

*Release 0.3.11*

**Colour Developers**

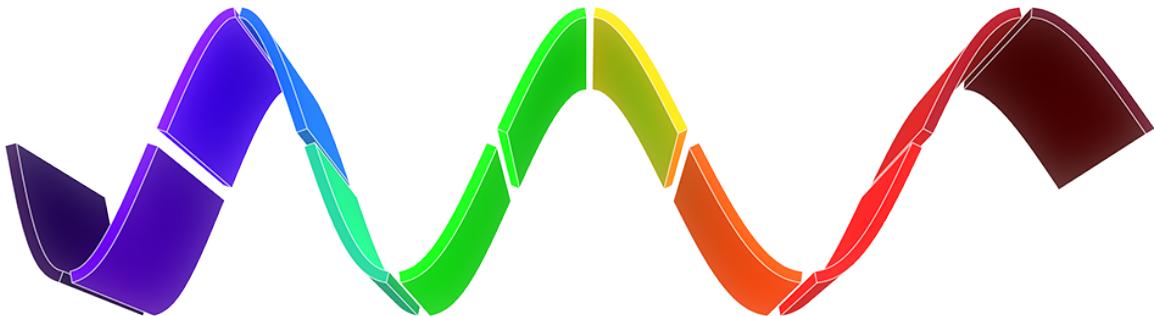
Feb 18, 2018



# CONTENTS

<b>1 Features</b>	<b>3</b>
<b>2 Installation</b>	<b>5</b>
<b>3 Usage</b>	<b>7</b>
3.1 Colour Manual . . . . .	7
3.1.1 Tutorial . . . . .	7
3.1.1.1 Overview . . . . .	8
3.1.1.2 From Spectral Power Distribution . . . . .	14
3.1.1.3 Convert to Tristimulus Values . . . . .	24
3.1.1.4 From <i>CIE XYZ</i> Colourspace . . . . .	25
3.1.1.5 Convert to Screen Colours . . . . .	25
3.1.1.6 Generate Colour Rendition Charts . . . . .	26
3.1.1.7 Convert to Chromaticity Coordinates . . . . .	29
3.1.1.8 And More... . . . . .	31
3.1.2 Reference . . . . .	31
3.1.2.1 Colour . . . . .	31
3.1.2.2 Indices and tables . . . . .	452
3.1.3 Bibliography . . . . .	452
3.1.3.1 Indirect References . . . . .	452
3.2 Examples . . . . .	453
<b>4 Contributing</b>	<b>461</b>
<b>5 Changes</b>	<b>463</b>
<b>6 Bibliography</b>	<b>465</b>
<b>7 See Also</b>	<b>467</b>
<b>8 About</b>	<b>469</b>
<b>Bibliography</b>	<b>471</b>
<b>Index</b>	<b>485</b>





Colour is a [Python](#) colour science package implementing a comprehensive number of colour theory transformations and algorithms.

It is open source and freely available under the [New BSD License](#) terms.



---

**CHAPTER  
ONE**

---

**FEATURES**

Colour features a rich dataset and collection of objects, please see the [features](#) page for more information.



## INSTALLATION

Anaconda from *Continuum Analytics* is the Python distribution we use to develop **Colour**: it ships all the scientific dependencies we require and is easily deployed cross-platform:

```
$ conda create -y -n python-colour
$ source activate python-colour
$ conda install -y -c conda-forge colour-science
```

**Colour** can be easily installed from the [Python Package Index](#) by issuing this command in a shell:

```
$ pip install colour-science
```

The detailed installation procedure is described in the [Installation Guide](#).



## USAGE

The two main references for `Colour` usage are the [Colour Manual](#) and the [Jupyter Notebooks](#) with detailed historical and theoretical context and images.

### 3.1 Colour Manual

#### 3.1.1 Tutorial

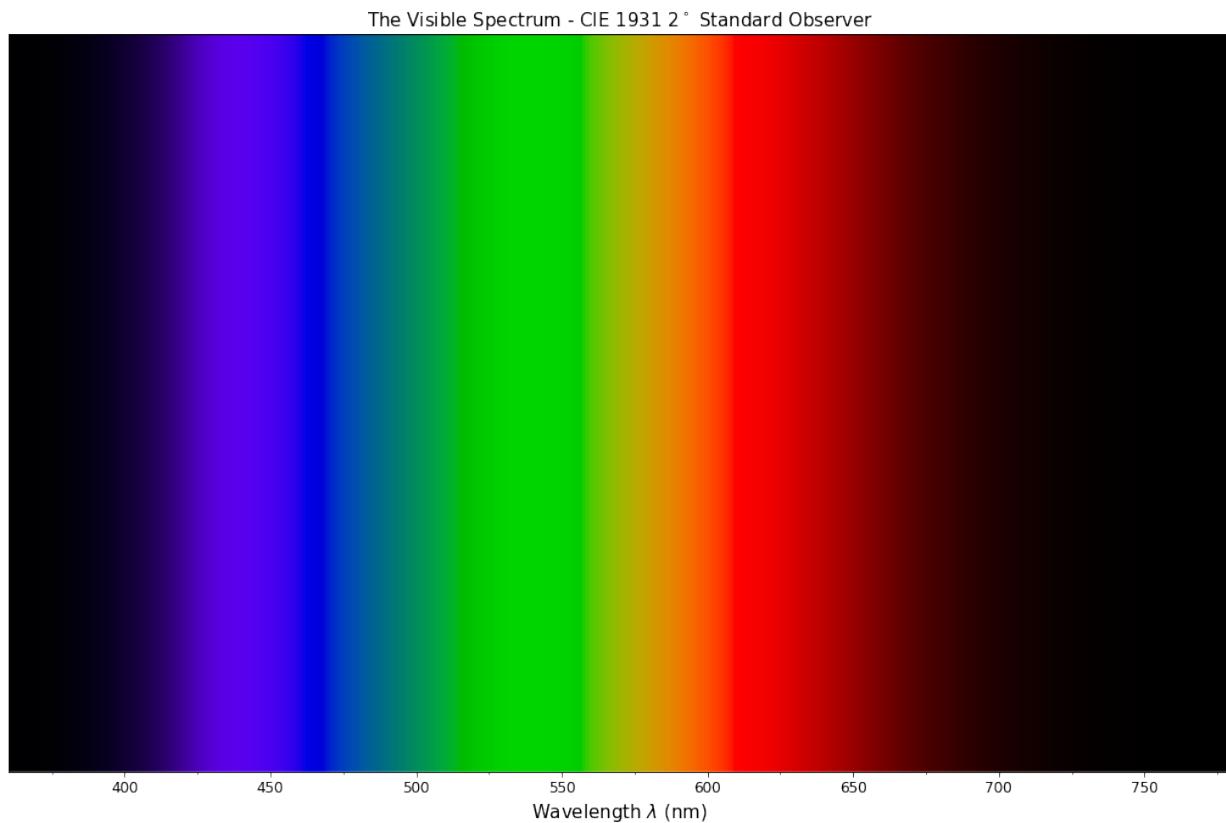
`Colour` spreads over various domains of Colour Science from colour models to optical phenomena, this tutorial will not give you a complete overview of the API but will still be a good introduction.

---

**Note:** A directory full of examples is available at this path in your `Colour` installation: `colour/examples`. You can also explore it directly on Github: <https://github.com/colour-science/colour/tree/master/colour/examples>

---

```
from colour.plotting import *
colour_plotting_defaults()
visible_spectrum_plot()
```



### 3.1.1.1 Overview

Colour is organised around various sub-packages:

- *adaptation*: Chromatic adaptation models and transformations.
- *algebra*: Algebra utilities.
- *appearance*: Colour appearance models.
- *biochemistry*: Biochemistry computations.
- *continuous*: Base objects for continuous data representation.
- *characterisation*: Colour fitting and camera characterisation.
- *colorimetry*: Core objects for colour computations.
- *constants*: CIE and CODATA constants.
- *corresponding*: Corresponding colour chromaticities computations.
- *difference*: Colour difference computations.
- *examples*: Examples for the sub-packages.
- *io*: Input / output objects for reading and writing data.
- *models*: Colour models.
- *notation*: Colour notation systems.
- *phenomena*: Computation of various optical phenomena.

- *plotting*: Diagrams, figures, etc...
- *quality*: Colour quality computation.
- *recovery*: Reflectance recovery.
- *temperature*: Colour temperature and correlated colour temperature computation.
- *utilities*: Various utilities and data structures.
- *volume*: Colourspace volumes computation and optimal colour stimuli.

Most of the public API is available from the root colour namespace:

```
import colour

print(colour.__all__[:5] + ['...'])
```

```
['handle_numpy_errors', 'ignore_numpy_errors', 'raise_numpy_errors', 'print_numpy_errors', 'warn_numpy_
←errors', '...']
```

The various sub-packages also expose their public API:

```
from pprint import pprint

import colour.plotting

for sub_package in ('adaptation', 'algebra', 'appearance', 'biochemistry',
                    'characterisation', 'colorimetry', 'constants',
                    'continuous', 'corresponding', 'difference', 'io',
                    'models', 'notation', 'phenomena', 'plotting', 'quality',
                    'recovery', 'temperature', 'utilities', 'volume'):
    print(sub_package.title())
    pprint(getattr(colour, sub_package).__all__[:5] + ['...'])
    print('\n')
```

```
Adaptation
['CHROMATIC_ADAPTATION_TRANSFORMS',
 'XYZ_SCALING_CAT',
 'VON_KRIES_CAT',
 'BRADFORD_CAT',
 'SHARP_CAT',
 '...']
```

```
Algebra
['cartesian_to_spherical',
 'spherical_to_cartesian',
 'cartesian_to_polar',
 'polar_to_cartesian',
 'cartesian_to_cylindrical',
 '...']
```

```
Appearance
['Hunt_InductionFactors',
 'HUNT_VIEWING_CONDITIONS',
 'Hunt_Specification',
 'XYZ_to_Hunt',
 'ATD95_Specification',
```

```
'...']  
  
Biochemistry  
['reaction_rate_MichealisMenten',  
 'substrate_concentration_MichealisMenten',  
 '...']  
  
Characterisation  
['RGB_SpectralSensitivities',  
 'RGB_DisplayPrimaries',  
 'CAMERAS_RGB_SPECTRAL_SENSITIVITIES',  
 'COLOURCHECKERS',  
 'COLOURCHECKER_INDEXES_TO_NAMES_MAPPING',  
 '...']  
  
Colorimetry  
['SpectralShape',  
 'SpectralPowerDistribution',  
 'MultiSpectralPowerDistribution',  
 'DEFAULT_SPECTRAL_SHAPE',  
 'constant_spd',  
 '...']  
  
Continuous  
['AbstractContinuousFunction', 'Signal', 'MultiSignal', '...']  
  
Constants  
['CIE_E', 'CIE_K', 'K_M', 'KP_M', 'AVOGADRO_CONSTANT', '...']  
  
Corresponding  
['BRENEMAN_EXPERIMENTS',  
 'BRENEMAN_EXPERIMENTS_PRIMARIES_CHROMATICITIES',  
 'corresponding_chromaticities_prediction_CIE1994',  
 'corresponding_chromaticities_prediction_CMCCAT2000',  
 'corresponding_chromaticities_prediction_Fairchild1990',  
 '...']  
  
Difference  
['DELTA_E_METHODS',  
 'delta_E',  
 'delta_E_CIE1976',  
 'delta_E_CIE1994',  
 'delta_E_CIE2000',  
 '...']  
  
Io  
['IES_TM2714_Spd',  
 'read_image',  
 'write_image',  
 'read_spectral_data_from_csv_file',
```

```
'read_spds_from_csv_file',
'...']

Models
['XYZ_to_xyY', 'xyY_to_XYZ', 'xy_to_xyY', 'xyY_to_xy', 'xy_to_XYZ', '...']

Notation
['MUNSELL_COLOURS_ALL',
'MUNSELL_COLOURS_1929',
'MUNSELL_COLOURS_REAL',
'MUNSELL_COLOURS',
'munsell_value',
'...']

Phenomena
['scattering_cross_section',
'rayleigh_optical_depth',
'rayleigh_scattering',
'rayleigh_scattering_spd',
'...']

Plotting
['ASTM_G_173_ETR',
'PLOTTING_RESOURCES_DIRECTORY',
'DEFAULT_FIGURE_ASPECT_RATIO',
'DEFAULT_FIGURE_WIDTH',
'DEFAULT_FIGURE_HEIGHT',
'...']

Quality
['TCS_SPDS',
'VS_SPDS',
'CRI_Specification',
'colour_rendering_index',
'CQS_Specification',
'...']

Recovery
['SMITS_1999_SPDS',
'XYZ_to_spectral_Meng2015',
'RGB_to_spectral_Smits1999',
'REFLECTANCE_RECOVERY_METHODS',
'XYZ_to_spectral',
'...']

Temperature
['CCT_TO_UV_METHODS',
'UV_TO_CCT_METHODS',
'CCT_to_uv',
'CCT_to_uv_Ohno2013',
'CCT_to_uv_Robertson1968',
```

```
'...']  
  
Utilities  
['handle_numpy_errors',  
'ignore_numpy_errors',  
'raise_numpy_errors',  
'print_numpy_errors',  
'warn_numpy_errors',  
'...']  
  
Volume  
['ILLUMINANTS_OPTIMAL_COLOUR_STIMULI',  
'is_within_macadam_limits',  
'is_within_mesh_volume',  
'is_within_pointer_gamut',  
'is_within_visible_spectrum',  
'...']
```

The code is documented and almost every docstrings have usage examples:

```
print(colour.temperature.CCT_to_uv_Ohno2013.__doc__)
```

```
Returns the *CIE UCS* colourspace *uv* chromaticity coordinates from given  
correlated colour temperature :math:`T_{cp}` , :math:`\Delta_{uv}` and  
colour matching functions using *Ohno (2013)* method.
```

#### Parameters

```
-----  
CCT : numeric  
    Correlated colour temperature :math:`T_{cp}` .  
D_uv : numeric, optional  
    :math:`\Delta_{uv}` .  
cmfs : XYZ_ColourMatchingFunctions, optional  
    Standard observer colour matching functions.
```

#### Returns

```
-----  
ndarray  
    *CIE UCS* colourspace *uv* chromaticity coordinates.
```

#### References

```
... [4] Ohno, Y. (2014). Practical Use and Calculation of CCT and Duv.  
    LEUKOS, 10(1), 4755. doi:10.1080/15502724.2014.839020
```

#### Examples

```
-----  
>>> from colour import STANDARD_OBSERVERS_CMFS  
>>> cmfs = STANDARD_OBSERVERS_CMFS['CIE 1931 2 Degree Standard Observer']  
>>> CCT = 6507.4342201047066  
>>> D_uv = 0.003223690901513  
>>> CCT_to_uv_Ohno2013(CCT, D_uv, cmfs) # doctest: +ELLIPSIS  
array([ 0.197799...,  0.3122004...])
```

At the core of Colour is the colour.colorimetry sub-package, it defines the objects needed for spectral related computations and many others:

```
import colour.colorimetry as colorimetry

pprint(colorimetry.__all__)
```

```
['SpectralShape',
 'SpectralPowerDistribution',
 'MultiSpectralPowerDistribution',
 'DEFAULT_SPECTRAL_SHAPE',
 'constant_spd',
 'zeros_spd',
 'ones_spd',
 'blackbody_spd',
 'blackbody_spectral_radiance',
 'planck_law',
 'LMS_ConeFundamentals',
 'RGB_ColourMatchingFunctions',
 'XYZ_ColourMatchingFunctions',
 'CMFS',
 'LMS_CMFS',
 'RGB_CMFS',
 'STANDARD_OBSERVERS_CMFS',
 'ILLUMINANTS',
 'D_ILLUMINANTS_S_SPDS',
 'HUNTERLAB_ILLUMINANTS',
 'ILLUMINANTS_RELATIVE_SPDS',
 'LIGHT_SOURCES',
 'LIGHT_SOURCES_RELATIVE_SPDS',
 'LEFS',
 'PHOTOPIC_LEFS',
 'SCOTOPIC_LEFS',
 'BANDPASS_CORRECTION_METHODS',
 'bandpass_correction',
 'bandpass_correction_Stearns1988',
 'D_illuminant_relative_spd',
 'CIE_standard_illuminant_A_function',
 'mesopic_luminous_efficiency_function',
 'mesopic_weighting_function',
 'LIGHTNESS_METHODS',
 'lightness',
 'lightness_Glasser1958',
 'lightness_Wyszecki1963',
 'lightness_CIE1976',
 'lightness_Fairchild2010',
 'lightness_Fairchild2011',
 'LUMINANCE_METHODS',
 'luminance',
 'luminance_Newhall1943',
 'luminance_ASTMD153508',
 'luminance_CIE1976',
 'luminance_Fairchild2010',
 'luminance_Fairchild2011',
 'dominant_wavelength',
 'complementary_wavelength',
 'excitation_purity',
 'colorimetric_purity',
 'luminous_flux',
 'luminous_efficiency',
 'luminous_efficacy',
```

```
'RGB_10_degree_cmfs_to_LMS_10_degree_cmfs',
'RGB_2_degree_cmfs_to_XYZ_2_degree_cmfs',
'RGB_10_degree_cmfs_to_XYZ_10_degree_cmfs',
'LMS_2_degree_cmfs_to_XYZ_2_degree_cmfs',
'LMS_10_degree_cmfs_to_XYZ_10_degree_cmfs',
'SPECTRAL_TO_XYZ_METHODS',
'spectral_to_XYZ',
'ASTME30815_PRACTISE_SHAPE',
'lagrange_coefficients_ASTME202211',
'tristimulus_weighting_factors_ASTME202211',
'adjust_tristimulus_weighting_factors_ASTME30815',
'spectral_to_XYZ_integration',
'spectral_to_XYZ_tristimulus_weighting_factors_ASTME30815',
'spectral_to_XYZ_ASTME30815',
'wavelength_to_XYZ',
'WHITENESS_METHODS',
'whiteness',
'whiteness_Berger1959',
'whiteness_Taube1960',
'whiteness_Stensby1968',
'whiteness_ASTME313',
'whiteness_Ganz1979',
'whiteness_CIE2004',
'YELLOWNESS_METHODS',
'yellowness',
'yellowness_ASTMD1925',
'yellowness_ASTME313']
```

Colour computations leverage a comprehensive dataset available in pretty much each sub-packages, for example `colour.colorimetry.dataset` defines the following data:

```
import colour.colorimetry.dataset as dataset

pprint(dataset.__all__)
```

```
['CMFS',
'LMS_CMFS',
'RGB_CMFS',
'STANDARD_OBSERVERS_CMFS',
'ILLUMINANTS',
'D_ILLUMINANTS_S_SPDS',
'HUNTERLAB_ILLUMINANTS',
'ILLUMINANTS_RELATIVE_SPDS',
'LIGHT_SOURCES',
'LIGHT_SOURCES_RELATIVE_SPDS',
'LEFS',
'PHOTOPIC_LEFS',
'SCOTOPIC_LEFS']
```

### 3.1.1.2 From Spectral Power Distribution

Whether it be a sample spectral power distribution, colour matching functions or illuminants, spectral data is manipulated using an object built with the `colour.SpectralPowerDistribution` class or based on it:

```
# Defining a sample spectral power distribution data.
sample_spd_data = {
```

```
380: 0.048,
385: 0.051,
390: 0.055,
395: 0.060,
400: 0.065,
405: 0.068,
410: 0.068,
415: 0.067,
420: 0.064,
425: 0.062,
430: 0.059,
435: 0.057,
440: 0.055,
445: 0.054,
450: 0.053,
455: 0.053,
460: 0.052,
465: 0.052,
470: 0.052,
475: 0.053,
480: 0.054,
485: 0.055,
490: 0.057,
495: 0.059,
500: 0.061,
505: 0.062,
510: 0.065,
515: 0.067,
520: 0.070,
525: 0.072,
530: 0.074,
535: 0.075,
540: 0.076,
545: 0.078,
550: 0.079,
555: 0.082,
560: 0.087,
565: 0.092,
570: 0.100,
575: 0.107,
580: 0.115,
585: 0.122,
590: 0.129,
595: 0.134,
600: 0.138,
605: 0.142,
610: 0.146,
615: 0.150,
620: 0.154,
625: 0.158,
630: 0.163,
635: 0.167,
640: 0.173,
645: 0.180,
650: 0.188,
655: 0.196,
660: 0.204,
665: 0.213,
```

```
670: 0.222,
675: 0.231,
680: 0.242,
685: 0.251,
690: 0.261,
695: 0.271,
700: 0.282,
705: 0.294,
710: 0.305,
715: 0.318,
720: 0.334,
725: 0.354,
730: 0.372,
735: 0.392,
740: 0.409,
745: 0.420,
750: 0.436,
755: 0.450,
760: 0.462,
765: 0.465,
770: 0.448,
775: 0.432,
780: 0.421}

spd = colour.SpectralPowerDistribution(sample_spd_data, name='Sample')
print(repr(spd))
```

```
SpectralPowerDistribution([[ 3.8000000e+02,   4.8000000e-02],
                           [ 3.8500000e+02,   5.1000000e-02],
                           [ 3.9000000e+02,   5.5000000e-02],
                           [ 3.9500000e+02,   6.0000000e-02],
                           [ 4.0000000e+02,   6.5000000e-02],
                           [ 4.0500000e+02,   6.8000000e-02],
                           [ 4.1000000e+02,   6.8000000e-02],
                           [ 4.1500000e+02,   6.7000000e-02],
                           [ 4.2000000e+02,   6.4000000e-02],
                           [ 4.2500000e+02,   6.2000000e-02],
                           [ 4.3000000e+02,   5.9000000e-02],
                           [ 4.3500000e+02,   5.7000000e-02],
                           [ 4.4000000e+02,   5.5000000e-02],
                           [ 4.4500000e+02,   5.4000000e-02],
                           [ 4.5000000e+02,   5.3000000e-02],
                           [ 4.5500000e+02,   5.3000000e-02],
                           [ 4.6000000e+02,   5.2000000e-02],
                           [ 4.6500000e+02,   5.2000000e-02],
                           [ 4.7000000e+02,   5.2000000e-02],
                           [ 4.7500000e+02,   5.3000000e-02],
                           [ 4.8000000e+02,   5.4000000e-02],
                           [ 4.8500000e+02,   5.5000000e-02],
                           [ 4.9000000e+02,   5.7000000e-02],
                           [ 4.9500000e+02,   5.9000000e-02],
                           [ 5.0000000e+02,   6.1000000e-02],
                           [ 5.0500000e+02,   6.2000000e-02],
                           [ 5.1000000e+02,   6.5000000e-02],
                           [ 5.1500000e+02,   6.7000000e-02],
                           [ 5.2000000e+02,   7.0000000e-02],
                           [ 5.2500000e+02,   7.2000000e-02],
                           [ 5.3000000e+02,   7.4000000e-02],
```

```

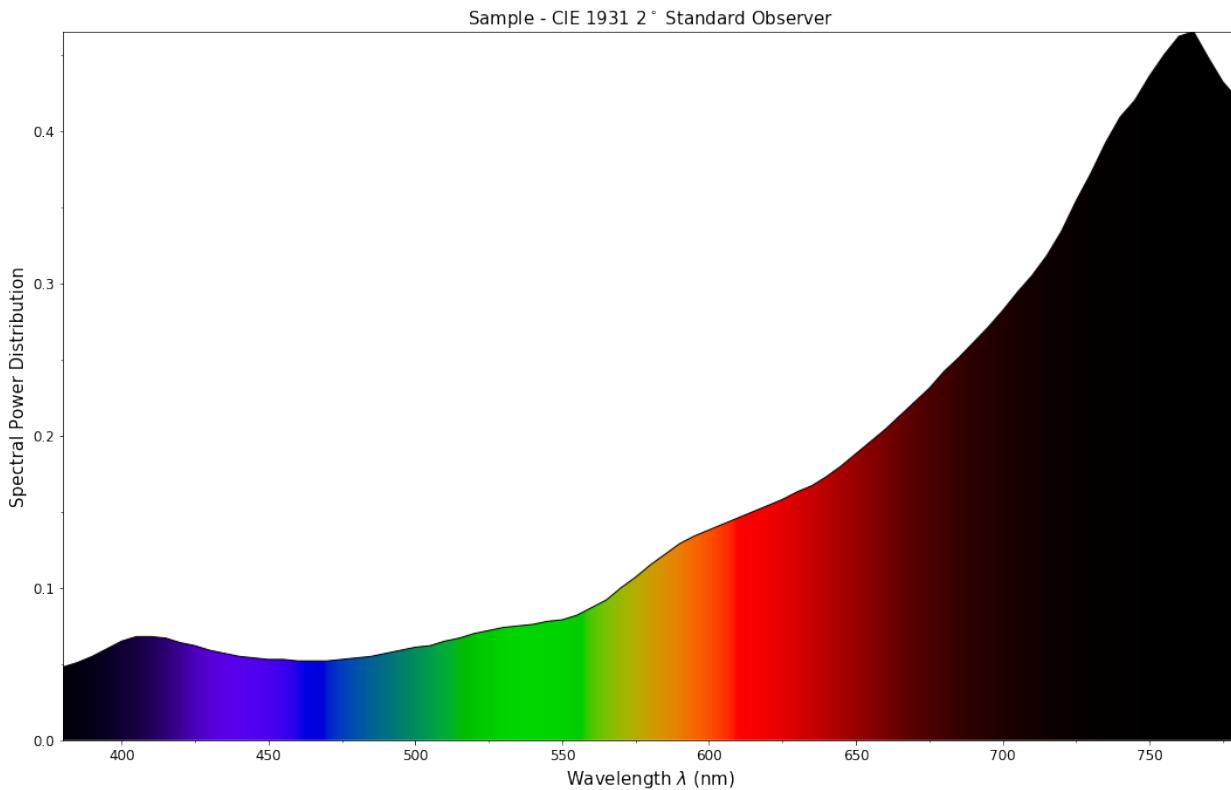
[ 5.35000000e+02, 7.5000000e-02],
[ 5.40000000e+02, 7.6000000e-02],
[ 5.45000000e+02, 7.8000000e-02],
[ 5.50000000e+02, 7.9000000e-02],
[ 5.55000000e+02, 8.2000000e-02],
[ 5.60000000e+02, 8.7000000e-02],
[ 5.65000000e+02, 9.2000000e-02],
[ 5.70000000e+02, 1.0000000e-01],
[ 5.75000000e+02, 1.0700000e-01],
[ 5.80000000e+02, 1.1500000e-01],
[ 5.85000000e+02, 1.2200000e-01],
[ 5.90000000e+02, 1.2900000e-01],
[ 5.95000000e+02, 1.3400000e-01],
[ 6.00000000e+02, 1.3800000e-01],
[ 6.05000000e+02, 1.4200000e-01],
[ 6.10000000e+02, 1.4600000e-01],
[ 6.15000000e+02, 1.5000000e-01],
[ 6.20000000e+02, 1.5400000e-01],
[ 6.25000000e+02, 1.5800000e-01],
[ 6.30000000e+02, 1.6300000e-01],
[ 6.35000000e+02, 1.6700000e-01],
[ 6.40000000e+02, 1.7300000e-01],
[ 6.45000000e+02, 1.8000000e-01],
[ 6.50000000e+02, 1.8800000e-01],
[ 6.55000000e+02, 1.9600000e-01],
[ 6.60000000e+02, 2.0400000e-01],
[ 6.65000000e+02, 2.1300000e-01],
[ 6.70000000e+02, 2.2200000e-01],
[ 6.75000000e+02, 2.3100000e-01],
[ 6.80000000e+02, 2.4200000e-01],
[ 6.85000000e+02, 2.5100000e-01],
[ 6.90000000e+02, 2.6100000e-01],
[ 6.95000000e+02, 2.7100000e-01],
[ 7.00000000e+02, 2.8200000e-01],
[ 7.05000000e+02, 2.9400000e-01],
[ 7.10000000e+02, 3.0500000e-01],
[ 7.15000000e+02, 3.1800000e-01],
[ 7.20000000e+02, 3.3400000e-01],
[ 7.25000000e+02, 3.5400000e-01],
[ 7.30000000e+02, 3.7200000e-01],
[ 7.35000000e+02, 3.9200000e-01],
[ 7.40000000e+02, 4.0900000e-01],
[ 7.45000000e+02, 4.2000000e-01],
[ 7.50000000e+02, 4.3600000e-01],
[ 7.55000000e+02, 4.5000000e-01],
[ 7.60000000e+02, 4.6200000e-01],
[ 7.65000000e+02, 4.6500000e-01],
[ 7.70000000e+02, 4.4800000e-01],
[ 7.75000000e+02, 4.3200000e-01],
[ 7.80000000e+02, 4.2100000e-01]],

interpolator=SpragueInterpolator,
interpolator_args={},
extrapolator=Extrapolator,
extrapolator_args={u'right': None, u'method': u'Constant', u'left': None})

```

The sample spectral power distribution can be easily plotted against the visible spectrum:

```
# Plotting the sample spectral power distribution.  
single_spd_plot(spd)
```



With the sample spectral power distribution defined, we can retrieve its shape:

```
# Displaying the sample spectral power distribution shape.  
print(spd.shape)
```

```
(380.0, 780.0, 5.0)
```

The shape returned is an instance of `colour.SpectralShape` class:

```
repr(spd.shape)
```

```
'SpectralShape(380.0, 780.0, 5.0)'
```

`colour.SpectralShape` is used throughout Colour to define spectral dimensions and is instantiated as follows:

```
# Using *colour.SpectralShape* with iteration.  
shape = colour.SpectralShape(start=0, end=10, interval=1)  
for wavelength in shape:  
    print(wavelength)  
  
# *colour.SpectralShape.range* method is providing the complete range of values.  
shape = colour.SpectralShape(0, 10, 0.5)  
shape.range()
```

```
0.0
1.0
2.0
3.0
4.0
5.0
6.0
7.0
8.0
9.0
10.0
```

```
array([ 0. ,  0.5,  1. ,  1.5,  2. ,  2.5,  3. ,  3.5,  4. ,
       4.5,  5. ,  5.5,  6. ,  6.5,  7. ,  7.5,  8. ,  8.5,
       9. ,  9.5, 10. ])
```

Colour defines three convenient objects to create constant spectral power distributions:

- colour.constant\_spd
- colour.zeros\_spd
- colour.ones\_spd

```
# Defining a constant spectral power distribution.
constant_spd = colour.constant_spd(100)
print("Constant Spectral Power Distribution")
print(constant_spd.shape)
print(constant_spd[400])

# Defining a zeros filled spectral power distribution.
print('\n"Zeros Filled Spectral Power Distribution"')
zeros_spd = colour.zeros_spd()
print(zeros_spd.shape)
print(zeros_spd[400])

# Defining a ones filled spectral power distribution.
print('\n"Ones Filled Spectral Power Distribution"')
ones_spd = colour.ones_spd()
print(ones_spd.shape)
print(ones_spd[400])
```

```
"Constant Spectral Power Distribution"
(360.0, 780.0, 1.0)
100.0

"Zeros Filled Spectral Power Distribution"
(360.0, 780.0, 1.0)
0.0

"Ones Filled Spectral Power Distribution"
(360.0, 780.0, 1.0)
1.0
```

By default the shape used by colour.constant\_spd, colour.zeros\_spd and colour.ones\_spd is the one defined by colour.DEFAULT\_SPECTRAL\_SHAPE attribute using the *CIE 1931 2° Standard Observer* shape.

```
print(repr(colour.DEFAULT_SPECTRAL_SHAPE))
```

```
SpectralShape(360, 780, 1)
```

A custom shape can be passed to construct a constant spectral power distribution with user defined dimensions:

```
colour.ones_spd(colour.SpectralShape(400, 700, 5))[450]
```

```
1.0
```

The colour.SpectralPowerDistribution class supports the following arithmetical operations:

- *addition*
- *subtraction*
- *multiplication*
- *division*

```
spd1 = colour.ones_spd()
print('"Ones Filled Spectral Power Distribution"')
print(spd1[400])

print('\n"x2 Constant Multiplied"')
print((spd1 * 2)[400])

print('\n"+ Spectral Power Distribution"')
print((spd1 + colour.ones_spd())[400])
```

```
"Ones Filled Spectral Power Distribution"
```

```
1.0
```

```
"x2 Constant Multiplied"
```

```
2.0
```

```
"+ Spectral Power Distribution"
```

```
2.0
```

Often interpolation of the spectral power distribution is needed, this is achieved with the colour.SpectralPowerDistribution.interpolate method. Depending on the wavelengths uniformity, the default interpolation method will differ. Following CIE 167:2005 recommendation: The method developed by Sprague (1880) should be used for interpolating functions having a uniformly spaced independent variable and a Cubic Spline method for non-uniformly spaced independent variable [CIET13805a].

We can check the uniformity of the sample spectral power distribution:

```
# Checking the sample spectral power distribution uniformity.
print(spd.is_uniform())
```

```
True
```

Since the sample spectral power distribution is uniform the interpolation will default to the colour.SpragueInterpolator interpolator.

---

**Note:** Interpolation happens in place and may alter your original data, use the colour.SpectralPowerDistribution.copy method to produce a copy of your spectral power distribution before interpolation.

---

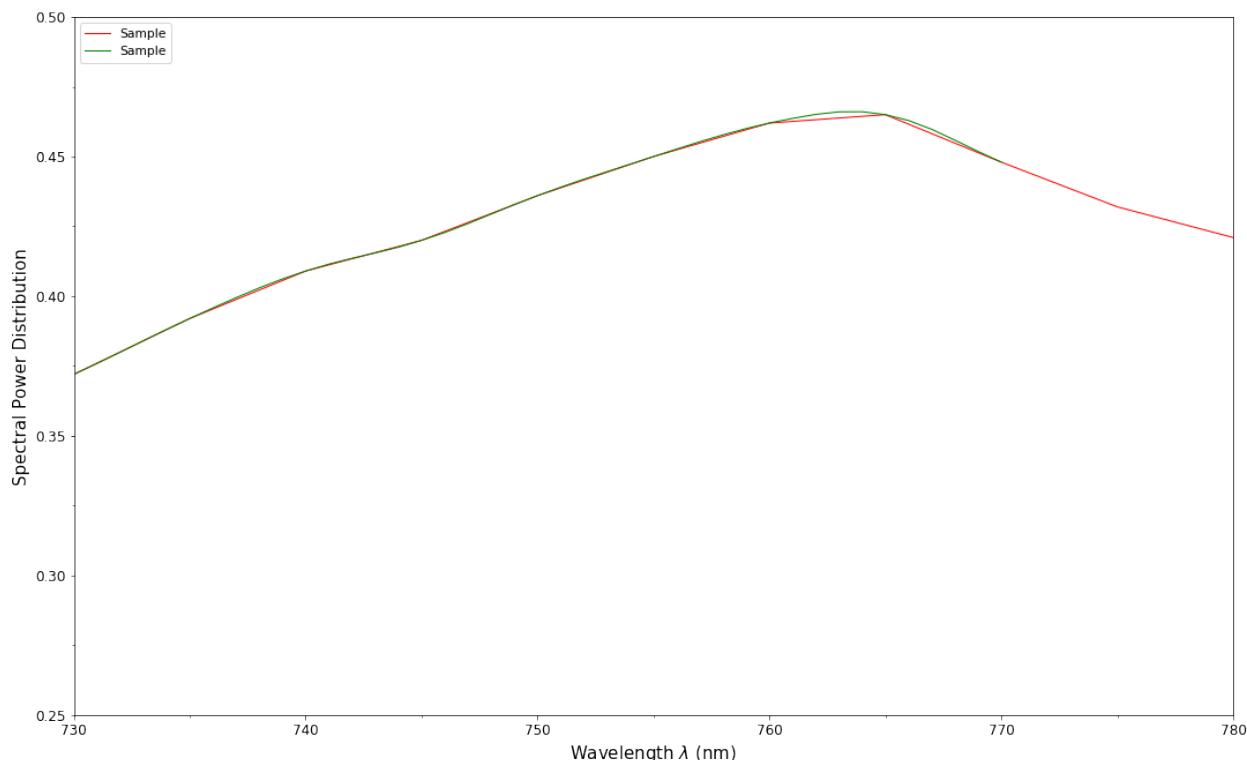
```
# *Colour* can emit a substantial amount of warnings, we filter them.
colour.filter_warnings()

# Copying the sample spectral power distribution.
spd_copy = spd.copy()

# Interpolating the copied sample spectral power distribution.
spd_copy.interpolate(colour.SpectralShape(400, 770, 1))
spd_copy[401]
```

```
0.06580959999999996
```

```
# Comparing the interpolated spectral power distribution with the original one.
multi_spd_plot([spd, spd_copy], bounding_box=[730, 780, 0.25, 0.5])
```



Extrapolation although dangerous can be used to help aligning two spectral power distributions together. CIE publication CIE 15:2004 “Colorimetry” recommends that unmeasured values may be set equal to the nearest measured value of the appropriate quantity in truncation [[CIET14804d](#)]:

```
# Extrapolating the copied sample spectral power distribution.
spd_copy.extrapolate(colour.SpectralShape(340, 830))
spd_copy[340], spd_copy[830]
```

```
(0.0650000000000002, 0.44800000000000018)
```

The underlying interpolator can be swapped for any of the [Colour](#) interpolators.

```
pprint([
    export for export in colour.algebra.interpolation.__all__
        if 'Interpolator' in export
])
```

```
[u'KernelInterpolator',
 u'LinearInterpolator',
 u'SpragueInterpolator',
 u'CubicSplineInterpolator',
 u'PchipInterpolator',
 u'NullInterpolator']
```

```
# Changing interpolator while trimming the copied spectral power distribution.
spd_copy.interpolate(
    colour.SpectralShape(400, 700, 10), interpolator=colour.LinearInterpolator)
```

```
SpectralPowerDistribution([[ 4.0000000e+02,  6.5000000e-02],
                           [ 4.1000000e+02,  6.8000000e-02],
                           [ 4.2000000e+02,  6.4000000e-02],
                           [ 4.3000000e+02,  5.9000000e-02],
                           [ 4.4000000e+02,  5.5000000e-02],
                           [ 4.5000000e+02,  5.3000000e-02],
                           [ 4.6000000e+02,  5.2000000e-02],
                           [ 4.7000000e+02,  5.2000000e-02],
                           [ 4.8000000e+02,  5.4000000e-02],
                           [ 4.9000000e+02,  5.7000000e-02],
                           [ 5.0000000e+02,  6.1000000e-02],
                           [ 5.1000000e+02,  6.5000000e-02],
                           [ 5.2000000e+02,  7.0000000e-02],
                           [ 5.3000000e+02,  7.4000000e-02],
                           [ 5.4000000e+02,  7.6000000e-02],
                           [ 5.5000000e+02,  7.9000000e-02],
                           [ 5.6000000e+02,  8.7000000e-02],
                           [ 5.7000000e+02,  1.0000000e-01],
                           [ 5.8000000e+02,  1.1500000e-01],
                           [ 5.9000000e+02,  1.2900000e-01],
                           [ 6.0000000e+02,  1.3800000e-01],
                           [ 6.1000000e+02,  1.4600000e-01],
                           [ 6.2000000e+02,  1.5400000e-01],
                           [ 6.3000000e+02,  1.6300000e-01],
                           [ 6.4000000e+02,  1.7300000e-01],
                           [ 6.5000000e+02,  1.8800000e-01],
                           [ 6.6000000e+02,  2.0400000e-01],
                           [ 6.7000000e+02,  2.2200000e-01],
                           [ 6.8000000e+02,  2.4200000e-01],
                           [ 6.9000000e+02,  2.6100000e-01],
                           [ 7.0000000e+02,  2.8200000e-01]],
                           interpolator=SpragueInterpolator,
                           interpolator_args={},
                           extrapolator=Extrapolator,
                           extrapolator_args={u'right': None, u'method': u'Constant', u'left': None})
```

The extrapolation behaviour can be changed for *Linear* method instead of the *Constant* default method or

even use arbitrary constant *left* and *right* values:

```
# Extrapolating the copied sample spectral power distribution with *Linear* method.
spd_copy.extrapolate(
    colour.SpectralShape(340, 830),
    extrapolator_args={'method': 'Linear',
                       'right': 0})
spd_copy[340], spd_copy[830]
```

```
(0.046999999999999348, 0.0)
```

Aligning a spectral power distribution is a convenient way to first interpolates the current data within its original bounds, then, if required, extrapolate any missing values to match the requested shape:

```
# Aligning the cloned sample spectral power distribution.
# We first trim the spectral power distribution as above.
spd_copy.interpolate(colour.SpectralShape(400, 700))
spd_copy.align(colour.SpectralShape(340, 830, 5))
spd_copy[340], spd_copy[830]
```

```
(0.06500000000000002, 0.2819999999999975)
```

The `colour.SpectralPowerDistribution` class also supports various arithmetic operations like *addition*, *subtraction*, *multiplication* or *division* with *numeric* and *array\_like* variables or other `colour.SpectralPowerDistribution` class instances:

```
spd = colour.SpectralPowerDistribution({
    410: 0.25,
    420: 0.50,
    430: 0.75,
    440: 1.0,
    450: 0.75,
    460: 0.50,
    480: 0.25
})

print((spd.copy() + 1).values)
print((spd.copy() * 2).values)
print((spd * [0.35, 1.55, 0.75, 2.55, 0.95, 0.65, 0.15]).values)
print((spd * colour.constant_spd(2, spd.shape) * colour.constant_spd(3, spd.shape)).values)
```

```
[ 1.25  1.5   1.75  2.    1.75  1.5   1.25]
[ 0.5   1.    1.5   2.    1.5   1.    0.5]
[ 0.0875  0.775   0.5625  2.55     0.7125  0.325   0.0375]
[ 1.5   3.    4.5   6.    4.5   3.    nan   1.5]
```

The spectral power distribution can be normalised with an arbitrary factor:

```
print(spd.normalise().values)
print(spd.normalise(100).values)
```

```
[ 0.25  0.5   0.75  1.    0.75  0.5   0.25]
[ 25.   50.   75.  100.   75.   50.   25.]
```

A the heart of the `colour.SpectralPowerDistribution` class is the `colour.continuous.Signal` class which implements the `colour.continuous.Signal.function` method.

Evaluating the function for any independent domain  $x \in \mathbb{R}$  variable returns a corresponding range  $y \in \mathbb{R}$  variable.

It adopts an interpolating function encapsulated inside an extrapolating function. The resulting function independent domain, stored as discrete values in the `colour.continuous.Signal.domain` attribute corresponds with the function dependent and already known range stored in the `colour.continuous.Signal.range` attribute.

Describing the `colour.continuous.Signal` class is beyond the scope of this tutorial but we can illustrate its core capability.

```
import numpy as np

range_ = np.linspace(10, 100, 10)
signal = colour.continuous.Signal(range_)
print(repr(signal))
```

```
Signal([[ 0.,  10.],
       [ 1.,  20.],
       [ 2.,  30.],
       [ 3.,  40.],
       [ 4.,  50.],
       [ 5.,  60.],
       [ 6.,  70.],
       [ 7.,  80.],
       [ 8.,  90.],
       [ 9., 100.]],
      interpolator=KernelInterpolator,
      interpolator_args={},
      extrapolator=Extrapolator,
      extrapolator_args={u'right': nan, u'method': u'Constant', u'left': nan})
```

```
# Returning the corresponding range *y* variable for any arbitrary independent domain **x** variable.
signal[np.random.uniform(0, 9, 10)]
```

```
array([ 55.91309735,  65.4172615 ,  65.54495059,  88.17819416,
       61.88860248,  10.53878826,  55.25130534,  46.14659783,
      86.41406136,  84.59897703])
```

### 3.1.1.3 Convert to Tristimulus Values

From a given spectral power distribution, *CIE XYZ* tristimulus values can be calculated:

```
spd = colour.SpectralPowerDistribution(sample_spd_data)
cmfs = colour.STANDARD_OBSERVERS_CMFS['CIE 1931 2 Degree Standard Observer']
illuminant = colour.ILLUMINANTS_RELATIVE_SPDS['D65']

# Calculating the sample spectral power distribution *CIE XYZ* tristimulus values.
XYZ = colour.spectral_to_XYZ(spd, cmfs, illuminant)
print(XYZ)
```

```
[ 10.97085572  9.70278591  6.05562778]
```

### 3.1.1.4 From CIE XYZ Colourspace

*CIE XYZ* is the central colourspace for Colour Science from which many computations are available, cascading to even more computations:

```
# Displaying objects interacting directly with the *CIE XYZ* colourspace.
pprint([name for name in colour.__all__ if name.startswith('XYZ_to')])
```

```
['XYZ_to_Hunt',
 'XYZ_to_ATD95',
 'XYZ_to_CIECAM02',
 'XYZ_to_LLAB',
 'XYZ_to_Nayatani95',
 'XYZ_to_RLAB',
 'XYZ_to_xyY',
 'XYZ_to_xy',
 'XYZ_to_Lab',
 'XYZ_to_Luv',
 'XYZ_to_UCS',
 'XYZ_to_UVW',
 'XYZ_to_hdr_CIELab',
 'XYZ_to_K_ab_HunterLab1966',
 'XYZ_to_Hunter_Lab',
 'XYZ_to_Hunter_Rdab',
 'XYZ_to_Hunter_Rdab',
 'XYZ_to_IPT',
 'XYZ_to_hdr_IPT',
 'XYZ_to_colourspace_model',
 'XYZ_to_RGB',
 'XYZ_to_sRGB',
 'XYZ_to_spectral_Meng2015',
 'XYZ_to_spectral']
```

### 3.1.1.5 Convert to Screen Colours

We can for instance converts the *CIE XYZ* tristimulus values into *sRGB* colourspace *RGB* values in order to display them on screen:

```
# The output domain of *colour.spectral_to_XYZ* is [0, 100] and the input
# domain of *colour.XYZ_to_sRGB* is [0, 1]. We need to take it in account and
# rescale the input *CIE XYZ* colourspace matrix.
RGB = colour.XYZ_to_sRGB(XYZ / 100)
print(RGB)
```

```
[ 0.45675795  0.30986982  0.24861924]
```

```
# Plotting the *sRGB* colourspace colour of the *Sample* spectral power distribution.
single_colour_swatch_plot(ColourSwatch('Sample', RGB), text_size=32)
```

## Sample

### 3.1.1.6 Generate Colour Rendition Charts

In the same way, we can compute values from a colour rendition chart sample.

---

**Note:** This is useful for render time checks in the VFX industry, where you can use a synthetic colour chart into your render and ensure the colour management is acting as expected.

---

The `colour.characterisation` sub-package contains the dataset for various colour rendition charts:

```
# Colour rendition charts chromaticity coordinates.  
print(sorted(colour.characterisation.COLOURCHECKERS.keys()))
```

```
# Colour rendition charts spectral power distributions.  
print(sorted(colour.characterisation.COLOURCHECKERS_SPDS.keys()))
```

```
[u'BabelColor Average', u'ColorChecker 1976', u'ColorChecker 2005', u'babel_average', u'cc2005']  
[u'BabelColor Average', u'ColorChecker N Ohta', u'babel_average', u'cc_ohta']
```

**Note:** The above `cc2005`, `babel_average` and `cc_ohta` keys are convenient aliases for respectively `ColorChecker 2005`, `BabelColor Average` and `ColorChecker N Ohta` keys.

```
# Plotting the *sRGB* colourspace colour of *neutral 5 (.70 D)* patch.  
patch_name = 'neutral 5 (.70 D)'  
patch_spd = colour.COLOURCHECKERS_SPDS['ColorChecker N Ohta'][patch_name]  
XYZ = colour.spectral_to_XYZ(patch_spd, cmfs, illuminant)  
RGB = colour.XYZ_to_sRGB(XYZ / 100)  
  
single_colour_swatch_plot(ColourSwatch(patch_name.title(), RGB), text_size=32)
```

## Neutral 5 (.70 D)

Colour defines a convenient plotting object to draw synthetic colour rendition charts figures:

```
colour_checker_plot(colour_checker='ColorChecker 2005', text_display=False)
```



### 3.1.1.7 Convert to Chromaticity Coordinates

Given a spectral power distribution, chromaticity coordinates  $xy$  can be computed using the `colour.XYZ_to_xy` definition:

```
# Computing *xy* chromaticity coordinates for the *neutral 5 (.70 D)* patch.
xy = colour.XYZ_to_xy(XYZ)
print(xy)
```

```
[ 0.31259787  0.32870029]
```

Chromaticity coordinates  $xy$  can be plotted into the *CIE 1931 Chromaticity Diagram*:

```
import pylab

# Plotting the *CIE 1931 Chromaticity Diagram*.
# The argument *standalone=False* is passed so that the plot doesn't get displayed
# and can be used as a basis for other plots.
chromaticity_diagram_plot_CIE1931(standalone=False)

# Plotting the *xy* chromaticity coordinates.
x, y = xy
pylab.plot(x, y, 'o-', color='white')
```

```
# Annotating the plot.  
pylab.annotate(patch_spd.name.title(),  
    xy=xy,  
    xytext=(-50, 30),  
    textcoords='offset points',  
    arrowprops=dict(arrowstyle='->', connectionstyle='arc3, rad=-0.2'))  
  
# Displaying the plot.  
render(  
    standalone=True,  
    limits=(-0.1, 0.9, -0.1, 0.9),  
    x_tighten=True,  
    y_tighten=True)
```



### 3.1.1.8 And More...

We hope that this small introduction has been useful and gave you the envy to see more, if you want to explore the API a good place to start is the [Jupyter Notebooks](#) page.

## 3.1.2 Reference

### 3.1.2.1 Colour

#### Chromatic Adaptation

- [Chromatic Adaptation](#)
- [Fairchild \(1990\)](#)
- [CIE 1994](#)
- [CMCCAT2000](#)
- [Von Kries](#)

#### Chromatic Adaptation

colour

<code>chromatic_adaptation(XYZ, XYZ_w, XYZ_wr[, ...])</code>	Adapts given stimulus from test viewing conditions to reference viewing conditions.
<code>CHROMATIC_ADAPTATION_METHODS</code>	Supported chromatic adaptation methods.
<code>CMCCAT2000_VIEWING_CONDITIONS</code>	Reference <code>CMCCAT2000</code> chromatic adaptation model viewing conditions.

#### colour.chromatic\_adaptation

`colour.chromatic_adaptation(XYZ, XYZ_w, XYZ_wr, method='Von Kries', **kwargs)`  
Adapts given stimulus from test viewing conditions to reference viewing conditions.

##### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values of stimulus to adapt.
- `XYZ_w` (array\_like) – Test viewing condition *CIE XYZ* tristimulus values of the white-point.
- `XYZ_wr` (array\_like) – Reference viewing condition *CIE XYZ* tristimulus values of the whitepoint.
- `method` (unicode, optional) – {‘Von Kries’, ‘CIE 1994’, ‘CMCCAT2000’, ‘Fairchild 1990’}, Computation method.

##### Other Parameters

- `E_o1` (numeric) – {`colour.adaptation.chromatic_adaptation_CIE1994()`}, Test illuminance  $E_{o1}$  in  $cd/m^2$ .

- **E\_o2** (*numeric*) – {`colour.adaptation.chromatic_adaptation_CIE1994()`}, Reference illuminance  $E_{o2}$  in  $cd/m^2$ .
- **L\_A1** (*numeric or array\_like*) – {`colour.adaptation.chromatic_adaptation_CMCCAT2000()`}, Luminance of test adapting field  $L_{A1}$  in  $cd/m^2$ .
- **L\_A2** (*numeric or array\_like*) – {`colour.adaptation.chromatic_adaptation_CMCCAT2000()`}, Luminance of reference adapting field  $L_{A2}$  in  $cd/m^2$ .
- **Y\_n** (*numeric or array\_like*) – {`colour.adaptation.chromatic_adaptation_Fairchild1990()`}, Luminance  $Y_n$  of test adapting stimulus in  $cd/m^2$ .
- **Y\_o** (*numeric*) – {`colour.adaptation.chromatic_adaptation_CIE1994()`}, Luminance factor  $Y_o$  of achromatic background as percentage in domain [18, 100].
- **direction** (*unicode, optional*) – {`colour.adaptation.chromatic_adaptation_CMCCAT2000()`}, {'Forward', 'Reverse'}, Chromatic adaptation direction.
- **discount\_illuminant** (*bool, optional*) – {`colour.adaptation.chromatic_adaptation_Fairchild1990()`}, Truth value indicating if the illuminant should be discounted.
- **n** (*numeric, optional*) – {`colour.adaptation.chromatic_adaptation_CIE1994()`}, Noise component in fundamental primary system.
- **surround** (*CMCCAT2000\_InductionFactors, optional*) – {`colour.adaptation.chromatic_adaptation_CMCCAT2000()`}, Surround viewing conditions induction factors.
- **transform** (*unicode, optional*) – {`colour.adaptation.chromatic_adaptation_VonKries()`}, {'CAT02', 'XYZ Scaling', 'Von Kries', 'Bradford', 'Sharp', 'Fairchild', 'CMCCAT97', 'CMCCAT2000', 'CAT02\_BRILL\_CAT', 'Bianco', 'Bianco PC'}, Chromatic adaptation transform.

**Returns** *CIE XYZ\_c* tristimulus values of the stimulus corresponding colour.

**Return type** ndarray

## References

- [*CIET13294*]
- [*Fai91*]
- [*Fai13c*]
- [*Fai13b*]
- [*LLRH02*]
- [*WRC12a*]

## Examples

*Von Kries* chromatic adaptation:

```
>>> import numpy as np
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_w = np.array([1.09846607, 1.00000000, 0.35582280])
>>> XYZ_wr = np.array([0.95042855, 1.00000000, 1.08890037])
>>> chromatic_adaptation(XYZ, XYZ_w, XYZ_wr)
...
array([ 0.0839746...,  0.1141321...,  0.2862554...])
```

CIE 1994 chromatic adaptation, requires extra *kwargs*:

```
>>> XYZ = np.array([0.2800, 0.2126, 0.0527])
>>> XYZ_w = np.array([1.09867452, 1.00000000, 0.35591556])
>>> XYZ_wr = np.array([0.95045593, 1.00000000, 1.08905775])
>>> Y_o = 20
>>> E_o = 1000
>>> chromatic_adaptation(
...     XYZ, XYZ_w, XYZ_wr, method='CIE 1994', Y_o=Y_o, E_o1=E_o, E_o2=E_o)
...
array([ 0.2403379...,  0.2115621...,  0.1764301...])
```

CMCCAT2000 chromatic adaptation, requires extra *kwargs*:

```
>>> XYZ = np.array([0.2248, 0.2274, 0.0854])
>>> XYZ_w = np.array([1.1115, 1.0000, 0.3520])
>>> XYZ_wr = np.array([0.9481, 1.0000, 1.0730])
>>> L_A = 200
>>> chromatic_adaptation(
...     XYZ, XYZ_w, XYZ_wr, method='CMCCAT2000', L_A1=L_A, L_A2=L_A)
...
array([ 0.1952698...,  0.2306834...,  0.2497175...])
```

Fairchild (1990) chromatic adaptation, requires extra *kwargs*:

```
>>> XYZ = np.array([0.1953, 0.2307, 0.2497])
>>> Y_n = 200
>>> chromatic_adaptation(
...     XYZ, XYZ_w, XYZ_wr, method='Fairchild 1990', Y_n=Y_n)
...
array([ 0.2332526...,  0.2332455...,  0.7611593...])
```

## colour.CHROMATIC\_ADAPTATION\_METHODS

```
colour.CHROMATIC_ADAPTATION_METHODS = CaseInsensitiveMapping({'Von Kries': ..., 'CMCCAT2000': ..., 'CIE 1994': ...})
```

Supported chromatic adaptation methods.

## References

- [CIET13294]
- [Fai91]
- [Fai13c]
- [Fai13b]
- [LLRH02]

- [\[WRC12a\]](#)

**CHROMATIC\_ADAPTATION\_METHODS** [CaseInsensitiveMapping] {‘CIE 1994’, ‘CMCCAT2000’, ‘Fairchild 1990’, ‘Von Kries’}

## colour.CMCCAT2000\_VIEWING\_CONDITIONS

```
colour.CMCCAT2000_VIEWING_CONDITIONS = CaseInsensitiveMapping({u'Dark': ..., u'Dim': ..., u'Average': ...})  
Reference CMCCAT2000 chromatic adaptation model viewing conditions.
```

### References

- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

**CMCCAT2000\_VIEWING\_CONDITIONS** [CaseInsensitiveMapping] (‘Average’, ‘Dim’, ‘Dark’)

### Dataset

colour

CHROMATIC_ADAPTATION_TRANSFORMS	Supported chromatic adaptation transforms.
---------------------------------	--

## colour.CHROMATIC\_ADAPTATION\_TRANSFORMS

```
colour.CHROMATIC_ADAPTATION_TRANSFORMS = CaseInsensitiveMapping({u'Bradford': ..., u'Fairchild': ..., u'CAT02': ...})  
Supported chromatic adaptation transforms.
```

### References

- [\[BS10\]](#)
- [\[BS08\]](#)
- [\[Fai\]](#)
- [\[LPLMartinezverdu07\]](#)
- [\[Lin09a\]](#)
- [\[WRC12b\]](#)
- [\[WRC12a\]](#)
- [\[Wikb\]](#)

**CHROMATIC\_ADAPTATION\_TRANSFORMS** [CaseInsensitiveMapping] {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}

## Fairchild (1990)

`colour.adaptation`

---

<code>chromatic_adaptation_Fairchild1990(XYZ_1, ...)</code>	Adapts given stimulus <i>CIE XYZ_1</i> tristimulus values from test viewing conditions to reference viewing conditions using <i>Fairchild (1990)</i> chromatic adaptation model.
---	--

---

### `colour.adaptation.chromatic_adaptation_Fairchild1990`

`colour.adaptation.chromatic_adaptation_Fairchild1990(XYZ_1, XYZ_n, XYZ_r, Y_n, discount_illuminant=False)`

Adapts given stimulus *CIE XYZ\_1* tristimulus values from test viewing conditions to reference viewing conditions using *Fairchild (1990)* chromatic adaptation model.

#### Parameters

- `XYZ_1` (`array_like`) – *CIE XYZ\_1* tristimulus values of test sample / stimulus in domain [0, 100].
- `XYZ_n` (`array_like`) – Test viewing condition *CIE XYZ\_n* tristimulus values of whitepoint.
- `XYZ_r` (`array_like`) – Reference viewing condition *CIE XYZ\_r* tristimulus values of whitepoint.
- `Y_n` (`numeric` or `array_like`) – Luminance  $Y_n$  of test adapting stimulus in  $cd/m^2$ .
- `discount_illuminant` (`bool`, optional) – Truth value indicating if the illuminant should be discounted.

**Returns** Adapted *CIE XYZ\_2* tristimulus values of stimulus.

**Return type** `ndarray`

**Warning:** The input domain and output range of that definition are non standard!

#### Notes

- Input *CIE XYZ\_1*, *CIE XYZ\_n* and *CIE XYZ\_r* tristimulus values are in domain [0, 100].
- Output *CIE XYZ\_2* tristimulus values are in range [0, 100].

#### References

- [\[Fai91\]](#)
- [\[Fai13c\]](#)

## Examples

```
>>> XYZ_1 = np.array([19.53, 23.07, 24.97])
>>> XYZ_n = np.array([111.15, 100.00, 35.20])
>>> XYZ_r = np.array([94.81, 100.00, 107.30])
>>> Y_n = 200
>>> chromatic_adaptation_Fairchild1990(XYZ_1, XYZ_n, XYZ_r, Y_n)
...
array([ 23.3252634...,  23.3245581...,  76.1159375...])
```

## CIE 1994

colour.adaptation

---

chromatic_adaptation_CIE1994(XYZ_1, xy_o1, ...)	Adapts given stimulus <i>CIE XYZ_1</i> tristimulus values from test viewing conditions to reference viewing conditions using <i>CIE 1994</i> chromatic adaptation model.
---	--

---

### colour.adaptation.chromatic\_adaptation\_CIE1994

colour.adaptation.chromatic\_adaptation\_CIE1994(*XYZ\_1*, *xy\_o1*, *xy\_o2*, *Y\_o*, *E\_o1*, *E\_o2*, *n=1*)  
Adapts given stimulus *CIE XYZ\_1* tristimulus values from test viewing conditions to reference viewing conditions using *CIE 1994* chromatic adaptation model.

#### Parameters

- **XYZ\_1** (array\_like) – *CIE XYZ* tristimulus values of test sample / stimulus in domain [0, 100].
- **xy\_o1** (array\_like) – Chromaticity coordinates  $x_{o1}$  and  $y_{o1}$  of test illuminant and background.
- **xy\_o2** (array\_like) – Chromaticity coordinates  $x_{o2}$  and  $y_{o2}$  of reference illuminant and background.
- **Y\_o** (numeric) – Luminance factor  $Y_o$  of achromatic background as percentage in domain [18, 100].
- **E\_o1** (numeric) – Test illuminance  $E_{o1}$  in  $cd/m^2$ .
- **E\_o2** (numeric) – Reference illuminance  $E_{o2}$  in  $cd/m^2$ .
- **n** (numeric, optional) – Noise component in fundamental primary system.

**Returns** Adapted *CIE XYZ\_2* tristimulus values of test stimulus.

**Return type** ndarray

**Warning:** The input domain and output range of that definition are non standard!

## Notes

- Input *CIE XYZ\_1* tristimulus values are in domain [0, 100].

- Output *CIE XYZ\_2* tristimulus values are in range [0, 100].

## References

- [\[CIET13294\]](#)

## Examples

```
>>> XYZ_1 = np.array([28.00, 21.26, 5.27])
>>> xy_o1 = np.array([0.4476, 0.4074])
>>> xy_o2 = np.array([0.3127, 0.3290])
>>> Y_o = 20
>>> E_o1 = 1000
>>> E_o2 = 1000
>>> chromatic_adaptation_CIE1994(XYZ_1, xy_o1, xy_o2, Y_o, E_o1, E_o2)
...
array([ 24.0337952..., 21.1562121..., 17.6430119...])
```

## CMCCAT2000

`colour.adaptation`

---

<code>chromatic_adaptation_CMCCAT2000(XYZ, XYZ_w, ...)</code>	Adapts given stimulus <i>CIE XYZ</i> tristimulus values using given viewing conditions.
<code>CMCCAT2000_VIEWING_CONDITIONS</code>	Reference <i>CMCCAT2000</i> chromatic adaptation model viewing conditions.

---

### `colour.adaptation.chromatic_adaptation_CMCCAT2000`

`colour.adaptation.chromatic_adaptation_CMCCAT2000(XYZ, XYZ_w, XYZ_wr, L_A1, L_A2, surround=CMCCAT2000_InductionFactors(F=1), direction=u'Forward')`

Adapts given stimulus *CIE XYZ* tristimulus values using given viewing conditions.

This definition is a convenient wrapper around `colour.adaptation.chromatic_adaptation_forward_CMCCAT2000()` and `colour.adaptation.chromatic_adaptation_reverse_CMCCAT2000()`.

#### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values of the stimulus to adapt.
- `XYZ_w` (`array_like`) – Source viewing condition *CIE XYZ* tristimulus values of the whitepoint.
- `XYZ_wr` (`array_like`) – Target viewing condition *CIE XYZ* tristimulus values of the whitepoint.
- `L_A1` (`numeric` or `array_like`) – Luminance of test adapting field  $L_{A1}$  in  $cd/m^2$ .
- `L_A2` (`numeric` or `array_like`) – Luminance of reference adapting field  $L_{A2}$  in  $cd/m^2$ .

- **surround** (`CMCCAT2000_InductionFactors`, optional) – Surround viewing conditions induction factors.
- **direction** (unicode, optional) – {‘Forward’, ‘Reverse’}, Chromatic adaptation direction.

**Returns** Adapted stimulus *CIE XYZ* tristimulus values.

**Return type** ndarray

**Warning:** The input domain and output range of that definition are non standard!

## Notes

- Input *CIE XYZ*, *CIE XYZ\_w* and *CIE XYZ\_wr* tristimulus values are in domain [0, 100].
- Output *CIE XYZ* tristimulus values are in range [0, 100].

## References

- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

## Examples

```
>>> XYZ = np.array([22.48, 22.74, 8.54])
>>> XYZ_w = np.array([111.15, 100.00, 35.20])
>>> XYZ_wr = np.array([94.81, 100.00, 107.30])
>>> L_A1 = 200
>>> L_A2 = 200
>>> chromatic_adaptation_CMCCAT2000(
...     XYZ, XYZ_w, XYZ_wr, L_A1, L_A2, direction='Forward')
...
array([ 19.5269832...,  23.0683396...,  24.9717522...])
```

Using the *CMCCAT2000* reverse model:

```
>>> XYZ = np.array([19.52698326, 23.06833960, 24.97175229])
>>> XYZ_w = np.array([111.15, 100.00, 35.20])
>>> XYZ_wr = np.array([94.81, 100.00, 107.30])
>>> L_A1 = 200
>>> L_A2 = 200
>>> chromatic_adaptation_CMCCAT2000(
...     XYZ, XYZ_w, XYZ_wr, L_A1, L_A2, direction='Reverse')
...
array([ 22.48,  22.74,   8.54])
```

## colour.adaptation.CMCCAT2000\_VIEWING\_CONDITIONS

`colour.adaptation.CMCCAT2000_VIEWING_CONDITIONS` = `CaseInsensitiveMapping({u'Dark': ..., u'Dim': ..., u'Average': ...})`  
Reference *CMCCAT2000* chromatic adaptation model viewing conditions.

## References

- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

**CMCCAT2000\_VIEWING\_CONDITIONS** [CaseInsensitiveMapping] ('Average', 'Dim', 'Dark')

## Ancillary Objects

colour.adaptation

<code>chromatic_adaptation_forward_CMCCAT2000(XYZ, ...)</code>	Adapts given stimulus <i>CIE XYZ</i> tristimulus values from test viewing conditions to reference viewing conditions using <i>CMCCAT2000</i> forward chromatic adaptation model.
<code>chromatic_adaptation_reverse_CMCCAT2000(...)</code>	Adapts given stimulus corresponding colour <i>CIE XYZ</i> tristimulus values from reference viewing conditions to test viewing conditions using <i>CMCCAT2000</i> reverse chromatic adaptation model.
<code>CMCCAT2000_InductionFactors</code>	<i>CMCCAT2000</i> chromatic adaptation model induction factors.

`colour.adaptation.chromatic_adaptation_forward_CMCCAT2000`

`colour.adaptation.chromatic_adaptation_forward_CMCCAT2000(XYZ, XYZ_w, XYZ_wr,  
L_A1, L_A2, surround=CMCCAT2000_InductionFactors(F=1))`

Adapts given stimulus *CIE XYZ* tristimulus values from test viewing conditions to reference viewing conditions using *CMCCAT2000* forward chromatic adaptation model.

### Parameters

- **XYZ** (array\_like) – *CIE XYZ* tristimulus values of the stimulus to adapt.
- **XYZ\_w** (array\_like) – Test viewing condition *CIE XYZ* tristimulus values of the white-point.
- **XYZ\_wr** (array\_like) – Reference viewing condition *CIE XYZ* tristimulus values of the whitepoint.
- **L\_A1** (numeric or array\_like) – Luminance of test adapting field  $L_{A1}$  in  $cd/m^2$ .
- **L\_A2** (numeric or array\_like) – Luminance of reference adapting field  $L_{A2}$  in  $cd/m^2$ .
- **surround** (*CMCCAT2000\_InductionFactors*, optional) – Surround viewing conditions induction factors.

**Returns** *CIE XYZ\_c* tristimulus values of the stimulus corresponding colour.

**Return type** ndarray

**Warning:** The input domain and output range of that definition are non standard!

## Notes

- Input  $CIE XYZ$ ,  $CIE XYZ_w$  and  $CIE XYZ_wr$  tristimulus values are in domain [0, 100].
- Output  $CIE XYZ_c$  tristimulus values are in range [0, 100].

## References

- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

## Examples

```
>>> XYZ = np.array([22.48, 22.74, 8.54])
>>> XYZ_w = np.array([111.15, 100.00, 35.20])
>>> XYZ_wr = np.array([94.81, 100.00, 107.30])
>>> L_A1 = 200
>>> L_A2 = 200
>>> chromatic_adaptation_forward_CMCCAT2000(XYZ, XYZ_w, XYZ_wr, L_A1, L_A2)
...
array([ 19.5269832...,  23.0683396...,  24.9717522...])
```

## colour.adaptation.chromatic\_adaptation\_reverse\_CMCCAT2000

```
colour.adaptation.chromatic_adaptation_reverse_CMCCAT2000(XYZ_c,           XYZ_w,           XYZ_wr,
                                                               L_A1,           L_A2,           surround=CMCCAT2000_InductionFactors(F=1))
```

Adapts given stimulus corresponding colour  $CIE XYZ$  tristimulus values from reference viewing conditions to test viewing conditions using  $CMCCAT2000$  reverse chromatic adaptation model.

### Parameters

- **XYZ\_c** (array\_like) –  $CIE XYZ$  tristimulus values of the stimulus to adapt.
- **XYZ\_w** (array\_like) – Test viewing condition  $CIE XYZ$  tristimulus values of the whitepoint.
- **XYZ\_wr** (array\_like) – Reference viewing condition  $CIE XYZ$  tristimulus values of the whitepoint.
- **L\_A1** (numeric or array\_like) – Luminance of test adapting field  $L_{A1}$  in  $cd/m^2$ .
- **L\_A2** (numeric or array\_like) – Luminance of reference adapting field  $L_{A2}$  in  $cd/m^2$ .
- **surround** ([CMCCAT2000\\_InductionFactors](#), optional) – Surround viewing conditions induction factors.

**Returns**  $CIE XYZ_c$  tristimulus values of the adapted stimulus.

**Return type** ndarray

**Warning:** The input domain and output range of that definition are non standard!

## Notes

- Input  $CIE XYZ_c$ ,  $CIE XYZ_w$  and  $CIE XYZ_wr$  tristimulus values are in domain [0, 100].
- Output  $CIE XYZ$  tristimulus values are in range [0, 100].

## References

- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

## Examples

```
>>> XYZ_c = np.array([19.53, 23.07, 24.97])
>>> XYZ_w = np.array([111.15, 100.00, 35.20])
>>> XYZ_wr = np.array([94.81, 100.00, 107.30])
>>> L_A1 = 200
>>> L_A2 = 200
>>> chromatic_adaptation_reverse_CMCCAT2000(XYZ_c, XYZ_w, XYZ_wr, L_A1,
...                                              L_A2)
...
array([ 22.4839876...,  22.7419485...,   8.5393392...])
```

## colour.adaptation.CMCCAT2000\_InductionFactors

**class** colour.adaptation.CMCCAT2000\_InductionFactors  
*CMCCAT2000* chromatic adaptation model induction factors.

**Parameters**  $F$  (numeric or array\_like) –  $F$  surround condition.

## References

- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

Create new instance of CMCCAT2000\_InductionFactors( $F$ )

**\_\_init\_\_()**  
 $x.\text{__init__}(\dots)$  initializes  $x$ ; see help(type( $x$ )) for signature

## Methods

---

count(...)

index((value, [start, ...])

Raises ValueError if the value is not present.

---

## Von Kries

colour.adaptation

chromatic_adaptation_VonKries(	XYZ,	XYZ_w,	Adapts given stimulus from test viewing conditions to reference viewing conditions.
CHROMATIC_ADAPTATION_TRANSFORMS	Supported chromatic adaptation transforms.		

## colour.adaptation.chromatic\_adaptation\_VonKries

colour.adaptation.chromatic\_adaptation\_VonKries(*XYZ*, *XYZ\_w*, *XYZ\_wr*, *transform*=u'CAT02')  
Adapts given stimulus from test viewing conditions to reference viewing conditions.

### Parameters

- **XYZ** (array\_like) – *CIE XYZ* tristimulus values of stimulus to adapt.
- **XYZ\_w** (array\_like) – Test viewing condition *CIE XYZ* tristimulus values of white-point.
- **XYZ\_wr** (array\_like) – Reference viewing condition *CIE XYZ* tristimulus values of whitepoint.
- **transform** (unicode, optional) – {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}, Chromatic adaptation transform.

**Returns** *CIE XYZ\_c* tristimulus values of the stimulus corresponding colour.

**Return type** ndarray

## References

- [Fai13b]

## Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_w = np.array([1.09846607, 1.00000000, 0.35582280])
>>> XYZ_wr = np.array([0.95042855, 1.00000000, 1.08890037])
>>> chromatic_adaptation_VonKries(XYZ, XYZ_w, XYZ_wr)
array([ 0.0839746...,  0.1141321...,  0.2862554...])
```

Using Bradford method:

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_w = np.array([1.09846607, 1.00000000, 0.35582280])
>>> XYZ_wr = np.array([0.95042855, 1.00000000, 1.08890037])
>>> transform = 'Bradford'
>>> chromatic_adaptation_VonKries(XYZ, XYZ_w, XYZ_wr, transform)
...
array([ 0.0854032...,  0.1140122...,  0.2972149...])
```

## colour.adaptation.CHROMATIC\_ADAPTATION\_TRANSFORMS

colour.adaptation.CHROMATIC\_ADAPTATION\_TRANSFORMS = CaseInsensitiveMapping({u'Bradford': ..., u'Fairchild': ...})  
Supported chromatic adaptation transforms.

## References

- [BS10]
  - [BS08]
  - [Fai]
  - [LPLMartinezverdu07]
  - [Lin09a]
  - [WRC12b]
  - [WRC12a]
  - [Wikb]

**CHROMATIC\_ADAPTATION\_TRANSFORMS** [CaseInsensitiveMapping] {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}

## Dataset

## colour.adaptation

BRADFORD\_CAT  
BS\_CAT  
BS\_PC\_CAT  
CAT02\_BRILL\_CAT  
CAT02\_CAT  
CMCCAT2000\_CAT  
CMCCAT97\_CAT  
FAIRCHILD\_CAT  
SHARP\_CAT  
VON\_KRIES\_CAT  
XYZ\_SCALING\_CAT

colour.adaptation.BRADFORD CAT

colour.adaptation.BS CAT

## colour.adaptation.BS PC CAT

```
colour.adaptation.BS_PC_CAT = array([[ 0.6489,  0.3915, -0.0404], [-0.3775,  1.3055,  0.072 ], [-0.0271,  0.0888,
```

colour.adaptation.CAT02\_BRILL\_CAT

```
colour.adaptation.CAT02_BRILL_CAT = array([[ 0.7328,  0.4296, -0.1624], [-0.7036,  1.6975,  0.0061], [ 0.,  0.,  0.]])
```

**colour.adaptation.CAT02\_CAT**

```
colour.adaptation.CAT02_CAT = array([[ 0.7328,  0.4296, -0.1624], [-0.7036,  1.6975,  0.0061], [ 0.003 ,  0.0136,
```

**colour.adaptation.CMCCAT2000\_CAT**

```
colour.adaptation.CMCCAT2000_CAT = array([[ 7.98200000e-01,  3.38900000e-01, -1.37100000e-01], [-5.91800000e-
```

**colour.adaptation.CMCCAT97\_CAT**

```
colour.adaptation.CMCCAT97_CAT = array([[ 0.8951, -0.7502,  0.0389], [ 0.2664,  1.7135,  0.0685], [-0.1614,  0.0357, -0.
```

**colour.adaptation.FAIRCHILD\_CAT**

```
colour.adaptation.FAIRCHILD_CAT = array([[ 0.8562,  0.3372, -0.1934], [-0.836 ,  1.8327,  0.0033], [ 0.0357, -0.
```

**colour.adaptation.SHARP\_CAT**

```
colour.adaptation.SHARP_CAT = array([[ 1.2694, -0.0988, -0.1706], [-0.8364,  1.8006,  0.0357], [ 0.0297, -0.0357,
```

**colour.adaptation.VON\_KRIES\_CAT**

```
colour.adaptation.VON_KRIES_CAT = array([[ 0.40024,  0.7076 , -0.08081], [-0.2263 ,  1.16532,  0.0457 ], [ 0. ,  0. ,  0.]])
```

**colour.adaptation.XYZ\_SCALING\_CAT**

```
colour.adaptation.XYZ_SCALING_CAT = array([[ 1.,  0.,  0.], [ 0.,  1.,  0.], [ 0.,  0.,  1.]])
```

**Ancillary Objects**

```
colour.adaptation
```

---

<code>chromatic_adaptation_matrix_VonKries(XYZ_w, ...)</code>	Computes the <i>chromatic adaptation</i> matrix from test viewing conditions to reference viewing conditions.
---	---

---

**colour.adaptation.chromatic\_adaptation\_matrix\_VonKries**

```
colour.adaptation.chromatic_adaptation_matrix_VonKries(XYZ_w, XYZ_wr, transform=u'CAT02')
```

Computes the *chromatic adaptation* matrix from test viewing conditions to reference viewing conditions.

**Parameters**

- `XYZ_w` (`array_like`) – Test viewing condition *CIE XYZ* tristimulus values of white-point.
- `XYZ_wr` (`array_like`) – Reference viewing condition *CIE XYZ* tristimulus values of whitepoint.

- **transform** (unicode, optional) – {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}, Chromatic adaptation transform.

**Returns** Chromatic adaptation matrix.

**Return type** ndarray

**Raises** `KeyError` – If chromatic adaptation method is not defined.

## References

- [Fai13b]

## Examples

```
>>> XYZ_w = np.array([1.09846607, 1.00000000, 0.35582280])
>>> XYZ_wr = np.array([0.95042855, 1.00000000, 1.08890037])
>>> chromatic_adaptation_matrix_VonKries(XYZ_w, XYZ_wr)
...
array([[ 0.8687653..., -0.1416539...,  0.3871961...],
       [-0.1030072...,  1.0584014...,  0.1538646...],
       [ 0.0078167...,  0.0267875...,  2.9608177...]])
```

Using Bradford method:

```
>>> XYZ_w = np.array([1.09846607, 1.00000000, 0.35582280])
>>> XYZ_wr = np.array([0.95042855, 1.00000000, 1.08890037])
>>> method = 'Bradford'
>>> chromatic_adaptation_matrix_VonKries(XYZ_w, XYZ_wr, method)
...
array([[ 0.8446794..., -0.1179355...,  0.3948940...],
       [-0.1366408...,  1.1041236...,  0.1291981...],
       [ 0.0798671..., -0.1349315...,  3.1928829...]])
```

## Algebra

- *Extrapolation*
- *Interpolation*
- *Coordinates*
- *Geometry*
- *Matrix*
- *Random*

## Extrapolation

colour

---

`Extrapolator([interpolator, method, left, ...])`

Extrapolates the 1-D function of given interpolator.

## colour.Extrapolator

```
class colour.Extrapolator(interpolator=None,    method=u'Linear',    left=None,    right=None,
                           dtype=<type 'numpy.float64'>)
```

Extrapolates the 1-D function of given interpolator.

The `colour.Extrapolator` class acts as a wrapper around a given *Colour* or *scipy* interpolator class instance with compatible signature. Two extrapolation methods are available:

- *Linear*: Linearly extrapolates given points using the slope defined by the interpolator boundaries ( $x_i[0], x_i[1]$ ) if  $x < x_i[0]$  and ( $x_i[-1], x_i[-2]$ ) if  $x > x_i[-1]$ .
- *Constant*: Extrapolates given points by assigning the interpolator boundaries values  $x_i[0]$  if  $x < x_i[0]$  and  $x_i[-1]$  if  $x > x_i[-1]$ .

Specifying the *left* and *right* arguments takes precedence on the chosen extrapolation method and will assign the respective *left* and *right* values to the given points.

### Parameters

- `interpolator` (object) – Interpolator object.
- `method` (unicode, optional) – {‘Linear’, ‘Constant’}, Extrapolation method.
- `left` (numeric, optional) – Value to return for  $x < x_i[0]$ .
- `right` (numeric, optional) – Value to return for  $x > x_i[-1]$ .
- `dtype` (type) – Data type used for internal conversions.

```
__class__()
```

### Notes

The interpolator must define *x* and *y* attributes.

### References

- [\[Sas\]](#)
- [\[WRC12d\]](#)

### Examples

Extrapolating a single numeric variable:

```
>>> from colour.algebra import LinearInterpolator
>>> x = np.array([3, 4, 5])
>>> y = np.array([1, 2, 3])
>>> interpolator = LinearInterpolator(x, y)
>>> extrapolator = Extrapolator(interpolator)
>>> extrapolator(1)
-1.0
```

Extrapolating an *array\_like* variable:

```
>>> extrapolator(np.array([6, 7, 8]))
array([ 4.,  5.,  6.])
```

Using the *Constant* extrapolation method:

```
>>> x = np.array([3, 4, 5])
>>> y = np.array([1, 2, 3])
>>> interpolator = LinearInterpolator(x, y)
>>> extrapolator = Extrapolator(interpolator, method='Constant')
>>> extrapolator(np.array([0.1, 0.2, 8, 9]))
array([ 1.,  1.,  3.,  3.])
```

Using defined *left* boundary and *Constant* extrapolation method:

```
>>> x = np.array([3, 4, 5])
>>> y = np.array([1, 2, 3])
>>> interpolator = LinearInterpolator(x, y)
>>> extrapolator = Extrapolator(interpolator, method='Constant', left=0)
>>> extrapolator(np.array([0.1, 0.2, 8, 9]))
array([ 0.,  0.,  3.,  3.])
```

`__init__(interpolator=None, method=u'Linear', left=None, right=None, dtype=<type 'numpy.float64'>)`

## Methods

---

`__init__([interpolator, method, left, ...])`

---

## Interpolation

colour

<code>KernelInterpolator(x, y[, window, kernel, ...])</code>	Kernel based interpolation of a 1-D function.
<code>LinearInterpolator(x, y[, dtype])</code>	Linearly interpolates a 1-D function.
<code>NullInterpolator(x, y[, absolute_tolerance, ...])</code>	Performs 1-D function null interpolation, i.e.
<code>PchipInterpolator(x, y, *args, **kwargs)</code>	Interpolates a 1-D function using Piecewise Cubic Hermite Interpolating Polynomial interpolation.
<code>SpragueInterpolator(x, y[, dtype])</code>	Constructs a fifth-order polynomial that passes through <i>y</i> dependent variable.
<code>lagrange_coefficients(r[, n])</code>	Computes the <i>Lagrange Coefficients</i> at given point <i>r</i> for degree <i>n</i> .

colour.KernelInterpolator

```
class colour.KernelInterpolator(x, y, window=3, kernel=<function kernel_lanczos>, kernel_args=None, padding_args=None, dtype=<type 'numpy.float64'>)
```

Kernel based interpolation of a 1-D function.

The reconstruction of a continuous signal can be described as a linear convolution operation. Inter-

polation can be expressed as a convolution of the given discrete function  $g(x)$  with some continuous interpolation kernel  $k(w)$ :

$$\hat{g}(w_0) = [k * g](w_0) = \sum_{x=-\infty}^{\infty} k(w_0 - x) \cdot g(x)$$

#### Parameters

- **x** (`array_like`) – Independent  $x$  variable values corresponding with  $y$  variable.
- **y** (`array_like`) – Dependent and already known  $y$  variable values to interpolate.
- **window** (`int`, optional) – Width of the window in samples on each side.
- **kernel** (`callable`, optional) – Kernel to use for interpolation.
- **kernel\_args** (`dict`, optional) – Arguments to use when calling the kernel.
- **padding\_args** (`dict`, optional) – Arguments to use when padding  $y$  variable values with the `np.pad()` definition.
- **dtype** (`type`) – Data type used for internal conversions.

```
x
y
window
kernel
kernel_args
padding_args
__call__()
```

#### References

- [\[BB09\]](#)
- [\[Wikm\]](#)

#### Examples

Interpolating a single numeric variable:

```
>>> y = np.array([5.9200, 9.3700, 10.8135, 4.5100,
...                 69.5900, 27.8007, 86.0500])
>>> x = np.arange(len(y))
>>> f = KernelInterpolator(x, y)
>>> f(0.5)
6.9411400...
```

Interpolating an *array\_like* variable:

```
>>> f([0.25, 0.75])
array([ 6.1806208...,  8.0823848...])
```

Using a different *lanczos* kernel:

```
>>> f = KernelInterpolator(x, y, kernel=kernel_sinc)
>>> f([0.25, 0.75])
array([ 6.5147317...,  8.3965466...])
```

Using a different window size:

```
>>> f = KernelInterpolator(
...     x,
...     y,
...     window=16,
...     kernel=kernel_lanczos,
...     kernel_args={'a': 16})
>>> f([0.25, 0.75])
array([ 5.396179...,  5.652109...])
```

```
__init__(x, y, window=3, kernel=<function kernel_lanczos>, kernel_args=None,
padding_args=None, dtype=<type 'numpy.float64'>)
```

## Methods

---

**`__init__(x, y[, window, kernel, ...])`**

---

`colour.LinearInterpolator`

**class colour.LinearInterpolator(x, y, dtype=<type 'numpy.float64'>)**  
Linearly interpolates a 1-D function.

### Parameters

- **x** (`array_like`) – Independent *x* variable values corresponding with *y* variable.
- **y** (`array_like`) – Dependent and already known *y* variable values to interpolate.
- **dtype** (`type`) – Data type used for internal conversions.

`x`

`y`

`__call__()`

### Notes

This class is a wrapper around `numpy.interp` definition.

### Examples

Interpolating a single numeric variable:

```
>>> y = np.array([5.9200, 9.3700, 10.8135, 4.5100,
...               69.5900, 27.8007, 86.0500])
>>> x = np.arange(len(y))
>>> f = LinearInterpolator(x, y)
>>> # Doctests ellipsis for Python 2.x compatibility.
```

```
>>> f(0.5)
7.64...
```

Interpolating an *array\_like* variable:

```
>>> f([0.25, 0.75])
array([ 6.7825,  8.5075])
```

```
__init__(x, y, dtype=<type 'numpy.float64'>)
```

## Methods

---

```
__init__(x, y[, dtype])
```

---

### colour.NullInterpolator

```
class colour.NullInterpolator(x, y, absolute_tolerance=1e-06, relative_tolerance=1e-06, default=nan, dtype=<type 'numpy.float64'>)
```

Performs 1-D function null interpolation, i.e. a call within given tolerances will return existing *y* variable values and default if outside tolerances.

#### Parameters

- **x** (ndarray) – Independent *x* variable values corresponding with *y* variable.
- **y** (ndarray) – Dependent and already known *y* variable values to interpolate.
- **absolute\_tolerance** (numeric, optional) – Absolute tolerance.
- **relative\_tolerance** (numeric, optional) – Relative tolerance.
- **default** (numeric, optional) – Default value for interpolation outside tolerances.
- **dtype** (type) – Data type used for internal conversions.

```
x
y
relative_tolerance
absolute_tolerance
default
__call__()
```

## Examples

```
>>> y = np.array([5.9200, 9.3700, 10.8135, 4.5100,
...                 69.5900, 27.8007, 86.0500])
>>> x = np.arange(len(y))
>>> f = NullInterpolator(x, y)
>>> f(0.5)
nan
>>> f(1.0)
9.3699999...
```

```
>>> f = NullInterpolator(x, y, absolute_tolerance=0.01)
>>> f(1.01)
9.3699999...
```

`__init__(x, y, absolute_tolerance=1e-06, relative_tolerance=1e-06, default=nan, dtype=<type 'numpy.float64'>)`

## Methods

---

`__init__(x, y[, absolute_tolerance, ...])`

---

### colour.PchipInterpolator

`class colour.PchipInterpolator(x, y, *args, **kwargs)`

Interpolates a 1-D function using Piecewise Cubic Hermite Interpolating Polynomial interpolation.

`y`

## Notes

- This class is a wrapper around `scipy.interpolate.PchipInterpolator` class.

`__init__(x, y, *args, **kwargs)`

## Methods

---

`__init__(x, y, *args, **kwargs)`

---

<code>antiderivative([nu])</code>	Construct a new piecewise polynomial representing the antiderivative.
<code>construct_fast(c, x[, extrapolate, axis])</code>	Construct the piecewise polynomial without making checks.
<code>derivative([nu])</code>	Construct a new piecewise polynomial representing the derivative.
<code>extend(c, x[, right])</code>	Add additional breakpoints and coefficients to the polynomial.
<code>from_derivatives(xi, yi[, orders, extrapolate])</code>	Construct a piecewise polynomial in the Bernstein basis, compatible with the specified values and derivatives at breakpoints.
<code>from_power_basis(pp[, extrapolate])</code>	Construct a piecewise polynomial in Bernstein basis from a power basis polynomial.
<code>integrate(a, b[, extrapolate])</code>	Compute a definite integral over a piecewise polynomial.
<code>roots()</code>	Return the roots of the interpolated function.

### colour.SpragueInterpolator

`class colour.SpragueInterpolator(x, y, dtype=<type 'numpy.float64'>)`

Constructs a fifth-order polynomial that passes through `y` dependent variable.

*Sprague (1880)* method is recommended by the *CIE* for interpolating functions having a uniformly spaced independent variable.

### Parameters

- **x** (`array_like`) – Independent  $x$  variable values corresponding with  $y$  variable.
- **y** (`array_like`) – Dependent and already known  $y$  variable values to interpolate.
- **dtype** (`type`) – Data type used for internal conversions.

`x`  
`y`  
`__call__()`

### Notes

The minimum number  $k$  of data points required along the interpolation axis is  $k = 6$ .

### References

- [\[CIET13805b\]](#)
- [\[WRC12e\]](#)

### Examples

Interpolating a single numeric variable:

```
>>> y = np.array([5.9200, 9.3700, 10.8135, 4.5100,
...                 69.5900, 27.8007, 86.0500])
>>> x = np.arange(len(y))
>>> f = SpragueInterpolator(x, y)
>>> f(0.5)
7.2185025...
```

Interpolating an *array\_like* variable:

```
>>> f([0.25, 0.75])
array([ 6.7295161...,  7.8140625...])
```

`__init__(x, y[, dtype=`<code>'numpy.float64'`)`

### Methods

---

`__init__(x, y[, dtype])`

---

### colour.lagrange\_coefficients

`colour.lagrange_coefficients(r, n=4)`

Computes the *Lagrange Coefficients* at given point  $r$  for degree  $n$ .

**Parameters**

- `r` (numeric) – Point to get the *Lagrange Coefficients* at.
- `n` (int, optional) – Degree of the *Lagrange Coefficients* being calculated.

**Returns****Return type** ndarray**References**

- [\[Fai85\]](#)
- [\[Wikl\]](#)

**Examples**

```
>>> lagrange_coefficients(0.1)
array([ 0.8265,  0.2755, -0.1305,  0.0285])
```

**Interpolation Kernels**

colour

<code>kernel_nearest_neighbour(x)</code>	Returns the <i>nearest-neighbour</i> kernel evaluated at given samples.
<code>kernel_linear(x)</code>	Returns the <i>linear</i> kernel evaluated at given samples.
<code>kernel_sinc(x[, a])</code>	Returns the <i>sinc</i> kernel evaluated at given samples.
<code>kernel_lanczos(x[, a])</code>	Returns the <i>lanczos</i> kernel evaluated at given samples.
<code>kernel_cardinal_spline(x[, a, b])</code>	Returns the <i>cardinal spline</i> kernel evaluated at given samples.

**colour.kernel\_nearest\_neighbour**`colour.kernel_nearest_neighbour(x)`Returns the *nearest-neighbour* kernel evaluated at given samples.**Parameters** `x` (array\_like) – Samples at which to evaluate the *nearest-neighbour* kernel.**Returns** The *nearest-neighbour* kernel evaluated at given samples.**Return type** ndarray**References**

- [\[BB09\]](#)

**Examples**

```
>>> kernel_nearest_neighbour(np.linspace(0, 1, 10))
array([1, 1, 1, 1, 1, 0, 0, 0, 0, 0])
```

## colour.kernel\_linear

```
colour.kernel_linear(x)
```

Returns the *linear* kernel evaluated at given samples.

**Parameters** `x` (`array_like`) – Samples at which to evaluate the *linear* kernel.

**Returns** The *linear* kernel evaluated at given samples.

**Return type** ndarray

## References

- [\[BB09\]](#)

## Examples

```
>>> kernel_linear(np.linspace(0, 1, 10))
array([ 1.          ,  0.8888888...,  0.7777777...,  0.6666666...,  0.5555555...,
       0.4444444...,  0.3333333...,  0.2222222...,  0.1111111...,  0.        ])
```

## colour.kernel\_sinc

```
colour.kernel_sinc(x, a=3)
```

Returns the *sinc* kernel evaluated at given samples.

**Parameters**

- `x` (`array_like`) – Samples at which to evaluate the *sinc* kernel.
- `a` (`int`, optional) – Size of the *sinc* kernel.

**Returns** The *sinc* kernel evaluated at given samples.

**Return type** ndarray

## References

- [\[BB09\]](#)

## Examples

```
>>> kernel_sinc(np.linspace(0, 1, 10))
array([ 1.000000...e+00,   9.7981553...e-01,   9.2072542...e-01,
       8.2699334...e-01,   7.0531659...e-01,   5.6425327...e-01,
       4.1349667...e-01,   2.6306440...e-01,   1.2247694...e-01,
       3.8981718...e-17])
```

## colour.kernel\_lanczos

```
colour.kernel_lanczos(x, a=3)
```

Returns the *lanczos* kernel evaluated at given samples.

### Parameters

- **x** (`array_like`) – Samples at which to evaluate the *lanczos* kernel.
- **a** (`int`, optional) – Size of the *lanczos* kernel.

**Returns** The *lanczos* kernel evaluated at given samples.

**Return type** ndarray

## References

- [\[Wikm\]](#)

## Examples

```
>>> kernel_lanczos(np.linspace(0, 1, 10))
array([ 1.0000000...e+00,   9.7760615...e-01,   9.1243770...e-01,
       8.1030092...e-01,   6.8012706...e-01,   5.3295773...e-01,
       3.8071690...e-01,   2.3492839...e-01,   1.0554054...e-01,
       3.2237621...e-17])
```

## colour.kernel\_cardinal\_spline

```
colour.kernel_cardinal_spline(x, a=0.5, b=0.0)
```

Returns the *cardinal spline* kernel evaluated at given samples.

Notable *cardinal spline* *a* and *b* parameterizations:

- *Catmull-Rom*: (*a* = 0.5, *b* = 0)
- *Cubic B-Spline*: (*a* = 0, *b* = 1)
- *Mitchell-Netravalli*: (*a* =  $\frac{1}{3}$ , *b* =  $\frac{1}{3}$ )

### Parameters

- **x** (`array_like`) – Samples at which to evaluate the *cardinal spline* kernel.
- **a** (`int`, optional) – *a* control parameter.
- **b** (`int`, optional) – *b* control parameter.

**Returns** The *cardinal spline* kernel evaluated at given samples.

**Return type** ndarray

## References

- [\[BB09\]](#)

## Examples

```
>>> kernel_cardinal_spline(np.linspace(0, 1, 10))
array([ 1.          ,  0.9711934...,  0.8930041...,  0.7777777...,  0.6378600...,
       0.4855967...,  0.3333333...,  0.1934156...,  0.0781893...,  0.          ])
```

## Coordinates

### colour.algebra

<code>cartesian_to_spherical(a)</code>	Transforms given Cartesian coordinates array $xyz$ to Spherical coordinates array $\rho\theta\phi$ (radial distance, inclination or elevation and azimuth).
<code>spherical_to_cartesian(a)</code>	Transforms given Spherical coordinates array $\rho\theta\phi$ (radial distance, inclination or elevation and azimuth) to Cartesian coordinates array $xyz$ .
<code>cartesian_to_polar(a)</code>	Transforms given Cartesian coordinates array $xy$ to Polar coordinates array $\rho\phi$ (radial coordinate, angular coordinate).
<code>polar_to_cartesian(a)</code>	Transforms given Polar coordinates array $\rho\phi$ (radial coordinate, angular coordinate) to Cartesian coordinates array $xy$ .
<code>cartesian_to_cylindrical(a)</code>	Transforms given Cartesian coordinates array $xyz$ to Cylindrical coordinates array $\rho\phi z$ (azimuth, radial distance and height).
<code>cylindrical_to_cartesian(a)</code>	Transforms given Cylindrical coordinates array $\rho\phi z$ (azimuth, radial distance and height) to Cartesian coordinates array $xyz$ .

### colour.algebra.cartesian\_to\_spherical

#### colour.algebra.cartesian\_to\_spherical(*a*)

Transforms given Cartesian coordinates array  $xyz$  to Spherical coordinates array  $\rho\theta\phi$  (radial distance, inclination or elevation and azimuth).

**Parameters** *a* (array\_like) – Cartesian coordinates array  $xyz$  to transform.

**Returns** Spherical coordinates array  $\rho\theta\phi$ .

**Return type** ndarray

## References

- [\[Wiko\]](#)
- [\[Wikk\]](#)

## Examples

```
>>> a = np.array([3, 1, 6])
>>> cartesian_to_spherical(a)
array([ 6.7823299...,  1.0857465...,  0.3217505...])
```

### colour.algebra.spherical\_to\_cartesian

`colour.algebra.spherical_to_cartesian(a)`

Transforms given Spherical coordinates array  $\rho\theta\phi$  (radial distance, inclination or elevation and azimuth) to Cartesian coordinates array  $xyz$ .

**Parameters** `a` (`array_like`) – Spherical coordinates array  $\rho\theta\phi$  to transform.

**Returns** Cartesian coordinates array  $xyz$ .

**Return type** `ndarray`

## References

- [\[Wiko\]](#)
- [\[Wikk\]](#)

## Examples

```
>>> a = np.array([6.78232998, 1.08574654, 0.32175055])
>>> spherical_to_cartesian(a)
array([ 3.        ,  0.9999999...,  6.        ])
```

### colour.algebra.cartesian\_to\_polar

`colour.algebra.cartesian_to_polar(a)`

Transforms given Cartesian coordinates array  $xy$  to Polar coordinates array  $\rho\phi$  (radial coordinate, angular coordinate).

**Parameters** `a` (`array_like`) – Cartesian coordinates array  $xy$  to transform.

**Returns** Polar coordinates array  $\rho\phi$ .

**Return type** `ndarray`

## References

- [\[Wiko\]](#)
- [\[Wikk\]](#)

## Examples

```
>>> a = np.array([3, 1])
>>> cartesian_to_polar(a)
array([ 3.1622776...,  0.3217505...])
```

### colour.algebra.polar\_to\_cartesian

colour.algebra.polar\_to\_cartesian(*a*)

Transforms given Polar coordinates array  $\rho\phi$  (radial coordinate, angular coordinate) to Cartesian coordinates array *xy*.

**Parameters** *a* (array\_like) – Polar coordinates array  $\rho\phi$  to transform.

**Returns** Cartesian coordinates array *xy*.

**Return type** ndarray

## References

- [\[Wiko\]](#)
- [\[Wikk\]](#)

## Examples

```
>>> a = np.array([3.16227766, 0.32175055])
>>> polar_to_cartesian(a)
array([ 3.       ,  0.9999999...])
```

### colour.algebra.cartesian\_to\_cylindrical

colour.algebra.cartesian\_to\_cylindrical(*a*)

Transforms given Cartesian coordinates array *xyz* to Cylindrical coordinates array  $\rho\phi z$  (azimuth, radial distance and height).

**Parameters** *a* (array\_like) – Cartesian coordinates array *xyz* to transform.

**Returns** Cylindrical coordinates array  $\rho\phi z$ .

**Return type** ndarray

## References

- [\[Wiko\]](#)
- [\[Wikk\]](#)

## Examples

```
>>> a = np.array([3, 1, 6])
>>> cartesian_to_cylindrical(a)
array([ 3.1622776...,  0.3217505...,  6.        ])
```

### colour.algebra.cylindrical\_to\_cartesian

`colour.algebra.cylindrical_to_cartesian(a)`

Transforms given Cylindrical coordinates array  $\rho\phi z$  (azimuth, radial distance and height) to Cartesian coordinates array  $xyz$ .

**Parameters** `a` (`array_like`) – Cylindrical coordinates array  $\rho\phi z$  to transform.

**Returns** Cartesian coordinates array  $xyz$ .

**Return type** `ndarray`

## References

- [\[Wiko\]](#)
- [\[Wikk\]](#)

## Examples

```
>>> a = np.array([3.16227766, 0.32175055, 6.00000000])
>>> cylindrical_to_cartesian(a)
array([ 3.        ,  0.999999...,  6.        ])
```

## Geometry

### colour.algebra

<code>normalise_vector(a)</code>	Normalises given vector $a$ .
<code>euclidean_distance(a, b)</code>	Returns the euclidean distance between point arrays $a$ and $b$ .
<code>extend_line_segment(a, b[, distance])</code>	Extends the line segment defined by point arrays $a$ and $b$ by given distance and return the new end point.
<code>intersect_line_segments(l_1, l_2)</code>	Computes $l_1$ line segments intersections with $l_2$ line segments.

### colour.algebra.normalise\_vector

`colour.algebra.normalise_vector(a)`

Normalises given vector  $a$ .

**Parameters** `a` (`array_like`) – Vector  $a$  to normalise.

**Returns** Normalised vector  $a$ .

**Return type** ndarray

### Examples

```
>>> a = np.array([0.07049534, 0.10080000, 0.09558313])
>>> normalise_vector(a)
array([ 0.4525410...,  0.6470802...,  0.6135908...])
```

## colour.algebra.euclidean\_distance

colour.algebra.euclidean\_distance(*a*, *b*)

Returns the euclidean distance between point arrays *a* and *b*.

### Parameters

- **a** (array\_like) – Point array *a*.
- **b** (array\_like) – Point array *b*.

**Returns** Euclidean distance.

**Return type** numeric or ndarray

### Examples

```
>>> a = np.array([100.00000000, 21.57210357, 272.22819350])
>>> b = np.array([100.00000000, 426.67945353, 72.39590835])
>>> euclidean_distance(a, b)
451.7133019...
```

## colour.algebra.extend\_line\_segment

colour.algebra.extend\_line\_segment(*a*, *b*, *distance*=1)

Extends the line segment defined by point arrays *a* and *b* by given distance and return the new end point.

### Parameters

- **a** (array\_like) – Point array *a*.
- **b** (array\_like) – Point array *b*.
- **distance** (numeric, optional) – Distance to extend the line segment.

**Returns** New end point.

**Return type** ndarray

### References

- [Sae]

## Notes

- Input line segment points coordinates are 2d coordinates.

## Examples

```
>>> a = np.array([0.95694934, 0.13720932])
>>> b = np.array([0.28382835, 0.60608318])
>>> extend_line_segment(a, b)
array([-0.5367248..., 1.1776534...])
```

## colour.algebra.intersect\_line\_segments

`colour.algebra.intersect_line_segments(l_1, l_2)`

Computes  $l_1$  line segments intersections with  $l_2$  line segments.

### Parameters

- `l_1` (array\_like) –  $l_1$  line segments array, each row is a line segment such as  $(x_1, y_1, x_2, y_2)$  where  $(x_1, y_1)$  and  $(x_2, y_2)$  are respectively the start and end points of  $l_1$  line segments.
- `l_2` (array\_like) –  $l_2$  line segments array, each row is a line segment such as  $(x_3, y_3, x_4, y_4)$  where  $(x_3, y_3)$  and  $(x_4, y_4)$  are respectively the start and end points of  $l_2$  line segments.

**Returns** Line segments intersections specification.

**Return type** `LineSegmentsIntersections_Specification`

## References

- [\[Bou\]](#)
- [\[Erda\]](#)

## Notes

- Input line segments points coordinates are 2d coordinates.

## Examples

```
>>> l_1 = np.array(
...     [[[0.15416284, 0.7400497],
...      [0.26331502, 0.53373939]],
...     [[0.01457496, 0.91874701],
...      [0.90071485, 0.03342143]]]
... )
>>> l_2 = np.array(
...     [[[0.95694934, 0.13720932],
...      [0.28382835, 0.60608318]],
```

```
...      [[0.94422514, 0.85273554],
...      [0.00225923, 0.52122603]],
...      [[0.55203763, 0.48537741],
...      [0.76813415, 0.16071675]]]
...
)
>>> s = intersect_line_segments(l_1, l_2)
>>> s.xy
array([[[],      nan,      nan],
       [ 0.2279184..., 0.6006430...],
       [      nan,      nan],

       [[ 0.4281451..., 0.5055568...],
       [ 0.3056055..., 0.6279838...],
       [ 0.7578749..., 0.1761301...]]])
>>> s.intersect
array([[False,  True, False],
       [ True,  True,  True]], dtype=bool)
>>> s.parallel
array([[False, False, False],
       [False, False, False]], dtype=bool)
>>> s.coincident
array([[False, False, False],
       [False, False, False]], dtype=bool)
```

## Ancillary Objects

colour.algebra

[LineSegmentsIntersections\\_Specification](#)

Defines the specification for intersection of line segments  $l_1$  and  $l_2$  returned by `colour.algebra.intersect_line_segments()` definition.

`colour.algebra.LineSegmentsIntersections_Specification`

`class colour.algebra.LineSegmentsIntersections_Specification`

Defines the specification for intersection of line segments  $l_1$  and  $l_2$  returned by `colour.algebra.intersect_line_segments()` definition.

### Parameters

- **xy** (array\_like) – Array of  $l_1$  and  $l_2$  line segments intersections coordinates. Non existing segments intersections coordinates are set with `np.nan`.
- **intersect** (array\_like) – Array of `bool` indicating if line segments  $l_1$  and  $l_2$  intersect.
- **parallel** (array\_like) – Array of `bool` indicating if line segments  $l_1$  and  $l_2$  are parallel.
- **coincident** (array\_like) – Array of `bool` indicating if line segments  $l_1$  and  $l_2$  are coincident.

Create new instance of `LineSegmentsIntersections_Specification(xy, intersect, parallel, coincident)`

`__init__()`

`x.__init__(...)` initializes `x`; see `help(type(x))` for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...])</code>	Raises ValueError if the value is not present.

---

## Matrix

`colour.algebra`

---

<code>is_identity(a[, n])</code>	Returns if <i>a</i> array is an identity matrix.
----------------------------------	--

---

`colour.algebra.is_identity`

`colour.algebra.is_identity(a, n=3)`  
Returns if *a* array is an identity matrix.

### Parameters

- `a` (array\_like, (N)) – Variable *a* to test.
- `n` (int, optional) – Matrix dimension.

**Returns** Is identity matrix.

**Return type** bool

## Examples

```
>>> is_identity(np.array([1, 0, 0, 0, 1, 0, 0, 0, 1]).reshape(3, 3))
True
>>> is_identity(np.array([1, 2, 0, 0, 1, 0, 0, 0, 1]).reshape(3, 3))
False
```

## Random

`colour.algebra`

---

<code>random_triplet_generator(size[, limits, ...])</code>	Returns a generator yielding random triplets.
--	---

---

`colour.algebra.random_triplet_generator`

`colour.algebra.random_triplet_generator(size, limits=array([[0, 1], [0, 1], [0, 1]]), random_state=<mtrand.RandomState object>)`

Returns a generator yielding random triplets.

### Parameters

- `size` (integer) – Generator size.
- `limits` (array\_like, (3, 2)) – Random values limits on each triplet axis.

- `random_state` (`RandomState`) – Mersenne Twister pseudo-random number generator.

**Returns** Random triplets generator.

**Return type** generator

## Notes

- The doctest is assuming that `np.random.RandomState()` definition will return the same sequence no matter which *OS* or *Python* version is used. There is however no formal promise about the *prng* sequence reproducibility of either *Python* or *Numpy* implementations: Laurent. (2012). Reproducibility of python pseudo-random numbers across systems and versions? Retrieved January 20, 2015, from <http://stackoverflow.com/questions/8786084/reproducibility-of-python-pseudo-random-numbers-across-systems-and-versions>

## Examples

```
>>> from pprint import pprint
>>> prng = np.random.RandomState(4)
>>> pprint(tuple(random_triplet_generator(10, random_state=prng)))
...
(array([ 0.9670298...,  0.5472322...,  0.9726843...]),
 array([ 0.7148159...,  0.6977288...,  0.2160895...]),
 array([ 0.9762744...,  0.0062302...,  0.2529823...]),
 array([ 0.4347915...,  0.7793829...,  0.1976850...]),
 array([ 0.8629932...,  0.9834006...,  0.1638422...]),
 array([ 0.5973339...,  0.0089861...,  0.3865712...]),
 array([ 0.0441600...,  0.9566529...,  0.4361466...]),
 array([ 0.9489773...,  0.7863059...,  0.8662893...]),
 array([ 0.1731654...,  0.0749485...,  0.6007427...]),
 array([ 0.1679721...,  0.7333801...,  0.4084438...]))
```

## Colour Appearance Models

- *ATD* (1995)
- *CIECAM02*
- *CAM16*
- *Hunt*
- *LLAB*(*l:c*)
- *Nayatani* (1995)
- *RLAB*

## ATD (1995)

colour

---

<code>XYZ_to_ATD95(XYZ, XYZ_0, Y_0, k_1, k_2[, sigma])</code>	Computes the <i>ATD (1995)</i> colour vision model correlates.
<a href="#">ATD95_Specification</a>	Defines the <i>ATD (1995)</i> colour vision model specification.

---

## colour.XYZ\_to\_ATD95

`colour.XYZ_to_ATD95(XYZ, XYZ_0, Y_0, k_1, k_2, sigma=300)`  
Computes the *ATD (1995)* colour vision model correlates.

### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values of test sample / stimulus in domain [0, 100].
- `XYZ_0` (array\_like) – *CIE XYZ* tristimulus values of reference white in domain [0, 100].
- `Y_0` (numeric or array\_like) – Absolute adapting field luminance in  $cd/m^2$ .
- `k_1` (numeric or array\_like) – Application specific weight  $k_1$ .
- `k_2` (numeric or array\_like) – Application specific weight  $k_2$ .
- `sigma` (numeric or array\_like, optional) – Constant  $\sigma$  varied to predict different types of data.

**Returns** *ATD (1995)* colour vision model specification.

**Return type** [ATD95\\_Specification](#)

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].
- Input *CIE XYZ\_0* tristimulus values are in domain [0, 100].
- For unrelated colors, there is only self-adaptation and  $k_1$  is set to 1.0 while  $k_2$  is set to 0.0. For related colors such as typical colorimetric applications,  $k_1$  is set to 0.0 and  $k_2$  is set to a value between 15 and 50 (*Guth, 1995*).

## References

- [\[Fai13a\]](#)
- [\[Gut95\]](#)

## Examples

```
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_0 = np.array([95.05, 100.00, 108.88])
>>> Y_0 = 318.31
>>> k_1 = 0.0
>>> k_2 = 50.0
>>> XYZ_to_ATD95(XYZ, XYZ_0, Y_0, k_1, k_2)
ATD95_Specification(h=1.9089869..., C=1.2064060..., Q=0.1814003..., A_1=0.1787931... T_1=0.
˓→0286942..., D_1=0.0107584..., A_2=0.0192182..., T_2=0.0205377..., D_2=0.0107584...)
```

## colour.ATD95\_Specification

### class colour.ATD95\_Specification

Defines the *ATD* (1995) colour vision model specification.

This specification has field names consistent with the remaining colour appearance models in colour.appearance but diverge from *Fairchild (2013)* reference.

#### Parameters

- **h** (numeric or array\_like) – Hue angle  $H$  in degrees.
- **C** (numeric or array\_like) – Correlate of saturation  $C$ . *Guth (1995)* incorrectly uses the terms saturation and chroma interchangeably. However,  $C$  is here a measure of saturation rather than chroma since it is measured relative to the achromatic response for the stimulus rather than that of a similarly illuminated white.
- **Q** (numeric or array\_like) – Correlate of brightness  $Br$ .
- **A\_1** (numeric or array\_like) – First stage  $A_1$  response.
- **T\_1** (numeric or array\_like) – First stage  $T_1$  response.
- **D\_1** (numeric or array\_like) – First stage  $D_1$  response.
- **A\_2** (numeric or array\_like) – Second stage  $A_2$  response.
- **T\_2** (numeric or array\_like) – Second stage  $A_2$  response.
- **D\_2** (numeric or array\_like) – Second stage  $D_2$  response.

#### Notes

- This specification is the one used in the current model implementation.

#### References

- [\[Fai13a\]](#)
- [\[Gut95\]](#)

Create new instance of ATD95\_Specification(h, C, Q, A\_1, T\_1, D\_1, A\_2, T\_2, D\_2)

**\_\_init\_\_()**  
x.\_\_init\_\_(...) initializes x; see help(type(x)) for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...])</code>	Raises <code>ValueError</code> if the value is not present.

---

## CIECAM02

### colour

---

<code>XYZ_to_CIECAM02(XYZ, XYZ_w, L_A, Y_b[, ...])</code>	Computes the <i>CIECAM02</i> colour appearance model correlates from given <i>CIE XYZ</i> tristimulus values.
<code>CIECAM02_to_XYZ(CIECAM02_specification, ...)</code>	Converts <i>CIECAM02</i> specification to <i>CIE XYZ</i> tristimulus values.
<code>CIECAM02_Specification</code>	Defines the <i>CIECAM02</i> colour appearance model specification.
<code>CIECAM02_VIEWING_CONDITIONS</code>	Reference <i>CIECAM02</i> colour appearance model viewing conditions.

---

### colour.XYZ\_to\_CIECAM02

`colour.XYZ_to_CIECAM02(XYZ, XYZ_w, L_A, Y_b, surround=CIECAM02_InductionFactors( $F=1, c=0.69, N_c=1$ ), discount_illuminant=False)`

Computes the *CIECAM02* colour appearance model correlates from given *CIE XYZ* tristimulus values.

This is the *forward* implementation.

#### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values of test sample / stimulus in domain [0, 100].
- `XYZ_w` (array\_like) – *CIE XYZ* tristimulus values of reference white in domain [0, 100].
- `L_A` (numeric or array\_like) – Adapting field luminance  $L_A$  in  $cd/m^2$ , (often taken to be 20% of the luminance of a white object in the scene).
- `Y_b` (numeric or array\_like) – Relative luminance of background  $Y_b$  in  $cd/m^2$ .
- `surround` (`CIECAM02_InductionFactors`, optional) – Surround viewing conditions induction factors.
- `discount_illuminant` (bool, optional) – Truth value indicating if the illuminant should be discounted.

**Returns** *CIECAM02* colour appearance model specification.

**Return type** `CIECAM02_Specification`

**Warning:** The input domain of that definition is non standard!

## Notes

- Input  $CIE XYZ$  tristimulus values are in domain [0, 100].
- Input  $CIE XYZ_w$  tristimulus values are in domain [0, 100].

## References

- [\[Fai04\]](#)
- [\[LL13\]](#)
- [\[MFH+02\]](#)
- [\[Wikf\]](#)

## Examples

```
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0
>>> surround = CIECAM02_VIEWING_CONDITIONS['Average']
>>> XYZ_to_CIECAM02(XYZ, XYZ_w, L_A, Y_b, surround)
CIECAM02_Specification(J=41.7310911..., C=0.1047077..., h=219.0484326..., s=2.3603053..., Q=195.
↪3713259..., M=0.1088421..., H=278.0607358..., HC=None)
```

## colour.CIECAM02\_to\_XYZ

colour.CIECAM02\_to\_XYZ(CIECAM02\_specification, XYZ\_w, L\_A, Y\_b, surround, round=CIECAM02\_InductionFactors(F=1, c=0.69, N\_c=1), discount\_illuminant=False)

Converts  $CIECAM02$  specification to  $CIE XYZ$  tristimulus values.

This is the *reverse* implementation.

### Parameters

- **CIECAM02\_specification** ([CIECAM02\\_Specification](#)) –  $CIECAM02$  colour appearance model specification. Correlate of *Lightness*  $J$ , correlate of *chroma*  $C$  or correlate of *colourfulness*  $M$  and *hue* angle  $h$  in degrees must be specified, e.g.  $JCh$  or  $JMh$ .
- **XYZ\_w** ([array\\_like](#)) –  $CIE XYZ$  tristimulus values of reference white.
- **L\_A** ([numeric](#) or [array\\_like](#)) – Adapting field *luminance*  $L_A$  in  $cd/m^2$ , (often taken to be 20% of the luminance of a white object in the scene).
- **Y\_b** ([numeric](#) or [array\\_like](#)) – Relative luminance of background  $Y_b$  in  $cd/m^2$ .
- **surround** ([CIECAM02\\_InductionFactors](#), optional) – Surround viewing conditions.
- **discount\_illuminant** ([bool](#), optional) – Discount the illuminant.

**Returns** **XYZ** –  $CIE XYZ$  tristimulus values.

**Return type** [ndarray](#)

**Raises** `ValueError` – If neither `C` or `M` correlates have been defined in the `CIECAM02_specification` argument.

**Warning:** The output range of that definition is non standard!

## Notes

- `CIECAM02_specification` can also be passed as a compatible argument `colour.utilities.as_namedtuple()` definition.
- Input *CIE XYZ\_w* tristimulus values are in domain [0, 100].
- Output *CIE XYZ* tristimulus values are in range [0, 100].

## References

- [\[Fai04\]](#)
- [\[LL13\]](#)
- [\[MFH+02\]](#)
- [\[Wikf\]](#)

## Examples

```
>>> specification = CIECAM02_Specification(J=41.731091132513917,
...                                         C=0.104707757171031,
...                                         h=219.048432658311780)
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0
>>> CIECAM02_to_XYZ(specification, XYZ_w, L_A, Y_b)
array([ 19.01...,  20... ,  21.78...])
```

## colour.CIECAM02\_Specification

### class colour.CIECAM02\_Specification

Defines the *CIECAM02* colour appearance model specification.

#### Parameters

- `J` (numeric or array\_like) – Correlate of *Lightness J*.
- `C` (numeric or array\_like) – Correlate of *chroma C*.
- `h` (numeric or array\_like) – *Hue angle h* in degrees.
- `s` (numeric or array\_like) – Correlate of *saturation s*.
- `Q` (numeric or array\_like) – Correlate of *brightness Q*.
- `M` (numeric or array\_like) – Correlate of *colourfulness M*.
- `H` (numeric or array\_like) – *Hue h quadrature H*.

- `HC` (numeric or array\_like) –  $Hue h$  composition  $H^C$ .

## References

- [\[Fai04\]](#)
- [\[LL13\]](#)
- [\[MFH+02\]](#)
- [\[Wikf\]](#)

Returns a new instance of the `colour.CIECAM02_Specification` class.

`__init__()`  
`x.__init__(...)` initializes `x`; see `help(type(x))` for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...))</code>	Raises <code>ValueError</code> if the value is not present.

---

## colour.CIECAM02\_VIEWING\_CONDITIONS

`colour.CIECAM02_VIEWING_CONDITIONS = CaseInsensitiveMapping({u'Dark': ..., u'Dim': ..., u'Average': ...})`  
Reference *CIECAM02* colour appearance model viewing conditions.

## References

- [\[Fai04\]](#)
- [\[LL13\]](#)
- [\[MFH+02\]](#)
- [\[Wikf\]](#)

`CIECAM02_VIEWING_CONDITIONS` [CaseInsensitiveMapping] {'Average', 'Dim', 'Dark'}

## Ancillary Objects

`colour.appearance`

---

<code>CIECAM02_InductionFactors</code>	<code>CIECAM02</code> colour appearance model induction factors.
--	--

---

## colour.appearance.CIECAM02\_InductionFactors

`class colour.appearance.CIECAM02_InductionFactors`  
*CIECAM02* colour appearance model induction factors.

### Parameters

- `F` (numeric or array\_like) – Maximum degree of adaptation  $F$ .
- `c` (numeric or array\_like) – Exponential non linearity  $c$ .
- `N_c` (numeric or array\_like) – Chromatic induction factor  $N_c$ .

## References

- [\[Fai04\]](#)
- [\[LL13\]](#)
- [\[MFH+02\]](#)
- [\[Wikf\]](#)

Create new instance of CIECAM02\_InductionFactors(`F`, `c`, `N_c`)

`__init__()`  
`x.__init__(...)` initializes `x`; see `help(type(x))` for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...])</code>	Raises <code>ValueError</code> if the value is not present.

---

## CAM16

### colour

---

<code>XYZ_to_CAM16(</code> <code>XYZ</code> , <code>XYZ_w</code> , <code>L_A</code> , <code>Y_b</code> [, ...])	Computes the <i>CAM16</i> colour appearance model correlates from given <i>CIE XYZ</i> tristimulus values.
<code>CAM16_to_XYZ(</code> <code>CAM16_specification</code> , <code>XYZ_w</code> , ...)	Converts <i>CAM16</i> specification to <i>CIE XYZ</i> tristimulus values.
<code>CAM16_Specification</code>	Defines the <i>CAM16</i> colour appearance model specification.
<code>CAM16_VIEWING_CONDITIONS</code>	Reference <i>CAM16</i> colour appearance model viewing conditions.

---

### colour.XYZ\_to\_CAM16

`colour.XYZ_to_CAM16(``XYZ`, `XYZ_w`, `L_A`, `Y_b`, `surround=CAM16_InductionFactors(F=1.0, c=0.69, N_c=1.0)`, `discount_illuminant=False`)

Computes the *CAM16* colour appearance model correlates from given *CIE XYZ* tristimulus values.

This is the *forward* implementation.

#### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values of test sample / stimulus in domain [0, 100].
- `XYZ_w` (array\_like) – *CIE XYZ* tristimulus values of reference white in domain [0, 100].

- `L_A` (numeric or array\_like) – Adapting field *luminance*  $L_A$  in  $cd/m^2$ , (often taken to be 20% of the luminance of a white object in the scene).
- `Y_b` (numeric or array\_like) – Relative luminance of background  $Y_b$  in  $cd/m^2$ .
- `surround` (`CAM16_InductionFactors`, optional) – Surround viewing conditions induction factors.
- `discount_illuminant` (bool, optional) – Truth value indicating if the illuminant should be discounted.

**Returns** *CAM16* colour appearance model specification.

**Return type** `CAM16_Specification`

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].
- Input *CIE XYZ\_w* tristimulus values are in domain [0, 100].

## References

- [\[LLW+17\]](#)

## Examples

```
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0
>>> surround = CAM16_VIEWING_CONDITIONS['Average']
>>> XYZ_to_CAM16(XYZ, XYZ_w, L_A, Y_b, surround)
CAM16_Specification(J=41.7180250..., C=11.9413446..., h=210.3838955..., s=25.3564036..., Q=193.
˓→0617673..., M=12.4128523..., H=267.0983345..., HC=None)
```

## colour.CAM16\_to\_XYZ

```
colour.CAM16_to_XYZ(CAM16_specification, XYZ_w, L_A, Y_b, N_c=1.0, dis-
```

*round=CAM16\_InductionFactors(F=1.0,*

*c=0.69,*

*Y\_b,*

*sur-*

*round*

*count\_illuminant=False)*

Converts *CAM16* specification to *CIE XYZ* tristimulus values.

This is the *reverse* implementation.

### Parameters

- `CAM16_specification` (`CAM16_Specification`) – *CAM16* colour appearance model specification. Correlate of *Lightness*  $J$ , correlate of *chroma*  $C$  or correlate of *colourfulness*  $M$  and *hue* angle  $h$  in degrees must be specified, e.g.  $JCh$  or  $JMh$ .

- `XYZ_w` (array\_like) – *CIE XYZ* tristimulus values of reference white.
- `L_A` (numeric or array\_like) – Adapting field *luminance*  $L_A$  in  $cd/m^2$ , (often taken to be 20% of the luminance of a white object in the scene).
- `Y_b` (numeric or array\_like) – Relative luminance of background  $Y_b$  in  $cd/m^2$ .
- `surround` (`CAM16_InductionFactors`, optional) – Surround viewing conditions.
- `discount_illuminant` (bool, optional) – Discount the illuminant.

**Returns** `XYZ` – *CIE XYZ* tristimulus values.

**Return type** ndarray

**Raises** `ValueError` – If neither `C` or `M` correlates have been defined in the `CAM16_specification` argument.

**Warning:** The output range of that definition is non standard!

## Notes

- `CAM16_specification` can also be passed as a compatible argument `colour.utilities.as_namedtuple()` definition.
- Input *CIE XYZ\_w* tristimulus values are in domain [0, 100].
- Output *CIE XYZ* tristimulus values are in range [0, 100].

## References

- [\[LLW+17\]](#)

## Examples

```
>>> specification = CAM16_Specification(J=41.718025051415616,
...                                         C=11.941344635245843,
...                                         h=210.38389558131118)
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0
>>> CAM16_to_XYZ(specification, XYZ_w, L_A, Y_b)
array([ 19.01...,  20... ,  21.78...])
```

## colour.CAM16\_Specification

### class colour.CAM16\_Specification

Defines the *CAM16* colour appearance model specification.

#### Parameters

- `J` (numeric or array\_like) – Correlate of *Lightness*  $J$ .
- `C` (numeric or array\_like) – Correlate of *chroma*  $C$ .

- **h** (numeric or array\_like) – Hue angle  $h$  in degrees.
- **s** (numeric or array\_like) – Correlate of saturation  $s$ .
- **Q** (numeric or array\_like) – Correlate of brightness  $Q$ .
- **M** (numeric or array\_like) – Correlate of colourfulness  $M$ .
- **H** (numeric or array\_like) – Hue  $h$  quadrature  $H$ .
- **HC** (numeric or array\_like) – Hue  $h$  composition  $H^C$ .

## References

- [\[LLW+17\]](#)

Returns a new instance of the `colour.CAM16_Specification` class.

`__init__()`  
`x.__init__(...)` initializes `x`; see `help(type(x))` for signature

## Methods

---

`count(...)`

`index((value, [start, ...])`

Raises `ValueError` if the value is not present.

---

## colour.CAM16\_VIEWING\_CONDITIONS

`colour.CAM16_VIEWING_CONDITIONS = CaseInsensitiveMapping({u'Dark': ..., u'Dim': ..., u'Average': ...})`  
Reference *CAM16* colour appearance model viewing conditions.

## References

- [\[LLW+17\]](#)

`CAM16_VIEWING_CONDITIONS` [CaseInsensitiveMapping] {'Average', 'Dim', 'Dark'}

## Ancillary Objects

`colour.appearance`

---

`CAM16_InductionFactors`

*CAM16* colour appearance model induction factors.

---

## colour.appearance.CAM16\_InductionFactors

`class colour.appearance.CAM16_InductionFactors`  
*CAM16* colour appearance model induction factors.

### Parameters

- **F** (numeric or array\_like) – Maximum degree of adaptation  $F$ .
- **c** (numeric or array\_like) – Exponential non linearity  $c$ .

- `N_c` (numeric or array\_like) – Chromatic induction factor  $N_c$ .

## References

- [LLW+17]

Create new instance of CAM16\_InductionFactors(`F`, `c`, `N_c`)

`__init__()`  
`x.__init__(...)` initializes `x`; see help(type(`x`)) for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...])</code>	Raises ValueError if the value is not present.

---

## Hunt

### colour

---

<code>XYZ_to_Hunt(</code> <code>XYZ</code> , <code>XYZ_w</code> , <code>XYZ_b</code> , <code>L_A</code> [, ...])	Computes the <i>Hunt</i> colour appearance model correlates.
<code>Hunt_Specification</code>	Defines the <i>Hunt</i> colour appearance model specification.
<code>HUNT_VIEWING_CONDITIONS</code>	Reference <i>Hunt</i> colour appearance model viewing conditions.

---

### colour.XYZ\_to\_Hunt

```
colour.XYZ_to_Hunt(XYZ, XYZ_w, XYZ_b, L_A, surround=Hunt_InductionFactors(N_c=1, N_b=75,
N_cb=None, N_bb=None), L_AS=None, CCT_w=None, XYZ_p=None, p=None,
S=None, S_w=None, helson_judd_effect=False, discount_illuminant=True)
```

Computes the *Hunt* colour appearance model correlates.

#### Parameters

- `XYZ` (array\_like) – CIE XYZ tristimulus values of test sample / stimulus in domain [0, 100].
- `XYZ_w` (array\_like) – CIE XYZ tristimulus values of reference white in domain [0, 100].
- `XYZ_b` (array\_like) – CIE XYZ tristimulus values of background in domain [0, 100].
- `L_A` (numeric or array\_like) – Adapting field luminance  $L_A$  in  $cd/m^2$ .
- `surround` (Hunt\_InductionFactors, optional) – Surround viewing conditions induction factors.
- `L_AS` (numeric or array\_like, optional) – Scotopic luminance  $L_{AS}$  of the illuminant, approximated if not specified.

- `CCT_w` (numeric or array\_like, optional) – Correlated color temperature  $T_{cp}$ : of the illuminant, needed to approximate  $L_{AS}$ .
- `XYZ_p` (array\_like, optional) – *CIE XYZ* tristimulus values of proximal field in domain [0, 100], assumed to be equal to background if not specified.
- `p` (numeric or array\_like, optional) – Simultaneous contrast / assimilation factor  $p$  with value in domain [-1, 0] when simultaneous contrast occurs and domain [0, 1] when assimilation occurs.
- `S` (numeric or array\_like, optional) – Scotopic response  $S$  to the stimulus, approximated using tristimulus values  $Y$  of the stimulus if not specified.
- `S_w` (numeric or array\_like, optional) – Scotopic response  $S_w$  for the reference white, approximated using the tristimulus values  $Y_w$  of the reference white if not specified.
- `helson_judd_effect` (bool, optional) – Truth value indicating whether the *Helson-Judd* effect should be accounted for.
- `discount_illuminant` (bool, optional) – Truth value indicating if the illuminant should be discounted.

**Returns** *Hunt* colour appearance model specification.

**Return type** *Hunt\_Specification*

**Raises** `ValueError` – If an illegal arguments combination is specified.

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].
- Input *CIE XYZ\_b* tristimulus values are in domain [0, 100].
- Input *CIE XYZ\_w* tristimulus values are in domain [0, 100].
- Input *CIE XYZ\_p* tristimulus values are in domain [0, 100].

## References

- [\[Fai13f\]](#)
- [\[Hun04\]](#)

## Examples

```
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> XYZ_b = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> surround = HUNT_VIEWING_CONDITIONS['Normal Scenes']
>>> CCT_w = 6504.0
>>> XYZ_to_Hunt(XYZ, XYZ_w, XYZ_b, L_A, surround, CCT_w=CCT_w)
```

```

...
Hunt_Specification(J=30.0462678..., C=0.1210508..., h=269.2737594..., s=0.0199093..., Q=22.
                     ↪2097654..., M=0.1238964..., H=None, HC=None)

```

## colour.Hunt\_Specification

### `class colour.Hunt_Specification`

Defines the *Hunt* colour appearance model specification.

This specification has field names consistent with the remaining colour appearance models in colour.appearance but diverge from *Fairchild (2013)* reference.

#### Parameters

- `J` (numeric or array\_like) – Correlate of *Lightness*  $J$ .
- `C` (numeric or array\_like) – Correlate of *chroma*  $C_{94}$ .
- `h` (numeric or array\_like) – *Hue* angle  $h_S$  in degrees.
- `s` (numeric or array\_like) – Correlate of *saturation*  $s$ .
- `Q` (numeric or array\_like) – Correlate of *brightness*  $Q$ .
- `M` (numeric or array\_like) – Correlate of *colourfulness*  $M_{94}$ .
- `H` (numeric or array\_like) – *Hue*  $h$  quadrature  $H$ .
- `HC` (numeric or array\_like) – *Hue*  $h$  composition  $H_C$ .

#### Notes

- This specification is the one used in the current model implementation.

#### References

- [\[Fai13f\]](#)
- [\[Hun04\]](#)

Create new instance of Hunt\_Specification(`J`, `C`, `h`, `s`, `Q`, `M`, `H`, `HC`)

`__init__()`  
`x.__init__(...)` initializes `x`; see `help(type(x))` for signature

#### Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...])</code>	Raises <code>ValueError</code> if the value is not present.

---

## colour.HUNT\_VIEWING\_CONDITIONS

```
colour.HUNT_VIEWING_CONDITIONS = CaseInsensitiveMapping({u'Large Transparencies On Light Boxes': ...,
                                                       u'Telecine': ...})
```

Reference *Hunt* colour appearance model viewing conditions.

## References

- [\[Fai13f\]](#)
- [\[Hun04\]](#)

**HUNT\_VIEWING\_CONDITIONS** [CaseInsensitiveMapping] {‘Small Areas, Uniform Background & Surrounds’, ‘Normal Scenes’, ‘Television & CRT, Dim Surrounds’, ‘Large Transparencies On Light Boxes’, ‘Projected Transparencies, Dark Surrounds’}

Aliases:

- ‘small\_uniform’: ‘Small Areas, Uniform Background & Surrounds’
- ‘normal’: ‘Normal Scenes’
- ‘tv\_dim’: ‘Television & CRT, Dim Surrounds’
- ‘light\_boxes’: ‘Large Transparencies On Light Boxes’
- ‘projected\_dark’: ‘Projected Transparencies, Dark Surrounds’

## LLAB(l:c)

colour

<a href="#">XYZ_to_LLAB</a> (XYZ, XYZ_0, Y_b, L[, surround, ...])	Computes the $LLAB(l:c)$ colour appearance model correlates.
<a href="#">LLAB_Specification</a>	Defines the $LLAB(l:c)$ colour appearance model specification.
<a href="#">LLAB_VIEWING_CONDITIONS</a>	Reference $*LLAB(l - c)*$ colour appearance model viewing conditions.

### colour.XYZ\_to\_LLAB

colour.XYZ\_to\_LLAB(XYZ, XYZ\_0, Y\_b, L, surround=LLAB\_InductionFactors(D=1, F\_S=3, F\_L=1, F\_C=1))

Computes the  $LLAB(l:c)$  colour appearance model correlates.

#### Parameters

- **XYZ** (array\_like) – CIE XYZ tristimulus values of test sample / stimulus in domain [0, 100].
- **XYZ\_0** (array\_like) – CIE XYZ tristimulus values of reference white in domain [0, 100].
- **Y\_b** (numeric or array\_like) – Luminance factor of the background in  $cd/m^2$ .
- **L** (numeric or array\_like) – Absolute luminance  $L$  of reference white in  $cd/m^2$ .
- **surround** ([LLAB\\_InductionFactors](#), optional) – Surround viewing conditions induction factors.

**Returns**  $LLAB(l:c)$  colour appearance model specification.

**Return type** [LLAB\\_Specification](#)

**Warning:** The output range of that definition is non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].
- Input *CIE XYZ\_0* tristimulus values are in domain [0, 100].

## References

- [\[Fai13e\]](#)
- [\[LLK96\]](#)
- [\[LM96\]](#)

## Examples

```
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_0 = np.array([95.05, 100.00, 108.88])
>>> Y_b = 20.0
>>> L = 318.31
>>> surround = LLAB_VIEWING_CONDITIONS['ref_average_4_minus']
>>> XYL_to_LLAB(XYZ, XYZ_0, Y_b, L, surround)
LLAB_Specification(J=37.3668650..., C=0.0089496..., h=270..., s=0.0002395..., M=0.0190185..., ↴HC=None, a=..., b=-0.0190185...)
```

## colour.LLAB\_Specification

### class colour.LLAB\_Specification

Defines the *LLAB(l:c)* colour appearance model specification.

This specification has field names consistent with the remaining colour appearance models in colour.appearance but diverge from *Fairchild (2013)* reference.

#### Parameters

- **J** (numeric or array\_like) – Correlate of *Lightness*  $L_L$ .
- **C** (numeric or array\_like) – Correlate of *chroma*  $Ch_L$ .
- **h** (numeric or array\_like) – *Hue* angle  $h_L$  in degrees.
- **s** (numeric or array\_like) – Correlate of *saturation*  $s_L$ .
- **M** (numeric or array\_like) – Correlate of *colourfulness*  $C_L$ .
- **HC** (numeric or array\_like) – *Hue*  $h$  composition  $H^C$ .
- **a** (numeric or array\_like) – Opponent signal  $A_L$ .
- **b** (numeric or array\_like) – Opponent signal  $B_L$ .

## Notes

- This specification is the one used in the current model implementation.

## References

- [\[Fai13e\]](#)
- [\[LLK96\]](#)
- [\[LM96\]](#)

Create new instance of LLAB\_Specification(J, C, h, s, M, HC, a, b)

`__init__()`  
x.`__init__`(...) initializes x; see help(type(x)) for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...])</code>	Raises ValueError if the value is not present.

---

## colour.LLAB\_VIEWING\_CONDITIONS

`colour.LLAB_VIEWING_CONDITIONS = CaseInsensitiveMapping({u'35mm Projection Transparency, Dark Surround': ...})`  
Reference \*LLAB(l - c)\* colour appearance model viewing conditions.

## References

- [\[Fai13e\]](#)
- [\[LLK96\]](#)
- [\[LM96\]](#)

`LLAB_VIEWING_CONDITIONS` [CaseInsensitiveMapping] {'Reference Samples & Images, Average Surround, Subtending > 4', 'Reference Samples & Images, Average Surround, Subtending < 4', 'Television & VDU Displays, Dim Surround', 'Cut Sheet Transparency, Dim Surround';, '35mm Projection Transparency, Dark Surround'}

Aliases:

- 'ref\_average\_4\_plus': 'Reference Samples & Images, Average Surround, Subtending > 4'
- 'ref\_average\_4\_minus': 'Reference Samples & Images, Average Surround, Subtending < 4'
- 'tv\_dim': 'Television & VDU Displays, Dim Surround'
- 'sheet\_dim': 'Cut Sheet Transparency, Dim Surround'
- 'projected\_dark': '35mm Projection Transparency, Dark Surround'

## Ancillary Objects

`colour.appearance`

---

LLAB_InductionFactors	<i>LLAB(l:c)</i> colour appearance model induction factors.
-----------------------	---

---

**colour.appearance.LLAB\_InductionFactors**

**class** colour.appearance.**LLAB\_InductionFactors**  
*LLAB(l:c)* colour appearance model induction factors.

**Parameters**

- **D** (numeric or array\_like) – *Discounting-the-Illuminant* factor  $D$  in domain [0, 1].
- **F\_S** (numeric or array\_like) – Surround induction factor  $F_S$ .
- **F\_L** (numeric or array\_like) – *Lightness* induction factor  $F_L$ .
- **F\_C** (numeric or array\_like) – *Chroma* induction factor  $F_C$ .

**References**

- [\[Fai13e\]](#)
- [\[LLK96\]](#)
- [\[LM96\]](#)

Create new instance of LLAB\_InductionFactors(D, F\_S, F\_L, F\_C)

**\_\_init\_\_()**  
x.\_\_init\_\_(...) initializes x; see help(type(x)) for signature

**Methods**


---

<b>count(...)</b>	
<b>index((value, [start, ...])</b>	Raises ValueError if the value is not present.

---

**Nayatani (1995)**

## colour

---

<b>XYZ_to_Nayatani95(XYZ, XYZ_n, Y_o, E_o, E_or)</b>	Computes the <i>Nayatani (1995)</i> colour appearance model correlates.
<b>Nayatani95_Specification</b>	Defines the <i>Nayatani (1995)</i> colour appearance model specification.

---

**colour.XYZ\_to\_Nayatani95**

**colour.XYZ\_to\_Nayatani95(XYZ, XYZ\_n, Y\_o, E\_o, E\_or, n=1)**  
Computes the *Nayatani (1995)* colour appearance model correlates.

**Parameters**

- **XYZ** (array\_like) – *CIE XYZ* tristimulus values of test sample / stimulus in domain

[0, 100].

- **XYZ\_n** (array\_like) – CIE XYZ tristimulus values of reference white in domain [0, 100].
- **Y\_o** (numeric or array\_like) – Luminance factor  $Y_o$  of achromatic background as percentage in domain [0.18, 1.0]
- **E\_o** (numeric or array\_like) – Illuminance  $E_o$  of the viewing field in lux.
- **E\_or** (numeric or array\_like) – Normalising illuminance  $E_{or}$  in lux usually in domain [1000, 3000]
- **n** (numeric or array\_like, optional) – Noise term used in the non linear chromatic adaptation model.

**Returns** Nayatani (1995) colour appearance model specification.

**Return type** *Nayatani95\_Specification*

**Warning:** The input domain of that definition is non standard!

## Notes

- Input CIE XYZ tristimulus values are in domain [0, 100].
- Input CIE XYZ\_n tristimulus values are in domain [0, 100].

## References

- [\[Fai13g\]](#)
- [\[NSY95\]](#)

## Examples

```
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_n = np.array([95.05, 100.00, 108.88])
>>> Y_o = 20.0
>>> E_o = 5000.0
>>> E_or = 1000.0
>>> XYZ_to_Nayatani95(XYZ, XYZ_n, Y_o, E_o, E_or)
Nayatani95_Specification(Lstar_P=49.9998829..., C=0.0133550..., h=257.5232268..., s=0.0133550..., u
↪Q=62.6266734..., M=0.0167262..., H=None, HC=None, Lstar_N=50.0039154...)
```

## colour.Nayatani95\_Specification

### class colour.Nayatani95\_Specification

Defines the Nayatani (1995) colour appearance model specification.

This specification has field names consistent with the remaining colour appearance models in colour.appearance but diverge from Fairchild (2013) reference.

#### Parameters

- `Lstar_P` (numeric or array\_like) – Correlate of *achromatic Lightness*  $L_p^*$ .
- `C` (numeric or array\_like) – Correlate of *chroma*  $C$ .
- `h` (numeric or array\_like) – *Hue* angle  $\theta$  in degrees.
- `s` (numeric or array\_like) – Correlate of *saturation*  $S$ .
- `Q` (numeric or array\_like) – Correlate of *brightness*  $B_r$ .
- `M` (numeric or array\_like) – Correlate of *colourfulness*  $M$ .
- `H` (numeric or array\_like) – *Hue h quadrature*  $H$ .
- `HC` (numeric or array\_like) – *Hue h composition*  $H_C$ .
- `Lstar_N` (numeric or array\_like) – Correlate of *normalised achromatic Lightness*  $L_n^*$ .

## Notes

- This specification is the one used in the current model implementation.

## References

- [\[Fai13g\]](#)
- [\[NSY95\]](#)

Create new instance of Nayatani95\_Specification(Lstar\_P, C, h, s, Q, M, H, HC, Lstar\_N)

`__init__()`  
`x.__init__(...)` initializes `x`; see `help(type(x))` for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...])</code>	Raises <code>ValueError</code> if the value is not present.

---

## RLAB

### colour

---

<code>XYZ_to_RLAB(XYZ, XYZ_n, Y_n[, sigma, D])</code>	Computes the <i>RLAB</i> model color appearance correlates.
<code>RLAB_D_FACTOR</code>	<i>RLAB</i> colour appearance model <i>Discounting-the-Illuminant</i> factor values.
<code>RLAB_Specification</code>	Defines the <i>RLAB</i> colour appearance model specification.
<code>RLAB_VIEWING_CONDITIONS</code>	Reference <i>RLAB</i> colour appearance model viewing conditions.

---

## colour.XYZ\_to\_RLAB

colour.XYZ\_to\_RLAB(XYZ, XYZ\_n, Y\_n, sigma=0.4347826086956522, D=1)

Computes the RLAB model color appearance correlates.

### Parameters

- **XYZ** (array\_like) – CIE XYZ tristimulus values of test sample / stimulus in domain [0, 100].
- **XYZ\_n** (array\_like) – CIE XYZ tristimulus values of reference white in domain [0, 100].
- **Y\_n** (numeric or array\_like) – Absolute adapting luminance in  $cd/m^2$ .
- **sigma** (numeric or array\_like, optional) – Relative luminance of the surround, see [colour.RLAB\\_VIEWING\\_CONDITIONS](#) for reference.
- **D** (numeric or array\_like, optional) – Discounting-the-Illuminant factor in domain [0, 1].

**Returns** RLAB colour appearance model specification.

**Return type** [RLAB\\_Specification](#)

**Warning:** The input domain of that definition is non standard!

### Notes

- Input CIE XYZ tristimulus values are in domain [0, 100].
- Input CIE XYZ\_n tristimulus values are in domain [0, 100].

### References

- [\[Fai96\]](#)
- [\[Fai13h\]](#)

### Examples

```
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_n = np.array([109.85, 100, 35.58])
>>> Y_n = 31.83
>>> sigma = RLAB_VIEWING_CONDITIONS['Average']
>>> D = RLAB_D_FACTOR['Hard Copy Images']
>>> XYZ_to_RLAB(XYZ, XYZ_n, Y_n, sigma, D)
RLAB_Specification(J=49.8347069..., C=54.8700585..., h=286.4860208..., s=1.1010410..., HC=None, ↴
a=15.5711021..., b=-52.6142956...)
```

## colour.RLAB\_D\_FACTOR

```
colour.RLAB_D_FACTOR = CaseInsensitiveMapping({u'Soft Copy Images': ..., u'Projected Transparencies, Dark Room': ...})
```

*RLAB* colour appearance model *Discounting-the-Illuminant* factor values.

### References

- [\[Fai96\]](#)
- [\[Fai13h\]](#)

**RLAB\_D\_FACTOR** [CaseInsensitiveMapping] {‘Hard Copy Images’, ‘Soft Copy Images’, ‘Projected Transparencies, Dark Room’}

Aliases:

- ‘hard\_cp\_img’: ‘Hard Copy Images’
- ‘soft\_cp\_img’: ‘Soft Copy Images’
- ‘projected\_dark’: ‘Projected Transparencies, Dark Room’

## colour.RLAB\_Specification

### class colour.RLAB\_Specification

Defines the *RLAB* colour appearance model specification.

This specification has field names consistent with the remaining colour appearance models in colour.appearance but diverge from *Fairchild (2013)* reference.

#### Parameters

- **J** (numeric or array\_like) – Correlate of *Lightness*  $L^R$ .
- **C** (numeric or array\_like) – Correlate of *achromatic chroma*  $C^R$ .
- **h** (numeric or array\_like) – *Hue* angle  $h^R$  in degrees.
- **s** (numeric or array\_like) – Correlate of *saturation*  $s^R$ .
- **HC** (numeric or array\_like) – *Hue h* composition  $H^C$ .
- **a** (numeric or array\_like) – Red-green chromatic response  $a^R$ .
- **b** (numeric or array\_like) – Yellow-blue chromatic response  $b^R$ .

### Notes

- This specification is the one used in the current model implementation.

### References

- [\[Fai96\]](#)
- [\[Fai13h\]](#)

Create new instance of RLAB\_Specification(J, C, h, s, HC, a, b)

`__init__()`  
x.`__init__`(...) initializes x; see help(type(x)) for signature

## Methods

---

<code>count(...)</code>	
<code>index((value, [start, ...))</code>	Raises ValueError if the value is not present.

---

## colour.RLAB\_VIEWING\_CONDITIONS

`colour.RLAB_VIEWING_CONDITIONS = CaseInsensitiveMapping({u'Dark': ..., u'Dim': ..., u'Average': ...})`  
Reference RLAB colour appearance model viewing conditions.

## References

- [\[Fai96\]](#)
- [\[Fai13h\]](#)

`RLAB_VIEWING_CONDITIONS` [CaseInsensitiveMapping] {'Average', 'Dim', 'Dark'}

## Biochemistry

- *Michaelis–Menten Kinetics*

### Michaelis–Menten Kinetics

#### colour.biochemistry

---

<code>reaction_rate_MichealisMenten(S, V_max, K_m)</code>	Describes the rate of enzymatic reactions, by relating reaction rate $v$ to concentration of a substrate $S$ .
<code>substrate_concentration_MichealisMenten(v, ...)</code>	Describes the rate of enzymatic reactions, by relating concentration of a substrate $S$ to reaction rate $v$ .

---

#### colour.biochemistry.reaction\_rate\_MichealisMenten

`colour.biochemistry.reaction_rate_MichealisMenten(S, V_max, K_m)`  
Describes the rate of enzymatic reactions, by relating reaction rate  $v$  to concentration of a substrate  $S$ .

##### Parameters

- `S` (array\_like) – Concentration of a substrate  $S$ .
- `V_max` (array\_like) – Maximum rate  $V_{max}$  achieved by the system, at saturating substrate concentration.
- `K_m` (array\_like) – Substrate concentration  $V_{max}$  at which the reaction rate is half

of  $V_{max}$ .

**Returns** Reaction rate  $v$ .

**Return type** array\_like

## References

- [\[Wikt\]](#)

## Examples

```
>>> reaction_rate_MichealisMenten(0.5, 2.5, 0.8)
0.9615384...
```

## colour.biochemistry.substrate\_concentration\_MichealisMenten

colour.biochemistry.**substrate\_concentration\_MichealisMenten**( $v$ ,  $V_{max}$ ,  $K_m$ )

Describes the rate of enzymatic reactions, by relating concentration of a substrate  $S$  to reaction rate  $v$ .

### Parameters

- $v$  (array\_like) – Reaction rate  $v$ .
- $V_{max}$  (array\_like) – Maximum rate  $V_{max}$  achieved by the system, at saturating substrate concentration.
- $K_m$  (array\_like) – Substrate concentration  $V_{max}$  at which the reaction rate is half of  $V_{max}$ .

**Returns** Concentration of a substrate  $S$ .

**Return type** array\_like

## References

- [\[Wikt\]](#)

## Examples

```
>>> substrate_concentration_MichealisMenten(0.961538461538461, 2.5, 0.8)
...
0.4999999...
```

## Colour Characterisation

- [Colour Fitting](#)
- [Colour Rendition Charts](#)

- *Cameras*
- *Displays*

## Colour Fitting

colour

---

`first_order_colour_fit(m_1, m_2)`

Performs a first order colour fit from given  $m_1$  colour array to  $m_2$  colour array.

---

`colour.first_order_colour_fit`

`colour.first_order_colour_fit(m_1, m_2)`

Performs a first order colour fit from given  $m_1$  colour array to  $m_2$  colour array. The resulting colour fitting matrix is computed using multiple linear regression.

The purpose of that object is for example the matching of two *ColorChecker* colour rendition charts together.

### Parameters

- `m_1` (array\_like, (3, n)) – Test array  $m_1$  to fit onto array  $m_2$ .
- `m_2` (array\_like, (3, n)) – Reference array the array  $m_1$  will be colour fitted against.

**Returns** Colour fitting matrix.

**Return type** ndarray (3, 3)

## Examples

```
>>> m_1 = np.array(  
...     [[0.17224810, 0.09170660, 0.06416938],  
...      [0.49189645, 0.27802050, 0.21923399],  
...      [0.10999751, 0.18658946, 0.29938611],  
...      [0.11666120, 0.14327905, 0.05713804],  
...      [0.18988879, 0.18227649, 0.36056247],  
...      [0.12501329, 0.42223442, 0.37027445],  
...      [0.64785606, 0.22396782, 0.03365194],  
...      [0.06761093, 0.11076896, 0.39779139],  
...      [0.49101797, 0.09448929, 0.11623839],  
...      [0.11622386, 0.04425753, 0.14469986],  
...      [0.36867946, 0.44545230, 0.06028681],  
...      [0.61632937, 0.32323906, 0.02437089],  
...      [0.03016472, 0.06153243, 0.29014596],  
...      [0.11103655, 0.30553067, 0.08149137],  
...      [0.41162190, 0.05816656, 0.04845934],  
...      [0.73339206, 0.53075188, 0.02475212],  
...      [0.47347718, 0.08834792, 0.30310315],  
...      [0.00000000, 0.25187016, 0.35062450],  
...      [0.76809639, 0.78486240, 0.77808297],  
...      [0.53822392, 0.54307997, 0.54710883],
```

```

...
[0.35458526, 0.35318419, 0.35524431],
...
[0.17976704, 0.18000531, 0.17991488],
...
[0.09351417, 0.09510603, 0.09675027],
...
[0.03405071, 0.03295077, 0.03702047]]
...
)
>>> m_2 = np.array(
...
[[0.15579559, 0.09715755, 0.07514556],
...
[0.39113140, 0.25943419, 0.21266708],
...
[0.12824821, 0.18463570, 0.31508023],
...
[0.12028974, 0.13455659, 0.07408400],
...
[0.19368988, 0.21158946, 0.37955964],
...
[0.19957425, 0.36085439, 0.40678123],
...
[0.48896605, 0.20691688, 0.05816533],
...
[0.09775522, 0.16710693, 0.47147724],
...
[0.39358649, 0.12233400, 0.10526425],
...
[0.10780332, 0.07258529, 0.16151473],
...
[0.27502671, 0.34705454, 0.09728099],
...
[0.43980441, 0.26880559, 0.05430533],
...
[0.05887212, 0.11126272, 0.38552469],
...
[0.12705825, 0.25787860, 0.13566464],
...
[0.35612929, 0.07933258, 0.05118732],
...
[0.48131976, 0.42082843, 0.07120612],
...
[0.34665585, 0.15170714, 0.24969804],
...
[0.08261116, 0.24588716, 0.48707733],
...
[0.66054904, 0.65941137, 0.66376412],
...
[0.48051509, 0.47870296, 0.48230082],
...
[0.33045354, 0.32904184, 0.33228886],
...
[0.18001305, 0.17978567, 0.18004416],
...
[0.10283975, 0.10424680, 0.10384975],
...
[0.04742204, 0.04772203, 0.04914226]]
...
)
>>> first_order_colour_fit(m_1, m_2)
array([[ 0.6982266...,  0.0307162...,  0.1621042...],
       [ 0.0689349...,  0.6757961...,  0.1643038...],
      [-0.0631495...,  0.0921247...,  0.9713415...]])

```

## Colour Rendition Charts

### Dataset

colour

COLOURCHECKERS	Aggregated <i>ColourCheckers</i> chromaticity coordinates.
COLOURCHECKERS_SPDS	Aggregated <i>ColourCheckers</i> spectral power distributions.

### colour.COLOURCHECKERS

```

colour.COLOURCHECKERS = CaseInsensitiveMapping({u'ColorChecker 2005': ...,
                                                u'babel_average': ...,
                                                u'ColorCh

```

Aggregated *ColourCheckers* chromaticity coordinates.

## References

- [\[Bab12b\]](#)
- [\[Bab12a\]](#)

**COLOURCHECKERS** [CaseInsensitiveMapping] {‘BabelColor Average’, ‘ColorChecker 2005’, ‘ColorChecker 1976’}

Aliases:

- ‘babel\_average’: ‘BabelColor Average’
- ‘cc2005’: ‘ColorChecker 2005’

## colour.COLOURCHECKERS\_SPDS

```
colour.COLOURCHECKERS_SPDS = CaseInsensitiveMapping({u'babel_average': ..., u'ColorChecker N Ohta': ...})
```

Aggregated *ColourCheckers* spectral power distributions.

## References

- [\[Oht97\]](#)
- [\[Bab12b\]](#)
- [\[Bab12a\]](#)
- [\[MunsellCSsciencea\]](#)

**COLOURCHECKERS** [CaseInsensitiveMapping] {‘BabelColor Average’, ‘ColorChecker N Ohta’}

Aliases:

- ‘babel\_average’: ‘BabelColor Average’
- ‘cc\_ohta’: ‘ColorChecker N Ohta’

## Cameras

### colour.characterisation

---

<code>RGB_SpectralSensitivities([data, domain, labels])</code>	Implements support for a camera <i>RGB</i> spectral sensitivities.
--	--

---

### colour.characterisation.RGB\_SpectralSensitivities

```
class colour.characterisation.RGB_SpectralSensitivities(data=None, domain=None, labels=None, **kwargs)
```

Implements support for a camera *RGB* spectral sensitivities.

#### Parameters

- `data` (Series or Dataframe or Signal or MultiSignal or

---

`MultiSpectralPowerDistribution` or `array_like` or `dict_like`, optional)  
– Data to be stored in the multi-spectral power distribution.

- `domain` (`array_like`, optional) – Values to initialise the multiple `colour.SpectralPowerDistribution` class instances `colour.continuous.Signal.wavelengths` attribute with. If both data and domain arguments are defined, the latter will be used to initialise the `colour.continuous.Signal.wavelengths` attribute.
- `labels` (`array_like`, optional) – Names to use for the `colour.SpectralPowerDistribution` class instances.

### Other Parameters

- `name` (`unicode`, optional) – Multi-spectral power distribution name.
- `interpolator` (`object`, optional) – Interpolator class type to use as interpolating function for the `colour.SpectralPowerDistribution` class instances.
- `interpolator_args` (`dict_like`, optional) – Arguments to use when instantiating the interpolating function of the `colour.SpectralPowerDistribution` class instances.
- `extrapolator` (`object`, optional) – Extrapolator class type to use as extrapolating function for the `colour.SpectralPowerDistribution` class instances.
- `extrapolator_args` (`dict_like`, optional) – Arguments to use when instantiating the extrapolating function of the `colour.SpectralPowerDistribution` class instances.
- `strict_labels` (`array_like`, optional) – Multi-spectral power distribution labels for figures, default to `colour.characterisation.RGB_SpectralSensitivities.labels` attribute value.

`__init__(data=None, domain=None, labels=None, **kwargs)`

### Methods

---

<code>__init__([data, domain, labels])</code>	
<code>align(shape[, interpolator, ...])</code>	Aligns the multi-spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <code>a</code> operand, the operation can be either performed on a copy or in-place.
<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <code>x</code> closest element.
<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the multi-spectral power distribution in-place accordingly to <i>CIE 15:2004</i> and <i>CIE 167:2005</i> recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <code>x</code> variable and corresponding range <code>y</code> variable using given method.
<code>get()</code>	

Continued on next page

Table 3.50 – continued from previous page

interpolate(shape[, interpolator, ...])	Interpolates the multi-spectral power distribution in-place accordingly to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
is_uniform()	Returns if independent domain <i>x</i> variable is uniform.
multi_signal_unpack_data([data, domain, ...])	Unpack given data for multi-continuous signal instantiation.
normalise([factor])	Normalises the multi-spectral power distribution with given normalization factor.
to_dataframe()	Converts the continuous signal to a <i>Pandas DataFrame</i> class instance.
trim(shape)	Trims the multi-spectral power distribution wavelengths to given shape.
trim_wavelengths(shape)	
zeros()	

## Dataset

colour

CAMERAS_RGB_SPECTRAL_SENSITIVITIES	Cameras <i>RGB</i> spectral sensitivities.
------------------------------------	--

colour.CAMERAS\_RGB\_SPECTRAL\_SENSITIVITIES

```
colour.CAMERAS_RGB_SPECTRAL_SENSITIVITIES = CaseInsensitiveMapping({u'Nikon 5100 (NPL)': ..., u'Sigma SDMerill (NPL)': ...})
```

Cameras *RGB* spectral sensitivities.

## References

- [DFGM15]

CAMERAS\_RGB\_SPECTRAL\_SENSITIVITIES [CaseInsensitiveMapping] {Nikon 5100 (NPL), Sigma SDMerill (NPL)}

## Displays

colour.characterisation

RGB_DisplayPrimaries([data, domain, labels])	Implements support for a <i>RGB</i> display (such as a <i>CRT</i> or <i>LCD</i> ) primaries multi-spectral power distributions.
--	---

colour.characterisation.RGB\_DisplayPrimaries

```
class colour.characterisation.RGB_DisplayPrimaries(data=None, domain=None, labels=None, **kwargs)
```

Implements support for a *RGB* display (such as a *CRT* or *LCD*) primaries multi-spectral power distributions.

### Parameters

- **data** (Series or Dataframe or Signal or MultiSignal or MultiSpectralPowerDistribution or array\_like or dict\_like, optional) – Data to be stored in the multi-spectral power distribution.
- **domain** (array\_like, optional) – Values to initialise the multiple colour.SpectralPowerDistribution class instances colour.continuous.Signal.wavelengths attribute with. If both data and domain arguments are defined, the latter will be used to initialise the colour.continuous.Signal.wavelengths attribute.
- **labels** (array\_like, optional) – Names to use for the colour.SpectralPowerDistribution class instances.

#### Other Parameters

- **name** (unicode, optional) – Multi-spectral power distribution name.
- **interpolator** (object, optional) – Interpolator class type to use as interpolating function for the colour.SpectralPowerDistribution class instances.
- **interpolator\_args** (dict\_like, optional) – Arguments to use when instantiating the interpolating function of the colour.SpectralPowerDistribution class instances.
- **extrapolator** (object, optional) – Extrapolator class type to use as extrapolating function for the colour.SpectralPowerDistribution class instances.
- **extrapolator\_args** (dict\_like, optional) – Arguments to use when instantiating the extrapolating function of the colour.SpectralPowerDistribution class instances.
- **strict\_labels** (array\_like, optional) – Multi-spectral power distribution labels for figures, default to colour.characterisation.RGB\_DisplayPrimaries.labels attribute value.

`__init__(data=None, domain=None, labels=None, **kwargs)`

#### Methods

<code>__init__([data, domain, labels])</code>	
<code>align(shape[, interpolator, ...])</code>	Aligns the multi-spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.
<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the multi-spectral power distribution in-place accordingly to CIE 15:2004 and CIE 167:2005 recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <i>x</i> variable and corresponding range <i>y</i> variable using given method.

Continued on next page

Table 3.53 – continued from previous page

<code>get()</code>	
<code>interpolate(shape[, interpolator, ...])</code>	Interpolates the multi-spectral power distribution in-place accordingly to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
<code>is_uniform()</code>	Returns if independent domain $x$ variable is uniform.
<code>multi_signal_unpack_data([data, domain, ...])</code>	Unpack given data for multi-continuous signal instantiation.
<code>normalise([factor])</code>	Normalises the multi-spectral power distribution with given normalization factor.
<code>to_dataframe()</code>	Converts the continuous signal to a <i>Pandas DataFrame</i> class instance.
<code>trim(shape)</code>	Trims the multi-spectral power distribution wavelengths to given shape.
<code>trim_wavelengths(shape)</code>	
<code>zeros()</code>	

## Dataset

`colour`

<code>DISPLAYS_RGB_PRIMARIES</code>	Displays <i>RGB primaries</i> multi-spectral power distributions.
-------------------------------------	---

## `colour.DISPLAYS_RGB_PRIMARIES`

`colour.DISPLAYS_RGB_PRIMARIES = CaseInsensitiveMapping({u'Typical CRT Brainard 1997': ..., u'Apple Studio Di...`  
Displays *RGB primaries* multi-spectral power distributions.

## References

- [\[FW98\]](#)
- [\[Mac10\]](#)

`DISPLAYS_RGB_PRIMARIES` [CaseInsensitiveMapping] {Apple Studio Display, Typical CRT Brainard 1997}

## Colorimetry

- *Spectral Data Structure*
- *Spectral Data Generation*
- *Conversion to Tristimulus Values*
  - *ASTM E308-15*
  - *Integration*
- *Spectral Bandpass Dependence Correction*

- Stearns and Stearns (1988)
- Colour Matching Functions
- Colour Matching Functions Transformations
- Illuminants and Light Sources
- Dominant Wavelength and Purity
- Luminous Efficiency Functions
- Lightness Computation
  - Glasser; Mckinney, Reilly and Schnelle (1958)
  - Wyszecki (1963)
  - CIE 1976
  - Fairchild and Wyble (2010)
  - Fairchild and Chen (2011)
- Luminance Computation
  - Newhall, Nickerson and Judd (1943)
  - CIE 1976
  - ASTM D1535-08e1
  - Fairchild and Wyble (2010)
  - Fairchild and Chen (2011)
- Whiteness Computation
  - Berger (1959)
  - Taube (1960)
  - Stensby (1968)
  - ASTM E313
  - Ganz and Griesser (1979)
  - CIE 2004
- Yellowness Computation
  - ASTM D1925
  - ASTM E313

## Spectral Data Structure

colour

<code>SpectralPowerDistribution([data, domain])</code>	Defines the spectral power distribution: the base object for spectral computations.
<code>MultiSpectralPowerDistribution([data, ...])</code>	Defines multi-spectral power distribution: the base object for multi spectral computations.

Continued on next page

Table 3.55 – continued from previous page

<code>SpectralShape([start, end, interval])</code>	Defines the base object for spectral power distribution shape.
<code>DEFAULT_SPECTRAL_SHAPE</code>	Default spectral shape according to <i>ASTM E308-15</i> practise shape.
<code>ASTME30815_PRACTISE_SHAPE</code>	<i>Shape for *ASTM E308-15 practise* – (360, 780, 1).</i>

## colour.SpectralPowerDistribution

```
class colour.SpectralPowerDistribution(data=None, domain=None, **kwargs)
```

Defines the spectral power distribution: the base object for spectral computations.

The spectral power distribution will be initialised according to *CIE 15:2004* recommendation: the method developed by *Sprague (1880)* will be used for interpolating functions having a uniformly spaced independent variable and the *Cubic Spline* method for non-uniformly spaced independent variable. Extrapolation is performed according to *CIE 167:2005* recommendation.

### Parameters

- `data` (`Series` or `Signal`, `SpectralPowerDistribution` or `array_like` or `dict_like`, optional) – Data to be stored in the spectral power distribution.
- `domain` (`array_like`, optional) – Values to initialise the `colour.SpectralPowerDistribution.wavelength` attribute with. If both `data` and `domain` arguments are defined, the latter will be used to initialise the `colour.SpectralPowerDistribution.wavelength` attribute.

### Other Parameters

- `name` (`unicode`, optional) – Spectral power distribution name.
- `interpolator` (`object`, optional) – Interpolator class type to use as interpolating function.
- `interpolator_args` (`dict_like`, optional) – Arguments to use when instantiating the interpolating function.
- `extrapolator` (`object`, optional) – Extrapolator class type to use as extrapolating function.
- `extrapolator_args` (`dict_like`, optional) – Arguments to use when instantiating the extrapolating function.
- `strict_name` (`unicode`, optional) – Spectral power distribution name for figures, default to `colour.SpectralPowerDistribution.name` attribute value.

`strict_name`

`wavelengths`

`values`

`shape`

`__init__()`

`extrapolate()`

`interpolate()`

`align()`

`trim()`

`normalise()`

## References

- [\[CIET13805a\]](#)
- [\[CIET13805c\]](#)
- [\[CIET14804g\]](#)

## Examples

Instantiating a spectral power distribution with a uniformly spaced independent variable:

```
>>> from colour.utilities import numpy_print_options
>>> data = {
...     500: 0.0651,
...     520: 0.0705,
...     540: 0.0772,
...     560: 0.0870,
...     580: 0.1128,
...     600: 0.1360
... }
>>> with numpy_print_options(suppress=True):
...     SpectralPowerDistribution(data)
SpectralPowerDistribution([[ 500.      ,    0.0651],
                           [ 520.      ,    0.0705],
                           [ 540.      ,    0.0772],
                           [ 560.      ,    0.087 ],
                           [ 580.      ,    0.1128],
                           [ 600.      ,    0.136 ]],
                           interpolator=SpragueInterpolator,
                           interpolator_args={},
                           extrapolator=Extrapolator,
                           extrapolator_args={...})
```

Instantiating a spectral power distribution with a non-uniformly spaced independent variable:

```
>>> data[510] = 0.31416
>>> with numpy_print_options(suppress=True):
...     SpectralPowerDistribution(data)
SpectralPowerDistribution([[ 500.      ,    0.0651 ],
                           [ 510.      ,    0.31416],
                           [ 520.      ,    0.0705 ],
                           [ 540.      ,    0.0772 ],
                           [ 560.      ,    0.087 ],
                           [ 580.      ,    0.1128 ],
                           [ 600.      ,    0.136 ]],
                           interpolator=CubicSplineInterpolator,
                           interpolator_args={},
                           extrapolator=Extrapolator,
                           extrapolator_args={...})
```

`__init__(data=None, domain=None, **kwargs)`

## Methods

<code>__init__([data, domain])</code>	
<code>align(shape[, interpolator, ...])</code>	Aligns the spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.
<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the spectral power distribution in-place according to <i>CIE 15:2004</i> and <i>CIE 167:2005</i> recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <i>x</i> variable and corresponding range <i>y</i> variable using given method.
<code>get()</code>	
<code>interpolate(shape[, interpolator, ...])</code>	Interpolates the spectral power distribution in-place according to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
<code>is_uniform()</code>	Returns if independent domain <i>x</i> variable is uniform.
<code>normalise([factor])</code>	Normalises the spectral power distribution using given normalization factor.
<code>signal_unpack_data([data, domain, dtype])</code>	Unpack given data for continuous signal instantiation.
<code>to_series()</code>	Converts the continuous signal to a <i>Pandas Series</i> class instance.
<code>trim(shape)</code>	Trims the spectral power distribution wavelengths to given spectral shape.
<code>trim_wavelengths(shape)</code>	
<code>zeros()</code>	

## colour.MultiSpectralPowerDistribution

```
class colour.MultiSpectralPowerDistribution(data=None,      domain=None,      labels=None,
                                            **kwargs)
```

Defines multi-spectral power distribution: the base object for multi spectral computations. It is used to model colour matching functions, display primaries, camera sensitivities, etc...

The multi-spectral power distribution will be initialised according to *CIE 15:2004* recommendation: the method developed by *Sprague (1880)* will be used for interpolating functions having a uniformly spaced independent variable and the *Cubic Spline* method for non-uniformly spaced independent variable. Extrapolation is performed according to *CIE 167:2005* recommendation.

### Parameters

- `data` (Series or Dataframe or Signal or MultiSignal or

`MultiSpectralPowerDistribution` or `array_like` or `dict_like`, optional)  
– Data to be stored in the multi-spectral power distribution.

- `domain` (`array_like`, optional) – Values to initialise the multiple `colour.SpectralPowerDistribution` class instances `colour.continuous.Signal.wavelengths` attribute with. If both data and domain arguments are defined, the latter will be used to initialise the `colour.continuous.Signal.wavelengths` attribute.
- `labels` (`array_like`, optional) – Names to use for the `colour.SpectralPowerDistribution` class instances.

### Other Parameters

- `name` (`unicode`, optional) – Multi-spectral power distribution name.
- `interpolator` (`object`, optional) – Interpolator class type to use as interpolating function for the `colour.SpectralPowerDistribution` class instances.
- `interpolator_args` (`dict_like`, optional) – Arguments to use when instantiating the interpolating function of the `colour.SpectralPowerDistribution` class instances.
- `extrapolator` (`object`, optional) – Extrapolator class type to use as extrapolating function for the `colour.SpectralPowerDistribution` class instances.
- `extrapolator_args` (`dict_like`, optional) – Arguments to use when instantiating the extrapolating function of the `colour.SpectralPowerDistribution` class instances.
- `strict_labels` (`array_like`, optional) – Multi-spectral power distribution labels for figures, default to `colour.MultiSpectralPowerDistribution.labels` attribute value.

`strict_name`  
`strict_labels`  
`wavelengths`  
`values`  
`shape`  
`extrapolate()`  
`interpolate()`  
`align()`  
`trim()`  
`normalise()`

### References

- [CIET13805a]
- [CIET13805c]
- [CIET14804g]

### Examples

Instantiating a multi-spectral power distribution with a uniformly spaced independent variable:

```
>>> from colour.utilities import numpy_print_options
>>> data = {
...     500: (0.004900, 0.323000, 0.272000),
...     510: (0.009300, 0.503000, 0.158200),
...     520: (0.063270, 0.710000, 0.078250),
...     530: (0.165500, 0.862000, 0.042160),
...     540: (0.290400, 0.954000, 0.020300),
...     550: (0.433450, 0.994950, 0.008750),
...     560: (0.594500, 0.995000, 0.003900)
... }
>>> labels = ('x_bar', 'y_bar', 'z_bar')
>>> with numpy_print_options(suppress=True):
...     MultiSpectralPowerDistribution(data, labels=labels)
...
MultiSpectral...([[ 500.      ,  0.0049 ,  0.323   ,  0.272   ],
...    ... [ 510.      ,  0.0093 ,  0.503   ,  0.1582 ],
...    ... [ 520.      ,  0.06327,  0.71    ,  0.07825],
...    ... [ 530.      ,  0.1655 ,  0.862   ,  0.04216],
...    ... [ 540.      ,  0.2904 ,  0.954   ,  0.0203 ],
...    ... [ 550.      ,  0.43345,  0.99495,  0.00875],
...    ... [ 560.      ,  0.5945 ,  0.995   ,  0.0039 ]],
... labels=[...'x_bar', ...'y_bar', ...'z_bar'],
... interpolator=SpragueInterpolator,
... interpolator_args={},
... extrapolator=Extrapolator,
... extrapolator_args={...})
```

Instantiating a spectral power distribution with a non-uniformly spaced independent variable:

```
>>> data[511] = (0.00314, 0.31416, 0.03142)
>>> with numpy_print_options(suppress=True):
...     MultiSpectralPowerDistribution(data, labels=labels)
...
MultiSpectral...([[ 500.      ,  0.0049 ,  0.323   ,  0.272   ],
...    ... [ 510.      ,  0.0093 ,  0.503   ,  0.1582 ],
...    ... [ 511.      ,  0.00314,  0.31416,  0.03142],
...    ... [ 520.      ,  0.06327,  0.71    ,  0.07825],
...    ... [ 530.      ,  0.1655 ,  0.862   ,  0.04216],
...    ... [ 540.      ,  0.2904 ,  0.954   ,  0.0203 ],
...    ... [ 550.      ,  0.43345,  0.99495,  0.00875],
...    ... [ 560.      ,  0.5945 ,  0.995   ,  0.0039 ]],
... labels=[...'x_bar', ...'y_bar', ...'z_bar'],
... interpolator=CubicSplineInterpolator,
... interpolator_args={},
... extrapolator=Extrapolator,
... extrapolator_args={...})
```

`__init__(data=None, domain=None, labels=None, **kwargs)`

## Methods

---

`__init__([data, domain, labels])`

Continued on next page

Table 3.57 – continued from previous page

<code>align(shape[, interpolator, ...])</code>	Aligns the multi-spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.
<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the multi-spectral power distribution in-place accordingly to <i>CIE 15:2004</i> and <i>CIE 167:2005</i> recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <i>x</i> variable and corresponding range <i>y</i> variable using given method.
<code>get()</code>	
<code>interpolate(shape[, interpolator, ...])</code>	Interpolates the multi-spectral power distribution in-place accordingly to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
<code>is_uniform()</code>	Returns if independent domain <i>x</i> variable is uniform.
<code>multi_signal_unpack_data([data, domain, ...])</code>	Unpack given data for multi-continuous signal instantiation.
<code>normalise([factor])</code>	Normalises the multi-spectral power distribution with given normalization factor.
<code>to_dataframe()</code>	Converts the continuous signal to a <i>Pandas DataFrame</i> class instance.
<code>trim(shape)</code>	Trims the multi-spectral power distribution wavelengths to given shape.
<code>trim_wavelengths(shape)</code>	
<code>zeros()</code>	

## colour.SpectralShape

`class colour.SpectralShape(start=None, end=None, interval=None)`

Defines the base object for spectral power distribution shape.

### Parameters

- `start` (numeric, optional) – Wavelength  $\lambda_i$  range start in nm.
- `end` (numeric, optional) – Wavelength  $\lambda_i$  range end in nm.
- `interval` (numeric, optional) – Wavelength  $\lambda_i$  range interval.

`start`

`end`

`interval`

`boundaries`

```
__str__()
__repr__()
__iter__()
__contains__()
__len__()
__eq__()
__ne__()
range()
```

## Examples

```
>>> SpectralShape(360, 830, 1)
SpectralShape(360, 830, 1)
```

```
__init__(start=None, end=None, interval=None)
```

## Methods

---

<code>__init__([start, end, interval])</code>	
<code>range([dtype])</code>	Returns an iterable range for the spectral shape.

---

## colour.DEFAULT\_SPECTRAL\_SHAPE

```
colour.DEFAULT_SPECTRAL_SHAPE = SpectralShape(360, 780, 1)
Default spectral shape according to ASTM E308-15 practise shape.

DEFAULT_SPECTRAL_SHAPE : SpectralShape
```

## colour.ASTM30815\_PRACTISE\_SHAPE

```
colour.ASTM30815_PRACTISE_SHAPE = SpectralShape(360, 780, 1)
Shape for *ASTM E308-15 practise* – (360, 780, 1).
```

## References

- [\[ASTMInternational15\]](#)

```
ASTM30815_PRACTISE_SHAPE : SpectralShape
```

## Spectral Data Generation

```
colour
```

<code>blackbody_spd(temperature[, shape, c1, c2, n])</code>	Returns the spectral power distribution of the planckian radiator for given temperature $T[K]$ .
<code>CIE_standard_illuminant_A_function(wl)</code>	<i>CIE Standard Illuminant A</i> is intended to represent typical, domestic,
<code>D_illuminant_relative_spd(xy)</code>	Returns the relative spectral power distribution of given <i>CIE Standard Illuminant D Series</i> using given $xy$ chromaticity coordinates.
<code>constant_spd(k[, shape, dtype])</code>	Returns a spectral power distribution of given spectral shape filled with constant $k$ values.
<code>ones_spd([shape])</code>	Returns a spectral power distribution of given spectral shape filled with ones.
<code>zeros_spd([shape])</code>	Returns a spectral power distribution of given spectral shape filled with zeros.

## colour.blackbody\_spd

```
colour.blackbody_spd(temperature,    shape=SpectralShape(360,    780,    1),    c1=3.741771e-16,
                     c2=0.014388, n=1)
```

Returns the spectral power distribution of the planckian radiator for given temperature  $T[K]$ .

### Parameters

- `temperature` (numeric) – Temperature  $T[K]$  in kelvin degrees.
- `shape` (`SpectralShape`, optional) – Spectral shape used to create the spectral power distribution of the planckian radiator.
- `c1` (numeric, optional) – The official value of  $c1$  is provided by the Committee on Data for Science and Technology (CODATA) and is  $c1 = 3,741771 \times 10.16 \text{ W/m}_2$  (*Mohr and Taylor, 2000*).
- `c2` (numeric, optional) – Since  $T$  is measured on the International Temperature Scale, the value of  $c2$  used in colorimetry should follow that adopted in the current International Temperature Scale (ITS-90) (*Preston-Thomas, 1990; Mielenz et al., 1991*), namely  $c2 = 1,4388 \times 10.2 \text{ m/K}$ .
- `n` (numeric, optional) – Medium index of refraction. For dry air at 15C and 101 325 Pa, containing 0,03 percent by volume of carbon dioxide, it is approximately 1,00028 throughout the visible region although *CIE 15:2004* recommends using  $n = 1$ .

**Returns** Blackbody spectral power distribution.

**Return type** `SpectralPowerDistribution`

### Examples

```
>>> from colour import STANDARD_OBSERVERS_CMFS
>>> from colour.utilities import numpy_print_options
>>> cmfs = STANDARD_OBSERVERS_CMFS['CIE 1931 2 Degree Standard Observer']
>>> with numpy_print_options(suppress=True):
...     blackbody_spd(5000, cmfs.shape)
SpectralPowerDistribution([[ 3.6000000e+02,   6.65427827e+12],
                           [ 3.6100000e+02,   6.70960528e+12],
                           [ 3.6200000e+02,   6.76482512e+12],
                           [ 3.6300000e+02,   6.81993308e+12],
```

[ 3.64000000e+02,	6.87492449e+12],
[ 3.65000000e+02,	6.92979475e+12],
[ 3.66000000e+02,	6.98453932e+12],
[ 3.67000000e+02,	7.03915372e+12],
[ 3.68000000e+02,	7.09363351e+12],
[ 3.69000000e+02,	7.14797433e+12],
[ 3.70000000e+02,	7.20217187e+12],
[ 3.71000000e+02,	7.25622190e+12],
[ 3.72000000e+02,	7.31012021e+12],
[ 3.73000000e+02,	7.36386268e+12],
[ 3.74000000e+02,	7.41744525e+12],
[ 3.75000000e+02,	7.47086391e+12],
[ 3.76000000e+02,	7.52411471e+12],
[ 3.77000000e+02,	7.57719377e+12],
[ 3.78000000e+02,	7.63009726e+12],
[ 3.79000000e+02,	7.68282141e+12],
[ 3.80000000e+02,	7.73536252e+12],
[ 3.81000000e+02,	7.78771695e+12],
[ 3.82000000e+02,	7.83988111e+12],
[ 3.83000000e+02,	7.89185148e+12],
[ 3.84000000e+02,	7.94362458e+12],
[ 3.85000000e+02,	7.99519703e+12],
[ 3.86000000e+02,	8.04656547e+12],
[ 3.87000000e+02,	8.09772662e+12],
[ 3.88000000e+02,	8.14867726e+12],
[ 3.89000000e+02,	8.19941421e+12],
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[ 7.7000000e+02,	1.07380198e+13],
[ 7.7100000e+02,	1.07216721e+13],
[ 7.7200000e+02,	1.07053078e+13],
[ 7.7300000e+02,	1.06889276e+13],
[ 7.7400000e+02,	1.06725317e+13],
[ 7.7500000e+02,	1.06561206e+13],
[ 7.7600000e+02,	1.06396947e+13],
[ 7.7700000e+02,	1.06232545e+13],
[ 7.7800000e+02,	1.06068004e+13],
[ 7.7900000e+02,	1.05903327e+13],
[ 7.8000000e+02,	1.05738520e+13],
[ 7.8100000e+02,	1.05573585e+13],
[ 7.8200000e+02,	1.05408527e+13],
[ 7.8300000e+02,	1.05243351e+13],
[ 7.8400000e+02,	1.05078060e+13],
[ 7.8500000e+02,	1.04912657e+13],
[ 7.8600000e+02,	1.04747147e+13],
[ 7.8700000e+02,	1.04581535e+13],
[ 7.8800000e+02,	1.04415822e+13],
[ 7.8900000e+02,	1.04250014e+13],
[ 7.9000000e+02,	1.04084115e+13],
[ 7.9100000e+02,	1.03918127e+13],
[ 7.9200000e+02,	1.03752055e+13],
[ 7.9300000e+02,	1.03585902e+13],
[ 7.9400000e+02,	1.03419672e+13],
[ 7.9500000e+02,	1.03253369e+13],
[ 7.9600000e+02,	1.03086995e+13],
[ 7.9700000e+02,	1.02920556e+13],
[ 7.9800000e+02,	1.02754053e+13],
[ 7.9900000e+02,	1.02587492e+13],
[ 8.0000000e+02,	1.02420874e+13],
[ 8.0100000e+02,	1.02254204e+13],
[ 8.0200000e+02,	1.02087485e+13],
[ 8.0300000e+02,	1.01920720e+13],
[ 8.0400000e+02,	1.01753913e+13],
[ 8.0500000e+02,	1.01587067e+13],
[ 8.0600000e+02,	1.01420186e+13],
[ 8.0700000e+02,	1.01253271e+13],
[ 8.0800000e+02,	1.01086328e+13],
[ 8.0900000e+02,	1.00919358e+13],
[ 8.1000000e+02,	1.00752366e+13],
[ 8.1100000e+02,	1.00585353e+13],
[ 8.1200000e+02,	1.00418324e+13],
[ 8.1300000e+02,	1.00251282e+13],
[ 8.1400000e+02,	1.00084228e+13],
[ 8.1500000e+02,	9.99171677e+12],
[ 8.1600000e+02,	9.97501023e+12],
[ 8.1700000e+02,	9.95830354e+12],
[ 8.1800000e+02,	9.94159699e+12],
[ 8.1900000e+02,	9.92489086e+12],
[ 8.2000000e+02,	9.90818544e+12],
[ 8.2100000e+02,	9.89148102e+12],
[ 8.2200000e+02,	9.87477789e+12],
[ 8.2300000e+02,	9.85807631e+12],
[ 8.2400000e+02,	9.84137658e+12],
[ 8.2500000e+02,	9.82467896e+12],
[ 8.2600000e+02,	9.80798374e+12],
[ 8.2700000e+02,	9.79129116e+12],

```
[ 8.2800000e+02,   9.77460152e+12],
[ 8.2900000e+02,   9.75791506e+12],
[ 8.3000000e+02,   9.74123205e+12]],  
interpolator=SpragueInterpolator,  
interpolator_args={},  
extrapolator=Extrapolator,  
extrapolator_args={...})
```

## colour.CIE\_standard\_illuminant\_A\_function

`colour.CIE_standard_illuminant_A_function(wl)`

*CIE Standard Illuminant A* is intended to represent typical, domestic, tungsten-filament lighting.

Its relative spectral power distribution is that of a Planckian radiator at a temperature of approximately 2856 K. *CIE Standard Illuminant A* should be used in all applications of colorimetry involving the use of incandescent lighting, unless there are specific reasons for using a different illuminant.

**Parameters** `wl` (`array_like`) – Wavelength to evaluate the function at.

**Returns** *CIE Standard Illuminant A* value at given wavelength.

**Return type** `ndarray`

## References

- [\[CIET14804a\]](#)

## Examples

```
>>> wl = np.array([560, 580, 581.5])
>>> CIE_standard_illuminant_A_function(wl)
array([ 100.          ,  114.4363383...,  115.5285063...])
```

## colour.D\_illuminant\_relative\_spd

`colour.D_illuminant_relative_spd(xy)`

Returns the relative spectral power distribution of given *CIE Standard Illuminant D Series* using given `xy` chromaticity coordinates.

## References

- [\[Lin07\]](#)
- [\[WS00b\]](#)

**Parameters** `xy` (`array_like`) – `xy` chromaticity coordinates.

**Returns** *CIE Standard Illuminant D Series* relative spectral power distribution.

**Return type** `SpectralPowerDistribution`

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> xy = np.array([0.34570, 0.35850])
>>> with numpy_print_options(suppress=True):
...     D_illuminant_relative_spd(xy)
SpectralPowerDistribution([[ 300.        ,  0.0193039...],
 [ 310.        ,  2.1265303...],
 [ 320.        ,  7.9867359...],
 [ 330.        , 15.1666959...],
 [ 340.        , 18.3413202...],
 [ 350.        , 21.3757973...],
 [ 360.        , 24.2528862...],
 [ 370.        , 26.2782171...],
 [ 380.        , 24.7348842...],
 [ 390.        , 30.0518667...],
 [ 400.        , 49.458942 ...],
 [ 410.        , 56.6929605...],
 [ 420.        , 60.1981682...],
 [ 430.        , 57.9390276...],
 [ 440.        , 74.9047554...],
 [ 450.        , 87.3151258...],
 [ 460.        , 90.6691236...],
 [ 470.        , 91.4109985...],
 [ 480.        , 95.1362798...],
 [ 490.        , 91.9956940...],
 [ 500.        , 95.7488852...],
 [ 510.        , 96.6315995...],
 [ 520.        , 97.1308377...],
 [ 530.        , 102.0961518...],
 [ 540.        , 100.7580555...],
 [ 550.        , 102.3164095...],
 [ 560.        , 100.        ...],
 [ 570.        , 97.7339937...],
 [ 580.        , 98.9175842...],
 [ 590.        , 93.5440898...],
 [ 600.        , 97.7548532...],
 [ 610.        , 99.3559831...],
 [ 620.        , 99.1396431...],
 [ 630.        , 95.8275899...],
 [ 640.        , 99.0028159...],
 [ 650.        , 95.8307955...],
 [ 660.        , 98.3850717...],
 [ 670.        , 103.2245516...],
 [ 680.        , 99.3672578...],
 [ 690.        , 87.5676019...],
 [ 700.        , 91.8218781...],
 [ 710.        , 93.0772354...],
 [ 720.        , 77.0098456...],
 [ 730.        , 86.6795856...],
 [ 740.        , 92.7570922...],
 [ 750.        , 78.3784557...],
 [ 760.        , 57.8075859...],
 [ 770.        , 83.0873522...],
 [ 780.        , 78.4245724...],
 [ 790.        , 79.7098456...],
 [ 800.        , 73.5435857...],
 [ 810.        , 64.0424558...],
```

```
[ 820.        ,  70.9121958...],
[ 830.        ,  74.5862223...]],  
interpolator=SpragueInterpolator,  
interpolator_args={},  
extrapolator=Extrapolator,  
extrapolator_args={...})
```

## colour.constant\_spd

colour.constant\_spd(*k*, *shape*=SpectralShape(360, 780, 1), *dtype*=<type 'numpy.float64'>)

Returns a spectral power distribution of given spectral shape filled with constant *k* values.

### Parameters

- ***k* (numeric)** – Constant *k* to fill the spectral power distribution with.
- ***shape* (SpectralShape, optional)** – Spectral shape used to create the spectral power distribution.
- ***dtype* (type)** – Data type used for the spectral power distribution.

**Returns** Constant *k* to filled spectral power distribution.

**Return type** SpectralPowerDistribution

### Notes

- By default, the spectral power distribution will use the shape given by colour.DEFAULT\_SPECTRAL\_SHAPE attribute.

### Examples

```
>>> spd = constant_spd(100)
>>> spd.shape
SpectralShape(360.0, 780.0, 1.0)
>>> spd[400]
100.0
```

## colour.ones\_spd

colour.ones\_spd(*shape*=SpectralShape(360, 780, 1))

Returns a spectral power distribution of given spectral shape filled with ones.

**Parameters** ***shape* (SpectralShape, optional)** – Spectral shape used to create the spectral power distribution.

**Returns** Ones filled spectral power distribution.

**Return type** SpectralPowerDistribution

## Notes

- By default, the spectral power distribution will use the shape given by `colour.DEFAULT_SPECTRAL_SHAPE` attribute.

## Examples

```
>>> spd = ones_spd()
>>> spd.shape
SpectralShape(360.0, 780.0, 1.0)
>>> spd[400]
1.0
```

## `colour.zeros_spd`

`colour.zeros_spd(shape=SpectralShape(360, 780, 1))`

Returns a spectral power distribution of given spectral shape filled with zeros.

**Parameters** `shape` (`SpectralShape`, optional) – Spectral shape used to create the spectral power distribution.

**Returns** Zeros filled spectral power distribution.

**Return type** `SpectralPowerDistribution`

## Notes

- By default, the spectral power distribution will use the shape given by `colour.DEFAULT_SPECTRAL_SHAPE` attribute.

## Examples

```
>>> spd = zeros_spd()
>>> spd.shape
SpectralShape(360.0, 780.0, 1.0)
>>> spd[400]
0.0
```

## `colour.colorimetry`

`blackbody_spectral_radiance(wavelength, ...)`

Returns the spectral radiance of a blackbody at thermodynamic temperature  $T[K]$  in a medium having index of refraction  $n$ .

`planck_law(wavelength, temperature[, c1, c2, n])`

Returns the spectral radiance of a blackbody at thermodynamic temperature  $T[K]$  in a medium having index of refraction  $n$ .

## colour.colorimetry.blackbody\_spectral\_radiance

```
colour.colorimetry.blackbody_spectral_radiance(wavelength, temperature, c1=3.741771e-16,
                                                c2=0.014388, n=1)
```

Returns the spectral radiance of a blackbody at thermodynamic temperature  $T[K]$  in a medium having index of refraction  $n$ .

### Parameters

- **wavelength** (numeric or array\_like) – Wavelength in meters.
- **temperature** (numeric or array\_like) – Temperature  $T[K]$  in kelvin degrees.
- **c1** (numeric or array\_like, optional) – The official value of  $c1$  is provided by the Committee on Data for Science and Technology (CODATA) and is  $c1 = 3,741771 \times 10.16 \text{ W/m}_2$  (*Mohr and Taylor, 2000*).
- **c2** (numeric or array\_like, optional) – Since  $T$  is measured on the International Temperature Scale, the value of  $c2$  used in colorimetry should follow that adopted in the current International Temperature Scale (ITS-90) (*Preston-Thomas, 1990; Mielcz et al., 1991*), namely  $c2 = 1,4388 \times 10.2 \text{ m/K}$ .
- **n** (numeric or array\_like, optional) – Medium index of refraction. For dry air at 15C and 101 325 Pa, containing 0,03 percent by volume of carbon dioxide, it is approximately 1,00028 throughout the visible region although *CIE 15:2004* recommends using  $n = 1$ .

**Returns** Radiance in *watts per steradian per square metre*.

**Return type** numeric or ndarray

### Notes

- The following form implementation is expressed in term of wavelength.
- The SI unit of radiance is *watts per steradian per square metre*.

### References

- [CIET14804c]

### Examples

```
>>> # Doctests ellipsis for Python 2.x compatibility.  
>>> planck_law(500 * 1e-9, 5500)  
20472701909806.5...
```

## colour.colorimetry.planck\_law

```
colour.colorimetry.planck_law(wavelength, temperature, c1=3.741771e-16, c2=0.014388, n=1)
```

Returns the spectral radiance of a blackbody at thermodynamic temperature  $T[K]$  in a medium having index of refraction  $n$ .

### Parameters

- **wavelength** (numeric or array\_like) – Wavelength in meters.
- **temperature** (numeric or array\_like) – Temperature  $T[K]$  in kelvin degrees.
- **c1** (numeric or array\_like, optional) – The official value of  $c1$  is provided by the Committee on Data for Science and Technology (CODATA) and is  $c1 = 3,741771 \times 10.16 \text{ W/m}_2$  (*Mohr and Taylor, 2000*).
- **c2** (numeric or array\_like, optional) – Since  $T$  is measured on the International Temperature Scale, the value of  $c2$  used in colorimetry should follow that adopted in the current International Temperature Scale (ITS-90) (*Preston-Thomas, 1990; Mielcz et al., 1991*), namely  $c2 = 1,4388 \times 10.2 \text{ m/K}$ .
- **n** (numeric or array\_like, optional) – Medium index of refraction. For dry air at 15C and 101 325 Pa, containing 0,03 percent by volume of carbon dioxide, it is approximately 1,00028 throughout the visible region although *CIE 15:2004* recommends using  $n = 1$ .

**Returns** Radiance in *watts per steradian per square metre*.

**Return type** numeric or ndarray

## Notes

- The following form implementation is expressed in term of wavelength.
- The SI unit of radiance is *watts per steradian per square metre*.

## References

- [CIET14804c]

## Examples

```
>>> # Doctests ellipsis for Python 2.x compatibility.
>>> planck_law(500 * 1e-9, 5500)
20472701909806.5...
```

## Conversion to Tristimulus Values

colour

spectral_to_XYZ(spd[, cmfs, illuminant, method])	Converts given spectral power distribution to <i>CIE XYZ</i> tristimulus values using given colour matching functions, illuminant and method.
SPECTRAL_TO_XYZ_METHODS	Supported spectral power distribution to <i>CIE XYZ</i> tristimulus values
wavelength_to_XYZ(wavelength[, cmfs])	Converts given wavelength $\lambda$ to <i>CIE XYZ</i> tristimulus values using given colour matching functions.

## colour.spectral\_to\_XYZ

```
colour.spectral_to_XYZ(spd, cmfs=XYZ_ColourMatchingFunctions(name='CIE 1931 2 Degree Standard Observer', ...), illuminant=SpectralPowerDistribution(name='1 Constant', ...), method=u'ASTM E308-15', **kwargs)
```

Converts given spectral power distribution to CIE XYZ tristimulus values using given colour matching functions, illuminant and method.

### Parameters

- **spd** (`SpectralPowerDistribution`) – Spectral power distribution.
- **cmfs** (`XYZ_ColourMatchingFunctions`) – Standard observer colour matching functions.
- **illuminant** (`SpectralPowerDistribution`, optional) – Illuminant spectral power distribution.
- **method** (unicode, optional) – {‘ASTM E308-15’, ‘Integration’}, Computation method.

### Other Parameters

- **use\_practice\_range** (bool, optional) – {`colour.colorimetry.spectral_to_XYZ_ASTME30815()`}, Practise ASTM E308-15 working wavelengths range is [360, 780], if *True* this argument will trim the colour matching functions appropriately.
- **mi\_5nm\_omission\_method** (bool, optional) – {`colour.colorimetry.spectral_to_XYZ_ASTME30815()`}, 5 nm measurement intervals spectral power distribution conversion to tristimulus values will use a 5 nm version of the colour matching functions instead of a table of tristimulus weighting factors.
- **mi\_20nm\_interpolation\_method** (bool, optional) – {`colour.colorimetry.spectral_to_XYZ_ASTME30815()`}, 20 nm measurement intervals spectral power distribution conversion to tristimulus values will use a dedicated interpolation method instead of a table of tristimulus weighting factors.

**Returns** CIE XYZ tristimulus values.

**Return type** ndarray, (3,)

**Warning:** The output range of that definition is non standard!

### Notes

- Output CIE XYZ tristimulus values are in range [0, 100].

### References

- [\[ASTMInternational11\]](#)
- [\[ASTMInternational15\]](#)
- [\[WS00d\]](#)

## Examples

```
>>> from colour import (
...     CMFS, ILLUMINANTS_RELATIVE_SPDS, SpectralPowerDistribution)
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> data = {
...     400: 0.0641,
...     420: 0.0645,
...     440: 0.0562,
...     460: 0.0537,
...     480: 0.0559,
...     500: 0.0651,
...     520: 0.0705,
...     540: 0.0772,
...     560: 0.0870,
...     580: 0.1128,
...     600: 0.1360,
...     620: 0.1511,
...     640: 0.1688,
...     660: 0.1996,
...     680: 0.2397,
...     700: 0.2852
... }
>>> spd = SpectralPowerDistribution(data)
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['D50']
>>> spectral_to_XYZ(spd, cmfs, illuminant)
...
array([ 11.5290265...,  9.9502091...,  4.7098882...])
>>> spectral_to_XYZ(spd, cmfs, illuminant, use_practice_range=False)
...
array([ 11.5291275...,  9.9502369...,  4.7098811...])
>>> spectral_to_XYZ(spd, cmfs, illuminant, method='Integration')
...
array([ 11.5296285...,  9.9499467...,  4.7066079...])
```

## colour.SPECTRAL\_TO\_XYZ\_METHODS

colour.SPECTRAL\_TO\_XYZ\_METHODS = CaseInsensitiveMapping({u'ASTM E308-15': ..., u'Integration': ..., u'astm2015': ...})  
Supported spectral power distribution to *CIE XYZ* tristimulus values conversion methods

## References

- [\[ASTMInternational11\]](#)
- [\[ASTMInternational15\]](#)
- [\[WSO0d\]](#)

**SPECTRAL\_TO\_XYZ\_METHODS** [CaseInsensitiveMapping] {‘ASTM E308-15’, ‘Integration’}

Aliases:

- ‘astm2015’: ‘ASTM E308-15’

## colour.wavelength\_to\_XYZ

```
colour.wavelength_to_XYZ(wavelength, cmfs=XYZ_ColourMatchingFunctions(name='CIE 1931 2 Degree Standard Observer', ...))
```

Converts given wavelength  $\lambda$  to *CIE XYZ* tristimulus values using given colour matching functions.

If the wavelength  $\lambda$  is not available in the colour matching function, its value will be calculated according to *CIE 15:2004* recommendation: the method developed by *Sprague (1880)* will be used for interpolating functions having a uniformly spaced independent variable and the *Cubic Spline* method for non-uniformly spaced independent variable.

### Parameters

- **wavelength** (numeric or array\_like) – Wavelength  $\lambda$  in nm.
- **cmfs** (XYZ\_ColourMatchingFunctions, optional) – Standard observer colour matching functions.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

**Raises** ValueError – If wavelength  $\lambda$  is not contained in the colour matching functions domain.

### Notes

- Output *CIE XYZ* tristimulus values are in range [0, 1].

### Examples

```
>>> from colour import CMFS
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> wavelength_to_XYZ(480, cmfs)
array([ 0.09564,  0.13902,  0.812950...])
>>> wavelength_to_XYZ(480.5, cmfs)
array([ 0.0914287...,  0.1418350...,  0.7915726...])
```

## ASTM E308-15

### colour.colorimetry

---

```
spectral_to_XYZ_ASTME30815(spd[, cmfs, ...])
```

Converts given spectral power distribution to *CIE XYZ* tristimulus values using given colour matching functions and illuminant according to practise *ASTM E308-15* method.

---

## colour.colorimetry.spectral\_to\_XYZ\_ASTME30815

```
colour.colorimetry.spectral_to_XYZ_ASTME30815(spd, cmfs=XYZ_ColourMatchingFunctions(name='CIE
1931 2 Degree Standard Observer', ...), illu-
minant=SpectralPowerDistribution(name='I
Constant', ...), use_practice_range=True,
mi_5nm_omission_method=True,
mi_20nm_interpolation_method=True)
```

Converts given spectral power distribution to *CIE XYZ* tristimulus values using given colour matching functions and illuminant according to practise *ASTM E308-15* method.

### Parameters

- **spd** (`SpectralPowerDistribution`) – Spectral power distribution.
- **cmfs** (`XYZ_ColourMatchingFunctions`) – Standard observer colour matching functions.
- **illuminant** (`SpectralPowerDistribution`, optional) – Illuminant spectral power distribution.
- **use\_practice\_range** (`bool`, optional) – Practise *ASTM E308-15* working wavelengths range is [360, 780], if *True* this argument will trim the colour matching functions appropriately.
- **mi\_5nm\_omission\_method** (`bool`, optional) – 5 nm measurement intervals spectral power distribution conversion to tristimulus values will use a 5 nm version of the colour matching functions instead of a table of tristimulus weighting factors.
- **mi\_20nm\_interpolation\_method** (`bool`, optional) – 20 nm measurement intervals spectral power distribution conversion to tristimulus values will use a dedicated interpolation method instead of a table of tristimulus weighting factors.

**Returns** *CIE XYZ* tristimulus values.

**Return type** `ndarray(3,)`

### Warning:

- The tables of tristimulus weighting factors are cached in `colour.colorimetry.tristimulus._TRISTIMULUS_WEIGHTING_FACTORS_CACHE` attribute. Their identifier key is defined by the colour matching functions and illuminant names along the current shape such as: *CIE 1964 10 Degree Standard Observer, A, (360.0, 830.0, 10.0)* Considering the above, one should be mindful that using similar colour matching functions and illuminant names but with different spectral data will lead to unexpected behaviour.
- The output range of that definition is non standard!

### Notes

- Output *CIE XYZ* tristimulus values are in range [0, 100].

### References

- *[ASTMInternational15]*

## Examples

```
>>> from colour import (
...     CMFS, ILLUMINANTS_RELATIVE_SPDS, SpectralPowerDistribution)
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> data = {
...     400: 0.0641,
...     420: 0.0645,
...     440: 0.0562,
...     460: 0.0537,
...     480: 0.0559,
...     500: 0.0651,
...     520: 0.0705,
...     540: 0.0772,
...     560: 0.0870,
...     580: 0.1128,
...     600: 0.1360,
...     620: 0.1511,
...     640: 0.1688,
...     660: 0.1996,
...     680: 0.2397,
...     700: 0.2852
... }
>>> spd = SpectralPowerDistribution(data)
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['D50']
>>> spectral_to_XYZ_ASTME30815(spd, cmfs, illuminant)
...
array([ 11.5290265...,   9.9502091...,   4.7098882...])
```

## Ancillary Objects

colour.colorimetry

`spectral_to_XYZ_tristimulus_weighting_factors_ASTME30815` Converts given spectral power distribution to *CIE XYZ* tristimulus values using given colour matching functions and illuminant using a table of tristimulus weighting factors according to practise *ASTM E308-15* method.

`adjust_tristimulus_weighting_factors_ASTME30815` Adjusts given table of tristimulus weighting factors to account for a shorter wavelengths range of the test spectral shape compared to the reference spectral shape using practise *ASTM E308-15* method: Weights at the wavelengths for which data are not available are added to the weights at the shortest and longest wavelength for which spectral data are available.

`lagrange_coefficients_ASTME202211([...])` Computes the *Lagrange Coefficients* for given interval size using practise *ASTM E2022-11* method.

`tristimulus_weighting_factors_ASTME202211(...)` Returns a table of tristimulus weighting factors for given colour matching functions and illuminant using practise *ASTM E2022-11* method.

**colour.colorimetry.spectral\_to\_XYZ\_tristimulus\_weighting\_factors\_ASTME30815**

```
colour.colorimetry.spectral_to_XYZ_tristimulus_weighting_factors_ASTME30815(spds,
    cmfs=XYZ_ColourMatchingFunctions(
        1931 2 Degree
        Standard
        Observer',
        ...), illuminant=SpectralPowerDistribution(name='Constant', ...))
```

Converts given spectral power distribution to *CIE XYZ* tristimulus values using given colour matching functions and illuminant using a table of tristimulus weighting factors according to practise *ASTM E308-15* method.

**Parameters**

- **spds** (`SpectralPowerDistribution`) – Spectral power distribution.
- **cmfs** (`XYZ_ColourMatchingFunctions`) – Standard observer colour matching functions.
- **illuminant** (`SpectralPowerDistribution`, optional) – Illuminant spectral power distribution.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray, (3,)

**Warning:** The output range of that definition is non standard!

**Notes**

- Output *CIE XYZ* tristimulus values are in range [0, 100].

**References**

- [\[ASTMInternational15\]](#)

**Examples**

```
>>> from colour import (
...     CMFS, ILLUMINANTS_RELATIVE_SPDS, SpectralPowerDistribution)
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> data = {
...     400: 0.0641,
...     420: 0.0645,
...     440: 0.0562,
...     460: 0.0537,
...     480: 0.0559,
...     500: 0.0651,
...     520: 0.0705,
...     540: 0.0772,
...     560: 0.0870,
```

```
...     580: 0.1128,
...     600: 0.1360,
...     620: 0.1511,
...     640: 0.1688,
...     660: 0.1996,
...     680: 0.2397,
...     700: 0.2852
...
... }
>>> spd = SpectralPowerDistribution(data)
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['D50']
>>> spectral_to_XYZ_tristimulus_weighting_factors_ASTME30815(
...     spd, cmfs, illuminant)
array([ 11.5296311...,   9.9505845...,   4.7098037...])
```

## colour.colorimetry.adjust\_tristimulus\_weighting\_factors\_ASTME30815

colour.colorimetry.adjust\_tristimulus\_weighting\_factors\_ASTME30815(*W, shape\_r, shape\_t*)

Adjusts given table of tristimulus weighting factors to account for a shorter wavelengths range of the test spectral shape compared to the reference spectral shape using practise *ASTM E308-15* method: Weights at the wavelengths for which data are not available are added to the weights at the shortest and longest wavelength for which spectral data are available.

### Parameters

- *W* (`array_like`) – Tristimulus weighting factors table.
- *shape\_r* (`SpectralShape`) – Reference spectral shape.
- *shape\_t* (`SpectralShape`) – Test spectral shape.

**Returns** Adjusted tristimulus weighting factors.

**Return type** ndarray

## References

- [ASTMInternational15]

## Examples

```
>>> from colour import (CMFS, CIE_standard_illuminant_A_function,
...     SpectralPowerDistribution, SpectralShape)
>>> from colour.utilities import numpy_print_options
>>> cmfs = CMFS['CIE 1964 10 Degree Standard Observer']
>>> wl = cmfs.shape.range()
>>> A = SpectralPowerDistribution(
...     dict(zip(wl, CIE_standard_illuminant_A_function(wl))),
...     name='A (360, 830, 1)')
>>> W = tristimulus_weighting_factors_ASTME202211(
...     cmfs, A, SpectralShape(360, 830, 20))
>>> with numpy_print_options(suppress=True):
...     adjust_tristimulus_weighting_factors_ASTME30815(
...         W, SpectralShape(360, 830, 20), SpectralShape(400, 700, 20))
...
array([[ 0.0509543...,   0.0040971...,   0.2144280...],
```

```
[ 0.7734225...,  0.0779839...,  3.6965732...],
[ 1.9000905...,  0.3037005...,  9.7554195...],
[ 1.9707727...,  0.8552809..., 11.4867325...],
[ 0.7183623...,  2.1457000...,  6.7845806...],
[ 0.0426667...,  4.8985328...,  2.3208000...],
[ 1.5223302...,  9.6471138...,  0.7430671...],
[ 5.6770329..., 14.4609708...,  0.1958194...],
[12.4451744..., 17.4742541...,  0.0051827...],
[20.5535772..., 17.5838219..., -0.0026512...],
[25.3315384..., 14.8957035...,  0.        ...],
[21.5711570..., 10.0796619...,  0.        ...],
[12.1785817...,  5.0680655...,  0.        ...],
[ 4.6675746...,  1.8303239...,  0.        ...],
[ 1.3236117...,  0.5129694...,  0.        ...],
[ 0.4171109...,  0.1618194...,  0.        ...]])
```

## colour.colorimetry.lagrange\_coefficients\_ASTME202211

`colour.colorimetry.lagrange_coefficients_ASTME202211(interval=10, interval_type=u'inner')`

Computes the *Lagrange Coefficients* for given interval size using practise *ASTM E2022-11* method.

### Parameters

- `interval` (`int`) – Interval size in nm.
- `interval_type` (`unicode`, optional) – {‘inner’, ‘boundary’}, If the interval is an *inner* interval *Lagrange Coefficients* are computed for degree 4. Degree 3 is used for a *boundary* interval.

`Returns` *Lagrange Coefficients*.

`Return type` `ndarray`

### References

- [\[ASTMInternational11\]](#)

### Examples

```
>>> lagrange_coefficients_ASTME202211(10, 'inner')
...
array([[ -0.028...,  0.940...,  0.104..., -0.016...],
       [-0.048...,  0.864...,  0.216..., -0.032...],
       [-0.059...,  0.773...,  0.331..., -0.045...],
       [-0.064...,  0.672...,  0.448..., -0.056...],
       [-0.062...,  0.562...,  0.562..., -0.062...],
       [-0.056...,  0.448...,  0.672..., -0.064...],
       [-0.045...,  0.331...,  0.773..., -0.059...],
       [-0.032...,  0.216...,  0.864..., -0.048...],
       [-0.016...,  0.104...,  0.940..., -0.028...]])
>>> lagrange_coefficients_ASTME202211(10, 'boundary')
...
array([[ 0.85...,  0.19..., -0.04...],
       [ 0.72...,  0.36..., -0.08...],
```

```
[ 0.59..., 0.51..., -0.10...],  
[ 0.48..., 0.64..., -0.12...],  
[ 0.37..., 0.75..., -0.12...],  
[ 0.28..., 0.84..., -0.12...],  
[ 0.19..., 0.91..., -0.10...],  
[ 0.12..., 0.96..., -0.08...],  
[ 0.05..., 0.99..., -0.04...]])
```

## colour.colorimetry.tristimulus\_weighting\_factors\_ASTME202211

`colour.colorimetry.tristimulus_weighting_factors_ASTME202211(cmfs, illuminant, shape)`

Returns a table of tristimulus weighting factors for given colour matching functions and illuminant using practise *ASTM E2022-11* method.

The computed table of tristimulus weighting factors should be used with spectral data that has been corrected for spectral bandpass dependence.

### Parameters

- `cmfs` ([XYZ\\_ColourMatchingFunctions](#)) – Standard observer colour matching functions.
- `illuminant` ([SpectralPowerDistribution](#)) – Illuminant spectral power distribution.
- `shape` ([SpectralShape](#)) – Shape used to build the table, only the interval is needed.

**Returns** Tristimulus weighting factors table.

**Return type** ndarray

**Raises** `ValueError` – If the colour matching functions or illuminant intervals are not equal to 1 nm.

### Warning:

- The tables of tristimulus weighting factors are cached in `colour.colorimetry.tristimulus._TRISTIMULUS_WEIGHTING_FACTORS_CACHE` attribute. Their identifier key is defined by the colour matching functions and illuminant names along the current shape such as: *CIE 1964 10 Degree Standard Observer; A, (360.0, 830.0, 10.0)* Considering the above, one should be mindful that using similar colour matching functions and illuminant names but with different spectral data will lead to unexpected behaviour.

### Notes

- Input colour matching functions and illuminant intervals are expected to be equal to 1 nm. If the illuminant data is not available at 1 nm interval, it needs to be interpolated using *CIE* recommendations: The method developed by *Sprague (1880)* should be used for interpolating functions having a uniformly spaced independent variable and a *Cubic Spline* method for non-uniformly spaced independent variable.

### References

- [\[ASTMInternational11\]](#)

## Examples

```
>>> from colour import (CMFS, CIE_standard_illuminant_A_function,
...     SpectralPowerDistribution, SpectralShape, numpy_print_options)
>>> cmfs = CMFS['CIE 1964 10 Degree Standard Observer']
>>> wl = cmfs.shape.range()
>>> A = SpectralPowerDistribution(
...     dict(zip(wl, CIE_standard_illuminant_A_function(wl))),
...     name='A (360, 830, 1)')
>>> with numpy_print_options(suppress=True):
...     tristimulus_weighting_factors_ASTME202211(
...         cmfs, A, SpectralShape(360, 830, 20))
...
array([[ -0.0002981..., -0.0000317..., -0.0013301...],
       [-0.0087155..., -0.0008915..., -0.0407436...],
       [ 0.0599679...,  0.0050203...,  0.2565018...],
       [ 0.7734225...,  0.0779839...,  3.6965732...],
       [ 1.9000905...,  0.3037005...,  9.7554195...],
       [ 1.9707727...,  0.8552809..., 11.4867325...],
       [ 0.7183623...,  2.1457000...,  6.7845806...],
       [ 0.0426667...,  4.8985328...,  2.3208000...],
       [ 1.5223302...,  9.6471138...,  0.7430671...],
       [ 5.6770329..., 14.4609708...,  0.1958194...],
       [12.4451744..., 17.4742541...,  0.0051827...],
       [20.5535772..., 17.5838219..., -0.0026512...],
       [25.3315384..., 14.8957035...,  0.        ...],
       [21.5711570..., 10.0796619...,  0.        ...],
       [12.1785817...,  5.0680655...,  0.        ...],
       [ 4.6675746...,  1.8303239...,  0.        ...],
       [ 1.3236117...,  0.5129694...,  0.        ...],
       [ 0.3175325...,  0.1230084...,  0.        ...],
       [ 0.0746341...,  0.0290243...,  0.        ...],
       [ 0.0182990...,  0.0071606...,  0.        ...],
       [ 0.0047942...,  0.0018888...,  0.        ...],
       [ 0.0013293...,  0.0005277...,  0.        ...],
       [ 0.0004254...,  0.0001704...,  0.        ...],
       [ 0.0000962...,  0.0000389...,  0.        ...]])
```

## Integration

`colour.colorimetry`

---

`spectral_to_XYZ_integration(spd[, cmfs, ...])`

Converts given spectral power distribution to *CIE XYZ* tristimulus values using given colour matching functions and illuminant according to classical integration method.

---

## colour.colorimetry.spectral\_to\_XYZ\_integration

```
colour.colorimetry.spectral_to_XYZ_integration(spd, cmfs=XYZ_ColourMatchingFunctions(name='CIE  
1931 2 Degree Standard Observer', ...), illuminant=SpectralPowerDistribution(name='I  
Constant', ...))
```

Converts given spectral power distribution to *CIE XYZ* tristimulus values using given colour matching functions and illuminant according to classical integration method.

### Parameters

- **spd** (`SpectralPowerDistribution`) – Spectral power distribution.
- **cmfs** (`XYZ_ColourMatchingFunctions`) – Standard observer colour matching functions.
- **illuminant** (`SpectralPowerDistribution`, optional) – Illuminant spectral power distribution.

**Returns** *CIE XYZ* tristimulus values.

**Return type** `ndarray` (3,)

**Warning:** The output range of that definition is non standard!

### Notes

- Output *CIE XYZ* tristimulus values are in range [0, 100].

### References

- [\[WS00d\]](#)

### Examples

```
>>> from colour import (  
...     CMFS, ILLUMINANTS_RELATIVE_SPDS, SpectralPowerDistribution)  
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']  
>>> data = {  
...     400: 0.0641,  
...     420: 0.0645,  
...     440: 0.0562,  
...     460: 0.0537,  
...     480: 0.0559,  
...     500: 0.0651,  
...     520: 0.0705,  
...     540: 0.0772,  
...     560: 0.0870,  
...     580: 0.1128,  
...     600: 0.1360,  
...     620: 0.1511,  
...     640: 0.1688,  
...     660: 0.1996,
```

```

...      680: 0.2397,
...      700: 0.2852
...
>>> spd = SpectralPowerDistribution(data)
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['D50']
>>> spectral_to_XYZ_integration(spd, cmfs, illuminant)
...
array([ 11.5296285...,   9.9499467...,   4.7066079...])

```

## Spectral Bandpass Dependence Correction

colour

<code>bandpass_correction(spd[, method])</code>	Implements spectral bandpass dependence correction on given spectral power distribution using given method.
<code>BANDPASS_CORRECTION_METHODS</code>	Supported spectral bandpass dependence correction methods.

### colour.bandpass\_correction

`colour.bandpass_correction(spd, method=u'Stearns 1988')`

Implements spectral bandpass dependence correction on given spectral power distribution using given method.

#### Parameters

- `spd` (`SpectralPowerDistribution`) – Spectral power distribution.
- `method` (unicode, optional) – ('Stearns 1988', ) Correction method.

**Returns** Spectral bandpass dependence corrected spectral power distribution.

**Return type** `SpectralPowerDistribution`

### colour.BANDPASS\_CORRECTION\_METHODS

`colour.BANDPASS_CORRECTION_METHODS = CaseInsensitiveMapping({u'Stearns 1988': ...})`

Supported spectral bandpass dependence correction methods.

`BANDPASS_CORRECTION_METHODS` [CaseInsensitiveMapping] {'Stearns 1988', }

### Stearns and Stearns (1988)

colour.colorimetry

<code>bandpass_correction_Stearns1988(spd)</code>	Implements spectral bandpass dependence correction on given spectral power distribution using <i>Stearns and Stearns (1988)</i> method.
---	---

## colour.colorimetry.bandpass\_correction\_Stearns1988

`colour.colorimetry.bandpass_correction_Stearns1988(spd)`

Implements spectral bandpass dependence correction on given spectral power distribution using *Stearns and Stearns (1988)* method.

**Parameters** `spd` (`SpectralPowerDistribution`) – Spectral power distribution.

**Returns** Spectral bandpass dependence corrected spectral power distribution.

**Return type** `SpectralPowerDistribution`

## References

- [\[SS88\]](#)
- [\[WRC12c\]](#)

## Examples

```
>>> from colour import SpectralPowerDistribution
>>> from colour.utilities import numpy_print_options
>>> data = {
...     500: 0.0651,
...     520: 0.0705,
...     540: 0.0772,
...     560: 0.0870,
...     580: 0.1128,
...     600: 0.1360
... }
>>> with numpy_print_options(suppress=True):
...     bandpass_correction_Stearns1988(
...         SpectralPowerDistribution(data))
...
SpectralPowerDistribution([[ 500.        ,  0.0646518...],
 [ 520.        ,  0.0704293...],
 [ 540.        ,  0.0769485...],
 [ 560.        ,  0.0856928...],
 [ 580.        ,  0.1129644...],
 [ 600.        ,  0.1379256...]],
 interpolator=SpragueInterpolator,
 interpolator_args={},
 extrapolator=Extrapolator,
 extrapolator_args={...})
```

## Colour Matching Functions

`colour.colorimetry`

<code>LMS_ConeFundamentals([data, domain, labels])</code>	Implements support for the Stockman and Sharpe <i>LMS</i> cone fundamentals colour matching functions.
<code>RGB_ColourMatchingFunctions([data, domain, ...])</code>	Implements support for the <i>CIE RGB</i> colour matching functions.

Continued on next page

Table 3.67 – continued from previous page

<code>XYZ_ColourMatchingFunctions([data, domain, ...])</code>	Implements support for the <i>CIE Standard Observers</i> XYZ colour matching functions.
---	--

## colour.colorimetry.LMS\_ConeFundamentals

```
class colour.colorimetry.LMS_ConeFundamentals(data=None,      domain=None,      labels=None,
                                              **kwargs)
```

Implements support for the Stockman and Sharpe *LMS* cone fundamentals colour matching functions.

### Parameters

- **data** (`Series` or `Dataframe` or `Signal` or `MultiSignal` or `MultiSpectralPowerDistribution` or `array_like` or `dict_like`, optional) – Data to be stored in the multi-spectral power distribution.
- **domain** (`array_like`, optional) – Values to initialise the multiple `colour.SpectralPowerDistribution` class instances `colour.continuous.Signal.wavelengths` attribute with. If both data and domain arguments are defined, the latter will be used to initialise the `colour.continuous.Signal.wavelengths` attribute.
- **labels** (`array_like`, optional) – Names to use for the `colour.SpectralPowerDistribution` class instances.

### Other Parameters

- **name** (`unicode`, optional) – Multi-spectral power distribution name.
- **interpolator** (`object`, optional) – Interpolator class type to use as interpolating function for the `colour.SpectralPowerDistribution` class instances.
- **interpolator\_args** (`dict_like`, optional) – Arguments to use when instantiating the interpolating function of the `colour.SpectralPowerDistribution` class instances.
- **extrapolator** (`object`, optional) – Extrapolator class type to use as extrapolating function for the `colour.SpectralPowerDistribution` class instances.
- **extrapolator\_args** (`dict_like`, optional) – Arguments to use when instantiating the extrapolating function of the `colour.SpectralPowerDistribution` class instances.
- **strict\_labels** (`array_like`, optional) – Multi-spectral power distribution labels for figures, default to `colour.colorimetry.LMS_ConeFundamentals.labels` attribute value.

`__init__(data=None, domain=None, labels=None, **kwargs)`

### Methods

<code>__init__([data, domain, labels])</code>	
<code>align(shape[, interpolator, ...])</code>	Aligns the multi-spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <code>a</code> operand, the operation can be either performed on a copy or in-place.

Continued on next page

Table 3.68 – continued from previous page

<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain $x$ closest element.
<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the multi-spectral power distribution in-place accordingly to <i>CIE 15:2004</i> and <i>CIE 167:2005</i> recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain $x$ variable and corresponding range $y$ variable using given method.
<code>get()</code>	
<code>interpolate(shape[, interpolator, ...])</code>	Interpolates the multi-spectral power distribution in-place accordingly to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
<code>is_uniform()</code>	Returns if independent domain $x$ variable is uniform.
<code>multi_signal_unpack_data([data, domain, ...])</code>	Unpack given data for multi-continuous signal instantiation.
<code>normalise([factor])</code>	Normalises the multi-spectral power distribution with given normalization factor.
<code>to_dataframe()</code>	Converts the continuous signal to a <i>Pandas DataFrame</i> class instance.
<code>trim(shape)</code>	Trims the multi-spectral power distribution wavelengths to given shape.
<code>trim_wavelengths(shape)</code>	
<code>zeros()</code>	

## colour.colorimetry.RGB\_ColourMatchingFunctions

```
class colour.colorimetry.RGB_ColourMatchingFunctions(data=None, domain=None, labels=None,
                                                       **kwargs)
```

Implements support for the *CIE RGB* colour matching functions.

### Parameters

- **data** (`Series` or `Dataframe` or `Signal` or `MultiSignal` or `MultiSpectralPowerDistribution` or `array_like` or `dict_like`, optional) – Data to be stored in the multi-spectral power distribution.
- **domain** (`array_like`, optional) – Values to initialise the multiple `colour.SpectralPowerDistribution` class instances `colour.continuous.Signal.wavelengths` attribute with. If both data and domain arguments are defined, the latter will be used to initialise the `colour.continuous.Signal.wavelengths` attribute.
- **labels** (`array_like`, optional) – Names to use for the `colour.SpectralPowerDistribution` class instances.

### Other Parameters

- **name** (`unicode`, optional) – Multi-spectral power distribution name.
- **interpolator** (`object`, optional) – Interpolator class type to use as interpolating function for the `colour.SpectralPowerDistribution` class instances.

- **interpolator\_args** (*dict\_like, optional*) – Arguments to use when instantiating the interpolating function of the `colour.SpectralPowerDistribution` class instances.
- **extrapolator** (*object, optional*) – Extrapolator class type to use as extrapolating function for the `colour.SpectralPowerDistribution` class instances.
- **extrapolator\_args** (*dict\_like, optional*) – Arguments to use when instantiating the extrapolating function of the `colour.SpectralPowerDistribution` class instances.
- **strict\_labels** (*array\_like, optional*) – Multi-spectral power distribution labels for figures, default to `colour.colorimetry.RGB_ColourMatchingFunctions.labels` attribute value.

`__init__(data=None, domain=None, labels=None, **kwargs)`

## Methods

<code>__init__([data, domain, labels])</code>	
<code>align(shape[, interpolator, ...])</code>	Aligns the multi-spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.
<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the multi-spectral power distribution in-place accordingly to <i>CIE 15:2004</i> and <i>CIE 167:2005</i> recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <i>x</i> variable and corresponding range <i>y</i> variable using given method.
<code>get()</code>	
<code>interpolate(shape[, interpolator, ...])</code>	Interpolates the multi-spectral power distribution in-place accordingly to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
<code>is_uniform()</code>	Returns if independent domain <i>x</i> variable is uniform.
<code>multi_signal_unpack_data([data, domain, ...])</code>	Unpack given data for multi-continuous signal instantiation.
<code>normalise([factor])</code>	Normalises the multi-spectral power distribution with given normalization factor.
<code>to_dataframe()</code>	Converts the continuous signal to a <i>Pandas DataFrame</i> class instance.
<code>trim(shape)</code>	Trims the multi-spectral power distribution wavelengths to given shape.
<code>trim_wavelengths(shape)</code>	
<code>zeros()</code>	

## colour.colorimetry.XYZ\_ColourMatchingFunctions

```
class colour.colorimetry.XYZ_ColourMatchingFunctions(data=None, domain=None, labels=None, **kwargs)
```

Implements support for the CIE Standard Observers XYZ colour matching functions.

### Parameters

- **data** (Series or Dataframe or Signal or MultiSignal or MultiSpectralPowerDistribution or array\_like or dict\_like, optional) – Data to be stored in the multi-spectral power distribution.
- **domain** (array\_like, optional) – Values to initialise the multiple colour.SpectralPowerDistribution class instances colour.continuous.Signal.wavelengths attribute with. If both data and domain arguments are defined, the latter will be used to initialise the colour.continuous.Signal.wavelengths attribute.
- **labels** (array\_like, optional) – Names to use for the colour.SpectralPowerDistribution class instances.

### Other Parameters

- **name** (unicode, optional) – Multi-spectral power distribution name.
- **interpolator** (object, optional) – Interpolator class type to use as interpolating function for the colour.SpectralPowerDistribution class instances.
- **interpolator\_args** (dict\_like, optional) – Arguments to use when instantiating the interpolating function of the colour.SpectralPowerDistribution class instances.
- **extrapolator** (object, optional) – Extrapolator class type to use as extrapolating function for the colour.SpectralPowerDistribution class instances.
- **extrapolator\_args** (dict\_like, optional) – Arguments to use when instantiating the extrapolating function of the colour.SpectralPowerDistribution class instances.
- **strict\_labels** (array\_like, optional) – Multi-spectral power distribution labels for figures, default to colour.colorimetry.XYZ\_ColourMatchingFunctions.labels attribute value.

```
__init__(data=None, domain=None, labels=None, **kwargs)
```

### Methods

---

<code>__init__([data, domain, labels])</code>	
<code>align(shape[, interpolator, ...])</code>	Aligns the multi-spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.

---

Continued on next page

Table 3.70 – continued from previous page

<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the multi-spectral power distribution in-place accordingly to <i>CIE 15:2004</i> and <i>CIE 167:2005</i> recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain $x$ variable and corresponding range $y$ variable using given method.
<code>get()</code>	
<code>interpolate(shape[, interpolator, ...])</code>	Interpolates the multi-spectral power distribution in-place accordingly to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
<code>is_uniform()</code>	Returns if independent domain $x$ variable is uniform.
<code>multi_signal_unpack_data([data, domain, ...])</code>	Unpack given data for multi-continuous signal instantiation.
<code>normalise([factor])</code>	Normalises the multi-spectral power distribution with given normalization factor.
<code>to_dataframe()</code>	Converts the continuous signal to a <i>Pandas DataFrame</i> class instance.
<code>trim(shape)</code>	Trims the multi-spectral power distribution wavelengths to given shape.
<code>trim_wavelengths(shape)</code>	
<code>zeros()</code>	

## Dataset

`colour`

<code>CMFS</code>	Aggregated colour matching functions.
<code>LMS_CMFS</code>	<i>LMS</i> colour matching functions.
<code>RGB_CMFS</code>	<i>CIE RGB</i> colour matching functions.
<code>STANDARD_OBSERVERS_CMFS</code>	<i>CIE Standard Observers XYZ</i> colour matching functions.

## `colour.CMFS`

```
colour.CMFS = CaseInsensitiveMapping({u'Stockman & Sharpe 2 Degree Cone Fundamentals': ..., u'Stiles & Burch'}
```

Aggregated colour matching functions.

## References

- [\[Bro09\]](#)
- [\[CVRe\]](#)
- [\[CVRf\]](#)
- [\[CVRg\]](#)
- [\[CVRc\]](#)
- [\[CVRh\]](#)
- [\[Mac10\]](#)

**CMFS** [CaseInsensitiveMapping] {'Stockman & Sharpe 10 Degree Cone Fundamentals', 'Stockman & Sharpe 2 Degree Cone Fundamentals', 'Wright & Guild 1931 2 Degree RGB CMFs', 'Stiles & Burch 1955 2 Degree RGB CMFs', 'Stiles & Burch 1959 10 Degree RGB CMFs', 'CIE 1931 2 Degree Standard Observer', 'CIE 1964 10 Degree Standard Observer', 'CIE 2012 2 Degree Standard Observer', 'CIE 2012 10 Degree Standard Observer'}

## colour.LMS\_CMFS

```
colour.LMS_CMFS = CaseInsensitiveMapping({u'Stockman & Sharpe 2 Degree Cone Fundamentals': ..., u'Smith & Po...  
LMS colour matching functions.
```

### References

- [\[CVRc\]](#)
- [\[Mac10\]](#)

**LMS\_CMFS** [CaseInsensitiveMapping] {'Stockman & Sharpe 2 Degree Cone Fundamentals', 'Stockman & Sharpe 10 Degree Cone Fundamentals', 'Smith & Pokorny 1975 Normal Trichromats'}

## colour.RGB\_CMFS

```
colour.RGB_CMFS = CaseInsensitiveMapping({u'Stiles & Burch 1959 10 Degree RGB CMFs': ..., u'Wright & Guild 1...  
CIE RGB colour matching functions.
```

### References

- [\[Bro09\]](#)
- [\[CVRg\]](#)
- [\[CVRh\]](#)

**RGB\_CMFS** [CaseInsensitiveMapping] {'Wright & Guild 1931 2 Degree RGB CMFs', 'Stiles & Burch 1955 2 Degree RGB CMFs', 'Stiles & Burch 1959 10 Degree RGB CMFs'}

## colour.STANDARD\_OBSERVERS\_CMFS

```
colour.STANDARD_OBSERVERS_CMFS = CaseInsensitiveMapping({u'CIE 1931 2 Degree Standard Observer': ..., u'cie_...  
CIE Standard Observers XYZ colour matching functions.
```

### References

- [\[CVRe\]](#)
- [\[CVRf\]](#)

**STANDARD\_OBSERVERS\_CMFS** [CaseInsensitiveMapping] {'CIE 1931 2 Degree Standard Observer', 'CIE 1964 10 Degree Standard Observer', 'CIE 2012 2 Degree Standard Observer', 'CIE 2012 10 Degree Standard Observer'}

Aliases:

- ‘cie\_2\_1931’: ‘CIE 1931 2 Degree Standard Observer’
- ‘cie\_10\_1964’: ‘CIE 1964 10 Degree Standard Observer’

## Colour Matching Functions Transformations

### Ancillary Objects

`colour.colorimetry`

<code>RGB_2_degree_cmfs_to_XYZ_2_degree_cmfs(...)</code>	Converts <i>Wright &amp; Guild 1931 2 Degree RGB CMFs</i> colour matching functions into the <i>CIE 1931 2 Degree Standard Observer</i> colour matching functions.
<code>RGB_10_degree_cmfs_to_XYZ_10_degree_cmfs(...)</code>	Converts <i>Stiles &amp; Burch 1959 10 Degree RGB CMFs</i> colour matching functions into the <i>CIE 1964 10 Degree Standard Observer</i> colour matching functions.
<code>RGB_10_degree_cmfs_to_LMS_10_degree_cmfs(...)</code>	Converts <i>Stiles &amp; Burch 1959 10 Degree RGB CMFs</i> colour matching functions into the <i>Stockman &amp; Sharpe 10 Degree Cone Fundamentals</i> spectral sensitivity functions.
<code>LMS_2_degree_cmfs_to_XYZ_2_degree_cmfs(...)</code>	Converts <i>Stockman &amp; Sharpe 2 Degree Cone Fundamentals</i> colour matching functions into the <i>CIE 2012 2 Degree Standard Observer</i> colour matching functions.
<code>LMS_10_degree_cmfs_to_XYZ_10_degree_cmfs(...)</code>	Converts <i>Stockman &amp; Sharpe 10 Degree Cone Fundamentals</i> colour matching functions into the <i>CIE 2012 10 Degree Standard Observer</i> colour matching functions.

`colour.colorimetry.RGB_2_degree_cmfs_to_XYZ_2_degree_cmfs`

`colour.colorimetry.RGB_2_degree_cmfs_to_XYZ_2_degree_cmfs(wavelength)`

Converts *Wright & Guild 1931 2 Degree RGB CMFs* colour matching functions into the *CIE 1931 2 Degree Standard Observer* colour matching functions.

**Parameters** `wavelength` (numeric or array\_like) – Wavelength  $\lambda$  in nm.

**Returns** *CIE 1931 2 Degree Standard Observer* spectral tristimulus values.

**Return type** ndarray

### Notes

- Data for the *CIE 1931 2 Degree Standard Observer* already exists, this definition is intended for educational purpose.

### References

- [WS00g]

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> with numpy_print_options(suppress=True):
...     RGB_2_degree_cmfs_to_XYZ_2_degree_cmfs(700)
array([ 0.0113577...,  0.004102  ,  0.         ])
```

### colour.colorimetry.RGB\_10\_degree\_cmfs\_to\_XYZ\_10\_degree\_cmfs

colour.colorimetry.RGB\_10\_degree\_cmfs\_to\_XYZ\_10\_degree\_cmfs(*wavelength*)

Converts *Stiles & Burch 1959 10 Degree RGB CMFs* colour matching functions into the *CIE 1964 10 Degree Standard Observer* colour matching functions.

**Parameters** *wavelength* (numeric or array\_like) – Wavelength  $\lambda$  in nm.

**Returns** *CIE 1964 10 Degree Standard Observer* spectral tristimulus values.

**Return type** ndarray

## Notes

- Data for the *CIE 1964 10 Degree Standard Observer* already exists, this definition is intended for educational purpose.

## References

- [WS00k]

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> with numpy_print_options(suppress=True):
...     RGB_10_degree_cmfs_to_XYZ_10_degree_cmfs(700)
array([ 0.0096432...,  0.0037526..., -0.0000041...])
```

### colour.colorimetry.RGB\_10\_degree\_cmfs\_to\_LMS\_10\_degree\_cmfs

colour.colorimetry.RGB\_10\_degree\_cmfs\_to\_LMS\_10\_degree\_cmfs(*wavelength*)

Converts *Stiles & Burch 1959 10 Degree RGB CMFs* colour matching functions into the *Stockman & Sharpe 10 Degree Cone Fundamentals* spectral sensitivity functions.

**Parameters** *wavelength* (numeric or array\_like) – Wavelength  $\lambda$  in nm.

**Returns** *Stockman & Sharpe 10 Degree Cone Fundamentals* spectral tristimulus values.

**Return type** ndarray

## Notes

- Data for the *Stockman & Sharpe 10 Degree Cone Fundamentals* already exists, this definition is intended for educational purpose.

## References

- [CIET13606]

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> with numpy_print_options(suppress=True):
...     RGB_10_degree_cmfs_to_LMS_10_degree_cmfs(700)
array([ 0.0052860...,  0.0003252...,  0.          ])
```

## colour.colorimetry.LMS\_2\_degree\_cmfs\_to\_XYZ\_2\_degree\_cmfs

`colour.colorimetry.LMS_2_degree_cmfs_to_XYZ_2_degree_cmfs(wavelength)`

Converts *Stockman & Sharpe 2 Degree Cone Fundamentals* colour matching functions into the *CIE 2012 2 Degree Standard Observer* colour matching functions.

**Parameters** `wavelength` (numeric or array\_like) – Wavelength  $\lambda$  in nm.

**Returns** *CIE 2012 2 Degree Standard Observer* spectral tristimulus values.

**Return type** ndarray

## Notes

- Data for the *CIE 2012 2 Degree Standard Observer* already exists, this definition is intended for educational purpose.

## References

- [CVRb]

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> with numpy_print_options(suppress=True):
...     LMS_2_degree_cmfs_to_XYZ_2_degree_cmfs(700)
array([ 0.0109677...,  0.0041959...,  0.          ])
```

## colour.colorimetry.LMS\_10\_degree\_cmfs\_to\_XYZ\_10\_degree\_cmfs

colour.colorimetry.LMS\_10\_degree\_cmfs\_to\_XYZ\_10\_degree\_cmfs(wavelength)

Converts Stockman & Sharpe 10 Degree Cone Fundamentals colour matching functions into the CIE 2012 10 Degree Standard Observer colour matching functions.

**Parameters** wavelength (numeric or array\_like) – Wavelength  $\lambda$  in nm.

**Returns** CIE 2012 10 Degree Standard Observer spectral tristimulus values.

**Return type** ndarray

### Notes

- Data for the CIE 2012 10 Degree Standard Observer already exists, this definition is intended for educational purpose.

### References

- [\[CVRa\]](#)

### Examples

```
>>> from colour.utilities import numpy_print_options
>>> with numpy_print_options(suppress=True):
...     LMS_10_degree_cmfs_to_XYZ_10_degree_cmfs(700)
array([ 0.0098162...,  0.0037761...,  0.         ])
```

## Illuminants and Light Sources

### Dataset

colour

ILLUMINANTS	Aggregated CIE illuminants chromaticity coordinates.
ILLUMINANTS_RELATIVE_SPDS	CIE illuminants relative spectral power distributions.
HUNTERLAB_ILLUMINANTS	Aggregated Hunter $L,a,b$ illuminant dataset.
LIGHT_SOURCES	Aggregated light sources chromaticity coordinates.
LIGHT_SOURCES_RELATIVE_SPDS	Aggregated light sources spectral power distributions.

### colour.ILLUMINANTS

```
colour.ILLUMINANTS = CaseInsensitiveMapping({u'CIE 1931 2 Degree Standard Observer': ...,
                                             u'CIE 1964 10 Degree Standard Observer': ...})
```

Aggregated CIE illuminants chromaticity coordinates.

### References

- [\[Wiky\]](#)

- [\[DigitalCInitiatives07\]](#)

**ILLUMINANTS** [CaseInsensitiveMapping] {‘CIE 1931 2 Degree Standard Observer’, ‘CIE 1964 10 Degree Standard Observer’}

Aliases:

- ‘cie\_2\_1931’: ‘CIE 1931 2 Degree Standard Observer’
- ‘cie\_10\_1964’: ‘CIE 1964 10 Degree Standard Observer’

## colour.ILLUMINANTS\_RELATIVE\_SPDS

```
colour.ILLUMINANTS_RELATIVE_SPDS = CaseInsensitiveMapping({u'FL3.8': ..., u'FL3.9': ..., u'D75': ..., u'FL100': ...})
CIE illuminants relative spectral power distributions.
```

### References

- [\[CIE04\]](#)
- [\[CIE\]](#)

ILLUMINANTS\_RELATIVE\_SPDS : CaseInsensitiveMapping

## colour.HUNTERLAB\_ILLUMINANTS

```
colour.HUNTERLAB_ILLUMINANTS = CaseInsensitiveMapping({u'CIE 1931 2 Degree Standard Observer': ..., u'CIE 1964 10 Degree Standard Observer': ...})
Aggregated Hunter L,a,b illuminant dataset.
```

### References

- [\[Hun08a\]](#)
- [\[Hun08b\]](#)

**HUNTERLAB\_ILLUMINANTS** [CaseInsensitiveMapping] {‘CIE 1931 2 Degree Standard Observer’, ‘CIE 1964 10 Degree Standard Observer’}

Aliases:

- ‘cie\_2\_1931’: ‘CIE 1931 2 Degree Standard Observer’
- ‘cie\_10\_1964’: ‘CIE 1964 10 Degree Standard Observer’

## colour.LIGHT\_SOURCES

```
colour.LIGHT_SOURCES = CaseInsensitiveMapping({u'CIE 1931 2 Degree Standard Observer': ..., u'CIE 1964 10 Degree Standard Observer': ...})
Aggregated light sources chromaticity coordinates.
```

**LIGHT\_SOURCES** [CaseInsensitiveMapping] {‘CIE 1931 2 Degree Standard Observer’, ‘CIE 1964 10 Degree Standard Observer’}

Aliases:

- ‘cie\_2\_1931’: ‘CIE 1931 2 Degree Standard Observer’
- ‘cie\_10\_1964’: ‘CIE 1964 10 Degree Standard Observer’

## colour.LIGHT\_SOURCES\_RELATIVE\_SPDS

```
colour.LIGHT_SOURCES_RELATIVE_SPDS = CaseInsensitiveMapping({u'Mercury': ... , u'F34T12WW/RS /EW (Warm White Aggregated light sources spectral power distributions.
```

```
LIGHT_SOURCES_RELATIVE_SPDS : CaseInsensitiveMapping
```

## Dominant Wavelength and Purity

colour

dominant_wavelength(xy, xy_n[, cmfs, reverse])	Returns the <i>dominant wavelength</i> $\lambda_d$ for given colour stimulus $xy$ and the related $xy_{wl}$ first and $xy_{cw}$ second intersection coordinates with the spectral locus.
complementary_wavelength(xy, xy_n[, cmfs])	Returns the <i>complementary wavelength</i> $\lambda_c$ for given colour stimulus $xy$ and the related $xy_{wl}$ first and $xy_{cw}$ second intersection coordinates with the spectral locus.
excitation_purity(xy, xy_n[, cmfs])	Returns the <i>excitation purity</i> $P_e$ for given colour stimulus $xy$ .
colorimetric_purity(xy, xy_n[, cmfs])	Returns the <i>colorimetric purity</i> $P_c$ for given colour stimulus $xy$ .

## colour.dominant\_wavelength

```
colour.dominant_wavelength(xy, xy_n, cmfs=XYZ_ColourMatchingFunctions(name='CIE 1931 2 Degree Standard Observer', ...), reverse=False)
```

Returns the *dominant wavelength*  $\lambda_d$  for given colour stimulus  $xy$  and the related  $xy_{wl}$  first and  $xy_{cw}$  second intersection coordinates with the spectral locus.

In the eventuality where the  $xy_{wl}$  first intersection coordinates are on the line of purples, the *complementary wavelength* will be computed in lieu.

The *complementary wavelength* is indicated by a negative sign and the  $xy_{cw}$  second intersection coordinates which are set by default to the same value than  $xy_{wl}$  first intersection coordinates will be set to the *complementary dominant wavelength* intersection coordinates with the spectral locus.

### Parameters

- **xy** (array\_like) – Colour stimulus  $xy$  chromaticity coordinates.
- **xy\_n** (array\_like) – Achromatic stimulus  $xy$  chromaticity coordinates.
- **cmfs** (XYZ\_ColourMatchingFunctions, optional) – Standard observer colour matching functions.
- **reverse** (bool, optional) – Reverse the computation direction to retrieve the *complementary wavelength*.

**Returns** Dominant wavelength, first intersection point  $xy$  chromaticity coordinates, second intersection point  $xy$  chromaticity coordinates.

**Return type** tuple

## References

- [CIET14804b]
- [Erdb]

## Examples

Dominant wavelength computation:

```
>>> from pprint import pprint
>>> xy = np.array([0.26415, 0.37770])
>>> xy_n = np.array([0.31270, 0.32900])
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> pprint(dominant_wavelength(xy, xy_n, cmfs))
(array(504...),
 array([ 0.0036969...,  0.6389577...]),
 array([ 0.0036969...,  0.6389577...]))
```

Complementary dominant wavelength is returned if the first intersection is located on the line of purples:

```
>>> xy = np.array([0.35000, 0.25000])
>>> pprint(dominant_wavelength(xy, xy_n, cmfs))
(array(-520...),
 array([ 0.4133314...,  0.1158663...]),
 array([ 0.0743553...,  0.8338050...]))
```

## colour.complementary\_wavelength

colour.complementary\_wavelength(xy, xy\_n, cmfs=XYZ\_ColourMatchingFunctions(name='CIE 1931 2  
Degree Standard Observer', ...))

Returns the *complementary wavelength*  $\lambda_c$  for given colour stimulus  $xy$  and the related  $xy_{wl}$  first and  $xy_{cw}$  second intersection coordinates with the spectral locus.

In the eventuality where the  $xy_{wl}$  first intersection coordinates are on the line of purples, the *dominant wavelength* will be computed in lieu.

The *dominant wavelength* is indicated by a negative sign and the  $xy_{cw}$  second intersection coordinates which are set by default to the same value than  $xy_{wl}$  first intersection coordinates will be set to the *dominant wavelength* intersection coordinates with the spectral locus.

### Parameters

- **xy** (array\_like) – Colour stimulus  $xy$  chromaticity coordinates.
- **xy\_n** (array\_like) – Achromatic stimulus  $xy$  chromaticity coordinates.
- **cmfs** (XYZ\_ColourMatchingFunctions, optional) – Standard observer colour matching functions.

**Returns** Complementary wavelength, first intersection point  $xy$  chromaticity coordinates, second intersection point  $xy$  chromaticity coordinates.

**Return type** tuple

## References

- [\[CIET14804b\]](#)
- [\[Erdb\]](#)

## Examples

Complementary wavelength computation:

```
>>> from pprint import pprint
>>> xy = np.array([0.35000, 0.25000])
>>> xy_n = np.array([0.31270, 0.32900])
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> pprint(complementary_wavelength(xy, xy_n, cmfs))
(array(520...),
 array([ 0.0743553...,  0.8338050...]),
 array([ 0.0743553...,  0.8338050...]))
```

Dominant wavelength is returned if the first intersection is located on the line of purples:

```
>>> xy = np.array([0.26415, 0.37770])
>>> pprint(complementary_wavelength(xy, xy_n, cmfs))
(array(-504...),
 array([ 0.4897494...,  0.1514035...]),
 array([ 0.0036969...,  0.6389577...]))
```

## colour.excitation\_purity

colour.**excitation\_purity**(*xy*, *xy\_n*, *cmfs*=XYZ\_ColourMatchingFunctions(*name*='CIE 1931 2 Degree Standard Observer', ...))

Returns the excitation purity  $P_e$  for given colour stimulus *xy*.

### Parameters

- **xy** (array\_like) – Colour stimulus *xy* chromaticity coordinates.
- **xy\_n** (array\_like) – Achromatic stimulus *xy* chromaticity coordinates.
- **cmfs** (XYZ\_ColourMatchingFunctions, optional) – Standard observer colour matching functions.

**Returns** Excitation purity  $P_e$ .

**Return type** numeric or array\_like

## References

- [\[CIET14804b\]](#)
- [\[Erdb\]](#)

## Examples

```
>>> xy = np.array([0.28350, 0.68700])
>>> xy_n = np.array([0.31270, 0.32900])
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> excitation_purity(xy, xy_n, cmfs)
0.9386035...
```

## colour.colorimetric\_purity

`colour.colorimetric_purity(xy, xy_n, cmfs=XYZ_ColourMatchingFunctions(name='CIE 1931 2 Degree Standard Observer', ...))`

Returns the *colorimetric purity*  $P_c$  for given colour stimulus  $xy$ .

### Parameters

- `xy` (array\_like) – Colour stimulus  $xy$  chromaticity coordinates.
- `xy_n` (array\_like) – Achromatic stimulus  $xy$  chromaticity coordinates.
- `cmfs` (XYZ\_ColourMatchingFunctions, optional) – Standard observer colour matching functions.

**Returns** Colorimetric purity  $P_c$ .

**Return type** numeric or array\_like

## References

- [CIET14804b]
- [Erdb]

## Examples

```
>>> xy = np.array([0.28350, 0.68700])
>>> xy_n = np.array([0.31270, 0.32900])
>>> cmfs = CMFS['CIE 1931 2 Degree Standard Observer']
>>> colorimetric_purity(xy, xy_n, cmfs)
0.9705976...
```

## Luminous Efficiency Functions

### colour

---

`luminous_efficiency(spd[, lef])`

Returns the *luminous efficacy* in  $lm \cdot W^{-1}$  of given spectral power distribution using given luminous efficiency function.

---

`luminous_efficiency(spd[, lef])`

Returns the *luminous efficiency* of given spectral power distribution using given luminous efficiency function.

Continued on next page

Table 3.75 – continued from previous page

<code>luminous_flux(spd[, lef, K_m])</code>	Returns the <i>luminous flux</i> for given spectral power distribution using given luminous efficiency function.
<code>mesopic_luminous_efficiency_function(Lp[, ...])</code>	Returns the mesopic luminous efficiency function $V_m(\lambda)$ for given photopic luminance $L_p$ .

## colour.luminous\_efficacy

`colour.luminous_efficacy(spd, lef=SpectralPowerDistribution(name='CIE 1924 Photopic Standard Observer', ...))`  
Returns the *luminous efficacy* in  $lm \cdot W^{-1}$  of given spectral power distribution using given luminous efficiency function.

### Parameters

- `spd` (`SpectralPowerDistribution`) – test spectral power distribution
- `lef` (`SpectralPowerDistribution`, optional) –  $V(\lambda)$  luminous efficiency function.

**Returns** Luminous efficacy in  $lm \cdot W^{-1}$ .

**Return type** numeric

## References

- [\[Wikr\]](#)

## Examples

```
>>> from colour import LIGHT_SOURCES_RELATIVE_SPDS
>>> spd = LIGHT_SOURCES_RELATIVE_SPDS['Neodium Incandescent']
>>> luminous_efficacy(spd)
136.2170803...
```

## colour.luminous\_efficiency

`colour.luminous_efficiency(spd, lef=SpectralPowerDistribution(name='CIE 1924 Photopic Standard Observer', ...))`  
Returns the *luminous efficiency* of given spectral power distribution using given luminous efficiency function.

### Parameters

- `spd` (`SpectralPowerDistribution`) – test spectral power distribution
- `lef` (`SpectralPowerDistribution`, optional) –  $V(\lambda)$  luminous efficiency function.

**Returns** Luminous efficiency.

**Return type** numeric

## References

- [\[Wikq\]](#)

## Examples

```
>>> from colour import LIGHT_SOURCES_RELATIVE_SPDS
>>> spd = LIGHT_SOURCES_RELATIVE_SPDS['Neodimium Incandescent']
>>> luminous_efficiency(spd)
0.1994393...
```

## colour.luminous\_flux

`colour.luminous_flux(spd, lef=SpectralPowerDistribution(name='CIE 1924 Photopic Standard Observer', ...), K_m=683.0)`

Returns the *luminous flux* for given spectral power distribution using given luminous efficiency function.

### Parameters

- `spd` (`SpectralPowerDistribution`) – test spectral power distribution
- `lef` (`SpectralPowerDistribution`, optional) –  $V(\lambda)$  luminous efficiency function.
- `K_m` (numeric, optional) –  $lm \cdot W^{-1}$  maximum photopic luminous efficiency

**Returns** Luminous flux.

**Return type** numeric

## References

- [\[Wikq\]](#)

## Examples

```
>>> from colour import LIGHT_SOURCES_RELATIVE_SPDS
>>> spd = LIGHT_SOURCES_RELATIVE_SPDS['Neodimium Incandescent']
>>> luminous_flux(spd)
23807.6555273...
```

## colour.mesopic\_luminous\_efficiency\_function

`colour.mesopic_luminous_efficiency_function(Lp, source=u'Blue Heavy', method=u'MOVE', photopic_lef=SpectralPowerDistribution(name='CIE 1924 Photopic Standard Observer', ...), scotopic_lef=SpectralPowerDistribution(name='CIE 1951 Scotopic Standard Observer', ...))`

Returns the mesopic luminous efficiency function  $V_m(\lambda)$  for given photopic luminance  $L_p$ .

### Parameters

- `Lp` (numeric) – Photopic luminance  $L_p$ .
- `source` (unicode, optional) – {‘Blue Heavy’, ‘Red Heavy’}, Light source colour temperature.

- **method** (unicode, optional) – {‘MOVE’, ‘LRC’}, Method to calculate the weighting factor.
- **photopic\_lef** (`SpectralPowerDistribution`, optional) –  $V(\lambda)$  photopic luminous efficiency function.
- **scotopic\_lef** (`SpectralPowerDistribution`, optional) –  $V'(\lambda)$  scotopic luminous efficiency function.

**Returns** Mesopic luminous efficiency function  $V_m(\lambda)$ .

**Return type** `SpectralPowerDistribution`

## References

- [Wiks]

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> with numpy_print_options(suppress=True):
...     mesopic_luminous_efficiency_function(0.2)
SpectralPowerDistribution([[ 380.        ,  0.000424 ...],
                           [ 381.        ,  0.0004781...],
                           [ 382.        ,  0.0005399...],
                           [ 383.        ,  0.0006122...],
                           [ 384.        ,  0.0006961...],
                           [ 385.        ,  0.0007929...],
                           [ 386.        ,  0.000907 ...],
                           [ 387.        ,  0.0010389...],
                           [ 388.        ,  0.0011923...],
                           [ 389.        ,  0.0013703...],
                           [ 390.        ,  0.0015771...],
                           [ 391.        ,  0.0018167...],
                           [ 392.        ,  0.0020942...],
                           [ 393.        ,  0.0024160...],
                           [ 394.        ,  0.0027888...],
                           [ 395.        ,  0.0032196...],
                           [ 396.        ,  0.0037222...],
                           [ 397.        ,  0.0042957...],
                           [ 398.        ,  0.0049531...],
                           [ 399.        ,  0.0057143...],
                           [ 400.        ,  0.0065784...],
                           [ 401.        ,  0.0075658...],
                           [ 402.        ,  0.0086912...],
                           [ 403.        ,  0.0099638...],
                           [ 404.        ,  0.0114058...],
                           [ 405.        ,  0.0130401...],
                           [ 406.        ,  0.0148750...],
                           [ 407.        ,  0.0169310...],
                           [ 408.        ,  0.0192211...],
                           [ 409.        ,  0.0217511...],
                           [ 410.        ,  0.0245342...],
                           [ 411.        ,  0.0275773...],
                           [ 412.        ,  0.0309172...],
                           [ 413.        ,  0.0345149...],
                           [ 414.        ,  0.0383998...],
```

[ 415.	,	0.0425744...],
[ 416.	,	0.0471074...],
[ 417.	,	0.0519322...],
[ 418.	,	0.0570541...],
[ 419.	,	0.0625466...],
[ 420.	,	0.0683463...],
[ 421.	,	0.0745255...],
[ 422.	,	0.0809440...],
[ 423.	,	0.0877344...],
[ 424.	,	0.0948915...],
[ 425.	,	0.1022731...],
[ 426.	,	0.109877 ...],
[ 427.	,	0.1178421...],
[ 428.	,	0.1260316...],
[ 429.	,	0.1343772...],
[ 430.	,	0.143017 ...],
[ 431.	,	0.1518128...],
[ 432.	,	0.1608328...],
[ 433.	,	0.1700088...],
[ 434.	,	0.1792726...],
[ 435.	,	0.1886934...],
[ 436.	,	0.1982041...],
[ 437.	,	0.2078032...],
[ 438.	,	0.2174184...],
[ 439.	,	0.2271147...],
[ 440.	,	0.2368196...],
[ 441.	,	0.2464623...],
[ 442.	,	0.2561153...],
[ 443.	,	0.2657160...],
[ 444.	,	0.2753387...],
[ 445.	,	0.2848520...],
[ 446.	,	0.2944648...],
[ 447.	,	0.3034902...],
[ 448.	,	0.3132347...],
[ 449.	,	0.3223257...],
[ 450.	,	0.3314513...],
[ 451.	,	0.3406129...],
[ 452.	,	0.3498117...],
[ 453.	,	0.3583617...],
[ 454.	,	0.3676377...],
[ 455.	,	0.3762670...],
[ 456.	,	0.3849392...],
[ 457.	,	0.3936540...],
[ 458.	,	0.4024077...],
[ 459.	,	0.4111965...],
[ 460.	,	0.4193298...],
[ 461.	,	0.4281803...],
[ 462.	,	0.4363804...],
[ 463.	,	0.4453117...],
[ 464.	,	0.4542949...],
[ 465.	,	0.4626509...],
[ 466.	,	0.4717570...],
[ 467.	,	0.4809300...],
[ 468.	,	0.4901776...],
[ 469.	,	0.4995075...],
[ 470.	,	0.5096145...],
[ 471.	,	0.5191293...],
[ 472.	,	0.5294259...],

[ 473.	,	0.5391316...],
[ 474.	,	0.5496217...],
[ 475.	,	0.5602103...],
[ 476.	,	0.5702197...],
[ 477.	,	0.5810207...],
[ 478.	,	0.5919093...],
[ 479.	,	0.6028683...],
[ 480.	,	0.6138806...],
[ 481.	,	0.6249373...],
[ 482.	,	0.6360619...],
[ 483.	,	0.6465989...],
[ 484.	,	0.6579538...],
[ 485.	,	0.6687841...],
[ 486.	,	0.6797939...],
[ 487.	,	0.6909887...],
[ 488.	,	0.7023827...],
[ 489.	,	0.7133032...],
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[ 732.	,	0.0002254...],
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[ 761.	,	0.0000281...],
[ 762.	,	0.0000262...],

```
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[ 775. , 0.0000106...],
[ 776. , 0.0000099...],
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[ 778. , 0.0000086...],
[ 779. , 0.0000080...],
[ 780. , 0.0000075...]],  

interpolator=SpragueInterpolator,  

interpolator_args={},  

extrapolator=Extrapolator,  

extrapolator_args={...})
```

## Dataset

colour

<code>LEFS</code>	Aggregated luminous efficiency functions.
<code>PHOTOPIC_LEFS</code>	Photopic luminous efficiency functions.
<code>SCOTOPIC_LEFS</code>	Scotopic luminous efficiency functions.

## colour.LEFS

```
colour.LEFS = CaseInsensitiveMapping({u'CIE 1924 Photopic Standard Observer': ..., u'CIE 1964 Photopic 10 Degree Standard Observer': ...}, name='LEFS')
Aggregated luminous efficiency functions.
```

## References

- [\[CVRd\]](#)
- [\[CVRf\]](#)
- [\[Wiks\]](#)

`LEFS` [CaseInsensitiveMapping] {‘CIE 1924 Photopic Standard Observer’, ‘Judd Modified CIE 1951 Photopic Standard Observer’, ‘Judd-Vos Modified CIE 1978 Photopic Standard Observer’, ‘CIE 1964 Photopic 10 Degree Standard Observer’, ‘CIE 2008 2 Degree Physiologically Relevant LEF’, ‘CIE 2008 10 Degree Physiologically Relevant LEF’, ‘CIE 1951 Scotopic Standard Observer’}

## colour.PHOTOPIC\_LEFS

```
colour.PHOTOPIC_LEFS = CaseInsensitiveMapping({u'CIE 1924 Photopic Standard Observer': ..., u'CIE 1964 Photopic luminous efficiency functions.
```

### References

- [\[CVRd\]](#)
- [\[CVRf\]](#)

**PHOTOPIC\_LEFS** [CaseInsensitiveMapping] {'CIE 1924 Photopic Standard Observer', 'Judd Modified CIE 1951 Photopic Standard Observer', 'Judd-Vos Modified CIE 1978 Photopic Standard Observer', 'CIE 1964 Photopic 10 Degree Standard Observer', 'CIE 2008 2 Degree Physiologically Relevant LEF', 'CIE 2008 10 Degree Physiologically Relevant LEF'}

Aliases:

- 'cie\_2\_1924': 'CIE 1931 2 Degree Standard Observer'
- 'cie\_10\_1964': 'CIE 1964 Photopic 10 Degree Standard Observer'

## colour.SCOTOPIC\_LEFS

```
colour.SCOTOPIC_LEFS = CaseInsensitiveMapping({u'cie_1951': ..., u'CIE 1951 Scotopic Standard Observer': ...})
```

### References

- [\[CVRf\]](#)

**SCOTOPIC\_LEFS** [CaseInsensitiveMapping] {'CIE 1951 Scotopic Standard Observer', }

Aliases:

- 'cie\_1951': 'CIE 1951 Scotopic Standard Observer'

## Lightness Computation

### colour

<code>lightness(Y[, method])</code>	Returns the <i>Lightness L</i> using given method.
<code>LIGHTNESS_METHODS</code>	Supported <i>Lightness</i> computations methods.

## colour.lightness

```
colour.lightness(Y, method=u'CIE 1976', **kwargs)
```

Returns the *Lightness L* using given method.

### Parameters

- `Y` (numeric or array\_like) – luminance  $Y$ .
- `method` (unicode, optional) – {‘CIE 1976’, ‘Glasser 1958’, ‘Wyszecki 1963’, ‘Fairchild 2010’, ‘Fairchild 2011’}, Computation method.

#### Other Parameters

- `Y_n` (numeric or array\_like, optional) – {`colour.colorimetry.lightness_CIE1976()`}, White reference luminance  $Y_n$ .
- `epsilon` (numeric or array\_like, optional) – {`colour.colorimetry.lightness_Fairchild2010()`, `colour.colorimetry.lightness_Fairchild2011()`},  $\epsilon$  exponent.

**Returns** Lightness  $L$ .

**Return type** numeric or array\_like

#### Notes

- Input luminance  $Y$  and optional  $Y_n$  are in domain [0, 100] or  $[0, \infty]$ .
- Output Lightness  $L$  is in range [0, 100].

#### References

- [\[FW10\]](#)
- [\[FC11\]](#)
- [\[GMRS58\]](#)
- [\[Lin03a\]](#)
- [\[Wikn\]](#)
- [\[Wys63\]](#)
- [\[WS00a\]](#)

#### Examples

```
>>> lightness(10.08)
37.9856290...
>>> lightness(10.08, Y_n=100)
37.9856290...
>>> lightness(10.08, Y_n=95)
38.9165987...
>>> lightness(10.08, method='Glasser 1958')
36.2505626...
>>> lightness(10.08, method='Wyszecki 1963')
37.0041149...
>>> lightness(10.08 / 100, epsilon=1.836, method='Fairchild 2010')
...
24.9022902...
```

## colour.LIGHTNESS\_METHODS

```
colour.LIGHTNESS_METHODS = CaseInsensitiveMapping({u'Lstar1976': ..., u'CIE 1976': ..., u'Fairchild 2011': ...})  
Supported Lightness computations methods.
```

### References

- [\[FW10\]](#)
- [\[FC11\]](#)
- [\[GMRS58\]](#)
- [\[Lin03a\]](#)
- [\[Wys63\]](#)
- [\[WS00a\]](#)

**LIGHTNESS\_METHODS** [CaseInsensitiveMapping] {'Glasser 1958', 'Wyszecki 1963', 'CIE 1976', 'Fairchild 2010', 'Fairchild 2011'}

Aliases:

- 'Lstar1976': 'CIE 1976'

## Glasser, Mckinney, Reilly and Schnelle (1958)

### colour.colorimetry

---

lightness_Glasser1958(Y)	Returns the Lightness $L$ of given luminance $Y$ using Glasser et alii (1958) method.
--------------------------	---

---

### colour.colorimetry.lightness\_Glasser1958

#### colour.colorimetry.lightness\_Glasser1958( $Y$ )

Returns the Lightness  $L$  of given luminance  $Y$  using Glasser et alii (1958) method.

**Parameters**  $Y$  (numeric or array\_like) – luminance  $Y$ .

**Returns** Lightness  $L$ .

**Return type** numeric or array\_like

### Notes

- Input luminance  $Y$  is in domain [0, 100].
- Output Lightness  $L$  is in range [0, 100].

### References

- [\[GMRS58\]](#)

## Examples

```
>>> lightness_Glasser1958(10.08)
36.2505626...
```

## Wyszecki (1963)

`colour.colorimetry`

---

<code>lightness_Wyszecki1963(Y)</code>	Returns the <i>Lightness W</i> of given <i>luminance Y</i> using <i>Wyszecki (1963)</i> method.
--	---

---

## `colour.colorimetry.lightness_Wyszecki1963`

`colour.colorimetry.lightness_Wyszecki1963(Y)`

Returns the *Lightness W* of given *luminance Y* using *Wyszecki (1963)* method.

**Parameters** `Y` (numeric or array\_like) – *luminance Y*.

**Returns** *Lightness W*.

**Return type** numeric or array\_like

## Notes

- Input *luminance Y* is in domain [0, 100].
- Output *Lightness W* is in range [0, 100].

## References

- [\[Wys63\]](#)

## Examples

```
>>> lightness_Wyszecki1963(10.08)
37.0041149...
```

## CIE 1976

`colour.colorimetry`

---

<code>lightness_CIE1976(Y[, Y_n])</code>	Returns the <i>Lightness L*</i> of given <i>luminance Y</i> using given reference white <i>luminance Y<sub>n</sub></i> as per <i>CIE 1976</i> recommendation.
--	---

---

Returns the *Lightness L\** of given *luminance Y* using given reference white *luminance Y<sub>n</sub>* as per *CIE 1976* recommendation.

## colour.colorimetry.lightness\_CIE1976

`colour.colorimetry.lightness_CIE1976(Y, Y_n=100)`

Returns the *Lightness*  $L^*$  of given *luminance*  $Y$  using given reference white *luminance*  $Y_n$  as per *CIE 1976* recommendation.

### Parameters

- `Y` (numeric or array\_like) – *luminance*  $Y$ .
- `Y_n` (numeric or array\_like, optional) – White reference *luminance*  $Y_n$ .

**Returns** *Lightness*  $L^*$ .

**Return type** numeric or array\_like

### Notes

- Input *luminance*  $Y$  and  $Y_n$  are in domain [0, 100].
- Output *Lightness*  $L^*$  is in range [0, 100].

### References

- [\[Lin03a\]](#)
- [\[WS00a\]](#)

### Examples

```
>>> lightness_CIE1976(10.08)
37.9856290...
```

## Fairchild and Wyble (2010)

`colour.colorimetry`

---

`lightness_Fairchild2010(Y[, epsilon])`

Computes *Lightness*  $L_{hdr}$  of given *luminance*  $Y$  using *Fairchild and Wyble (2010)* method according to *Michealis-Menten* kinetics.

---

## colour.colorimetry.lightness\_Fairchild2010

`colour.colorimetry.lightness_Fairchild2010(Y, epsilon=1.836)`

Computes *Lightness*  $L_{hdr}$  of given *luminance*  $Y$  using *Fairchild and Wyble (2010)* method according to *Michealis-Menten* kinetics.

### Parameters

- `Y` (array\_like) – *luminance*  $Y$ .
- `epsilon` (numeric or array\_like, optional) –  $\epsilon$  exponent.

**Returns** *Lightness  $L_{hdr}$ .*

**Return type** array\_like

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *luminance  $Y$*  is in domain  $[0, \infty]$ .

## References

- [FW10]

## Examples

```
>>> lightness_Fairchild2010(10.08 / 100)
24.9022902...
```

## Fairchild and Chen (2011)

colour.colorimetry

`lightness_Fairchild2011(Y[, epsilon, method])`

Computes *Lightness  $L_{hdr}$*  of given *luminance  $Y$*  using *Fairchild and Chen (2011)* method accordingly to *Michealis-Menten* kinetics.

## colour.colorimetry.lightness\_Fairchild2011

`colour.colorimetry.lightness_Fairchild2011(Y, epsilon=0.71, method=u'hdr-CIELAB')`

Computes *Lightness  $L_{hdr}$*  of given *luminance  $Y$*  using *Fairchild and Chen (2011)* method accordingly to *Michealis-Menten* kinetics.

### Parameters

- `Y` (array\_like) – *luminance  $Y$ .*
- `epsilon` (numeric or array\_like, optional) –  $\epsilon$  exponent.
- `method` (unicode, optional) – {‘hdr-CIELAB’, ‘hdr-IPT’}, *Lightness  $L_{hdr}$*  computation method.

**Returns** *Lightness  $L_{hdr}$ .*

**Return type** array\_like

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *luminance*  $Y$  is in domain  $[0, \infty]$ .

## References

- [FC11]

## Examples

```
>>> lightness_Fairchild2011(10.08 / 100)
26.45950981...
>>> lightness_Fairchild2011(10.08 / 100, method='hdr-IPT')
...
26.3524672...
```

## Luminance Computation

### colour

luminance(LV[, method])	Returns the <i>luminance</i> $Y$ of given <i>Lightness</i> $L^*$ or given <i>Munsell</i> value $V$ .
LUMINANCE_METHODS	Supported <i>luminance</i> computations methods.

### colour.luminance

colour.luminance(*LV*, *method=u'CIE 1976'*, \*\**kwargs*)  
Returns the *luminance*  $Y$  of given *Lightness*  $L^*$  or given *Munsell* value  $V$ .

#### Parameters

- **LV** (*numeric* or *array\_like*) – *Lightness*  $L^*$  or *Munsell* value  $V$ .
- **method** (*unicode*, *optional*) – {‘CIE 1976’, ‘Newhall 1943’, ‘ASTM D1535-08’, ‘Fairchild 2010’, ‘Fairchild 2011’}, Computation method.

#### Other Parameters

- **Y\_n** (*numeric* or *array\_like*, *optional*) – {colour.colorimetry.luminance\_CIE1976()}, White reference *luminance*  $Y_n$ .
- **epsilon** (*numeric* or *array\_like*, *optional*) – {colour.colorimetry.lightness\_Fairchild2010(), colour.colorimetry.lightness\_Fairchild2011()},  $\epsilon$  exponent.

**Returns** *luminance*  $Y$ .

**Return type** *numeric* or *array\_like*

## Notes

- Input *LV* is in domain  $[0, 100]$ ,  $[0, 10]$  or  $[0, 1]$  and optional *luminance*  $Y_n$  is in domain  $[0, 100]$ .

- Output *luminance*  $Y$  is in range [0, 100] or [0,  $\text{math:infy}$ ].

## References

- [\[ASTMInternational08\]](#)
- [\[FW10\]](#)
- [\[FC11\]](#)
- [\[Lin03a\]](#)
- [\[NNJ43\]](#)
- [\[Wikp\]](#)
- [\[WS00a\]](#)

## Examples

```
>>> luminance(37.98562910)
10.080000...
>>> luminance(37.98562910, Y_n=100)
10.080000...
>>> luminance(37.98562910, Y_n=95)
9.576000...
>>> luminance(3.74629715, method='Newhall 1943')
10.408974...
>>> luminance(3.74629715, method='ASTM D1535-08')
10.1488096...
>>> luminance(24.902290269546651, epsilon=1.836, method='Fairchild 2010')
...
0.1007999...
```

## colour.LUMINANCE\_METHODS

`colour.LUMINANCE_METHODS = CaseInsensitiveMapping({u'cie1976': ..., u'Newhall 1943': ..., u'ASTM D1535-08': ...})`  
Supported *luminance* computations methods.

## References

- [\[ASTMInternational08\]](#)
- [\[FW10\]](#)
- [\[FC11\]](#)
- [\[Lin03a\]](#)
- [\[NNJ43\]](#)
- [\[WS00a\]](#)

`LUMINANCE_METHODS [CaseInsensitiveMapping] {'Newhall 1943', 'ASTM D1535-08', 'CIE 1976', 'Fairchild 2010'}`

Aliases:

- ‘astm2008’: ‘ASTM D1535-08’
- ‘cie1976’: ‘CIE 1976’

## Newhall, Nickerson and Judd (1943)

colour.colorimetry

luminance\_Newhall1943(V)

Returns the  $luminance R_Y$  of given *Munsell* value  $V$  using *Newhall et alii (1943)* method.

---

colour.colorimetry.luminance\_Newhall1943

colour.colorimetry.luminance\_Newhall1943(V)

Returns the  $luminance R_Y$  of given *Munsell* value  $V$  using *Newhall et alii (1943)* method.

**Parameters**  $V$  (numeric or array\_like) – *Munsell* value  $V$ .

**Returns**  $luminance R_Y$ .

**Return type** numeric or array\_like

### Notes

- Input *Munsell* value  $V$  is in domain [0, 10].
- Output  $luminance R_Y$  is in range [0, 100].

### References

- [\[NNJ43\]](#)

### Examples

```
>>> luminance_Newhall1943(3.74629715382)
10.4089874...
```

## CIE 1976

colour.colorimetry

luminance\_CIE1976(Lstar[, Y\_n])

Returns the  $luminance Y$  of given *Lightness*  $L^*$  with given reference white  $luminance Y_n$ .

---

## colour.colorimetry.luminance\_CIE1976

`colour.colorimetry.luminance_CIE1976(Lstar, Y_n=100)`

Returns the *luminance*  $Y$  of given *Lightness*  $L^*$  with given reference white *luminance*  $Y_n$ .

### Parameters

- `Lstar` (numeric or array\_like) – *Lightness*  $L^*$
- `Y_n` (numeric or array\_like) – White reference *luminance*  $Y_n$ .

**Returns** *luminance*  $Y$ .

**Return type** numeric or array\_like

### Notes

- Input *Lightness*  $L^*$  and reference white *luminance*  $Y_n$  are in domain [0, 100].
- Output *luminance*  $Y$  is in range [0, 100].

### References

- [\[Lin03a\]](#)
- [\[WS00a\]](#)

### Examples

```
>>> luminance_CIE1976(37.98562910)
10.080000...
>>> luminance_CIE1976(37.98562910, 95)
9.576000...
```

## ASTM D1535-08e1

`colour.colorimetry`

---

`luminance_ASTMD153508(V)`

Returns the *luminance*  $Y$  of given *Munsell* value  $V$  using *ASTM D1535-08e1* method.

---

## colour.colorimetry.luminance\_ASTMD153508

`colour.colorimetry.luminance_ASTMD153508(V)`

Returns the *luminance*  $Y$  of given *Munsell* value  $V$  using *ASTM D1535-08e1* method.

**Parameters** `V` (numeric or array\_like) – *Munsell* value  $V$ .

**Returns** *luminance*  $Y$ .

**Return type** numeric or array\_like

## Notes

- Input *Munsell* value  $V$  is in domain [0, 10].
- Output *luminance*  $Y$  is in range [0, 100].

## References

- [\[ASTMInternational08\]](#)

## Examples

```
>>> luminance_ASTMD153508(3.74629715382)
10.1488096...
```

## Fairchild and Wyble (2010)

`colour.colorimetry`

`luminance_Fairchild2010(L_hdr[, epsilon])`

Computes *luminance*  $Y$  of given *Lightness*  $L_{hdr}$  using *Fairchild and Wyble (2010)* method according to *Michealis-Menten* kinetics.

`colour.colorimetry.luminance_Fairchild2010`

`colour.colorimetry.luminance_Fairchild2010(L_hdr, epsilon=1.836)`

Computes *luminance*  $Y$  of given *Lightness*  $L_{hdr}$  using *Fairchild and Wyble (2010)* method according to *Michealis-Menten* kinetics.

### Parameters

- `L_hdr` (`array_like`) – *Lightness*  $L_{hdr}$ .
- `epsilon` (`numeric` or `array_like`, optional) –  $\epsilon$  exponent.

**Returns** *luminance*  $Y$ .

**Return type** `array_like`

**Warning:** The output range of that definition is non standard!

## Notes

- Output *luminance*  $Y$  is in range [0, `math:infty`].

## References

- [\[FW10\]](#)

## Examples

```
>>> luminance_Fairchild2010(24.902290269546651, 1.836)
...
0.1007999...
```

## Fairchild and Chen (2011)

`colour.colorimetry`

<code>luminance_Fairchild2011(L_hdr[, epsilon, method])</code>	Computes <i>luminance Y</i> of given <i>Lightness L<sub>hdr</sub></i> using <i>Fairchild and Chen (2011)</i> method accordingly to <i>Michealis-Menten</i> kinetics.
--	--

## `colour.colorimetry.luminance_Fairchild2011`

`colour.colorimetry.luminance_Fairchild2011(L_hdr, epsilon=0.71, method=u'hdr-CIELAB')`  
 Computes *luminance Y* of given *Lightness L<sub>hdr</sub>* using *Fairchild and Chen (2011)* method accordingly to *Michealis-Menten* kinetics.

### Parameters

- `L_hdr` (`array_like`) – *Lightness L<sub>hdr</sub>*.
- `epsilon` (`numeric` or `array_like`, optional) –  $\epsilon$  exponent.
- `method` (`unicode`, optional) – {‘hdr-CIELAB’, ‘hdr-IPT’}, *Lightness L<sub>hdr</sub>* computation method.

**Returns** *luminance Y*.

**Return type** `array_like`

**Warning:** The output range of that definition is non standard!

## Notes

- Output *luminance Y* is in range [0,  $\infty$ ].

## References

- [FC11]

## Examples

```
>>> luminance_Fairchild2011(26.459509817572265)
0.1007999...
>>> luminance_Fairchild2011(26.352467267703549, method='hdr-IPT')
```

```
...  
0.1007999...
```

## Whiteness Computation

colour

<code>whiteness([method])</code>	Returns the <i>whiteness W</i> using given method.
<code>WHITENESS_METHODS</code>	Supported <i>whiteness</i> computations methods.

colour.whiteness

`colour.whiteness(method=u'CIE 2004', **kwargs)`

Returns the *whiteness W* using given method.

**Parameters** `method` (unicode, optional) – {‘CIE 2004’, ‘Berger 1959’, ‘Taube 1960’, ‘Stensby 1968’, ‘ASTM E313’, ‘Ganz 1979’, ‘CIE 2004’}, Computation method.

### Other Parameters

- `XYZ` (*array\_like*) – {`colour.colorimetry.whiteness_Berger1959()`, `colour.colorimetry.whiteness_Taube1960()`, `colour.colorimetry.whiteness_ASTME313()`}, *CIE XYZ* tristimulus values of sample.
- `XYZ_0` (*array\_like*) – {`colour.colorimetry.whiteness_Berger1959()`, `colour.colorimetry.whiteness_Taube1960()`}, *CIE XYZ* tristimulus values of reference white.
- `Lab` (*array\_like*) – {`colour.colorimetry.whiteness_Stensby1968()`}, *CIE L\*a\*b\** colourspace array of sample.
- `xy` (*array\_like*) – {`colour.colorimetry.whiteness_Ganz1979()`, `colour.colorimetry.whiteness_CIE2004()`}, Chromaticity coordinates *xy* of sample.
- `Y` (*numeric or array\_like*) – {`colour.colorimetry.whiteness_Ganz1979()`, `colour.colorimetry.whiteness_CIE2004()`}, Tristimulus *Y* value of sample.
- `xy_n` (*array\_like*) – {`colour.colorimetry.whiteness_CIE2004()`}, Chromaticity coordinates *xy\_n* of perfect diffuser.
- `observer` (unicode, optional) – {`colour.colorimetry.whiteness_CIE2004()`}, {‘CIE 1931 2 Degree Standard Observer’, ‘CIE 1964 10 Degree Standard Observer’}, *CIE Standard Observer* used for computations, *tint T* or *T<sub>10</sub>* value is dependent on viewing field angular subtense.

**Returns** *whiteness W*.

**Return type** numeric or ndarray

## References

- [CIET14804h]
- [WS00i]
- [XRP12]

- [\[Wikz\]](#)

## Examples

```
>>> xy = np.array([0.3167, 0.3334])
>>> Y = 100
>>> xy_n = np.array([0.3139, 0.3311])
>>> whiteness(xy=xy, Y=Y, xy_n=xy_n)
array([ 93.85..., -1.305...])
>>> XYZ = np.array([95.0000000, 100.0000000, 105.0000000])
>>> XYZ_0 = np.array([94.80966767, 100.0000000, 107.30513595])
>>> method = 'Taube 1960'
>>> whiteness(XYZ=XYZ, XYZ_0=XYZ_0, method=method)
91.4071738...
```

## colour.WHITENESS\_METHODS

`colour.WHITENESS_METHODS = CaseInsensitiveMapping({u'ASTM E313': ... , u'Ganz 1979': ... , u'Stensby 1968': ...})`  
Supported whiteness computations methods.

## References

- [\[CIET14804h\]](#)
- [\[XRP12\]](#)

`WHITENESS_METHODS [CaseInsensitiveMapping] {'CIE 2004', 'Berger 1959', 'Taube 1960', 'Stensby 1968', 'ASTM E313', 'Ganz 1979', 'CIE 2004'}`

Aliases:

- ‘cie2004’: ‘CIE 2004’

## Berger (1959)

### colour.colorimetry

<code>whiteness_Berger1959(XYZ, XYZ_0)</code>	Returns the whiteness index $WI$ of given sample <i>CIE XYZ</i> tristimulus values using Berger (1959) method.
---	--

### colour.colorimetry.whiteness\_Berger1959

`colour.colorimetry.whiteness_Berger1959(XYZ, XYZ_0)`

Returns the whiteness index  $WI$  of given sample *CIE XYZ* tristimulus values using Berger (1959) method.

#### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values of sample.
- `XYZ_0` (array\_like) – *CIE XYZ* tristimulus values of reference white.

**Returns** Whiteness  $WI$ .

**Return type** numeric or ndarray

## Notes

- Input  $CIE XYZ$  and  $CIE XYZ_0$  tristimulus values are in domain [0, 100].
- Whiteness  $WI$  values larger than 33.33 indicate a bluish white and values smaller than 33.33 indicate a yellowish white.

**Warning:** The input domain of that definition is non standard!

## References

- [\[XRP12\]](#)

## Examples

```
>>> XYZ = np.array([95.0000000, 100.0000000, 105.0000000])
>>> XYZ_0 = np.array([94.80966767, 100.0000000, 107.30513595])
>>> whiteness_Berger1959(XYZ, XYZ_0)
30.3638017...
```

## Taube (1960)

colour.colorimetry

`whiteness_Taube1960(XYZ, XYZ_0)`

Returns the whiteness index  $WI$  of given sample  $CIE XYZ$  tristimulus values using *Taube (1960)* method.

colour.colorimetry.whiteness\_Taube1960

`colour.colorimetry.whiteness_Taube1960(XYZ, XYZ_0)`

Returns the whiteness index  $WI$  of given sample  $CIE XYZ$  tristimulus values using *Taube (1960)* method.

### Parameters

- `XYZ` (array\_like) –  $CIE XYZ$  tristimulus values of sample.
- `XYZ_0` (array\_like) –  $CIE XYZ$  tristimulus values of reference white.

**Returns** Whiteness  $WI$ .

**Return type** numeric or ndarray

## Notes

- Input  $CIE XYZ$  and  $CIE XYZ_0$  tristimulus values are in domain [0, 100].
- Whiteness  $WI$  values larger than 100 indicate a bluish white and values smaller than 100 indicate a yellowish white.

## References

- [\[XRP12\]](#)

## Examples

```
>>> XYZ = np.array([95.0000000, 100.0000000, 105.0000000])
>>> XYZ_0 = np.array([94.80966767, 100.0000000, 107.30513595])
>>> whiteness_Taube1960(XYZ, XYZ_0)
91.4071738...
```

## Stensby (1968)

`colour.colorimetry`

`whiteness_Stensby1968(Lab)`

Returns the whiteness index  $WI$  of given sample  $CIE L^*a^*b^*$  colourspace array using Stensby (1968) method.

`colour.colorimetry.whiteness_Stensby1968`

`colour.colorimetry.whiteness_Stensby1968(Lab)`

Returns the whiteness index  $WI$  of given sample  $CIE L^*a^*b^*$  colourspace array using Stensby (1968) method.

**Parameters** `Lab` (`array_like`) –  $CIE L^*a^*b^*$  colourspace array of sample.

**Returns** Whiteness  $WI$ .

**Return type** numeric or ndarray

## Notes

- Input  $CIE L^*a^*b^*$  colourspace array is in domain [0, 100].
- Whiteness  $WI$  values larger than 100 indicate a bluish white and values smaller than 100 indicate a yellowish white.

## References

- [\[XRP12\]](#)

## Examples

```
>>> Lab = np.array([100.0000000, -2.46875131, -16.72486654])
>>> whiteness_Stensby1968(Lab)
142.7683456...
```

## ASTM E313

colour.colorimetry

whiteness\_ASTME313(XYZ)

Returns the *whiteness* index *WI* of given sample *CIE XYZ* tristimulus values using *ASTM E313* method.

### colour.colorimetry.whiteness\_ASTME313

colour.colorimetry.whiteness\_ASTME313(XYZ)

Returns the *whiteness* index *WI* of given sample *CIE XYZ* tristimulus values using *ASTM E313* method.

**Parameters** `XYZ` (array\_like) – *CIE XYZ* tristimulus values of sample.

**Returns** Whiteness *WI*.

**Return type** numeric or ndarray

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].

## References

- [\[XRP12\]](#)

## Examples

```
>>> XYZ = np.array([95.0000000, 100.0000000, 105.0000000])
>>> whiteness_ASTME313(XYZ)
55.740000...
```

## Ganz and Griesser (1979)

colour.colorimetry

---

<code>whiteness_Ganz1979(xy, Y)</code>	Returns the <i>whiteness</i> index $W$ and <i>tint</i> $T$ of given sample $xy$ chromaticity coordinates using <i>Ganz and Griesser (1979)</i> method.
--	--

---

## colour.colorimetry.whiteness\_Ganz1979

`colour.colorimetry.whiteness_Ganz1979(xy, Y)`

Returns the *whiteness* index  $W$  and *tint*  $T$  of given sample  $xy$  chromaticity coordinates using *Ganz and Griesser (1979)* method.

### Parameters

- `xy` (array\_like) – Chromaticity coordinates  $xy$  of sample.
- `Y` (numeric or array\_like) – Tristimulus  $Y$  value of sample.

**Returns** Whiteness  $W$  and tint  $T$ .

**Return type** ndarray

**Warning:** The input domain of that definition is non standard!

### Notes

- Input tristimulus  $Y$  value is in domain [0, 100].
- The formula coefficients are valid for *CIE Standard Illuminant D Series D65* and *CIE 1964 10 Degree Standard Observer*.
- Positive output *tint*  $T$  values indicate a greener tint while negative values indicate a redder tint.
- Whiteness differences of less than 5 Ganz units appear to be indistinguishable to the human eye.
- Tint differences of less than 0.5 Ganz units appear to be indistinguishable to the human eye.

### References

- [\[XRP12\]](#)

### Examples

```
>>> xy = np.array([0.3167, 0.3334])
>>> whiteness_Ganz1979(xy, 100)
array([ 85.6003766...,  0.6789003...])
```

### CIE 2004

`colour.colorimetry`

<code>whiteness_CIE2004(xy, Y, xy_n[, observer])</code>	Returns the <i>whiteness</i> $W$ or $W_{10}$ and <i>tint</i> $T$ or $T_{10}$ of given sample $xy$ chromaticity coordinates using <i>CIE 2004</i> method.
---	--

## colour.colorimetry.whiteness\_CIE2004

`colour.colorimetry.whiteness_CIE2004(xy, Y, xy_n, observer=u'CIE 1931 2 Degree Standard Observer')`

Returns the *whiteness*  $W$  or  $W_{10}$  and *tint*  $T$  or  $T_{10}$  of given sample  $xy$  chromaticity coordinates using *CIE 2004* method.

### Parameters

- `xy` (array\_like) – Chromaticity coordinates  $xy$  of sample.
- `Y` (numeric or array\_like) – Tristimulus  $Y$  value of sample.
- `xy_n` (array\_like) – Chromaticity coordinates  $xy_n$  of perfect diffuser.
- `observer` (unicode, optional) – {'CIE 1931 2 Degree Standard Observer', 'CIE 1964 10 Degree Standard Observer'}, *CIE Standard Observer* used for computations, *tint*  $T$  or  $T_{10}$  value is dependent on viewing field angular subtense.

**Returns** Whiteness  $W$  or  $W_{10}$  and *tint*  $T$  or  $T_{10}$  of given sample.

**Return type** ndarray

**Warning:** The input domain of that definition is non standard!

## Notes

- Input tristimulus  $Y$  value is in domain [0, 100].
- This method may be used only for samples whose values of  $W$  or  $W_{10}$  lie within the following limits: greater than 40 and less than 5Y - 280, or 5Y10 - 280.
- This method may be used only for samples whose values of  $T$  or  $T_{10}$  lie within the following limits: greater than -4 and less than +2.
- Output *whiteness*  $W$  or  $W_{10}$  values larger than 100 indicate a bluish white while values smaller than 100 indicate a yellowish white.
- Positive output *tint*  $T$  or  $T_{10}$  values indicate a greener tint while negative values indicate a redder tint.

## References

- [CIET14804h]

## Examples

```
>>> xy = np.array([0.3167, 0.3334])
>>> xy_n = np.array([0.3139, 0.3311])
>>> whiteness_CIE2004(xy, 100, xy_n)
array([ 93.85..., -1.305...])
```

## Yellowness Computation

### colour

<code>yellowness(XYZ[, method])</code>	Returns the <i>yellowness</i> $W$ using given method.
<code>YELLOWNESS_METHODS</code>	Supported <i>yellowness</i> computations methods.

### colour.yellowness

`colour.yellowness(XYZ, method=u'ASTM E313')`  
Returns the *yellowness*  $W$  using given method.

#### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values of sample.
- `method` (unicode, optional) – {‘ASTM E313’, ‘ASTM D1925’}, Computation method.

**Returns** *yellowness*  $Y$ .

**Return type** numeric or ndarray

## References

- [XRP12]

## Examples

```
>>> import numpy as np
>>> XYZ = np.array([95.0000000, 100.0000000, 105.0000000])
>>> yellowness(XYZ)
11.065000...
>>> method = 'ASTM D1925'
>>> yellowness(XYZ, method=method)
10.299999...
```

### colour.YELLOWNESS\_METHODS

`colour.YELLOWNESS_METHODS = CaseInsensitiveMapping({u'ASTM E313': ... , u'ASTM D1925': ...})`  
Supported *yellowness* computations methods.

## References

- [\[XRP12\]](#)

**YELLOWNESS\_METHODS** [CaseInsensitiveMapping] {'ASTM E313', 'ASTM D1925'}

## ASTM D1925

colour.colorimetry

---

yellowness_ASTMD1925(XYZ)	Returns the <i>yellowness</i> index $YI$ of given sample <i>CIE XYZ</i> tristimulus values using <i>ASTM D1925</i> method.
---------------------------	--

---

### colour.colorimetry.yellowness\_ASTMD1925

colour.colorimetry.**yellowness\_ASTMD1925**(XYZ)

Returns the *yellowness* index  $YI$  of given sample *CIE XYZ* tristimulus values using *ASTM D1925* method.

*ASTM D1925* has been specifically developed for the definition of the *Yellowness* of homogeneous, non-fluorescent, almost neutral-transparent, white-scattering or opaque plastics as they will be reviewed under daylight condition. It can be other materials as well, as long as they fit into this description.

**Parameters** `XYZ` (array\_like) – *CIE XYZ* tristimulus values of sample.

**Returns** Whiteness  $YI$ .

**Return type** numeric or ndarray

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].

## References

- [\[XRP12\]](#)

## Examples

```
>>> import numpy as np
>>> XYZ = np.array([95.0000000, 100.0000000, 105.0000000])
>>> yellowness_ASTMD1925(XYZ)
10.299999...
```

## ASTM E313

`colour.colorimetry`

`yellowness_ASTME313(XYZ)`

Returns the *yellowness* index  $YI$  of given sample *CIE XYZ* tristimulus values using *ASTM E313* method.

### `colour.colorimetry.yellowness_ASTME313`

`colour.colorimetry.yellowness_ASTME313(XYZ)`

Returns the *yellowness* index  $YI$  of given sample *CIE XYZ* tristimulus values using *ASTM E313* method.

*ASTM E313* has successfully been used for a variety of white or near white materials. This includes coatings, Plastics, Textiles.

**Parameters** `XYZ` (`array_like`) – *CIE XYZ* tristimulus values of sample.

**Returns** *Whiteness*  $YI$ .

**Return type** numeric or ndarray

**Warning:** The input domain of that definition is non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].

## References

- [\[XRP12\]](#)

## Examples

```
>>> import numpy as np
>>> XYZ = np.array([95.0000000, 100.0000000, 105.0000000])
>>> yellowness_ASTME313(XYZ)
11.065000...
```

## Constants

- [CIE](#)
- [CODATA](#)
- [Common](#)

## CIE

colour.constants

CIE_E	<i>CIE</i> $\epsilon$ constant.
CIE_K	<i>CIE</i> $\kappa$ constant.
K_M	Rounded maximum photopic luminous efficiency $K_m$ value in $lm \cdot W^{-1}$ .
KP_M	Rounded maximum scotopic luminous efficiency $K'_m$ value in $lm \cdot W^{-1}$ .

colour.constants.CIE\_E

colour.constants.CIE\_E = 0.008856451679035631

*CIE*  $\epsilon$  constant.

CIE\_E : numeric

### Notes

- The original *CIE* value for  $\epsilon$  is  $\epsilon = 0.008856$ , Lindbloom (2003) has shown that this value is causing a discontinuity at the junction point of the two functions grafted together to create the *Lightness L\** function.

That discontinuity can be avoided by using the rational representation as follows:  $\epsilon = 216 / 24389$ .

## References

- [Lin03a]

colour.constants.CIE\_K

colour.constants.CIE\_K = 903.2962962962963

*CIE*  $\kappa$  constant.

CIE\_K : numeric

### Notes

- The original *CIE* value for  $\kappa$  is  $\kappa = 903.3$ , Lindbloom (2003) has shown that this value is causing a discontinuity at the junction point of the two functions grafted together to create the *Lightness L\** function.

That discontinuity can be avoided by using the rational representation as follows:  $k = 24389 / 27$ .

## References

- [Lin03a]

## colour.constants.K\_M

`colour.constants.K_M = 683.0`

Rounded maximum photopic luminous efficiency  $K_m$  value in  $lm \cdot W^{-1}$ .

`K_M` : numeric

### Notes

- To be adequate for all practical applications the  $K_m$  value has been rounded from the original 683.002 value.

### References

- [\[WS00e\]](#)

## colour.constants.KP\_M

`colour.constants.KP_M = 1700.0`

Rounded maximum scotopic luminous efficiency  $K'_m$  value in  $lm \cdot W^{-1}$ .

`KP_M` : numeric

### Notes

- To be adequate for all practical applications the  $K'_m$  value has been rounded from the original 1700.06 value.

### References

- [\[WS00e\]](#)

## CODATA

### colour.constants

<code>AVOGADRO_CONSTANT</code>	Avogadro constant.
<code>BOLTZMANN_CONSTANT</code>	Boltzmann constant.
<code>LIGHT_SPEED</code>	Speed of light in vacuum.
<code>PLANCK_CONSTANT</code>	Planck constant.

## colour.constants.AVOGADRO\_CONSTANT

`colour.constants.AVOGADRO_CONSTANT = 6.02214179e+23`

Avogadro constant.

`AVOGADRO_CONSTANT` : numeric

**colour.constants.BOLTZMANN\_CONSTANT**

```
colour.constants.BOLTZMANN_CONSTANT = 1.38065e-23
```

Boltzmann constant.

BOLTZMANN\_CONSTANT : numeric

**colour.constants.LIGHT\_SPEED**

```
colour.constants.LIGHT_SPEED = 299792458.0
```

Speed of light in vacuum.

LIGHT\_SPEED : numeric

**colour.constants.PLANCK\_CONSTANT**

```
colour.constants.PLANCK_CONSTANT = 6.62607e-34
```

Planck constant.

PLANCK\_CONSTANT : numeric

**Common****colour.constants**

DEFAULT_FLOAT_DTYPE	alias of float64
EPSILON	
FLOATING_POINT_NUMBER_PATTERN	unicode(object=') -> unicode object
INTEGER_THRESHOLD	Integer threshold value.

**colour.constants.DEFAULT\_FLOAT\_DTYPE**

```
colour.constants.DEFAULT_FLOAT_DTYPE
```

alias of float64

**colour.constants.EPSILON**

```
colour.constants.EPSILON = 2.2204460492503131e-16
```

**colour.constants.FLOATING\_POINT\_NUMBER\_PATTERN**

```
colour.constants.FLOATING_POINT_NUMBER_PATTERN = u'[0-9]*\\.?[0-9]+([eE][+-]?[0-9]+)?'
```

unicode(object=') -> unicode object  
unicode(string[, encoding[, errors]]) -> unicode object

Create a new Unicode object from the given encoded string. encoding defaults to the current default string encoding. errors can be 'strict', 'replace' or 'ignore' and defaults to 'strict'.

**colour.constants.INTEGER\_THRESHOLD**

```
colour.constants.INTEGER_THRESHOLD = 0.001
```

Integer threshold value.

INTEGER\_THRESHOLD : numeric

**Continuous Signal**

- *Continuous Signal*

**Continuous Signal**

colour.continuous

<code>AbstractContinuousFunction([name])</code>	Defines the base class for abstract continuous function.
<code>Signal([data, domain])</code>	Defines the base class for continuous signal.
<code>MultiSignal([data, domain, labels])</code>	Defines the base class for multi-continuous signal, a container for multiple <code>colour.continuous.Signal</code> sub-class instances.

**colour.continuous.AbstractContinuousFunction**

```
class colour.continuous.AbstractContinuousFunction(name=None)
```

Defines the base class for abstract continuous function.

This is an ABCMeta abstract class that must be inherited by sub-classes.

The sub-classes are expected to implement the `colour.continuous.AbstractContinuousFunction.function()` method so that evaluating the function for any independent domain  $x \in \mathbb{R}$  variable returns a corresponding range  $y \in \mathbb{R}$  variable. A conventional implementation adopts an interpolating function encapsulated inside an extrapolating function. The resulting function independent domain, stored as discrete values in the `colour.continuous.AbstractContinuousFunction.domain` attribute corresponds with the function dependent and already known range stored in the `colour.continuous.AbstractContinuousFunction.range` attribute.

**Parameters** `name` (unicode, optional) – Continuous function name.

`name`  
`domain`  
`range`  
`interpolator`  
`interpolator_args`  
`extrapolator`  
`extrapolator_args`  
`function`

```
__str__()
__repr__()
__hash__()
__getitem__()
__setitem__()
__contains__()
__len__()
__eq__()
__ne__()
__iadd__()
__add__()
__isub__()
__sub__()
__imul__()
__mul__()
__idiv__()
__div__()
__ipow__()
__pow__()

arithmetical_operation()
fill_nan()
domain_distance()
is_uniform()
copy()
__init__(name=None)
```

## Methods

---

<code>__init__([name])</code>	
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place, must be reimplemented by sub-classes.
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.

---

Continued on next page

Table 3.103 – continued from previous page

<code>fill_nan([method, default])</code>	Fill NaNs in independent domain $x$ variable and corresponding range $y$ variable using given method, must be reimplemented by sub-classes.
<code>is_uniform()</code>	Returns if independent domain $x$ variable is uniform.

**colour.continuous.Signal**

```
class colour.continuous.Signal(data=None, domain=None, **kwargs)
```

Defines the base class for continuous signal.

The class implements the `Signal.function()` method so that evaluating the function for any independent domain  $x \in \mathbb{R}$  variable returns a corresponding range  $y \in \mathbb{R}$  variable. It adopts an interpolating function encapsulated inside an extrapolating function. The resulting function independent domain, stored as discrete values in the `colour.continuous.Signal.domain` attribute corresponds with the function dependent and already known range stored in the `colour.continuous.Signal.range` attribute.

**Parameters**

- `data` (`Series` or `Signal` or `array_like` or `dict_like`, optional) – Data to be stored in the continuous signal.
- `domain` (`array_like`, optional) – Values to initialise the `colour.continuous.Signal.domain` attribute with. If both `data` and `domain` arguments are defined, the latter will be used to initialise the `colour.continuous.Signal.domain` attribute.

**Other Parameters**

- `name` (`unicode`, optional) – Continuous signal name.
- `dtype` (`type`, optional) – `{np.float16, np.float32, np.float64, np.float128}`, Floating point data type.
- `interpolator` (`object`, optional) – Interpolator class type to use as interpolating function.
- `interpolator_args` (`dict_like`, optional) – Arguments to use when instantiating the interpolating function.
- `extrapolator` (`object`, optional) – Extrapolator class type to use as extrapolating function.
- `extrapolator_args` (`dict_like`, optional) – Arguments to use when instantiating the extrapolating function.

**dtype****domain****range****interpolator****interpolator\_args****extrapolator****extrapolator\_args****function****\_\_str\_\_()**

```
__repr__()
__getitem__()
__setitem__()
__contains__()
__eq__()
__ne__()
arithmetical_operation()
signal_unpack_data()
fill_nan()
to_series()
```

## Examples

Instantiation with implicit *domain*:

```
>>> range_ = np.linspace(10, 100, 10)
>>> print(Signal(range_))
[[ 0.  10.]
 [ 1.  20.]
 [ 2.  30.]
 [ 3.  40.]
 [ 4.  50.]
 [ 5.  60.]
 [ 6.  70.]
 [ 7.  80.]
 [ 8.  90.]
 [ 9. 100.]]
```

Instantiation with explicit *domain*:

```
>>> domain = np.arange(100, 1100, 100)
>>> print(Signal(range_, domain))
[[ 100.   10.]
 [ 200.   20.]
 [ 300.   30.]
 [ 400.   40.]
 [ 500.   50.]
 [ 600.   60.]
 [ 700.   70.]
 [ 800.   80.]
 [ 900.   90.]
 [1000.  100.]]
```

Instantiation with a *dict*:

```
>>> print(Signal(dict(zip(domain, range_))))
[[ 100.   10.]
 [ 200.   20.]
 [ 300.   30.]
 [ 400.   40.]
 [ 500.   50.]]
```

```
[ 600.  60.]
[ 700.  70.]
[ 800.  80.]
[ 900.  90.]
[1000. 100.]]
```

Instantiation with a *Pandas Series*:

```
>>> if is_pandas_installed():
...     from pandas import Series
...     print(Signal(
...         Series(dict(zip(domain, range_)))))

[[ 100.  10.]
[ 200.  20.]
[ 300.  30.]
[ 400.  40.]
[ 500.  50.]
[ 600.  60.]
[ 700.  70.]
[ 800.  80.]
[ 900.  90.]
[1000. 100.]]
```

Retrieving domain *y* variable for arbitrary range *x* variable:

```
>>> x = 150
>>> range_ = np.sin(np.linspace(0, 1, 10))
>>> Signal(range_, domain)[x]
0.0359701...
>>> x = np.linspace(100, 1000, 3)
>>> Signal(range_, domain)[x]
array([-..., 4.7669395...e-01, 8.4147098...e-01])
```

Using an alternative interpolating function:

```
>>> x = 150
>>> from colour.algebra import CubicSplineInterpolator
>>> Signal(
...     range_,
...     domain,
...     interpolator=CubicSplineInterpolator)[x]
0.0555274...
>>> x = np.linspace(100, 1000, 3)
>>> Signal(
...     range_,
...     domain,
...     interpolator=CubicSplineInterpolator)[x]
array([ 0.        ,  0.4794253...,  0.8414709...])
```

`__init__(data=None, domain=None, **kwargs)`

## Methods

---

`__init__([data, domain])`

Continued on next page

Table 3.104 – continued from previous page

<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <i>x</i> variable and corresponding range <i>y</i> variable using given method.
<code>is_uniform()</code>	Returns if independent domain <i>x</i> variable is uniform.
<code>signal_unpack_data([data, domain, dtype])</code>	Unpack given data for continuous signal instantiation.
<code>to_series()</code>	Converts the continuous signal to a <i>Pandas Series</i> class instance.

## colour.continuous.MultiSignal

**class** `colour.continuous.MultiSignal(data=None, domain=None, labels=None, **kwargs)`

Defines the base class for multi-continuous signal, a container for multiple `colour.continuous.Signal` sub-class instances.

### Parameters

- **data** (Series or Dataframe or `Signal` or `MultiSignal` or array\_like or dict\_like, optional) – Data to be stored in the multi-continuous signal.
- **domain** (array\_like, optional) – Values to initialise the multiple `colour.continuous.Signal` sub-class instances `colour.continuous.Signal.domain` attribute with. If both data and domain arguments are defined, the latter will be used to initialise the `colour.continuous.Signal.domain` attribute.
- **labels** (array\_like, optional) – Names to use for the `colour.continuous.Signal` sub-class instances.

### Other Parameters

- **name** (unicode, optional) – Multi-continuous signal name.
- **dtype** (type, optional) – {`np.float16`, `np.float32`, `np.float64`, `np.float128`}, Floating point data type.
- **interpolator** (object, optional) – Interpolator class type to use as interpolating function for the `colour.continuous.Signal` sub-class instances.
- **interpolator\_args** (dict\_like, optional) – Arguments to use when instantiating the interpolating function of the `colour.continuous.Signal` sub-class instances.
- **extrapolator** (object, optional) – Extrapolator class type to use as extrapolating function for the `colour.continuous.Signal` sub-class instances.
- **extrapolator\_args** (dict\_like, optional) – Arguments to use when instantiating the extrapolating function of the `colour.continuous.Signal` sub-class instances.
- **signal\_type** (type, optional) – The `colour.continuous.Signal` sub-class type used for instances.

### dtype

---

```

domain
range
interpolator
interpolator_args
extrapolator
extrapolator_args
function
signals
labels
__str__()
__repr__()
__getitem__()
__setitem__()
__contains__()
__eq__()
__ne__()
arithmetical_operation()
multi_signal_unpack_data()
fill_nan()
to_dataframe()

```

## Examples

Instantiation with implicit *domain* and a single signal:

```

>>> range_ = np.linspace(10, 100, 10)
>>> print(MultiSignal(range_))
[[ 0.  10.]
 [ 1.  20.]
 [ 2.  30.]
 [ 3.  40.]
 [ 4.  50.]
 [ 5.  60.]
 [ 6.  70.]
 [ 7.  80.]
 [ 8.  90.]
 [ 9. 100.]]

```

Instantiation with explicit *domain* and a single signal:

```

>>> domain = np.arange(100, 1100, 100)
>>> print(MultiSignal(range_, domain))
[[ 100.   10.]
 [ 200.   20.]
 [ 300.   30.]]

```

```
[ 400.  40.]  
[ 500.  50.]  
[ 600.  60.]  
[ 700.  70.]  
[ 800.  80.]  
[ 900.  90.]  
[1000. 100.]]
```

Instantiation with multiple signals:

```
>>> range_ = tstack([np.linspace(10, 100, 10) * 3)  
>>> range_ += np.array([0, 10, 20])  
>>> print(MultiSignal(range_, domain))  
[[ 100.   10.   20.   30.]  
 [ 200.   20.   30.   40.]  
 [ 300.   30.   40.   50.]  
 [ 400.   40.   50.   60.]  
 [ 500.   50.   60.   70.]  
 [ 600.   60.   70.   80.]  
 [ 700.   70.   80.   90.]  
 [ 800.   80.   90.  100.]  
 [ 900.   90.  100.  110.]  
 [1000.  100.  110.  120.]]
```

Instantiation with a *dict*:

```
>>> print(MultiSignal(dict(zip(domain, range_))))  
[[ 100.   10.   20.   30.]  
 [ 200.   20.   30.   40.]  
 [ 300.   30.   40.   50.]  
 [ 400.   40.   50.   60.]  
 [ 500.   50.   60.   70.]  
 [ 600.   60.   70.   80.]  
 [ 700.   70.   80.   90.]  
 [ 800.   80.   90.  100.]  
 [ 900.   90.  100.  110.]  
 [1000.  100.  110.  120.]]
```

Instantiation using a *Signal* sub-class:

```
>>> class NotSignal(Signal):  
...     pass  
  
>>> multi_signal = MultiSignal(range_, domain, signal_type=NotSignal)  
>>> print(multi_signal)  
[[ 100.   10.   20.   30.]  
 [ 200.   20.   30.   40.]  
 [ 300.   30.   40.   50.]  
 [ 400.   40.   50.   60.]  
 [ 500.   50.   60.   70.]  
 [ 600.   60.   70.   80.]  
 [ 700.   70.   80.   90.]  
 [ 800.   80.   90.  100.]  
 [ 900.   90.  100.  110.]  
 [1000.  100.  110.  120.]]  
>>> type(multi_signal.signals[0])  
<class 'multi_signal.NotSignal'>
```

Instantiation with a *Pandas Series*:

```
>>> if is_pandas_installed():
...     from pandas import Series
...     print(MultiSignal(
...         Series(dict(zip(domain, np.linspace(10, 100, 10))))))
[[ 100.   10.]
 [ 200.   20.]
 [ 300.   30.]
 [ 400.   40.]
 [ 500.   50.]
 [ 600.   60.]
 [ 700.   70.]
 [ 800.   80.]
 [ 900.   90.]
 [1000.  100.]]
```

Instantiation with a *Pandas Dataframe*:

```
>>> if is_pandas_installed():
...     from pandas import DataFrame
...     data = dict(zip(['a', 'b', 'c'], tsplit(range_)))
...     print(MultiSignal(
...         DataFrame(data, domain)))
[[ 100.   10.   20.   30.]
 [ 200.   20.   30.   40.]
 [ 300.   30.   40.   50.]
 [ 400.   40.   50.   60.]
 [ 500.   50.   60.   70.]
 [ 600.   60.   70.   80.]
 [ 700.   70.   80.   90.]
 [ 800.   80.   90.  100.]
 [ 900.   90.  100.  110.]
 [1000.  100.  110.  120.]]
```

Retrieving domain *y* variable for arbitrary range *x* variable:

```
>>> x = 150
>>> range_ = tstack([np.sin(np.linspace(0, 1, 10))] * 3)
>>> range_ += np.array([0.0, 0.25, 0.5])
>>> MultiSignal(range_, domain)[x]
array([ 0.0359701...,  0.2845447...,  0.5331193...])
>>> x = np.linspace(100, 1000, 3)
>>> MultiSignal(range_, domain)[x]
array([[ 4.4085384...e-20,  2.5000000...e-01,  5.0000000...e-01],
       [ 4.7669395...e-01,  7.2526859...e-01,  9.7384323...e-01],
       [ 8.4147098...e-01,  1.0914709...e+00,  1.3414709...e+00]])
```

Using an alternative interpolating function:

```
>>> x = 150
>>> from colour.algebra import CubicSplineInterpolator
>>> MultiSignal(
...     range_,
...     domain,
...     interpolator=CubicSplineInterpolator)[x]
array([ 0.0555274...,  0.3055274...,  0.5555274...])
>>> x = np.linspace(100, 1000, 3)
>>> MultiSignal(
```

```

...     range_,
...     domain,
...     interpolator=CubicSplineInterpolator)[x]
array([[ 0.           ...,  0.25           ...,  0.5           ...],
       [ 0.4794253...,  0.7294253...,  0.9794253...],
       [ 0.8414709...,  1.0914709...,  1.3414709...]])

```

`__init__(data=None, domain=None, labels=None, **kwargs)`

## Methods

<code>__init__([data, domain, labels])</code>	
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <i>x</i> variable and corresponding range <i>y</i> variable using given method.
<code>is_uniform()</code>	Returns if independent domain <i>x</i> variable is uniform.
<code>multi_signal_unpack_data([data, domain, ...])</code>	Unpack given data for multi-continuous signal instantiation.
<code>to_dataframe()</code>	Converts the continuous signal to a <i>Pandas DataFrame</i> class instance.

## Corresponding Chromaticities

- *Prediction*
  - *Fairchild (1990)*
  - *CIE 1994*
  - *CMCCAT2000*
  - *Von Kries*

## Prediction

### colour

<code>corresponding_chromaticities_prediction([...])</code>	Returns the corresponding chromaticities prediction for given chromatic adaptation model.
<code>CORRESPONDING_CHROMATICITIES_PREDICTION_MODELS</code>	Aggregated corresponding chromaticities prediction models.

## colour.corresponding\_chromaticities\_prediction

`colour.corresponding_chromaticities_prediction(experiment=1, model=u'Von Kries', **kwargs)`  
Returns the corresponding chromaticities prediction for given chromatic adaptation model.

### Parameters

- `experiment` (integer, optional) – {1, 2, 3, 4, 6, 8, 9, 11, 12} *Breneman (1987)* experiment number.
- `model` (unicode, optional) – {'Von Kries', 'CIE 1994', 'CMCCAT2000', 'Fairchild 1990'}, Chromatic adaptation model.

**Other Parameters** `transform` (unicode, optional) – {`colour.corresponding_chromaticities_prediction_VonKries()`}, {'CAT02', 'XYZ Scaling', 'Von Kries', 'Bradford', 'Sharp', 'Fairchild', 'CMCCAT97', 'CMCCAT2000', 'CAT02\_BRILL\_CAT', 'Bianco', 'Bianco PC'}, Chromatic adaptation transform.

**Returns** Corresponding chromaticities prediction.

**Return type** tuple

## References

- [Bre87]
- [CIET13294]
- [Fai91]
- [Fai13c]
- [Fai13b]
- [LLRH02]
- [WRC12a]

## Examples

```
>>> from pprint import pprint
>>> pr = corresponding_chromaticities_prediction(2, 'CMCCAT2000')
>>> pr = [(p.uvp_m, p.uvp_p) for p in pr]
>>> pprint(pr)
[((0.207, 0.486), (0.2083210..., 0.4727168...)),
 ((0.449, 0.511), (0.4459270..., 0.5077735...)),
 ((0.263, 0.505), (0.2640262..., 0.4955361...)),
 ((0.322, 0.545), (0.3316884..., 0.5431580...)),
 ((0.316, 0.537), (0.3222624..., 0.5357624...)),
 ((0.265, 0.553), (0.2710705..., 0.5501997...)),
 ((0.221, 0.538), (0.2261826..., 0.5294740...)),
 ((0.135, 0.532), (0.1439693..., 0.5190984...)),
 ((0.145, 0.472), (0.1494835..., 0.4556760...)),
 ((0.163, 0.331), (0.1563172..., 0.3164151...)),
 ((0.176, 0.431), (0.1763199..., 0.4127589...)),
 ((0.244, 0.349), (0.2287638..., 0.3499324...)])
```

## colour.CORRESPONDING\_CHROMATICITIES\_PREDICTION\_MODELS

```
colour.CORRESPONDING_CHROMATICITIES_PREDICTION_MODELS = CaseInsensitiveMapping({u'vonkries': ..., u'Von Kries': ...})
```

Aggregated corresponding chromaticities prediction models.

### References

- [\[Bre87\]](#)
- [\[CIET13294\]](#)
- [\[Fai91\]](#)
- [\[Fai13c\]](#)
- [\[Fai13b\]](#)
- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

**CORRESPONDING\_CHROMATICITIES\_PREDICTION\_MODELS** [CaseInsensitiveMapping] {‘CIE 1994’, ‘CMCCAT2000’, ‘Fairchild 1990’, ‘Von Kries’}

Aliases:

- ‘vonkries’: ‘Von Kries’

### Dataset

colour

BRENEMAN_EXPERIMENTS	<i>Breneman (1987)</i> experiments.
BRENEMAN_EXPERIMENTS_PRIMARIES_CHROMATICITIES	<i>Breneman (1987)</i> experiments primaries chromaticities.

## colour.BRENEMAN\_EXPERIMENTS

```
colour.BRENEMAN_EXPERIMENTS = {1: (BrenemanExperimentResult(name=u'Illuminant', uvp_t=array([ 0.259, 0.526]),
```

## References

- [Bre87]

BRENEMAN\_EXPERIMENTS\_PRIMARIES\_CHROMATICITIES : dict

### Fairchild (1990)

colour.corresponding

---

`corresponding_chromaticities_prediction_Fairchild1990` Returns the corresponding chromaticities prediction for Fairchild (1990) chromatic adaptation model.

---

`colour.corresponding.corresponding_chromaticities_prediction_Fairchild1990`

`colour.corresponding.corresponding_chromaticities_prediction_Fairchild1990(experiment=1)`  
Returns the corresponding chromaticities prediction for Fairchild (1990) chromatic adaptation model.

**Parameters** `experiment` (integer, optional) – {1, 2, 3, 4, 6, 8, 9, 11, 12} Breneman (1987)  
experiment number.

**Returns** Corresponding chromaticities prediction.

**Return type** tuple

## References

- [Bre87]
- [Fai91]
- [Fai13c]

## Examples

```
>>> from pprint import pprint
>>> pr = corresponding_chromaticities_prediction_Fairchild1990(2)
>>> pr = [(p.uvp_m, p.uvp_p) for p in pr]
>>> pprint(pr)
[((0.207, 0.486), (0.2089528..., 0.4724034...)),
 ((0.449, 0.511), (0.4375652..., 0.5121030...)),
 ((0.263, 0.505), (0.2621362..., 0.4972538...)),
 ((0.322, 0.545), (0.3235312..., 0.5475665...)),
 ((0.316, 0.537), (0.3151390..., 0.5398333...)),
 ((0.265, 0.553), (0.2634745..., 0.5544335...)),
 ((0.221, 0.538), (0.2211595..., 0.5324470...)),
 ((0.135, 0.532), (0.1396949..., 0.5207234...)),
 ((0.145, 0.472), (0.1512288..., 0.4533041...)),
 ((0.163, 0.331), (0.1715691..., 0.3026264...)),
 ((0.176, 0.431), (0.1825792..., 0.4077892...)),
 ((0.244, 0.349), (0.2418904..., 0.3413401...)])
```

## CIE 1994

colour.corresponding

---

`corresponding_chromaticities_prediction_CIE1994([.Re])` Returns the corresponding chromaticities prediction for *CIE 1994* chromatic adaptation model.

---

`colour.corresponding.corresponding_chromaticities_prediction_CIE1994`

`colour.corresponding.corresponding_chromaticities_prediction_CIE1994(experiment=1)`

Returns the corresponding chromaticities prediction for *CIE 1994* chromatic adaptation model.

**Parameters** `experiment` (integer, optional) – {1, 2, 3, 4, 6, 8, 9, 11, 12} *Breneman (1987)*  
experiment number.

**Returns** Corresponding chromaticities prediction.

**Return type** tuple

## References

- [Bre87]
- [CIET13294]

## Examples

```
>>> from pprint import pprint
>>> pr = corresponding_chromaticities_prediction_CIE1994(2)
>>> pr = [(p.uvp_m, p.uvp_p) for p in pr]
>>> pprint(pr)
[((0.207, 0.486), (0.2133909..., 0.4939794...)),
 ((0.449, 0.511), (0.4450345..., 0.5120939...)),
 ((0.263, 0.505), (0.2693262..., 0.5083212...)),
 ((0.322, 0.545), (0.3308593..., 0.5443940...)),
 ((0.316, 0.537), (0.3225195..., 0.5377826...)),
 ((0.265, 0.553), (0.2709737..., 0.5513666...)),
 ((0.221, 0.538), (0.2280786..., 0.5351592...)),
 ((0.135, 0.532), (0.1439436..., 0.5303576...)),
 ((0.145, 0.472), (0.1500743..., 0.4842895...)),
 ((0.163, 0.331), (0.1559955..., 0.3772379...)),
 ((0.176, 0.431), (0.1806318..., 0.4518475...)),
 ((0.244, 0.349), (0.2454445..., 0.4018004...)])
```

## CMCCAT2000

colour.corresponding

---

`corresponding_chromaticities_prediction_CMCCAT2000([.Re])` Returns the corresponding chromaticities prediction for *CMCCAT2000* chromatic adaptation model.

---

`colour.corresponding.corresponding_chromaticities_prediction_CMCCAT2000`

`colour.corresponding.corresponding_chromaticities_prediction_CMCCAT2000(experiment=1)`  
Returns the corresponding chromaticities prediction for *CMCCAT2000* chromatic adaptation model.

**Parameters** `experiment` (integer, optional) – {1, 2, 3, 4, 6, 8, 9, 11, 12} *Breneman (1987)*  
experiment number.

**Returns** Corresponding chromaticities prediction.

**Return type** tuple

## References

- [\[Bre87\]](#)
- [\[LLRH02\]](#)
- [\[WRC12a\]](#)

## Examples

```
>>> from pprint import pprint
>>> pr = corresponding_chromaticities_prediction_CMCCAT2000(2)
>>> pr = [(p.uvp_m, p.uvp_p) for p in pr]
>>> pprint(pr)
[((0.207, 0.486), (0.2083210..., 0.4727168...)),
 ((0.449, 0.511), (0.4459270..., 0.5077735...)),
 ((0.263, 0.505), (0.2640262..., 0.4955361...)),
 ((0.322, 0.545), (0.3316884..., 0.5431580...)),
 ((0.316, 0.537), (0.3222624..., 0.5357624...)),
 ((0.265, 0.553), (0.2710705..., 0.5501997...)),
 ((0.221, 0.538), (0.2261826..., 0.5294740...)),
 ((0.135, 0.532), (0.1439693..., 0.5190984...)),
 ((0.145, 0.472), (0.1494835..., 0.4556760...)),
 ((0.163, 0.331), (0.1563172..., 0.3164151...)),
 ((0.176, 0.431), (0.1763199..., 0.4127589...)),
 ((0.244, 0.349), (0.2287638..., 0.3499324...)])
```

## Von Kries

`colour.corresponding`


---

`corresponding_chromaticities_prediction_VonKries([Return])` Returns the corresponding chromaticities prediction for *Von Kries* chromatic adaptation model using given transform.

---

`colour.corresponding.corresponding_chromaticities_prediction_VonKries`

`colour.corresponding.corresponding_chromaticities_prediction_VonKries(experiment=1, transform=u'CAT02')`

Returns the corresponding chromaticities prediction for *Von Kries* chromatic adaptation model using

given transform.

#### Parameters

- **experiment** (integer, optional) – {1, 2, 3, 4, 6, 8, 9, 11, 12} *Breneman (1987)* experiment number.
- **transform** (unicode, optional) – {'CAT02', 'XYZ Scaling', 'Von Kries', 'Bradford', 'Sharp', 'Fairchild', 'CMCCAT97', 'CMCCAT2000', 'CAT02\_BRILL\_CAT', 'Bianco', 'Bianco PC'}, Chromatic adaptation transform.

**Returns** Corresponding chromaticities prediction.

**Return type** tuple

#### References

- [\[Bre87\]](#)
- [\[Fai13b\]](#)

#### Examples

```
>>> from pprint import pprint
>>> pr = corresponding_chromaticities_prediction_VonKries(2, 'Bradford')
>>> pr = [(p.uvp_m, p.uvp_p) for p in pr]
>>> pprint(pr)
[((0.207, 0.486), (0.2082014..., 0.4722922...)),
 ((0.449, 0.511), (0.4489102..., 0.5071602...)),
 ((0.263, 0.505), (0.2643545..., 0.4959631...)),
 ((0.322, 0.545), (0.3348730..., 0.5471220...)),
 ((0.316, 0.537), (0.3248758..., 0.5390589...)),
 ((0.265, 0.553), (0.2733105..., 0.5555028...)),
 ((0.221, 0.538), (0.2271480..., 0.5331317...)),
 ((0.135, 0.532), (0.1442730..., 0.5226804...)),
 ((0.145, 0.472), (0.1498745..., 0.4550785...)),
 ((0.163, 0.331), (0.1564975..., 0.3148795...)),
 ((0.176, 0.431), (0.1760593..., 0.4103772...)),
 ((0.244, 0.349), (0.2259805..., 0.3465291...))]
```

#### Colour Difference

- *Delta E*
- *CIE 1976*
- *CIE 1994*
- *CIE 2000*
- *CMC*
- *Luo, Cui and Li (2006)*
- *Li, Li, Wang, Zu, Luo, Cui, Melgosa, Brill and Pointer (2017)*

## Delta E

colour

<code>delta_E(a, b[, method])</code>	Returns the difference $\Delta E_{ab}$ between two given $CIE L^*a^*b^*$ or $J'a'b'$ colourspace arrays using given method.
<code>DELTA_E_METHODS</code>	Supported $\Delta E_{ab}$ computations methods.

### colour.delta\_E

`colour.delta_E(a, b, method='CIE 2000', **kwargs)`

Returns the difference  $\Delta E_{ab}$  between two given  $CIE L^*a^*b^*$  or  $J'a'b'$  colourspace arrays using given method.

#### Parameters

- `a` (`array_like`) –  $CIE L^*a^*b^*$  or  $J'a'b'$  colourspace array  $a$ .
- `b` (`array_like`) –  $CIE L^*a^*b^*$  or  $J'a'b'$  colourspace array  $b$ .
- `method` (`unicode, optional`) – {'CIE 2000', 'CIE 1976', 'CIE 1994', 'CMC', 'CAM02-LCD', 'CAM02-SCD', 'CAM02-UCS', 'CAM16-LCD', 'CAM16-SCD', 'CAM16-UCS'} Computation method.

#### Other Parameters

- `textiles` (`bool, optional`) – {`colour.difference.delta_E_CIE1994()`, `colour.difference.delta_E_CIE2000()`}, Textiles application specific parametric factors  $k_L = 2$ ,  $k_C = k_H = 1$ ,  $k_1 = 0.048$ ,  $k_2 = 0.014$  weights are used instead of  $k_L = k_C = k_H = 1$ ,  $k_1 = 0.045$ ,  $k_2 = 0.015$ .
- `l` (`numeric, optional`) – {`colour.difference.delta_E_CIE2000()`}, Lightness weighting factor.
- `c` (`numeric, optional`) – {`colour.difference.delta_E_CIE2000()`}, Chroma weighting factor.

**Returns** Colour difference  $\Delta E_{ab}$ .

**Return type** numeric or ndarray

## Examples

```
>>> import numpy as np
>>> a = np.array([100.0000000, 21.57210357, 272.22819350])
>>> b = np.array([100.0000000, 426.67945353, 72.39590835])
>>> delta_E(a, b)
94.0356490...
>>> delta_E(a, b, method='CIE 2000')
94.0356490...
>>> delta_E(a, b, method='CIE 1976')
451.7133019...
>>> delta_E(a, b, method='CIE 1994')
83.7792255...
>>> delta_E(a, b, method='CIE 1994', textiles=False)
...
```

```
83.7792255...
>>> a = np.array([54.90433134, -0.08450395, -0.06854831])
>>> b = np.array([54.90433134, -0.08442362, -0.06848314])
>>> delta_E(a, b, method='CAM02-UCS')
0.0001034...
>>> delta_E(a, b, method='CAM16-LCD')
0.0001034...
```

## colour.DELTA\_E\_METHODS

colour.DELTA\_E\_METHODS = CaseInsensitiveMapping({'cie1994': ..., 'CIE 1994': ..., 'CAM02-LCD': ..., 'CMC': ...})  
Supported  $\Delta E_{ab}$  computations methods.

### References

- [\[Lin03b\]](#)
- [\[Lin11\]](#)
- [\[Lin09b\]](#)
- [\[Lin09c\]](#)
- [\[LCL06\]](#)
- [\[Mel13\]](#)
- [\[Wikh\]](#)

DELTA\_E\_METHODS [CaseInsensitiveMapping] {'CIE 1976', 'CIE 1994', 'CIE 2000', 'CMC', 'CAM02-LCD', 'CAM02-SCD', 'CAM02-UCS', 'CAM16-LCD', 'CAM16-SCD', 'CAM16-UCS'}

Aliases:

- 'cie1976': 'CIE 1976'
- 'cie1994': 'CIE 1994'
- 'cie2000': 'CIE 2000'

## CIE 1976

### colour.difference

---

delta\_E\_CIE1976(Lab\_1, Lab\_2)

Returns the difference  $\Delta E_{ab}$  between two given CIE  $L^*a^*b^*$  colourspace arrays using CIE 1976 recommendation.

---

### colour.difference.delta\_E\_CIE1976

colour.difference.delta\_E\_CIE1976(Lab\_1, Lab\_2)

Returns the difference  $\Delta E_{ab}$  between two given CIE  $L^*a^*b^*$  colourspace arrays using CIE 1976 recommendation.

## Parameters

- **Lab\_1** (array\_like) –  $CIE L^*a^*b^*$  colourspace array 1.
- **Lab\_2** (array\_like) –  $CIE L^*a^*b^*$  colourspace array 2.

**Returns** Colour difference  $\Delta E_{ab}$ .

**Return type** numeric or ndarray

## References

- [Lin03b]

## Examples

```
>>> Lab_1 = np.array([100.0000000, 21.57210357, 272.22819350])
>>> Lab_2 = np.array([100.0000000, 426.67945353, 72.39590835])
>>> delta_E_CIE1976(Lab_1, Lab_2)
451.7133019...
```

## CIE 1994

### colour.difference

`delta_E_CIE1994(Lab_1, Lab_2[, textiles])`

Returns the difference  $\Delta E_{ab}$  between two given  $CIE L^*a^*b^*$  colourspace arrays using  $CIE 1994$  recommendation.

### colour.difference.delta\_E\_CIE1994

`colour.difference.delta_E_CIE1994(Lab_1, Lab_2, textiles=False)`

Returns the difference  $\Delta E_{ab}$  between two given  $CIE L^*a^*b^*$  colourspace arrays using  $CIE 1994$  recommendation.

## Parameters

- **Lab\_1** (array\_like) –  $CIE L^*a^*b^*$  colourspace array 1.
- **Lab\_2** (array\_like) –  $CIE L^*a^*b^*$  colourspace array 2.
- **textiles** (bool, optional) – Textiles application specific parametric factors  $k_L = 2$ ,  $k_C = k_H = 1$ ,  $k_1 = 0.048$ ,  $k_2 = 0.014$  weights are used instead of  $k_L = k_C = k_H = 1$ ,  $k_1 = 0.045$ ,  $k_2 = 0.015$ .

**Returns** Colour difference  $\Delta E_{ab}$ .

**Return type** numeric or ndarray

## Notes

- $CIE 1994$  colour differences are not symmetrical: difference between `Lab_1` and `Lab_2` may not be the same as difference between `Lab_2` and `Lab_1` thus one colour must be understood to be the

reference against which a sample colour is compared.

## References

- [\[Lin11\]](#)

## Examples

```
>>> Lab_1 = np.array([100.0000000, 21.57210357, 272.22819350])
>>> Lab_2 = np.array([100.0000000, 426.67945353, 72.39590835])
>>> delta_E_CIE1994(Lab_1, Lab_2)
83.7792255...
>>> delta_E_CIE1994(Lab_1, Lab_2, textiles=True)
88.3355530...
```

## CIE 2000

`colour.difference`

`delta_E_CIE2000(Lab_1, Lab_2[, textiles])`

Returns the difference  $\Delta E_{ab}$  between two given *CIE L\*a\*b\** colourspace arrays using *CIE 2000* recommendation.

`colour.difference.delta_E_CIE2000`

`colour.difference.delta_E_CIE2000(Lab_1, Lab_2, textiles=False)`

Returns the difference  $\Delta E_{ab}$  between two given *CIE L\*a\*b\** colourspace arrays using *CIE 2000* recommendation.

### Parameters

- `Lab_1` (`array_like`) – *CIE L\*a\*b\** colourspace array 1.
- `Lab_2` (`array_like`) – *CIE L\*a\*b\** colourspace array 2.
- `textiles` (`bool`, optional) – Textiles application specific parametric factors  $k_L = 2$ ,  $k_C = k_H = 1$  weights are used instead of  $k_L = k_C = k_H = 1$ .

**Returns** Colour difference  $\Delta E_{ab}$ .

**Return type** numeric or ndarray

## Notes

- *CIE 2000* colour differences are not symmetrical: difference between `Lab_1` and `Lab_2` may not be the same as difference between `Lab_2` and `Lab_1` thus one colour must be understood to be the reference against which a sample colour is compared.
- Parametric factors  $k_L = k_C = k_H = 1$  weights under *reference conditions*:
  - Illumination: D65 source
  - Illuminance: 1000 lx

- Observer: Normal colour vision
- Background field: Uniform, neutral gray with  $L^* = 50$
- Viewing mode: Object
- Sample size: Greater than 4 degrees
- Sample separation: Direct edge contact
- Sample colour-difference magnitude: Lower than 5.0  $\Delta E_{ab}$
- Sample structure: Homogeneous (without texture)

## References

- [\[Lin09b\]](#)
- [\[Mel13\]](#)

## Examples

```
>>> Lab_1 = np.array([100.0000000, 21.57210357, 272.22819350])
>>> Lab_2 = np.array([100.0000000, 426.67945353, 72.39590835])
>>> delta_E_CIE2000(Lab_1, Lab_2)
94.0356490...
>>> Lab_2 = np.array([50.0000000, 426.67945353, 72.39590835])
>>> delta_E_CIE2000(Lab_1, Lab_2)
100.8779470...
>>> delta_E_CIE2000(Lab_1, Lab_2, textiles=True)
95.7920535...
```

## CMC

### colour.difference

---

`delta_E_CMC(Lab_1, Lab_2[, l, c])`

Returns the difference  $\Delta E_{ab}$  between two given CIE  $L^*a^*b^*$  colourspace arrays using Colour Measurement Committee recommendation.

---

### colour.difference.delta\_E\_CMC

`colour.difference.delta_E_CMC(Lab_1, Lab_2, l=2, c=1)`

Returns the difference  $\Delta E_{ab}$  between two given CIE  $L^*a^*b^*$  colourspace arrays using Colour Measurement Committee recommendation.

The quasimetric has two parameters: *Lightness* (*l*) and *chroma* (*c*), allowing the users to weight the difference based on the ratio of *l:c*. Commonly used values are 2:1 for acceptability and 1:1 for the threshold of imperceptibility.

#### Parameters

- `Lab_1` (array\_like) – CIE  $L^*a^*b^*$  colourspace array 1.
- `Lab_2` (array\_like) – CIE  $L^*a^*b^*$  colourspace array 2.

- `l` (numeric, optional) – Lightness weighting factor.
- `c` (numeric, optional) – Chroma weighting factor.

**Returns** Colour difference  $\Delta E_{ab}$ .

**Return type** numeric or ndarray

## References

- [\[Lin09c\]](#)

## Examples

```
>>> Lab_1 = np.array([100.0000000, 21.57210357, 272.22819350])
>>> Lab_2 = np.array([100.0000000, 426.67945353, 72.39590835])
>>> delta_E_CMC(Lab_1, Lab_2)
172.7047712...
```

## Luo, Cui and Li (2006)

### colour.difference

<code>delta_E_CAM02LCD(Jpapbp_1, Jpapbp_2)</code>	Returns the difference $\Delta E'$ between two given <i>Li et alii (2017) CAM16-LCD</i> colourspaces $J'a'b'$ arrays.
<code>delta_E_CAM02SCD(Jpapbp_1, Jpapbp_2)</code>	Returns the difference $\Delta E'$ between two given <i>Li et alii (2017) CAM16-SCD</i> colourspaces $J'a'b'$ arrays.
<code>delta_E_CAM02UCS(Jpapbp_1, Jpapbp_2)</code>	Returns the difference $\Delta E'$ between two given <i>Li et alii (2017) CAM16-UCS</i> colourspaces $J'a'b'$ arrays.

### colour.difference.delta\_E\_CAM02LCD

#### colour.difference.`delta_E_CAM02LCD(Jpapbp_1, Jpapbp_2)`

Returns the difference  $\Delta E'$  between two given *Li et alii (2017) CAM16-LCD* colourspaces  $J'a'b'$  arrays.

##### Parameters

- `Jpapbp_1` (array\_like) – Standard / reference *Li et alii (2017) CAM16-LCD* colourspaces  $J'a'b'$  array.
- `Jpapbp_2` (array\_like) – Sample / test *Li et alii (2017) CAM16-LCD* colourspaces  $J'a'b'$  array.

**Returns** Colour difference  $\Delta E'$ .

**Return type** numeric or ndarray

## References

- [\[LLW+17\]](#)

## Notes

- This docstring is automatically generated, please refer to `colour.delta_E_CAM02LCD()` definition for an usage example.

### `colour.difference.delta_E_CAM02SCD`

`colour.difference.delta_E_CAM02SCD(Jpapbp_1, Jpapbp_2)`

Returns the difference  $\Delta E'$  between two given *Li et alii (2017) CAM16-SCD* colourspaces  $J'a'b'$  arrays.

#### Parameters

- `Jpapbp_1` (array\_like) – Standard / reference *Li et alii (2017) CAM16-SCD* colourspaces  $J'a'b'$  array.
- `Jpapbp_2` (array\_like) – Sample / test *Li et alii (2017) CAM16-SCD* colourspaces  $J'a'b'$  array.

**Returns** Colour difference  $\Delta E'$ .

**Return type** numeric or ndarray

## References

- [\[LLW+17\]](#)

## Notes

- This docstring is automatically generated, please refer to `colour.delta_E_CAM02SCD()` definition for an usage example.

### `colour.difference.delta_E_CAM02UCS`

`colour.difference.delta_E_CAM02UCS(Jpapbp_1, Jpapbp_2)`

Returns the difference  $\Delta E'$  between two given *Li et alii (2017) CAM16-UCS* colourspaces  $J'a'b'$  arrays.

#### Parameters

- `Jpapbp_1` (array\_like) – Standard / reference *Li et alii (2017) CAM16-UCS* colourspaces  $J'a'b'$  array.
- `Jpapbp_2` (array\_like) – Sample / test *Li et alii (2017) CAM16-UCS* colourspaces  $J'a'b'$  array.

**Returns** Colour difference  $\Delta E'$ .

**Return type** numeric or ndarray

## References

- [\[LLW+17\]](#)

## Notes

- This docstring is automatically generated, please refer to `colour.delta_E_CAM02UCS()` definition for an usage example.

**Li, Li, Wang, Zu, Luo, Cui, Melgosa, Brill and Pointer (2017)**

`colour.difference`

<code>delta_E_CAM16LCD(Jpapbp_1, Jpapbp_2)</code>	Returns the difference $\Delta E'$ between two given <i>Li et alii (2017) CAM16-LCD</i> colourspaces $J'a'b'$ arrays.
<code>delta_E_CAM16SCD(Jpapbp_1, Jpapbp_2)</code>	Returns the difference $\Delta E'$ between two given <i>Li et alii (2017) CAM16-SCD</i> colourspaces $J'a'b'$ arrays.
<code>delta_E_CAM16UCS(Jpapbp_1, Jpapbp_2)</code>	Returns the difference $\Delta E'$ between two given <i>Li et alii (2017) CAM16-UCS</i> colourspaces $J'a'b'$ arrays.

`colour.difference.delta_E_CAM16LCD`

`colour.difference.delta_E_CAM16LCD(Jpapbp_1, Jpapbp_2)`

Returns the difference  $\Delta E'$  between two given *Li et alii (2017) CAM16-LCD* colourspaces  $J'a'b'$  arrays.

### Parameters

- `Jpapbp_1` (`array_like`) – Standard / reference *Li et alii (2017) CAM16-LCD* colourspaces  $J'a'b'$  array.
- `Jpapbp_2` (`array_like`) – Sample / test *Li et alii (2017) CAM16-LCD* colourspaces  $J'a'b'$  array.

**Returns** Colour difference  $\Delta E'$ .

**Return type** numeric or ndarray

### References

- [LLW+17]

## Notes

- This docstring is automatically generated, please refer to `colour.delta_E_CAM02LCD()` definition for an usage example.

`colour.difference.delta_E_CAM16SCD`

`colour.difference.delta_E_CAM16SCD(Jpapbp_1, Jpapbp_2)`

Returns the difference  $\Delta E'$  between two given *Li et alii (2017) CAM16-SCD* colourspaces  $J'a'b'$  arrays.

### Parameters

- `Jpapbp_1` (`array_like`) – Standard / reference *Li et alii (2017) CAM16-SCD* colourspaces  $J'a'b'$  array.

- **Jpapbp\_2** (array\_like) – Sample / test *Li et alii (2017)* *CAM16-SCD* colourespaces  $J'a'b'$  array.

**Returns** Colour difference  $\Delta E'$ .

**Return type** numeric or ndarray

## References

- [\[LLW+17\]](#)

## Notes

- This docstring is automatically generated, please refer to `colour.delta_E_CAM02SCD()` definition for an usage example.

## colour.difference.delta\_E\_CAM16UCS

`colour.difference.delta_E_CAM16UCS(Jpapbp_1, Jpapbp_2)`

Returns the difference  $\Delta E'$  between two given *Li et alii (2017)* *CAM16-UCS* colourespaces  $J'a'b'$  arrays.

### Parameters

- **Jpapbp\_1** (array\_like) – Standard / reference *Li et alii (2017)* *CAM16-UCS* colourespaces  $J'a'b'$  array.
- **Jpapbp\_2** (array\_like) – Sample / test *Li et alii (2017)* *CAM16-UCS* colourespaces  $J'a'b'$  array.

**Returns** Colour difference  $\Delta E'$ .

**Return type** numeric or ndarray

## References

- [\[LLW+17\]](#)

## Notes

- This docstring is automatically generated, please refer to `colour.delta_E_CAM02UCS()` definition for an usage example.

## Input and Output

- *Image Data*
- *CSV Tabular Data*
- *IES TM-27-14 Data*
- *X-Rite Data*

## Image Data

colour

<code>read_image(path[, bit_depth])</code>	Reads given image using <i>OpenImageIO</i> .
<code>write_image(image, path[, bit_depth])</code>	Writes given image using <i>OpenImageIO</i> .

### colour.read\_image

`colour.read_image(path, bit_depth=u'float32')`  
Reads given image using *OpenImageIO*.

#### Parameters

- **path** (unicode) – Image path.
- **bit\_depth** (unicode, optional) – {‘float32’, ‘uint8’, ‘uint16’, ‘float16’}, Image bit\_depth.

**Returns** Image as a ndarray.

**Return type** ndarray

#### Notes

- For convenience, single channel images are squeezed to 2d arrays.

### Examples

```
>>> import os
>>> path = os.path.join('tests', 'resources', 'CMSTestPattern.exr')
>>> image = read_image(path)
```

### colour.write\_image

`colour.write_image(image, path, bit_depth=u'float32')`  
Writes given image using *OpenImageIO*.

#### Parameters

- **image** (array\_like) – Image data.
- **path** (unicode) – Image path.
- **bit\_depth** (unicode, optional) – {‘float32’, ‘uint8’, ‘uint16’, ‘float16’}, Image bit\_depth.

**Returns** Definition success.

**Return type** bool

## Examples

```
>>> import os
>>> path = os.path.join('tests', 'resources', 'CMSTestPattern.exr')
>>> image = read_image(path)
>>> path = os.path.join('tests', 'resources', 'CMSTestPattern.png')
>>> write_image(image, path, 'uint8')
True
```

## CSV Tabular Data

colour

<code>read_spds_from_csv_file(path[, delimiter, ...])</code>	Reads the spectral data from given CSV file and return its content as an <i>OrderedDict</i> of <code>colour.SpectralPowerDistribution</code> classes.
<code>read_spectral_data_from_csv_file(path[, ...])</code>	Reads the spectral data from given CSV file in the following form:
<code>write_spds_to_csv_file(spds, path[, ...])</code>	Writes the given spectral power distributions to given CSV file.

`colour.read_spds_from_csv_file`

`colour.read_spds_from_csv_file(path, delimiter=u',', fields=None, default=0)`

Reads the spectral data from given CSV file and return its content as an *OrderedDict* of `colour.SpectralPowerDistribution` classes.

### Parameters

- `path` (unicode) – Absolute CSV file path.
- `delimiter` (unicode, optional) – CSV file content delimiter.
- `fields` (array\_like, optional) – CSV file spectral data fields names. If no value is provided the first line of the file will be used for as spectral data fields names.
- `default` (numeric) – Default value for fields row with missing value.

`Returns` `colour.SpectralPowerDistribution` classes of given CSV file.

`Return type` `OrderedDict`

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> import os
>>> csv_file = os.path.join(os.path.dirname(__file__), 'tests',
...                         'resources', 'colorchecker_n_ohta.csv')
>>> spds = read_spds_from_csv_file(csv_file)
>>> print(tuple(spds.keys()))
('1', '2', '3', '4', '5', '6', '7', '8', '9', '10', '11', '12', '13', '14', '15', '16', '17', '18',
 ↵, '19', '20', '21', '22', '23', '24')
>>> with numpy_print_options(suppress=True):
```

```
...     spds['1']
SpectralPowerDistribution([[ 380. , 0.048],
                           [ 385. , 0.051],
                           [ 390. , 0.055],
                           [ 395. , 0.06 ],
                           [ 400. , 0.065],
                           [ 405. , 0.068],
                           [ 410. , 0.068],
                           [ 415. , 0.067],
                           [ 420. , 0.064],
                           [ 425. , 0.062],
                           [ 430. , 0.059],
                           [ 435. , 0.057],
                           [ 440. , 0.055],
                           [ 445. , 0.054],
                           [ 450. , 0.053],
                           [ 455. , 0.053],
                           [ 460. , 0.052],
                           [ 465. , 0.052],
                           [ 470. , 0.052],
                           [ 475. , 0.053],
                           [ 480. , 0.054],
                           [ 485. , 0.055],
                           [ 490. , 0.057],
                           [ 495. , 0.059],
                           [ 500. , 0.061],
                           [ 505. , 0.062],
                           [ 510. , 0.065],
                           [ 515. , 0.067],
                           [ 520. , 0.07 ],
                           [ 525. , 0.072],
                           [ 530. , 0.074],
                           [ 535. , 0.075],
                           [ 540. , 0.076],
                           [ 545. , 0.078],
                           [ 550. , 0.079],
                           [ 555. , 0.082],
                           [ 560. , 0.087],
                           [ 565. , 0.092],
                           [ 570. , 0.1 ],
                           [ 575. , 0.107],
                           [ 580. , 0.115],
                           [ 585. , 0.122],
                           [ 590. , 0.129],
                           [ 595. , 0.134],
                           [ 600. , 0.138],
                           [ 605. , 0.142],
                           [ 610. , 0.146],
                           [ 615. , 0.15 ],
                           [ 620. , 0.154],
                           [ 625. , 0.158],
                           [ 630. , 0.163],
                           [ 635. , 0.167],
                           [ 640. , 0.173],
                           [ 645. , 0.18 ],
                           [ 650. , 0.188],
                           [ 655. , 0.196],
                           [ 660. , 0.204],
```

```
[ 665. , 0.213],
[ 670. , 0.222],
[ 675. , 0.231],
[ 680. , 0.242],
[ 685. , 0.251],
[ 690. , 0.261],
[ 695. , 0.271],
[ 700. , 0.282],
[ 705. , 0.294],
[ 710. , 0.305],
[ 715. , 0.318],
[ 720. , 0.334],
[ 725. , 0.354],
[ 730. , 0.372],
[ 735. , 0.392],
[ 740. , 0.409],
[ 745. , 0.42 ],
[ 750. , 0.436],
[ 755. , 0.45 ],
[ 760. , 0.462],
[ 765. , 0.465],
[ 770. , 0.448],
[ 775. , 0.432],
[ 780. , 0.421]],
interpolator=SpragueInterpolator,
interpolator_args={},
extrapolator=Extrapolator,
extrapolator_args={...})
```

**colour.read\_spectral\_data\_from\_csv\_file****colour.read\_spectral\_data\_from\_csv\_file(path, delimiter=u',', fields=None, default=0)**

Reads the spectral data from given CSV file in the following form:

390, 4.15003E-04, 3.68349E-04, 9.54729E-03 395, 1.05192E-03, 9.58658E-04, 2.38250E-02 400,  
2.40836E-03, 2.26991E-03, 5.66498E-02 ... 830, 9.74306E-07, 9.53411E-08, 0.00000and returns it as an *OrderedDict* of *dict* as follows:

OrderedDict([ ('field', {'wavelength': 'value', ..., 'wavelength': 'value'}), ..., ('field', {'wavelength': 'value', ..., 'wavelength': 'value'})])

**Parameters**

- **path** (unicode) – Absolute CSV file path.
- **delimiter** (unicode, optional) – CSV file content delimiter.
- **fields** (array\_like, optional) – CSV file spectral data fields names. If no value is provided the first line of the file will be used as spectral data fields names.
- **default** (numeric, optional) – Default value for fields row with missing value.

**Returns** CSV file content.**Return type** OrderedDict**Raises** RuntimeError – If the CSV spectral data file doesn't define the appropriate fields.

## Notes

- A CSV spectral data file should define at least define two fields: one for the wavelengths and one for the associated values of one spectral power distribution.
- If no value is provided for the fields names, the first line of the file will be used as spectral data fields names.

## Examples

```
>>> import os
>>> from pprint import pprint
>>> csv_file = os.path.join(os.path.dirname(__file__), 'tests',
...                         'resources', 'colorchecker_n_ohta.csv')
>>> spds_data = read_spectral_data_from_csv_file(csv_file)
>>> pprint(list(spds_data.keys()))
['1',
 '2',
 '3',
 '4',
 '5',
 '6',
 '7',
 '8',
 '9',
 '10',
 '11',
 '12',
 '13',
 '14',
 '15',
 '16',
 '17',
 '18',
 '19',
 '20',
 '21',
 '22',
 '23',
 '24']
```

## colour.write\_spds\_to\_csv\_file

colour.**write\_spds\_to\_csv\_file**(*spds*, *path*, *delimiter*=u',', *fields*=None)

Writes the given spectral power distributions to given CSV file.

### Parameters

- **spds** (`dict`) – Spectral power distributions to write.
- **path** (`unicode`) – Absolute CSV file path.
- **delimiter** (`unicode`, optional) – CSV file content delimiter.
- **fields** (`array_like`, optional) – CSV file spectral data fields names. If no value is provided the order of fields will be the one defined by the sorted spectral power

distributions *dict*.

**Returns** Definition success.

**Return type** `bool`

**Raises** `RuntimeError` – If the given spectral power distributions have different shapes.

## IES TM-27-14 Data

colour

---

`IES_TM2714_Spd([path, header, ...])`

Defines a *IES TM-27-14* spectral power distribution.

---

### colour.IES\_TM2714\_Spd

```
class colour.IES_TM2714_Spd(path=None, header=None, spectral_quantity=None, reflection_geometry=None, transmission_geometry=None, width_FWHM=None, bandwidth_corrected=None)
```

Defines a *IES TM-27-14* spectral power distribution.

This class can read and write *IES TM-27-14* spectral data XML files.

#### Parameters

- **path** (`unicode`, optional) – Spectral data XML file path.
- **header** (`IES_TM2714_Header`, optional) – *IES TM-27-14* spectral power distribution header.
- **spectral\_quantity** (`unicode`, optional) – {‘flux’, ‘absorptance’, ‘transmittance’, ‘reflectance’, ‘intensity’, ‘irradiance’, ‘radiance’, ‘exitance’, ‘R-Factor’, ‘T-Factor’, ‘relative’, ‘other’}, Quantity of measurement for each element of the spectral data.
- **reflection\_geometry** (`unicode`, optional) – {‘di:8’, ‘de:8’, ‘8:di’, ‘8:de’, ‘d:d’, ‘d:0’, ‘45a:0’, ‘45c:0’, ‘0:45a’, ‘45x:0’, ‘0:45x’, ‘other’}, Spectral reflectance factors geometric conditions.
- **transmission\_geometry** (`unicode`, optional) – {‘0:0’, ‘di:0’, ‘de:0’, ‘0:di’, ‘0:de’, ‘d:d’, ‘other’}, Spectral transmittance factors geometric conditions.
- **bandwidth\_FWHM** (`numeric`, optional) – Spectroradiometer full-width half-maximum bandwidth in nanometers.
- **bandwidth\_corrected** (`bool`, optional) – Specifies if bandwidth correction has been applied to the measured data.

#### Notes

##### Reflection Geometry

- di:8: Diffuse / eight-degree, specular component included.
- de:8: Diffuse / eight-degree, specular component excluded.
- 8:di: Eight-degree / diffuse, specular component included.
- 8:de: Eight-degree / diffuse, specular component excluded.

- d:d: Diffuse / diffuse.
- d:0: Alternative diffuse.
- 45a:0: Forty-five degree annular / normal.
- 45c:0: Forty-five degree circumferential / normal.
- 0:45a: Normal / forty-five degree annular.
- 45x:0: Forty-five degree directional / normal.
- 0:45x: Normal / forty-five degree directional.
- other: User-specified in comments.

#### *Transmission Geometry*

- 0:0: Normal / normal.
- di:0: Diffuse / normal, regular component included.
- de:0: Diffuse / normal, regular component excluded.
- 0:di: Normal / diffuse, regular component included.
- 0:de: Normal / diffuse, regular component excluded.
- d:d: Diffuse / diffuse.
- other: User-specified in comments.

**mapping**

**path**

**header**

**spectral\_quantity**

**reflection\_geometry**

**transmission\_geometry**

**bandwidth\_FWHM**

**bandwidth\_corrected**

**read()**

**write()**

## References

- [\[IESCCommitteeTM2714WGroup14\]](#)

## Examples

```
>>> from os.path import dirname, join
>>> directory = join(dirname(__file__), 'tests', 'resources')
>>> spd = IES_TM2714_Spd(join(directory, 'Fluorescent.spdx'))
>>> spd.read()
True
>>> spd.header.manufacturer
```

```
'Unknown'
>>> # Doctests ellipsis for Python 2.x compatibility.
>>> spd[501.7]
0.0950000...
```

`__init__(path=None, header=None, spectral_quantity=None, reflection_geometry=None, transmission_geometry=None, bandwidth_FWHM=None, bandwidth_corrected=None)`

## Methods

<code>__init__([path, header, spectral_quantity, ...])</code>	
<code>align(shape[, interpolator, ...])</code>	Aligns the spectral power distribution in-place to given spectral shape: Interpolates first then extrapolates to fit the given range.
<code>arithmetical_operation(a, operation[, in_place])</code>	Performs given arithmetical operation with <i>a</i> operand, the operation can be either performed on a copy or in-place.
<code>clone()</code>	
<code>copy()</code>	Returns a copy of the sub-class instance, must be reimplemented by sub-classes.
<code>domain_distance(a)</code>	Returns the euclidean distance between given array and independent domain <i>x</i> closest element.
<code>extrapolate(shape[, extrapolator, ...])</code>	Extrapolates the spectral power distribution in-place according to <i>CIE 15:2004</i> and <i>CIE 167:2005</i> recommendations or given extrapolation arguments.
<code>fill_nan([method, default])</code>	Fill NaNs in independent domain <i>x</i> variable and corresponding range <i>y</i> variable using given method.
<code>get()</code>	
<code>interpolate(shape[, interpolator, ...])</code>	Interpolates the spectral power distribution in-place according to <i>CIE 167:2005</i> recommendation or given interpolation arguments.
<code>is_uniform()</code>	Returns if independent domain <i>x</i> variable is uniform.
<code>normalise([factor])</code>	Normalises the spectral power distribution using given normalization factor.
<code>read()</code>	Reads and parses the spectral data XML file path.
<code>signal_unpack_data([data, domain, dtype])</code>	Unpack given data for continuous signal instantiation.
<code>to_series()</code>	Converts the continuous signal to a <i>Pandas Series</i> class instance.
<code>trim(shape)</code>	Trims the spectral power distribution wavelengths to given spectral shape.
<code>trim_wavelengths(shape)</code>	
<code>write()</code>	Write the spd spectral data to XML file path.
<code>zeros()</code>	

## X-Rite Data

colour

---

```
read_spds_from_xrite_file(path)
```

Reads the spectral data from given *X-Rite* file and returns it as an *OrderedDict* of `colour.SpectralPowerDistribution` classes.

---

`colour.read_spds_from_xrite_file`

```
colour.read_spds_from_xrite_file(path)
```

Reads the spectral data from given *X-Rite* file and returns it as an *OrderedDict* of `colour.SpectralPowerDistribution` classes.

**Parameters** `path` (unicode) – Absolute *X-Rite* file path.

**Returns** `colour.SpectralPowerDistribution` classes of given *X-Rite* file.

**Return type** `OrderedDict`

## Notes

- This parser is minimalistic and absolutely not bullet proof.

## Examples

```
>>> import os
>>> from pprint import pprint
>>> xrite_file = os.path.join(os.path.dirname(__file__), 'tests',
...                             'resources',
...                             'xrite_digital_colour_checker.txt')
>>> spds_data = read_spds_from_xrite_file(xrite_file)
>>> pprint(list(spds_data.keys()))
['X1', 'X2', 'X3', 'X4', 'X5', 'X6', 'X7', 'X8', 'X9', 'X10']
```

## Colour Models

- *Tristimulus Values, CIE xyY Colourspace and Chromaticity Coordinates*
- *CIE L\*a\*b\* Colourspace*
- *CIE L\*u\*v\* Colourspace*
- *CIE UCS Colourspace*
- *CIE 1964 U\*V\*W\* Colourspace*
- *Hunter L,a,b Colour Scale*
- *Hunter Rd,a,b Colour Scale*
- *Luo, Cui and Li (2006)*
- *Li, Li, Wang, Zu, Luo, Cui, Melgosa, Brill and Pointer (2017)*
- *IPT Colourspace*

- *hdr-IPT Colourspace*
- *hdr-CIELAB Colourspace*
- *RGB Colourspace and Transformations*
  - *RGB Colourspace Derivation*
  - *RGB Colourspaces*
  - *Opto-Electronic Transfer Functions*
  - *Electro-Optical Transfer Functions*
  - *Opto-Optical Transfer Functions*
  - *Log Encoding and Decoding Curves*
  - *Colour Encodings*
    - \* *Y'CbCr Colour Encoding*
    - \* *IC<sub>T</sub>C<sub>P</sub> Colour Encoding*
  - *RGB Representations*
    - \* *Prismatic Colourspace*
    - \* *HSV Colourspace*
    - \* *HSL Colourspace*
    - \* *CMY Colourspace*
- *Pointer's Gamut*

## Tristimulus Values, CIE xyY Colourspace and Chromaticity Coordinates

### colour

<code>XYZ_to_xyY(XYZ[, illuminant])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>CIE xyY</i> colourspace and reference <i>illuminant</i> .
<code>xyY_to_XYZ(xyY)</code>	Converts from <i>CIE xyY</i> colourspace to <i>CIE XYZ</i> tristimulus values.
<code>XYZ_to_xy(XYZ[, illuminant])</code>	Returns the <i>xy</i> chromaticity coordinates from given <i>CIE XYZ</i> tristimulus values.
<code>xy_to_XYZ(xy)</code>	Returns the <i>CIE XYZ</i> tristimulus values from given <i>xy</i> chromaticity coordinates.
<code>xyY_to_xy(xyY)</code>	Converts from <i>CIE xyY</i> colourspace to <i>xy</i> chromaticity coordinates.
<code>xy_to_xyY(xy[, Y])</code>	Converts from <i>xy</i> chromaticity coordinates to <i>CIE xyY</i> colourspace by extending the array last dimension with <i>Y</i> Luminance.

### colour.XYZ\_to\_xyY

`colour.XYZ_to_xyY(XYZ, illuminant=array([ 0.3457, 0.3585]))`

Converts from *CIE XYZ* tristimulus values to *CIE xyY* colourspace and reference *illuminant*.

### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values.
- `illuminant` (array\_like, optional) – Reference *illuminant* chromaticity coordinates.

**Returns** *CIE xyY* colourspace array.

**Return type** ndarray

### Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].
- Output *CIE xyY* colourspace array is in range [0, 1].

### References

- [\[Lin03c\]](#)
- [\[Wikc\]](#)

### Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_to_xyY(XYZ)
array([ 0.2641477...,  0.3777000...,  0.1008     ])
```

## colour.xyY\_to\_XYZ

`colour.xyY_to_XYZ(xyY)`

Converts from *CIE xyY* colourspace to *CIE XYZ* tristimulus values.

**Parameters** `xyY` (array\_like) – *CIE xyY* colourspace array.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

### Notes

- Input *CIE xyY* colourspace array is in domain [0, 1].
- Output *CIE XYZ* tristimulus values are in range [0, 1].

### References

- [\[Lin09d\]](#)
- [\[Wikc\]](#)

## Examples

```
>>> xyY = np.array([0.26414772, 0.37770001, 0.10080000])
>>> xyY_to_XYZ(xyY)
array([ 0.0704953...,  0.1008      ,  0.0955831...])
```

## colour.XYZ\_to\_xy

`colour.XYZ_to_xy(XYZ, illuminant=array([ 0.3457, 0.3585]))`

Returns the *xy* chromaticity coordinates from given *CIE XYZ* tristimulus values.

### Parameters

- **XYZ** (array\_like) – *CIE XYZ* tristimulus values.
- **illuminant** (array\_like, optional) – Reference *illuminant* chromaticity coordinates.

**Returns** *xy* chromaticity coordinates.

**Return type** ndarray

### Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].
- Output *xy* chromaticity coordinates are in range [0, 1].

## References

- [\[Wikc\]](#)

## Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_to_xy(XYZ)
array([ 0.2641477...,  0.3777000...])
```

## colour.xy\_to\_XYZ

`colour.xy_to_XYZ(xy)`

Returns the *CIE XYZ* tristimulus values from given *xy* chromaticity coordinates.

**Parameters** **xy** (array\_like) – *xy* chromaticity coordinates.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

## Notes

- Input  $xy$  chromaticity coordinates are in domain [0, 1].
- Output  $CIE XYZ$  tristimulus values are in range [0, 1].

## References

- [\[Wikc\]](#)

## Examples

```
>>> xy = np.array([0.26414772, 0.37770001])
>>> xy_to_XYZ(xy)
array([ 0.6993585...,  1.          ,  0.9482453...])
```

## colour.xyY\_to\_xy

`colour.xyY_to_xy(xyY)`

Converts from  $CIE xyY$  colourspace to  $xy$  chromaticity coordinates.

$xyY$  argument with last dimension being equal to 2 will be assumed to be a  $xy$  chromaticity coordinates argument and will be returned directly by the definition.

**Parameters** `xyY` (`array_like`) –  $CIE xyY$  colourspace array or  $xy$  chromaticity coordinates.

**Returns**  $xy$  chromaticity coordinates.

**Return type** `ndarray`

## Notes

- Input  $CIE xyY$  colourspace array is in domain [0, 1].
- Output  $xy$  chromaticity coordinates are in range [0, 1].

## References

- [\[Wikc\]](#)

## Examples

```
>>> xyY = np.array([0.26414772, 0.37770001, 0.10080000])
>>> xyY_to_xy(xyY)
array([ 0.2641477...,  0.3777000...])
>>> xy = np.array([0.26414772, 0.37770001])
>>> xyY_to_xy(xy)
array([ 0.2641477...,  0.3777000...])
```

## colour.xy\_to\_xyY

`colour.xy_to_xyY(xy, Y=1)`

Converts from  $xy$  chromaticity coordinates to *CIE xyY* colourspace by extending the array last dimension with  $Y$  Luminance.

$xy$  argument with last dimension being equal to 3 will be assumed to be a *CIE xyY* colourspace array argument and will be returned directly by the definition.

### Parameters

- $xy$  (array\_like) –  $xy$  chromaticity coordinates or *CIE xyY* colourspace array.
- $Y$  (numeric, optional) – Optional  $Y$  Luminance value used to construct the *CIE xyY* colourspace array, otherwise the  $Y$  Luminance will be set to 1.

**Returns** *CIE xyY* colourspace array.

**Return type** ndarray

### Notes

- This definition is a convenient object provided to implement support of illuminant argument *luminance* value in various colour.models package objects such as `colour.Lab_to_XYZ()` or `colour.Luv_to_XYZ()`.
- Input  $xy$  chromaticity coordinates are in domain [0, 1].
- Output *CIE xyY* colourspace array is in range [0, 1].

### References

- [\[Wikc\]](#)

### Examples

```
>>> xy = np.array([0.26414772, 0.37770001])
>>> xy_to_xyY(xy)
array([ 0.2641477...,  0.3777000...,  1.        ])
>>> xy = np.array([0.26414772, 0.37770001, 0.10080000])
>>> xy_to_xyY(xy)
array([ 0.2641477...,  0.3777000...,  0.1008...])
>>> xy = np.array([0.26414772, 0.37770001])
>>> xy_to_xyY(xy, 100)
array([ 0.2641477...,  0.3777000...,  100.       ])
```

## CIE L\*a\*b\* Colourspace

colour

`XYZ_to_Lab(XYZ[, illuminant])`

Converts from *CIE XYZ* tristimulus values to *CIE L\*a\*b\** colourspace.

Continued on next page

Table 3.125 – continued from previous page

<code>Lab_to_XYZ(Lab[, illuminant])</code>	Converts from <i>CIE L*a*b*</i> colourspace to <i>CIE XYZ</i> tristimulus values.
<code>Lab_to_LCHab(Lab)</code>	Converts from <i>CIE L*a*b*</i> colourspace to <i>CIE L*C*Hab</i> colourspace.
<code>LCHab_to_Lab(LCHab)</code>	Converts from <i>CIE L*C*Hab</i> colourspace to <i>CIE L*a*b*</i> colourspace.

## colour.XYZ\_to\_Lab

`colour.XYZ_to_Lab(XYZ, illuminant=array([ 0.3457, 0.3585]))`  
Converts from *CIE XYZ* tristimulus values to *CIE L\*a\*b\** colourspace.

### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `illuminant` (`array_like`, optional) – Reference *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.

**Returns** *CIE L\*a\*b\** colourspace array.

**Return type** `ndarray`

### Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].
- Input *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Output *Lightness L\** is in range [0, 100].

### References

- [\[CIET14804f\]](#)

### Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_to_Lab(XYZ)
array([ 37.9856291..., -23.6290768..., -4.4174661...])
```

## colour.Lab\_to\_XYZ

`colour.Lab_to_XYZ(Lab, illuminant=array([ 0.3457, 0.3585]))`  
Converts from *CIE L\*a\*b\** colourspace to *CIE XYZ* tristimulus values.

### Parameters

- `Lab` (`array_like`) – *CIE L\*a\*b\** colourspace array.
- `illuminant` (`array_like`, optional) – Reference *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

## Notes

- Input *Lightness L\** is in domain [0, 100].
- Input *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Output *CIE XYZ* tristimulus values are in range [0, 1].

## References

- [\[CIET14804f\]](#)

## Examples

```
>>> Lab = np.array([37.98562910, -23.62907688, -4.41746615])
>>> Lab_to_XYZ(Lab)
array([ 0.0704953...,  0.1008     ,  0.0955831...])
```

## colour.Lab\_to\_LCHab

`colour.Lab_to_LCHab(Lab)`

Converts from *CIE L\*a\*b\** colourspace to *CIE L\*C\*Hab* colourspace.

**Parameters** `Lab` (array\_like) – *CIE L\*a\*b\** colourspace array.

**Returns** *CIE L\*C\*Hab* colourspace array.

**Return type** ndarray

## Notes

- *Lightness L\** is in domain [0, 100].

## References

- [\[CIET14804f\]](#)

## Examples

```
>>> Lab = np.array([37.98562910, -23.62907688, -4.41746615])
>>> Lab_to_LCHab(Lab)
array([ 37.9856291...,  24.0384542..., 190.5892337...])
```

## colour.LCHab\_to\_Lab

`colour.LCHab_to_Lab(LCHab)`

Converts from *CIE L\*C\*Hab* colourspace to *CIE L\*a\*b\** colourspace.

**Parameters** `LCHab` (`array_like`) – *CIE L\*C\*Hab* colourspace array.

**Returns** *CIE L\*a\*b\** colourspace array.

**Return type** `ndarray`

### Notes

- *Lightness L\** is in domain [0, 100].

### References

- [\[CIET14804f\]](#)

### Examples

```
>>> LCHab = np.array([37.98562910, 24.03845422, 190.58923377])
>>> LCHab_to_Lab(LCHab)
array([-37.9856291..., -23.6290768..., -4.4174661...])
```

## CIE L\*u\*v\* Colourspace

`colour`

<code>XYZ_to_Luv(XYZ[, illuminant])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>CIE L*u*v*</i> colourspace.
<code>Luv_to_XYZ(Luv[, illuminant])</code>	Converts from <i>CIE L*u*v*</i> colourspace to <i>CIE XYZ</i> tristimulus values.
<code>Luv_to_LChuv(Luv)</code>	Converts from <i>CIE L*u*v*</i> colourspace to <i>CIE L*C*Huv</i> colourspace.
<code>LChuv_to_Luv(LChuv)</code>	Converts from <i>CIE L*C*Huv</i> colourspace to <i>CIE L*u*v*</i> colourspace.
<code>Luv_to_uv(Luv[, illuminant])</code>	Returns the $uv^p$ chromaticity coordinates from given <i>CIE L*u*v*</i> colourspace array.
<code>Luv_uv_to_xy(uv)</code>	Returns the $xy$ chromaticity coordinates from given <i>CIE L*u*v*</i> colourspace $uv^p$ chromaticity coordinates.

## colour.XYZ\_to\_Luv

`colour.XYZ_to_Luv(XYZ, illuminant=array([ 0.3457, 0.3585]))`

Converts from *CIE XYZ* tristimulus values to *CIE L\*u\*v\** colourspace.

**Parameters**

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.

- **illuminant** (array\_like, optional) – Reference *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.

**Returns** *CIE L\*u\*v\** colourspace array.

**Return type** ndarray

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].
- Input *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Output  $L^*$  is in range [0, 100].

## References

- [\[CIET14804f\]](#)
- [\[Wikg\]](#)

## Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_to_Luv(XYZ)
array([ 37.9856291..., -28.8021959..., -1.3580070...])
```

## colour.Luv\_to\_XYZ

colour.Luv\_to\_XYZ(*Luv*, *illuminant*=array([0.3457, 0.3585]))

Converts from *CIE L\*u\*v\** colourspace to *CIE XYZ* tristimulus values.

### Parameters

- **Luv** (array\_like) – *CIE L\*u\*v\** colourspace array.
- **illuminant** (array\_like, optional) – Reference *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

## Notes

- Input  $L^*$  is in domain [0, 100].
- Input *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Output *CIE XYZ* tristimulus values are in range [0, 1].

## References

- [\[CIET14804f\]](#)
- [\[Wikg\]](#)

## Examples

```
>>> Luv = np.array([37.9856291, -28.80219593, -1.35800706])
>>> Luv_to_XYZ(Luv)
array([ 0.0704953...,  0.1008     ,  0.0955831...])
```

## colour.Luv\_to\_LChuv

### colour.Luv\_to\_LChuv(Luv)

Converts from  $CIE L^*u^*v^*$  colourspace to  $CIE L^*C^*Huv$  colourspace.

**Parameters** `Luv` (`array_like`) –  $CIE L^*u^*v^*$  colourspace array.

**Returns**  $CIE L^*C^*Huv$  colourspace array.

**Return type** ndarray

## Notes

- Input / output  $L^*$  is in domain / range [0, 100].

## References

- [\[CIET14804f\]](#)

## Examples

```
>>> Luv = np.array([37.9856291, -28.80219593, -1.35800706])
>>> Luv_to_LChuv(Luv)
array([ 37.9856291...,  28.8341927...,  182.6994640...])
```

## colour.LChuv\_to\_Luv

### colour.LChuv\_to\_Luv(LChuv)

Converts from  $CIE L^*C^*Huv$  colourspace to  $CIE L^*u^*v^*$  colourspace.

**Parameters** `LChuv` (`array_like`) –  $CIE L^*C^*Huv$  colourspace array.

**Returns**  $CIE L^*u^*v^*$  colourspace array.

**Return type** ndarray

## Notes

- Input / output  $L^*$  is in domain / range [0, 100].

## References

- [\[CIET14804f\]](#)

## Examples

```
>>> LCHuv = np.array([37.98562910, 28.83419279, 182.69946404])
>>> LCHuv_to_Luv(LCHuv)
array([ 37.9856291..., -28.8021959..., -1.3580070...])
```

## colour.Luv\_to\_uv

`colour.Luv_to_uv(Luv, illuminant=array([ 0.3457, 0.3585]))`

Returns the  $uv^p$  chromaticity coordinates from given CIE  $L^*u^*v^*$  colourspace array.

### Parameters

- `Luv` (array\_like) – CIE  $L^*u^*v^*$  colourspace array.
- `illuminant` (array\_like, optional) – Reference illuminant  $xy$  chromaticity coordinates or CIE  $xyY$  colourspace array.

**Returns**  $uv^p$  chromaticity coordinates.

**Return type** ndarray

## Notes

- Input  $L^*$  is in domain [0, 100].
- Input `illuminant`  $xy$  chromaticity coordinates or CIE  $xyY$  colourspace array are in domain  $[0, \infty]$ .
- Output  $uv^p$  chromaticity coordinates are in range [0, 1].

## References

- [\[CIET14804e\]](#)

## Examples

```
>>> Luv = np.array([37.9856291 , -28.80219593, -1.35800706])
>>> Luv_to_uv(Luv)
array([ 0.1508531...,  0.4853297...])
```

## colour.Luv\_uv\_to\_xy

`colour.Luv_uv_to_xy(uv)`

Returns the  $xy$  chromaticity coordinates from given  $CIE L^*u^*v^*$  colourspace  $uv^p$  chromaticity coordinates.

**Parameters** `uv` (`array_like`) –  $CIE L^*u^*v^*$   $u''v''$  chromaticity coordinates.

**Returns**  $xy$  chromaticity coordinates.

**Return type** `ndarray`

### Notes

- Input  $uv^p$  chromaticity coordinates are in domain [0, 1].
- Output  $xy$  is in range [0, 1].

### References

- [\[Wikx\]](#)

### Examples

```
>>> uv = np.array([0.150853098829857, 0.485329708543180])
>>> Luv_uv_to_xy(uv)
array([ 0.2641477...,  0.3777000...])
```

## CIE UCS Colourspace

### colour

<code>XYZ_to_UCS(XYZ)</code>	Converts from $CIE XYZ$ tristimulus values to $CIE UCS$ colourspace.
<code>UCS_to_XYZ(UVW)</code>	Converts from $CIE UCS$ colourspace to $CIE XYZ$ tristimulus values.
<code>UCS_to_uv(UVW)</code>	Returns the $uv$ chromaticity coordinates from given $CIE UCS$ colourspace array.
<code>UCS_uv_to_xy(uv)</code>	Returns the $xy$ chromaticity coordinates from given $CIE UCS$ colourspace $uv$ chromaticity coordinates.

### colour.XYZ\_to\_UCS

`colour.XYZ_to_UCS(XYZ)`

Converts from  $CIE XYZ$  tristimulus values to  $CIE UCS$  colourspace.

**Parameters** `XYZ` (`array_like`) –  $CIE XYZ$  tristimulus values.

**Returns**  $CIE UCS$  colourspace array.

**Return type** `ndarray`

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].
- Output *CIE UCS* colourspace array is in range [0, 1].

## References

- [\[Wikv\]](#)
- [\[Wikd\]](#)

## Examples

```
>>> import numpy as np
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_to_UCS(XYZ)
array([ 0.0469968...,  0.1008      ,  0.1637439...])
```

## colour.UCS\_to\_XYZ

`colour.UCS_to_XYZ(UVW)`

Converts from *CIE UCS* colourspace to *CIE XYZ* tristimulus values.

**Parameters** `UVW` (`array_like`) – *CIE UCS* colourspace array.

**Returns** *CIE XYZ* tristimulus values.

**Return type** `ndarray`

## Notes

- Input *CIE UCS* colourspace array is in domain [0, 1].
- Output *CIE XYZ* tristimulus values are in range [0, 1].

## References

- [\[Wikv\]](#)
- [\[Wikd\]](#)

## Examples

```
>>> import numpy as np
>>> UVW = np.array([0.04699689, 0.10080000, 0.16374390])
>>> UCS_to_XYZ(UVW)
array([ 0.0704953...,  0.1008      ,  0.0955831...])
```

## colour.UCS\_to\_uv

`colour.UCS_to_uv(UVW)`

Returns the  $uv$  chromaticity coordinates from given *CIE UCS* colourspace array.

**Parameters** `UVW` (`array_like`) – *CIE UCS* colourspace array.

**Returns**  $uv$  chromaticity coordinates.

**Return type** `ndarray`

### Notes

- Input *CIE UCS* colourspace array is in domain [0, 1].
- Output  $uv$  chromaticity coordinates are in range [0, 1].

### References

- [\[Wikv\]](#)

### Examples

```
>>> import numpy as np
>>> UCS = np.array([0.04699689, 0.10080000, 0.16374390])
>>> UCS_to_uv(UCS)
array([ 0.1508530...,  0.3235531...])
```

## colour.UCS\_uv\_to\_xy

`colour.UCS_uv_to_xy(uv)`

Returns the  $xy$  chromaticity coordinates from given *CIE UCS* colourspace  $uv$  chromaticity coordinates.

**Parameters** `uv` (`array_like`) – *CIE UCS*  $uv$  chromaticity coordinates.

**Returns**  $xy$  chromaticity coordinates.

**Return type** `ndarray`

### Notes

- Input  $uv$  chromaticity coordinates are in domain [0, 1].
- Output  $xy$  chromaticity coordinates are in range [0, 1].

### References

- [\[Wikv\]](#)

## Examples

```
>>> import numpy as np
>>> uv = np.array([0.150853087327666, 0.323553137295440])
>>> UCS_uv_to_xy(uv)
array([ 0.2641477...,  0.3777000...])
```

## CIE 1964 U\*V\*W\* Colourspace

colour

<code>XYZ_to_UVW(XYZ[, illuminant])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>CIE 1964 U*V*W*</i> colourspace.
--	--

### colour.XYZ\_to\_UVW

`colour.XYZ_to_UVW(XYZ, illuminant=array([ 0.3457, 0.3585]))`  
Converts from *CIE XYZ* tristimulus values to *CIE 1964 U\*V\*W\** colourspace.

#### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `illuminant` (`array_like`, optional) – Reference *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.

**Returns** *CIE 1964 U\*V\*W\** colourspace array.

**Return type** `ndarray`

**Warning:** The input domain and output range of that definition are non standard!

## Notes

- Input *CIE XYZ* tristimulus values are in domain  $[0, 100]$ .
- Input *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain  $[0, \infty]$ .
- Output *CIE 1964 U\*V\*W\** colourspace array is in range  $[0, 100]$ .

## References

- [\[Wike\]](#)

## Examples

```
>>> import numpy as np
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313]) * 100
>>> XYZ_to_UVW(XYZ)
array([-28.0579733..., -0.8819449..., 37.0041149...])
```

## Hunter L,a,b Colour Scale

### colour

<code>XYZ_to_Hunter_Lab(XYZ[, XYZ_n, K_ab])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>Hunter L,a,b</i> colour scale.
<code>Hunter_Lab_to_XYZ(Lab[, XYZ_n, K_ab])</code>	Converts from <i>Hunter L,a,b</i> colour scale to <i>CIE XYZ</i> tristimulus values.
<code>XYZ_to_K_ab_HunterLab1966(XYZ)</code>	Converts from <i>whitepoint CIE XYZ</i> tristimulus values to <i>Hunter L,a,b</i> $K_a$ and $K_b$ chromaticity coefficients.

### colour.XYZ\_to\_Hunter\_Lab

`colour.XYZ_to_Hunter_Lab(XYZ, XYZ_n=array([ 96.38, 100., 82.45]), K_ab=array([-173.51, 58.48]))`

Converts from *CIE XYZ* tristimulus values to *Hunter L,a,b* colour scale.

#### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `XYZ_n` (`array_like`, optional) – Reference *illuminant* tristimulus values.
- `K_ab` (`array_like`, optional) – Reference *illuminant* chromaticity coefficients, if `K_ab` is set to `None` it will be computed using `colour.XYZ_to_K_ab_HunterLab1966()`.

**Returns** *Hunter L,a,b* colour scale array.

**Return type** `ndarray`

#### Notes

- Input *CIE XYZ* and reference *illuminant* tristimulus values are in domain [0, 100].
- Output *Lightness L\** is in range [0, 100].

#### References

- [\[Hun08a\]](#)

#### Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313]) * 100
>>> D50 = HUNTERLAB_ILLUMINANTS[
...     'CIE 1931 2 Degree Standard Observer']['D50']
>>> XYZ_to_Hunter_Lab(XYZ, D50.XYZ_n, D50.K_ab)
array([-31.7490157..., -15.1146262..., -2.7866075...])
```

## colour.Hunter\_Lab\_to\_XYZ

```
colour.Hunter_Lab_to_XYZ(Lab, XYZ_n=array([ 96.38, 100., 82.45]), K_ab=array([-173.51,
-58.48]))
```

Converts from *Hunter L,a,b* colour scale to *CIE XYZ* tristimulus values.

### Parameters

- **Lab** (array\_like) – *Hunter L,a,b* colour scale array.
- **XYZ\_n** (array\_like, optional) – Reference *illuminant* tristimulus values.
- **K\_ab** (array\_like, optional) – Reference *illuminant* chromaticity coefficients, if **K\_ab** is set to *None* it will be computed using `colour.XYZ_to_K_ab_HunterLab1966()`.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

### Notes

- Input *Lightness L\** is in domain [0, 100].
- Input *CIE XYZ* and reference *illuminant* tristimulus values are in domain [0, 100].
- Output *CIE XYZ* tristimulus values are in range [0, 100].

### References

- [\[Hun08a\]](#)

### Examples

```
>>> Lab = np.array([31.74901573, -15.11462629, -2.78660758])
>>> D50 = HUNTERLAB_ILLUMINANTS[
...     'CIE 1931 2 Degree Standard Observer']['D50']
>>> Hunter_Lab_to_XYZ(Lab, D50.XYZ_n, D50.K_ab)
array([-7.049534, 10.08      ,  9.558313])
```

## colour.XYZ\_to\_K\_ab\_HunterLab1966

```
colour.XYZ_to_K_ab_HunterLab1966(XYZ)
```

Converts from *whitepoint CIE XYZ* tristimulus values to *Hunter L,a,b*  $K_a$  and  $K_b$  chromaticity coefficients.

**Parameters** **XYZ** (array\_like) – *Whitepoint CIE XYZ* tristimulus values.

**Returns** *Hunter L,a,b*  $K_a$  and  $K_b$  chromaticity coefficients.

**Return type** ndarray

### Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 100].

## References

- [\[Hun08b\]](#)

## Examples

```
>>> XYZ = np.array([109.850, 100.000, 35.585])
>>> XYZ_to_K_ab_HunterLab1966(XYZ)
array([ 185.2378721..., 38.4219142...])
```

## Hunter Rd,a,b Colour Scale

colour

<code>XYZ_to_Hunter_Rdab(XYZ[, XYZ_n, K_ab])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>Hunter Rd,a,b</i> colour scale.
---	---

### colour.XYZ\_to\_Hunter\_Rdab

`colour.XYZ_to_Hunter_Rdab(XYZ, XYZ_n=array([ 96.38, 100., 82.45]), K_ab=array([ 173.51, 58.48]))`

Converts from *CIE XYZ* tristimulus values to *Hunter Rd,a,b* colour scale.

#### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `XYZ_n` (`array_like`, optional) – Reference *illuminant* tristimulus values.
- `K_ab` (`array_like`, optional) – Reference *illuminant* chromaticity coefficients, if `K_ab` is set to `None` it will be computed using `colour.XYZ_to_K_ab_HunterLab1966()`.

**Returns** *Hunter Rd,a,b* colour scale array.

**Return type** ndarray

## Notes

- Input *CIE XYZ* and reference *illuminant* tristimulus values are in domain [0, 100].

## References

- [\[Hun12\]](#)

## Examples

```
>>> import numpy as np
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313]) * 100
>>> D50 = HUNTERLAB_ILLUMINANTS[
```

```

...      'CIE 1931 2 Degree Standard Observer']['D50']
>>> XYZ_to_Hunter_Rdab(XYZ, D50.XYZ_n, D50.K_ab)
...
array([ 10.08       , -18.6765376..., -3.4432992...])

```

## Luo, Cui and Li (2006)

### colour

JMh_CIECAM02_to_CAM02LCD(JMh)	Converts from <i>CIECAM02 JMh</i> correlates array to <i>Luo et alii (2006) CAM02-LCD</i> colourspace $J'a'b'$ array.
CAM02LCD_to_JMh_CIECAM02(Jpapbp)	Converts from <i>Luo et alii (2006) CAM02-LCD</i> colourspace $J'a'b'$ array to <i>CIECAM02 JMh</i> correlates array.
JMh_CIECAM02_to_CAM02SCD(JMh)	Converts from <i>CIECAM02 JMh</i> correlates array to <i>Luo et alii (2006) CAM02-SCD</i> colourspace $J'a'b'$ array.
CAM02SCD_to_JMh_CIECAM02(Jpapbp)	Converts from <i>Luo et alii (2006) CAM02-SCD</i> colourspace $J'a'b'$ array to <i>CIECAM02 JMh</i> correlates array.
JMh_CIECAM02_to_CAM02UCS(JMh)	Converts from <i>CIECAM02 JMh</i> correlates array to <i>Luo et alii (2006) CAM02-UCS</i> colourspace $J'a'b'$ array.
CAM02UCS_to_JMh_CIECAM02(Jpapbp)	Converts from <i>Luo et alii (2006) CAM02-UCS</i> colourspace $J'a'b'$ array to <i>CIECAM02 JMh</i> correlates array.

### colour.JMh\_CIECAM02\_to\_CAM02LCD

colour.JMh\_CIECAM02\_to\_CAM02LCD(*JMh*)

Converts from *CIECAM02 JMh* correlates array to *Luo et alii (2006) CAM02-LCD* colourspace  $J'a'b'$  array.

**Parameters** *JMh* (*array\_like*) – *CIECAM02* correlates array *JMh*.

**Returns** *Luo et alii (2006) CAM02-LCD* colourspace  $J'a'b'$  array.

**Return type** *ndarray*

### References

- [LCL06]

### Examples

```

>>> from colour.appearance import (
...     CIECAM02_VIEWING_CONDITIONS,
...     XYZ_to_CIECAM02)
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0

```

```
>>> surround = CIECAM02_VIEWING_CONDITIONS['Average']
>>> specification = XYZ_to_CIECAM02(
...     XYZ, XYZ_w, L_A, Y_b, surround)
>>> JMh = (specification.J, specification.M, specification.h)
>>> JMh_CIECAM02_to_CAM02LCD(JMh)
array([ 54.9043313..., -0.0845039..., -0.0685483...])
```

## colour.CAM02LCD\_to\_JMh\_CIECAM02

colour.CAM02LCD\_to\_JMh\_CIECAM02(*Jpapbp*)

Converts from *Luo et alii (2006)* CAM02-LCD colourspace  $J'a'b'$  array to *CIECAM02*  $JMh$  correlates array.

**Parameters** **Jpapbp** (array\_like) – *Luo et alii (2006)* CAM02-LCD colourspace  $J'a'b'$  array.

**Returns** *CIECAM02* correlates array  $JMh$ .

**Return type** ndarray

### References

- [LCL06]

### Examples

```
>>> Jpapbp = np.array([54.90433134, -0.08450395, -0.06854831])
>>> CAM02LCD_to_JMh_CIECAM02(Jpapbp)
array([ 4.1731091...e+01,  1.0884217...e-01,  2.1904843...e+02])
```

## colour.JMh\_CIECAM02\_to\_CAM02SCD

colour.JMh\_CIECAM02\_to\_CAM02SCD(*JMh*)

Converts from *CIECAM02*  $JMh$  correlates array to *Luo et alii (2006)* CAM02-SCD colourspace  $J'a'b'$  array.

**Parameters** **JMh** (array\_like) – *CIECAM02* correlates array  $JMh$ .

**Returns** *Luo et alii (2006)* CAM02-SCD colourspace  $J'a'b'$  array.

**Return type** ndarray

### References

- [LCL06]

### Examples

```
>>> from colour.appearance import (
...     CIECAM02_VIEWING_CONDITIONS,
...     XYZ_to_CIECAM02)
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0
>>> surround = CIECAM02_VIEWING_CONDITIONS['Average']
>>> specification = XYZ_to_CIECAM02(
...     XYZ, XYZ_w, L_A, Y_b, surround)
>>> JMh = (specification.J, specification.M, specification.h)
>>> JMh_CIECAM02_to_CAM02SCD(JMh)
array([ 54.9043313..., -0.0843617..., -0.0684329...])
```

## colour.CAM02SCD\_to\_JMh\_CIECAM02

`colour.CAM02SCD_to_JMh_CIECAM02(Jpapbp)`

Converts from *Luo et alii (2006) CAM02-SCD* colourspace  $J'a'b'$  array to *CIECAM02 JMh* correlates array.

**Parameters** `Jpapbp` (`array_like`) – *Luo et alii (2006) CAM02-SCD* colourspace  $J'a'b'$  array.

**Returns** *CIECAM02* correlates array  $JMh$ .

**Return type** `ndarray`

## References

- [LCL06]

## Examples

```
>>> Jpapbp = np.array([54.90433134, -0.08436178, -0.06843298])
>>> CAM02SCD_to_JMh_CIECAM02(Jpapbp)
array([ 4.1731091...e+01,   1.0884217...e-01,   2.1904843...e+02])
```

## colour.JMh\_CIECAM02\_to\_CAM02UCS

`colour.JMh_CIECAM02_to_CAM02UCS(JMh)`

Converts from *CIECAM02 JMh* correlates array to *Luo et alii (2006) CAM02-UCS* colourspace  $J'a'b'$  array.

**Parameters** `JMh` (`array_like`) – *CIECAM02* correlates array  $JMh$ .

**Returns** *Luo et alii (2006) CAM02-UCS* colourspace  $J'a'b'$  array.

**Return type** `ndarray`

## References

- [LCL06]

## Examples

```
>>> from colour.appearance import (
...     CIECAM02_VIEWING_CONDITIONS,
...     XYZ_to_CIECAM02)
>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0
>>> surround = CIECAM02_VIEWING_CONDITIONS['Average']
>>> specification = XYZ_to_CIECAM02(
...     XYZ, XYZ_w, L_A, Y_b, surround)
>>> JMh = (specification.J, specification.M, specification.h)
>>> JMh_CIECAM02_to_CAM02UCS(JMh)
array([ 54.9043313..., -0.0844236..., -0.0684831...])
```

## colour.CAM02UCS\_to\_JMh\_CIECAM02

colour.CAM02UCS\_to\_JMh\_CIECAM02(*Jpapbp*)

Converts from *Luo et alii (2006)* CAM02-UCS colourspace  $J'a'b'$  array to *CIECAM02 JMh* correlates array.

**Parameters** **Jpapbp** (array\_like) – *Luo et alii (2006)* CAM02-UCS colourspace  $J'a'b'$  array.

**Returns** *CIECAM02* correlates array  $JMh$ .

**Return type** ndarray

## References

- [\[LCL06\]](#)

## Examples

```
>>> Jpapbp = np.array([54.90433134, -0.08442362, -0.06848314])
>>> CAM02UCS_to_JMh_CIECAM02(Jpapbp)
array([-4.1731091...e+01,  1.0884217...e-01,  2.1904843...e+02])
```

## Li, Li, Wang, Zu, Luo, Cui, Melgosa, Brill and Pointer (2017)

colour

JMh_CAM16_to_CAM16LCD	Converts from <i>CAM16 JMh</i> correlates array to <i>Li et alii (2017)</i> CAM16-LCD colourspace $J'a'b'$ array.
CAM16LCD_to_JMh_CAM16	Converts from <i>Li et alii (2017)</i> CAM16-LCD colourspace $J'a'b'$ array to <i>CAM16 JMh</i> correlates array.
JMh_CAM16_to_CAM16SCD	Converts from <i>CAM16 JMh</i> correlates array to <i>Li et alii (2017)</i> CAM16-SCD colourspace $J'a'b'$ array.

Continued on next page

Table 3.132 – continued from previous page

CAM16SCD_to_JMh_CAM16	Converts from <i>Li et alii (2017) CAM16-SCD</i> colourspace $J'a'b'$ array to <i>CAM16 JMh</i> correlates array.
JMh_CAM16_to_CAM16UCS	Converts from <i>CAM16 JMh</i> correlates array to <i>Li et alii (2017) CAM16-UCS</i> colourspace $J'a'b'$ array.
CAM16UCS_to_JMh_CAM16	Converts from <i>Li et alii (2017) CAM16-UCS</i> colourspace $J'a'b'$ array to <i>CAM16 JMh</i> correlates array.

## colour.JMh\_CAM16\_to\_CAM16LCD

`colour.JMh_CAM16_to_CAM16LCD = <functools.partial object>`

Converts from *CAM16 JMh* correlates array to *Li et alii (2017) CAM16-LCD* colourspace  $J'a'b'$  array.

**Parameters** `JMh` (`array_like`) – *CAM16* correlates array *JMh*.

**Returns** *Li et alii (2017) CAM16-LCD* colourspace  $J'a'b'$  array.

**Return type** `ndarray`

### References

- [\[LLW+17\]](#)

### Notes

- This docstring is automatically generated, please refer to `colour.JMh_CIECAM02_to_CAM02LCD()` definition for an usage example.

## colour.CAM16LCD\_to\_JMh\_CAM16

`colour.CAM16LCD_to_JMh_CAM16 = <functools.partial object>`

Converts from *Li et alii (2017) CAM16-LCD* colourspace  $J'a'b'$  array to *CAM16 JMh* correlates array.

**Parameters** `Jpapbp` (`array_like`) – *Li et alii (2017) CAM16-LCD* colourspace  $J'a'b'$  array.

**Returns** *CAM16* correlates array *JMh*.

**Return type** `ndarray`

### References

- [\[LLW+17\]](#)

### Notes

- This docstring is automatically generated, please refer to `colour.CAM02LCD_to_JMh_CIECAM02()` definition for an usage example.

## colour.JMh\_CAM16\_to\_CAM16SCD

`colour.JMh_CAM16_to_CAM16SCD = <functools.partial object>`

Converts from *CAM16 JMh* correlates array to *Li et alii (2017) CAM16-SCD* colourspace  $J'a'b'$  array.

**Parameters** `JMh` (`array_like`) – *CAM16* correlates array *JMh*.

**Returns** *Li et alii (2017) CAM16-SCD* colourspace  $J'a'b'$  array.

**Return type** `ndarray`

### References

- [\[LLW+17\]](#)

### Notes

- This docstring is automatically generated, please refer to `colour.JMh_CIECAM02_to_CAM02SCD()` definition for an usage example.

## colour.CAM16SCD\_to\_JMh\_CAM16

`colour.CAM16SCD_to_JMh_CAM16 = <functools.partial object>`

Converts from *Li et alii (2017) CAM16-SCD* colourspace  $J'a'b'$  array to *CAM16 JMh* correlates array.

**Parameters** `Jpapbp` (`array_like`) – *Li et alii (2017) CAM16-SCD* colourspace  $J'a'b'$  array.

**Returns** *CAM16* correlates array *JMh*.

**Return type** `ndarray`

### References

- [\[LLW+17\]](#)

### Notes

- This docstring is automatically generated, please refer to `colour.CAM02SCD_to_JMh_CIECAM02()` definition for an usage example.

## colour.JMh\_CAM16\_to\_CAM16UCS

`colour.JMh_CAM16_to_CAM16UCS = <functools.partial object>`

Converts from *CAM16 JMh* correlates array to *Li et alii (2017) CAM16-UCS* colourspace  $J'a'b'$  array.

**Parameters** `JMh` (`array_like`) – *CAM16* correlates array *JMh*.

**Returns** *Li et alii (2017) CAM16-UCS* colourspace  $J'a'b'$  array.

**Return type** `ndarray`

## References

- [\[LLW+17\]](#)

## Notes

- This docstring is automatically generated, please refer to `colour.JMh_CIECAM02_to_CAM02UCS()` definition for an usage example.

### `colour.CAM16UCS_to_JMh_CAM16`

`colour.CAM16UCS_to_JMh_CAM16 = <functools.partial object>`

Converts from *Li et alii (2017) CAM16-UCS* colourspace  $J'a'b'$  array to *CAM16 JMh* correlates array.

**Parameters** `Jpapbp` (`array_like`) – *Li et alii (2017) CAM16-UCS* colourspace  $J'a'b'$  array.

**Returns** *CAM16* correlates array  $JMh$ .

**Return type** `ndarray`

## References

- [\[LLW+17\]](#)

## Notes

- This docstring is automatically generated, please refer to `colour.CAM02UCS_to_JMh_CIECAM02()` definition for an usage example.

### IPT Colourspace

`colour`

<code>XYZ_to_IPT(XYZ)</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>IPT</i> colourspace.
<code>IPT_to_XYZ(IPT)</code>	Converts from <i>IPT</i> colourspace to <i>CIE XYZ</i> tristimulus values.
<code>IPT_hue_angle(IPT)</code>	Computes the hue angle in degrees from <i>IPT</i> colourspace.

### `colour.XYZ_to_IPT`

`colour.XYZ_to_IPT(XYZ)`

Converts from *CIE XYZ* tristimulus values to *IPT* colourspace.

**Parameters** `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.

**Returns** *IPT* colourspace array.

**Return type** `ndarray`

## Notes

- Input *CIE XYZ* tristimulus values needs to be adapted for *CIE Standard Illuminant D Series D65*.

## References

- [Fai13d]

## Examples

```
>>> XYZ = np.array([0.96907232, 1.00000000, 1.12179215])
>>> XYZ_to_IPT(XYZ)
array([ 1.0030082...,  0.0190691..., -0.0136929...])
```

## colour.IPT\_to\_XYZ

colour.IPT\_to\_XYZ(*IPT*)

Converts from *IPT* colourspace to *CIE XYZ* tristimulus values.

**Parameters** *IPT* (array\_like) – *IPT* colourspace array.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

## References

- [Fai13d]

## Examples

```
>>> IPT = np.array([1.00300825, 0.01906918, -0.01369292])
>>> IPT_to_XYZ(IPT)
array([ 0.9690723...,  1.          ,  1.1217921...])
```

## colour.IPT\_hue\_angle

colour.IPT\_hue\_angle(*IPT*)

Computes the hue angle in degrees from *IPT* colourspace.

**Parameters** *IPT* (array\_like) – *IPT* colourspace array.

**Returns** Hue angle in degrees.

**Return type** numeric or ndarray

## References

- [Fai13d]

## Examples

```
>>> IPT = np.array([0.96907232, 1, 1.12179215])
>>> IPT_hue_angle(IPT)
48.2852074...
```

## hdr-IPT Colourspace

### colour

<code>XYZ_to_hdr_IPT(XYZ[, Y_s, Y_abs, method])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>hdr-IPT</i> colourspace.
<code>hdr_IPT_to_XYZ(IPT_hdr[, Y_s, Y_abs, method])</code>	Converts from <i>hdr-IPT</i> colourspace to <i>CIE XYZ</i> tristimulus values.
<code>HDR_IPT_METHODS</code>	Supported <i>hdr-IPT</i> colourspace computation methods.

### colour.XYZ\_to\_hdr\_IPT

`colour.XYZ_to_hdr_IPT(XYZ, Y_s=0.2, Y_abs=100, method=u'Fairchild 2011')`  
Converts from *CIE XYZ* tristimulus values to *hdr-IPT* colourspace.

#### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `Y_s` (`numeric` or `array_like`) – Relative luminance  $Y_s$  of the surround in domain  $[0, 1]$ .
- `Y_abs` (`numeric` or `array_like`) – Absolute luminance  $Y_{abs}$  of the scene diffuse white in  $cd/m^2$ .
- `method` (`unicode`, `optional`) – {‘Fairchild 2011’, ‘Fairchild 2010’}, Computation method.

**Returns** *hdr-IPT* colourspace array.

**Return type** `ndarray`

## Notes

- Input *CIE XYZ* tristimulus values needs to be adapted for *CIE Standard Illuminant D Series D65*.

## References

- [FW10]
- [FC11]

## Examples

```
>>> XYZ = np.array([0.96907232, 1.00000000, 1.12179215])
>>> XYZ_to_hdr_IPT(XYZ)
array([ 93.5317473..., 1.8564156..., -1.3292254...])
>>> XYZ_to_hdr_IPT(XYZ, method='Fairchild 2010')
array([ 94.6592917..., 0.3804177..., -0.2673118...])
```

## colour.hdr\_IPT\_to\_XYZ

`colour.hdr_IPT_to_XYZ(IPT_hdr, Y_s=0.2, Y_abs=100, method=u'Fairchild 2011')`

Converts from *hdr-IPT* colourspace to *CIE XYZ* tristimulus values.

### Parameters

- `IPT_hdr` (`array_like`) – *hdr-IPT* colourspace array.
- `Y_s` (`numeric` or `array_like`) – Relative luminance  $Y_s$  of the surround in domain  $[0, 1]$ .
- `Y_abs` (`numeric` or `array_like`) – Absolute luminance  $Y_{abs}$  of the scene diffuse white in  $cd/m^2$ .
- `method` (`unicode`, `optional`) – {‘Fairchild 2011’, ‘Fairchild 2010’}, Computation method.

**Returns** *CIE XYZ* tristimulus values.

**Return type** `ndarray`

## References

- [\[FW10\]](#)
- [\[FC11\]](#)

## Examples

```
>>> IPT_hdr = np.array([93.53174734, 1.85641567, -1.32922546])
>>> hdr_IPT_to_XYZ(IPT_hdr)
array([ 0.9690723..., 1.           , 1.1217921...])
>>> IPT_hdr = np.array([94.65929175, 0.38041773, -0.26731187])
>>> hdr_IPT_to_XYZ(IPT_hdr, method='Fairchild 2010')
...
array([ 0.9690723..., 1.           , 1.1217921...])
```

## colour.HDR\_IPT\_METHODS

`colour.HDR_IPT_METHODS = (u'Fairchild 2010', u'Fairchild 2011')`

Supported *hdr-IPT* colourspace computation methods.

## References

- [\[FW10\]](#)
- [\[FC11\]](#)

**HDR\_IPT\_METHODS** [tuple] {‘Fairchild 2011’, ‘Fairchild 2010’}

## hdr-CIELAB Colourspace

colour

<code>XYZ_to_hdr_CIELab(XYZ[, illuminant, Y_s, ...])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>hdr-CIELAB</i> colourspace.
<code>hdr_CIELab_to_XYZ(Lab_hdr[, illuminant, ...])</code>	Converts from <i>hdr-CIELAB</i> colourspace to <i>CIE XYZ</i> tristimulus values.
<code>HDR_CIELAB_METHODS</code>	Supported <i>hdr-CIELAB</i> colourspace computation methods.

### colour.XYZ\_to\_hdr\_CIELab

`colour.XYZ_to_hdr_CIELab(XYZ, illuminant=array([ 0.3457, 0.3585]), Y_s=0.2, Y_abs=100, method=u'Fairchild 2011')`

Converts from *CIE XYZ* tristimulus values to *hdr-CIELAB* colourspace.

#### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values.
- `illuminant` (array\_like, optional) – Reference *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.
- `Y_s` (numeric or array\_like) – Relative luminance  $Y_s$  of the surround in domain [0, 1].
- `Y_abs` (numeric or array\_like) – Absolute luminance  $Y_{abs}$  of the scene diffuse white in  $cd/m^2$ .
- `method` (unicode, optional) – {‘Fairchild 2011’, ‘Fairchild 2010’}, Computation method.

**Returns** *hdr-CIELAB* colourspace array.

**Return type** ndarray

#### Notes

- Conversion to polar coordinates to compute the *chroma*  $C_{hdr}$  and *hue*  $h_{hdr}$  correlates can be safely performed with `colour.Lab_to_LChab()` definition.
- Conversion to cartesian coordinates from the *Lightness*  $L_{hdr}$ , *chroma*  $C_{hdr}$  and *hue*  $h_{hdr}$  correlates can be safely performed with `colour.LChab_to_Lab()` definition.
- Input *CIE XYZ* tristimulus values are in domain [0, math:*infty*].
- Input *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].

## References

- [\[FW10\]](#)
- [\[FC11\]](#)

## Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_to_hdr_CIELab(XYZ)
array([ 26.4646106..., -24.613326 ..., -4.8479681...])
>>> XYZ_to_hdr_CIELab(XYZ, method='Fairchild 2010')
array([ 24.9020664..., -46.8312760..., -10.1427484...])
```

## colour.hdr\_CIELab\_to\_XYZ

`colour.hdr_CIELab_to_XYZ(Lab_hdr, illuminant=array([ 0.3457, 0.3585]), Y_s=0.2, Y_abs=100, method=u'Fairchild 2011')`

Converts from *hdr-CIELAB* colourspace to *CIE XYZ* tristimulus values.

### Parameters

- **Lab\_hdr** (array\_like) – *hdr-CIELAB* colourspace array.
- **illuminant** (array\_like, optional) – Reference *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.
- **Y\_s** (numeric or array\_like) – Relative luminance  $Y_s$  of the surround in domain [0, 1].
- **Y\_abs** (numeric or array\_like) – Absolute luminance  $Y_{abs}$  of the scene diffuse white in  $cd/m^2$ .
- **method** (unicode, optional) – {'Fairchild 2011', 'Fairchild 2010'}, Computation method.

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

## Notes

- Input *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain  $[0, \infty]$ .
- Output *CIE XYZ* tristimulus values are in range  $[0, \text{math:}infty]$ .

## References

- [\[FW10\]](#)
- [\[FC11\]](#)

## Examples

```
>>> Lab_hdr = np.array([26.46461067, -24.613326, -4.84796811])
>>> hdr_CIELab_to_XYZ(Lab_hdr)
array([ 0.0704953...,  0.1008     ,  0.0955831...])
>>> Lab_hdr = np.array([24.90206646, -46.83127607, -10.14274843])
>>> hdr_CIELab_to_XYZ(Lab_hdr, method='Fairchild 2010')
...
array([ 0.0704953...,  0.1008     ,  0.0955831...])
```

## colour.HDR\_CIELAB\_METHODS

`colour.HDR_CIELAB_METHODS = (u'Fairchild 2010', u'Fairchild 2011')`  
Supported *hdr-CIELAB* colourspace computation methods.

## References

- [\[FW10\]](#)
- [\[FC11\]](#)

`HDR_CIELAB_METHODS` [tuple] {‘Fairchild 2011’, ‘Fairchild 2010’}

## RGB Colourspace and Transformations

### colour

<code>XYZ_to_RGB(XYZ, illuminant_XYZ, ...[, ...])</code>	Converts from <i>CIE XYZ</i> tristimulus values to given <i>RGB</i> colourspace.
<code>RGB_to_XYZ(RGB, illuminant_RGB, ...[, ...])</code>	Converts from given <i>RGB</i> colourspace to <i>CIE XYZ</i> tristimulus values.
<code>RGB_to_RGB(RGB, input_colourspace, ...[, ...])</code>	Converts from given input <i>RGB</i> colourspace to output <i>RGB</i> colourspace using given <i>chromatic adaptation</i> method.
<code>RGB_to_RGB_matrix(input_colourspace, ...[, ...])</code>	Computes the matrix $M$ converting from given input <i>RGB</i> colourspace to output <i>RGB</i> colourspace using given <i>chromatic adaptation</i> method.

## colour.XYZ\_to\_RGB

`colour.XYZ_to_RGB(XYZ, illuminant_XYZ, illuminant_RGB, XYZ_to_RGB_matrix, chromatic_adaptation_transform=u'CAT02', encoding_cctf=None)`  
Converts from *CIE XYZ* tristimulus values to given *RGB* colourspace.

### Parameters

- `XYZ` (array\_like) – *CIE XYZ* tristimulus values.
- `illuminant_XYZ` (array\_like) – *CIE XYZ* tristimulus values *illuminant xy* chromaticity coordinates or *CIE xyY* colourspace array.

- `illuminant_RGB` (`array_like`) – *RGB colourspace illuminant xy chromaticity coordinates or CIE xyY colourspace array.*
- `XYZ_to_RGB_matrix` (`array_like`) – *Normalised primary matrix.*
- `chromatic_adaptation_transform` (`unicode, optional`) – {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}, *Chromatic adaptation transform.*
- `encoding_cctf` (`object, optional`) – *Encoding colour component transfer function (Encoding CCTF) or opto-electronic transfer function (OETF / OECF).*

**Returns** *RGB colourspace array.*

**Return type** `ndarray`

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].
- Input *illuminant\_XYZ* *xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Input *illuminant\_RGB* *xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Output *RGB* colourspace array is in range [0, 1].

## Examples

```
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> illuminant_XYZ = np.array([0.34570, 0.35850])
>>> illuminant_RGB = np.array([0.31270, 0.32900])
>>> chromatic_adaptation_transform = 'Bradford'
>>> XYZ_to_RGB_matrix = np.array(
...     [[3.24062548, -1.53720797, -0.49862860],
...      [-0.96893071, 1.87575606, 0.04151752],
...      [0.05571012, -0.20402105, 1.05699594]])
...
>>> XYZ_to_RGB(XYZ, illuminant_XYZ, illuminant_RGB, XYZ_to_RGB_matrix,
...             chromatic_adaptation_transform)
array([ 0.0110015...,  0.1273504...,  0.1163271...])
```

## colour.RGB\_to\_XYZ

```
colour.RGB_to_XYZ(RGB, illuminant_RGB, illuminant_XYZ, RGB_to_XYZ_matrix, chromatic_adaptation_transform=u'CAT02', decoding_cctf=None)
```

Converts from given *RGB* colourspace to *CIE XYZ* tristimulus values.

### Parameters

- `RGB` (`array_like`) – *RGB colourspace array.*
- `illuminant_RGB` (`array_like`) – *RGB colourspace illuminant chromaticity coordinates or CIE xyY colourspace array.*

- `illuminant_XYZ` (array\_like) – *CIE XYZ* tristimulus values *illuminant* chromaticity coordinates or *CIE xyY* colourspace array.
- `RGB_to_XYZ_matrix` (array\_like) – *Normalised primary matrix*.
- `chromatic_adaptation_transform` (unicode, optional) – {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}, *Chromatic adaptation transform*.
- `decoding_cctf` (object, optional) – Decoding colour component transfer function (Decoding CCTF) or electro-optical transfer function (EOTF / EOCF).

**Returns** *CIE XYZ* tristimulus values.

**Return type** ndarray

## Notes

- Input *RGB* colourspace array is in domain [0, 1].
- Input *illuminant\_RGB* *xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Input *illuminant\_XYZ* *xy* chromaticity coordinates or *CIE xyY* colourspace array are in domain [0,  $\infty$ ].
- Output *CIE XYZ* tristimulus values are in range [0, 1].

## Examples

```
>>> RGB = np.array([0.01100154, 0.12735048, 0.11632713])
>>> illuminant_RGB = np.array([0.31270, 0.32900])
>>> illuminant_XYZ = np.array([0.34570, 0.35850])
>>> chromatic_adaptation_transform = 'Bradford'
>>> RGB_to_XYZ_matrix = np.array(
...     [[0.41240000, 0.35760000, 0.18050000],
...      [0.21260000, 0.71520000, 0.07220000],
...      [0.01930000, 0.11920000, 0.95050000]])
...
>>> RGB_to_XYZ(RGB, illuminant_RGB, illuminant_XYZ, RGB_to_XYZ_matrix,
...               chromatic_adaptation_transform)
array([ 0.0704953...,  0.1008     ,  0.0955831...])
```

## colour.RGB\_to\_RGB

```
colour.RGB_to_RGB(RGB, input_colourspace, output_colourspace, chromatic_adaptation_transform=u'CAT02', apply_decoding_cctf=False, apply_encoding_cctf=False)
```

Converts from given input *RGB* colourspace to output *RGB* colourspace using given *chromatic adaptation* method.

### Parameters

- `RGB` (array\_like) – *RGB* colourspace array.
- `input_colourspace` ([RGB\\_Colourspace](#)) – *RGB* input colourspace.

- **output\_colourspace** (`RGB_Colourspace`) – *RGB output colourspace.*
- **chromatic\_adaptation\_transform** (`unicode`, `optional`) – `{'CAT02', 'XYZ Scaling', 'Von Kries', 'Bradford', 'Sharp', 'Fairchild', 'CMCCAT97', 'CMCCAT2000', 'CAT02_BRILL_CAT', 'Bianco', 'Bianco PC'}`, *Chromatic adaptation transform.*
- **apply\_decoding\_cctf** (`bool`, `optional`) – *Apply input colourspace decoding colour component transfer function / electro-optical transfer function.*
- **apply\_encoding\_cctf** (`bool`, `optional`) – *Apply output colourspace encoding colour component transfer function / opto-electronic transfer function.*

**Returns** *RGB colourspace array.*

**Return type** `ndarray`

## Notes

- Input / output *RGB* colourspace arrays are in domain / range [0, 1].
- Input / output *RGB* colourspace arrays are assumed to be representing linear light values.

## Examples

```
>>> from colour.models import sRGB_COLOURSPACE, PROPHOTO_RGB_COLOURSPACE
>>> RGB = np.array([0.01103742, 0.12734226, 0.11632971])
>>> RGB_to_RGB(RGB, sRGB_COLOURSPACE, PROPHOTO_RGB_COLOURSPACE)
...
array([ 0.0643561...,  0.1157331...,  0.1158069...])
```

## colour.RGB\_to\_RGB\_matrix

```
colour.RGB_to_RGB_matrix(input_colourspace, output_colourspace, chro-
                           matic_adaptation_transform=u'CAT02')
```

Computes the matrix  $M$  converting from given input *RGB* colourspace to output *RGB* colourspace using given *chromatic adaptation* method.

### Parameters

- **input\_colourspace** (`RGB_Colourspace`) – *RGB input colourspace.*
- **output\_colourspace** (`RGB_Colourspace`) – *RGB output colourspace.*
- **chromatic\_adaptation\_transform** (`unicode`, `optional`) – `{'CAT02', 'XYZ Scaling', 'Von Kries', 'Bradford', 'Sharp', 'Fairchild', 'CMCCAT97', 'CMCCAT2000', 'CAT02_BRILL_CAT', 'Bianco', 'Bianco PC'}`, *Chromatic adaptation transform.*

**Returns** Conversion matrix  $M$ .

**Return type** `ndarray`

## Examples

```
>>> from colour.models import sRGB_COLOURSPACE, PROPHOTO_RGB_COLOURSPACE
>>> RGB_to_RGB_matrix(sRGB_COLOURSPACE, PROPHOTO_RGB_COLOURSPACE)
...
array([[ 0.5288241...,  0.3340609...,  0.1373616...],
       [ 0.0975294...,  0.8790074...,  0.0233981...],
       [ 0.0163599...,  0.1066124...,  0.8772485...]])
```

## Ancillary Objects

colour

<code>XYZ_to_sRGB(XYZ[, illuminant, ...])</code>	Converts from <i>CIE XYZ</i> tristimulus values to <i>sRGB</i> colourspace.
<code>sRGB_to_XYZ(RGB[, illuminant, ...])</code>	Converts from <i>sRGB</i> colourspace to <i>CIE XYZ</i> tristimulus values.

## colour.XYZ\_to\_sRGB

`colour.XYZ_to_sRGB(XYZ, illuminant=array([ 0.3127, 0.329 ...]), chromatic_adaptation_transform=u'CAT02', apply_encoding_cctf=True)`  
Converts from *CIE XYZ* tristimulus values to *sRGB* colourspace.

### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `illuminant` (`array_like`, optional) – Source illuminant chromaticity coordinates.
- `chromatic_adaptation_transform` (`unicode`, optional) – {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}, *Chromatic adaptation transform*.
- `apply_encoding_cctf` (`bool`, optional) – Apply *sRGB* encoding colour component transfer function / opto-electronic transfer function.

**Returns** *sRGB* colour array.

**Return type** `ndarray`

### Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].

### Examples

```
>>> import numpy as np
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> XYZ_to_sRGB(XYZ)
array([ 0.1749881...,  0.3881947...,  0.3216031...])
```

## colour.sRGB\_to\_XYZ

```
colour.sRGB_to_XYZ(RGB,      illuminant=array([    0.3127,      0.329      J),      chro-
                           matic_adaptation_method=u'CAT02', apply_decoding_cctf=True)
Converts from sRGB colourspace to CIE XYZ tristimulus values.
```

### Parameters

- `RGB` (`array_like`) – *sRGB* colourspace array.
- `illuminant` (`array_like`, optional) – Source illuminant chromaticity coordinates.
- `chromatic_adaptation_method` (`unicode`, optional) – {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}, *Chromatic adaptation* method.
- `apply_decoding_cctf` (`bool`, optional) – Apply *sRGB* decoding colour component transfer function / electro-optical transfer function.

**Returns** *CIE XYZ* tristimulus values.

**Return type** `ndarray`

### Notes

- Input *RGB* colourspace array is in domain [0, 1].

### Examples

```
>>> import numpy as np
>>> RGB = np.array([0.17498172, 0.38818743, 0.32159978])
>>> sRGB_to_XYZ(RGB)
array([ 0.0704953...,  0.1008...,  0.0955831...])
```

## RGB Colourspace Derivation

### colour

<code>normalised_primary_matrix(primaries, whitepoint)</code>	Returns the <i>normalised primary matrix</i> using given <i>primaries</i> and <i>whitepoint xy</i> chromaticity coordinates.
<code>chromatically_adapted_primaries(primaries, ...)</code>	Chromatically adapts given <i>primaries xy</i> chromaticity coordinates from test <i>whitepoint_t</i> to reference <i>whitepoint_r</i> .
<code>primaries_whitepoint(npm)</code>	Returns the <i>primaries</i> and <i>whitepoint xy</i> chromaticity coordinates using given <i>normalised primary matrix</i> .
<code>RGB_luminance(RGB, primaries, whitepoint)</code>	Returns the <i>luminance Y</i> of given <i>RGB</i> components from given <i>primaries</i> and <i>whitepoint</i> .
<code>RGB_luminance_equation(primaries, whitepoint)</code>	Returns the <i>luminance equation</i> from given <i>primaries</i> and <i>whitepoint</i> .

## colour.normalised\_primary\_matrix

`colour.normalised_primary_matrix(primaries, whitepoint)`

Returns the *normalised primary matrix* using given *primaries* and *whitepoint* *xy* chromaticity coordinates.

### Parameters

- **primaries** (array\_like, (3, 2)) – Primaries *xy* chromaticity coordinates.
- **whitepoint** (array\_like) – Illuminant / whitepoint *xy* chromaticity coordinates.

**Returns** Normalised primary matrix.

**Return type** ndarray, (3, 3)

## References

- [SocietyoMPATEngineers93]

## Examples

```
>>> p = np.array([0.73470, 0.26530, 0.00000, 1.00000, 0.00010, -0.07700])
>>> w = np.array([0.32168, 0.33767])
>>> normalised_primary_matrix(p, w)
array([[ 9.5255239...e-01,    0.000000...e+00,    9.3678631...e-05],
       [ 3.4396645...e-01,    7.2816609...e-01,   -7.2132546...e-02],
       [ 0.000000...e+00,    0.000000...e+00,    1.0088251...e+00]])
```

## colour.chromatically\_adapted\_primaries

`colour.chromatically_adapted_primaries(primaries, whitepoint_t, whitepoint_r, chromatic_adaptation_transform='CAT02')`

Chromatically adapts given *primaries* *xy* chromaticity coordinates from test *whitepoint\_t* to reference *whitepoint\_r*.

### Parameters

- **primaries** (array\_like, (3, 2)) – Primaries *xy* chromaticity coordinates.
- **whitepoint\_t** (array\_like) – Test illuminant / whitepoint *xy* chromaticity coordinates.
- **whitepoint\_r** (array\_like) – Reference illuminant / whitepoint *xy* chromaticity coordinates.
- **chromatic\_adaptation\_transform** (unicode, optional) – {'CAT02', 'XYZ Scaling', 'Von Kries', 'Bradford', 'Sharp', 'Fairchild', 'CMCCAT97', 'CMCCAT2000', 'CAT02\_BRILL\_CAT', 'Bianco', 'Bianco PC'}, Chromatic adaptation transform.

**Returns** Chromatically adapted primaries *xy* chromaticity coordinates.

**Return type** ndarray

## Examples

```
>>> p = np.array([0.64, 0.33, 0.30, 0.60, 0.15, 0.06])
>>> w_t = np.array([0.31270, 0.32900])
>>> w_r = np.array([0.34570, 0.35850])
>>> chromatic_adaptation_transform = 'Bradford'
>>> chromatically_adapted_primaries(p, w_t, w_r,
...                                     chromatic_adaptation_transform)
...
...
array([[ 0.6484414...,  0.3308533...],
       [ 0.3211951...,  0.5978443...],
       [ 0.1558932...,  0.0660492...]])
```

## colour.primaries\_whitepoint

`colour.primaries_whitepoint(npm)`

Returns the *primaries* and *whitepoint* *xy* chromaticity coordinates using given *normalised primary matrix*.

**Parameters** `npm` (`array_like`, `(3, 3)`) – *Normalised primary matrix*.

**Returns** *Primaries* and *whitepoint* *xy* chromaticity coordinates.

**Return type** `tuple`

## References

- [\[Tri15\]](#)

## Examples

```
>>> npm = np.array([[9.52552396e-01, 0.00000000e+00, 9.36786317e-05],
...                  [3.43966450e-01, 7.28166097e-01, -7.21325464e-02],
...                  [0.00000000e+00, 0.00000000e+00, 1.00882518e+00]])
>>> p, w = primaries_whitepoint(npm)
>>> p
array([[ 7.3470000...e-01,   2.6530000...e-01],
       [ 0.0000000...e+00,   1.0000000...e+00],
       [ 1.0000000...e-04,  -7.7000000...e-02]])
>>> w
array([ 0.32168,  0.33767])
```

## colour.RGB\_luminance

`colour.RGB_luminance(RGB, primaries, whitepoint)`

Returns the *luminance* *Y* of given *RGB* components from given *primaries* and *whitepoint*.

**Parameters**

- `RGB` (`array_like`) – *RGB* chromaticity coordinate matrix.
- `primaries` (`array_like`, `(3, 2)`) – *Primaries* chromaticity coordinate matrix.

- **whitepoint** (array\_like) – Illuminant / whitepoint chromaticity coordinates.

**Returns** Luminance  $Y$ .

**Return type** numeric or ndarray

## Examples

```
>>> RGB = np.array([40.6, 4.2, 67.4])
>>> p = np.array([0.73470, 0.26530, 0.00000, 1.00000, 0.00010, -0.07700])
>>> whitepoint = np.array([0.32168, 0.33767])
>>> RGB_luminance(RGB, p, whitepoint)
12.1616018...
```

## colour.RGB\_luminance\_equation

`colour.RGB_luminance_equation(primaries, whitepoint)`

Returns the *luminance equation* from given *primaries* and *whitepoint*.

### Parameters

- **primaries** (array\_like, (3, 2)) – Primaries chromaticity coordinates.
- **whitepoint** (array\_like) – Illuminant / whitepoint chromaticity coordinates.

**Returns** Luminance equation.

**Return type** unicode

## Examples

```
>>> p = np.array([0.73470, 0.26530, 0.00000, 1.00000, 0.00010, -0.07700])
>>> whitepoint = np.array([0.32168, 0.33767])
>>> # Doctests skip for Python 2.x compatibility.
>>> RGB_luminance_equation(p, whitepoint)
'Y = 0.3439664...(R) + 0.7281660...(G) + -0.0721325...(B)'
```

## RGB Colourspaces

`colour`

<code>RGB_Colourspace(name, primaries, whitepoint)</code>	Implements support for the <i>RGB</i> colourspaces dataset from <code>colour.models.dataset.aces_rgb</code> , etc....
<code>RGB_COLOURSPACES</code>	Aggregated <i>RGB</i> colourspaces.

## colour.RGB\_Colourspace

```
class colour.RGB_Colourspace(name, primaries, whitepoint, illuminant=None, RGB_to_XYZ_matrix=None, XYZ_to_RGB_matrix=None, encoding_cctf=None, decoding_cctf=None, use_derived_RGB_to_XYZ_matrix=False, use_derived_XYZ_to_RGB_matrix=False)
```

Implements support for the *RGB* colourspaces dataset from `colour.models.dataset.aces_rgb`, etc....

Colour science literature related to *RGB* colourspaces and encodings defines their dataset using different degree of precision or rounding. While instances where a whitepoint is being defined with a value different than its canonical agreed one are rare, it is however very common to have normalised primary matrices rounded at different decimals. This can yield large discrepancies in computations.

Such an occurrence is the *V-Gamut* colourspace white paper, that defines the *V-Gamut* to *ITU-R BT.709* conversion matrix as follows:

```
[[ 1.806576 -0.695697 -0.110879]
 [-0.170090  1.305955 -0.135865]
 [-0.025206 -0.154468  1.179674]]
```

Computing this matrix using *ITU-R BT.709* colourspace derived normalised primary matrix yields:

```
[[ 1.8065736 -0.6956981 -0.1108786]
 [-0.1700890  1.3059548 -0.1358648]
 [-0.0252057 -0.1544678  1.1796737]]
```

The latter matrix is almost equals with the former, however performing the same computation using *IEC 61966-2-1:1999 sRGB* colourspace normalised primary matrix introduces severe disparities:

```
[[ 1.8063853 -0.6956147 -0.1109453]
 [-0.1699311  1.3058387 -0.1358616]
 [-0.0251630 -0.1544899  1.1797117]]
```

In order to provide support for both literature defined dataset and accurate computations enabling transformations without loss of precision, the `colour.RGB_Colourspace` class provides two sets of transformation matrices:

- Instantiation transformation matrices
- Derived transformation matrices

Upon instantiation, the `colour.RGB_Colourspace` class stores the given `RGB_to_XYZ_matrix` and `XYZ_to_RGB_matrix` arguments and also computes their derived counterpart using the `primaries` and `whitepoint` arguments.

Whether the initialisation or derived matrices are used in subsequent computations is dependent on the `colour.RGB_Colourspace.use_derived_RGB_to_XYZ_matrix` and `colour.RGB_Colourspace.use_derived_XYZ_to_RGB_matrix` attributes values.

### Parameters

- `name` (unicode) – *RGB* colourspace name.
- `primaries` (array\_like) – *RGB* colourspace primaries.
- `whitepoint` (array\_like) – *RGB* colourspace whitepoint.
- `illuminant` (unicode, optional) – *RGB* colourspace whitepoint name as illuminant.

- `RGB_to_XYZ_matrix` (`array_like`, optional) – Transformation matrix from colourspace to *CIE XYZ* tristimulus values.
- `XYZ_to_RGB_matrix` (`array_like`, optional) – Transformation matrix from *CIE XYZ* tristimulus values to colourspace.
- `encoding_cctf` (`object`, optional) – Encoding colour component transfer function (Encoding CCTF) / opto-electronic transfer function (OETF / OECF) that maps estimated tristimulus values in a scene to  $R'G'B'$  video component signal value.
- `decoding_cctf` (`object`, optional) – Decoding colour component transfer function (Decoding CCTF) / electro-optical transfer function (EOTF / EOCF) that maps an  $R'G'B'$  video component signal value to tristimulus values at the display.
- `use_derived_RGB_to_XYZ_matrix` (`bool`, optional) – Whether to use the instantiation time normalised primary matrix or to use a computed derived normalised primary matrix.
- `use_derived_XYZ_to_RGB_matrix` (`bool`, optional) – Whether to use the instantiation time inverse normalised primary matrix or to use a computed derived inverse normalised primary matrix.

```
name
primaries
whitepoint
illuminant
RGB_to_XYZ_matrix
XYZ_to_RGB_matrix
encoding_cctf
decoding_cctf
use_derived_RGB_to_XYZ_matrix
use_derived_XYZ_to_RGB_matrix
__str__()
__repr__()
use_derived_transformation_matrices()
```

## Notes

- The normalised primary matrix defined by `colour.RGB_Colourspace.RGB_to_XYZ_matrix` attribute is treated as the prime matrix from which the inverse will be calculated as required by the internal derivation mechanism. This behaviour has been chosen in accordance with literature where commonly a *RGB* colourspace is defined by its normalised primary matrix as it is directly computed from the chosen primaries and whitepoint.

## References

- [InternationalECCommission99]
- [Pan14]

## Examples

```
>>> p = np.array([0.73470, 0.26530, 0.00000, 1.00000, 0.00010, -0.07700])
>>> whitepoint = np.array([0.32168, 0.33767])
>>> RGB_to_XYZ_matrix = np.identity(3)
>>> XYZ_to_RGB_matrix = np.identity(3)
>>> colourspace = RGB_Colourspace('RGB_Colourspace', p, whitepoint, 'D60',
...                                 RGB_to_XYZ_matrix, XYZ_to_RGB_matrix)
>>> colourspace.RGB_to_XYZ_matrix
array([[ 1.,  0.,  0.],
       [ 0.,  1.,  0.],
       [ 0.,  0.,  1.]])
>>> colourspace.XYZ_to_RGB_matrix
array([[ 1.,  0.,  0.],
       [ 0.,  1.,  0.],
       [ 0.,  0.,  1.]])
>>> colourspace.use_derived_transformation_matrices(True)
True
>>> colourspace.RGB_to_XYZ_matrix
array([[ 9.5255239...e-01,  0.0000000...e+00,  9.3678631...e-05],
       [ 3.4396645...e-01,  7.2816609...e-01, -7.2132546...e-02],
       [ 0.0000000...e+00,  0.0000000...e+00,  1.0088251...e+00]])
>>> colourspace.XYZ_to_RGB_matrix
array([[ 1.0498110...e+00,  0.0000000...e+00, -9.7484540...e-05],
       [-4.9590302...e-01,  1.3733130...e+00,  9.8240036...e-02],
       [ 0.0000000...e+00,  0.0000000...e+00,  9.9125201...e-01]])
>>> colourspace.use_derived_RGB_to_XYZ_matrix = False
>>> colourspace.RGB_to_XYZ_matrix
array([[ 1.,  0.,  0.],
       [ 0.,  1.,  0.],
       [ 0.,  0.,  1.]])
>>> colourspace.use_derived_XYZ_to_RGB_matrix = False
>>> colourspace.XYZ_to_RGB_matrix
array([[ 1.,  0.,  0.],
       [ 0.,  1.,  0.],
       [ 0.,  0.,  1.]])
```

```
__init__(name, primaries, whitepoint, illuminant=None, RGB_to_XYZ_matrix=None,
        XYZ_to_RGB_matrix=None, encoding_cctf=None, decoding_cctf=None,
        use_derived_RGB_to_XYZ_matrix=False, use_derived_XYZ_to_RGB_matrix=False)
```

## Methods

---

<code>__init__(name, primaries, whitepoint[...])</code>	
<code>use_derived_transformation_matrices([usage])</code>	Enables or disables usage of both derived transformations matrices, the normalised primary matrix and its inverse in subsequent computations.

---

## colour.RGB\_COLOURSPACES

```
colour.RGB_COLOURSPACES = CaseInsensitiveMapping({u'ACEScc': ..., u'ACEScg': ..., u'Protune Native': ...,
                                                 Aggregated RGB colourespaces.})
```

RGB\_COLOURSPACES : CaseInsensitiveMapping

## Aliases:

- ‘aces’: ACES\_2065\_1\_COLOURSPACE.name
- ‘adobe1998’: ADOBE\_RGB\_1998\_COLOURSPACE.name
- ‘prophoto’: PROPHOTO\_RGB\_COLOURSPACE.name

colour.models

ACES_2065_1_COLOURSPACE	ACES2065-1 colourspace, base encoding, used for exchange of full fidelity
ACES_CC_COLOURSPACE	ACEScc colourspace, a working space for color correctors, target for ASC-CDL
ACES_CCT_COLOURSPACE	ACESct colourspace, an alternative working space for colour correctors,
ACES_PROXY_COLOURSPACE	ACESproxy colourspace, a lightweight encoding for transmission over HD-SDI
ACES_CG_COLOURSPACE	ACEScg colourspace, a working space for paint/compositor applications that
ADOBE_RGB_1998_COLOURSPACE	Adobe RGB (1998) colourspace.
ADOBE_WIDE_GAMUT_RGB_COLOURSPACE	Adobe Wide Gamut RGB colourspace.
ALEXA_WIDE_GAMUT_COLOURSPACE	ALEXA Wide Gamut colourspace.
APPLE_RGB_COLOURSPACE	Apple RGB colourspace.
BEST_RGB_COLOURSPACE	Best RGB colourspace.
BETA_RGB_COLOURSPACE	Beta RGB colourspace.
BT470_525_COLOURSPACE	ITU-R BT.470 - 525 colourspace.
BT470_625_COLOURSPACE	ITU-R BT.470 - 625 colourspace.
BT709_COLOURSPACE	ITU-R BT.709 colourspace.
BT2020_COLOURSPACE	ITU-R BT.2020 colourspace.
CIE_RGB_COLOURSPACE	CIE RGB colourspace.
CINEMA_GAMUT_COLOURSPACE	Cinema Gamut colourspace.
COLOR_MATCH_RGB_COLOURSPACE	ColorMatch RGB colourspace.
DCI_P3_COLOURSPACE	DCI-P3 colourspace.
DCI_P3_P_COLOURSPACE	DCI-P3+ colourspace.
DON_RGB_4_COLOURSPACE	Don RGB 4 colourspace.
ECI_RGB_V2_COLOURSPACE	ECI RGB v2 colourspace.
EKTA_SPACE_PS_5_COLOURSPACE	Ekta Space PS 5 colourspace.
PROTUNE_NATIVE_COLOURSPACE	Protune Native colourspace.
MAX_RGB_COLOURSPACE	Max RGB colourspace.
NTSC_COLOURSPACE	NTSC colourspace.
PAL_SECAM_COLOURSPACE	Pal/Secam colourspace.
RED_COLOR_COLOURSPACE	REDcolor colourspace.
RED_COLOR_2_COLOURSPACE	REDcolor2 colourspace.
RED_COLOR_3_COLOURSPACE	REDcolor3 colourspace.
RED_COLOR_4_COLOURSPACE	REDcolor4 colourspace.
RED_WIDE_GAMUT_RGB_COLOURSPACE	REDWideGamutRGB colourspace.
DRAGON_COLOR_COLOURSPACE	DRAGONcolor colourspace.
DRAGON_COLOR_2_COLOURSPACE	DRAGONcolor2 colourspace.
ROMM_RGB_COLOURSPACE	ROMM RGB colourspace.
RIMM_RGB_COLOURSPACE	RIMM RGB colourspace. In cases in which it is necessary to identify a
ERIMM_RGB_COLOURSPACE	ERIMM RGB colourspace.

Continued on next page

Table 3.141 – continued from previous page

PROPHOTO_RGB_COLOURSPACE	<i>ProPhoto RGB</i> colourspace, an alias colourspace for <i>ROMM RGB</i> .
RUSSELL_RGB_COLOURSPACE	<i>Russell RGB</i> colourspace.
SMPTE_240M_COLOURSPACE	<i>SMPTE 240M</i> colourspace.
S_GAMUT_COLOURSPACE	<i>S-Gamut</i> colourspace.
S_GAMUT3_COLOURSPACE	<i>S-Gamut3</i> colourspace.
S_GAMUT3_CINE_COLOURSPACE	<i>S-Gamut3.Cine</i> colourspace.
sRGB_COLOURSPACE	<i>sRGB</i> colourspace.
V_GAMUT_COLOURSPACE	<i>V-Gamut</i> colourspace.
XTREME_RGB_COLOURSPACE	<i>Xtreme RGB</i> colourspace.

colour.models.ACES\_2065\_1\_COLOURSPACE

colour.models.ACES\_2065\_1\_COLOURSPACE = RGB\_Colourspace(ACES2065-1, [[ 7.34700000e-01, 2.65300000e-01], [ 0.0, 1.0, 0.0], [ 0.0, 0.0, 1.0]], "ACES2065-1 colourspace, base encoding, used for exchange of full fidelity images and archiving.

## References

- [The AoMPAa Sciences Science at Council Academy CESAC ESP Subcommittee 14c]
  - [The AoMPAa Sciences Science at Council Academy CESAC ESP Subcommittee 14d]
  - [The AoMPAa Sciences Science at Council Academy CESAC ESP Subcommittee ]

ACES 2065 1 COLOURSPACE : RGB Colourspace

## colour.models.ACES\_CC\_COLOURSPACE

`colour.models.ACES_CC_COLOURSPACE = RGB_Colourspace(ACEScc, [[ 0.713, 0.293], [ 0.165, 0.83 ], [ 0.128, 0.04 ]])`  
ACEScc colourspace, a working space for color correctors, target for ASC-CDL values created on-set.

## References

- [The AoMPAa Sciences Science aTC Council Academy CESAC ESP Subcommittee 14c]
  - [The AoMPAa Sciences Science aTC Council Academy CESAC ESP Subcommittee 14d]
  - [The AoMPAa Sciences Science aTC Council Academy CESAC ESP Subcommittee 14b]
  - [The AoMPAa Sciences Science aTC Council Academy CESAC ESP Subcommittee 1]

## ACES CC COLOURSPACE : RGB Colourspace

colour.models.ACES\_CCT\_COLOURSPACE

`colour.models.ACES_CCT_COLOURSPACE = RGB_Colourspace(ACESct, [[ 0.713, 0.293], [ 0.165, 0.83 ], [ 0.128, 0.05 ]])`

*ACESct* colourspace, an alternative working space for colour correctors, intended to be transient and internal to software or hardware systems, and is specifically not intended for interchange or archiving.

## References

- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 14c\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 14d\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Project 16\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee\]](#)

ACES\_CCT\_COLOURSPACE : RGB\_Colourspace

### colour.models.ACES\_PROXY\_COLOURSPACE

`colour.models.ACES_PROXY_COLOURSPACE = RGB_Colourspace(ACESproxy, [[ 0.713, 0.293], [ 0.165, 0.83 ], [ 0.128,`  
ACESproxy colourspace, a lightweight encoding for transmission over HD-SDI (or other production  
transmission schemes), onset look management. Not intended to be stored or used in production  
imagery or for final colour grading / mastering.

## References

- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 14c\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 14d\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 14a\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee\]](#)

ACES\_PROXY\_COLOURSPACE : RGB\_Colourspace

### colour.models.ACES(CG)\_COLOURSPACE

`colour.models.ACES(CG)_COLOURSPACE = RGB_Colourspace(ACEScg, [[ 0.713, 0.293], [ 0.165, 0.83 ], [ 0.128, 0.044,`  
ACEScg colourspace, a working space for paint/compositor applications that don't support ACES2065-1  
or ACEScc.

## References

- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 14c\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 14d\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee 15\]](#)
- [\[The AoMPAa Sciences Science aTC Council Academy CESACES Subcommittee\]](#)

ACES(CG)\_COLOURSPACE : RGB\_Colourspace

### colour.models.ADOBE\_RGB\_1998\_COLOURSPACE

`colour.models.ADOBE_RGB_1998_COLOURSPACE = RGB_Colourspace(Adobe RGB (1998), [[ 0.64, 0.33], [ 0.21, 0.71], [ 0.17,`  
Adobe RGB (1998) colourspace.

## References

- [\[AdobeSystems05\]](#)

ADOBECOLORMODEL\_ADOBE\_RGB\_1998\_COLOURSPACE : RGB\_Colourspace

### colour.models.ADOBE\_WIDE\_GAMUT\_RGB\_COLOURSPACE

```
colour.models.ADOBE_WIDE_GAMUT_RGB_COLOURSPACE = RGB_Colourspace(Adobe Wide Gamut RGB, [[ 0.7347, 0.2653], [ 0.2653, 0.7347], [ 0.485, 0.485]], Adobe Wide Gamut RGB colourspace.
```

## References

- [\[Wik-1\]](#)

ADOBECOLORMODEL\_ADOBE\_WIDE\_GAMUT\_RGB\_COLOURSPACE : RGB\_Colourspace

### colour.models.ALEXA\_WIDE\_GAMUT\_COLOURSPACE

```
colour.models.ALEXA_WIDE_GAMUT_COLOURSPACE = RGB_Colourspace(ALEXA Wide Gamut, [[ 0.684, 0.313], [ 0.221, 0.684], [ 0.313, 0.221]], ALEXA Wide Gamut colourspace.
```

## References

- [\[ARR12\]](#)

ALEXACOLORMODEL\_ALEXA\_WIDE\_GAMUT\_COLOURSPACE : RGB\_Colourspace

### colour.models.APPLE\_RGB\_COLOURSPACE

```
colour.models.APPLE_RGB_COLOURSPACE = RGB_Colourspace(Apple RGB, [[ 0.625, 0.34], [ 0.28, 0.595], [ 0.155, 0.28]], Apple RGB colourspace.
```

## References

- [\[SBS99\]](#)

APPLECOLORMODEL\_APPLE\_RGB\_COLOURSPACE : RGB\_Colourspace

### colour.models.BEST\_RGB\_COLOURSPACE

```
colour.models.BEST_RGB_COLOURSPACE = RGB_Colourspace(Best RGB, [[ 0.73519164, 0.26480836], [ 0.21533613, 0.73519164], [ 0.485, 0.485]], Best RGB colourspace.
```

## References

- [\[Huta\]](#)

BEST\_RGB\_COLOURSPACE : RGB\_Colourspace

### colour.models.BETA\_RGB\_COLOURSPACE

```
colour.models.BETA_RGB_COLOURSPACE = RGB_Colourspace(Beta RGB, [[ 0.6888, 0.3112], [ 0.1986, 0.7551], [ 0.1200, 0.0498], [ 0.0498, 0.1200], [ 0.0498, 0.0498]], "Beta RGB colourspace.")
```

## References

- [\[Lin14\]](#)

BETA\_RGB\_COLOURSPACE : RGB\_Colourspace

### colour.models.BT470\_525\_COLOURSPACE

```
colour.models.BT470_525_COLOURSPACE = RGB_Colourspace(ITU-R BT.470 - 525, [[ 0.67, 0.33], [ 0.21, 0.71], [ 0.15, 0.06], [ 0.06, 0.15], [ 0.06, 0.06]], "ITU-R BT.470 - 525 colourspace.")
```

## References

- [\[InternationalTUnion98\]](#)

BT470\_525\_COLOURSPACE : RGB\_Colourspace

### colour.models.BT470\_625\_COLOURSPACE

```
colour.models.BT470_625_COLOURSPACE = RGB_Colourspace(ITU-R BT.470 - 625, [[ 0.64, 0.33], [ 0.29, 0.6 ], [ 0.15, 0.06], [ 0.06, 0.15], [ 0.06, 0.06]], "ITU-R BT.470 - 625 colourspace.")
```

## References

- [\[InternationalTUnion98\]](#)

BT470\_625\_COLOURSPACE : RGB\_Colourspace

### colour.models.BT709\_COLOURSPACE

```
colour.models.BT709_COLOURSPACE = RGB_Colourspace(ITU-R BT.709, [[ 0.64, 0.33], [ 0.3 , 0.6 ], [ 0.15, 0.06], [ 0.06, 0.15], [ 0.06, 0.06]], "ITU-R BT.709 colourspace.")
```

## References

- [\[InternationalTUnion15b\]](#)

BT709\_COLOURSPACE : RGB\_Colourspace

### colour.models.BT2020\_COLOURSPACE

```
colour.models.BT2020_COLOURSPACE = RGB_Colourspace(ITU-R BT.2020, [[ 0.708, 0.292], [ 0.17 , 0.797], [ 0.131,
```

*ITU-R BT.2020 colourspace.*

## References

- [\[InternationalTUnion15a\]](#)

BT2020\_COLOURSPACE : RGB\_Colourspace

### colour.models.CIE\_RGB\_COLOURSPACE

```
colour.models.CIE_RGB_COLOURSPACE = RGB_Colourspace(CIE RGB, [[ 0.73474284, 0.26525716], [ 0.27377903, 0.7174,
```

*CIE RGB colourspace.*

## References

- [\[FBH97\]](#)

CIE\_RGB\_COLOURSPACE : RGB\_Colourspace

### colour.models.CINEMA\_GAMUT\_COLOURSPACE

```
colour.models.CINEMA_GAMUT_COLOURSPACE = RGB_Colourspace(Cinema Gamut, [[ 0.74, 0.27], [ 0.17, 1.14], [ 0.08,
```

*Cinema Gamut colourspace.*

## References

- [\[Can14\]](#)

CINEMA\_GAMUT\_COLOURSPACE : RGB\_Colourspace

### colour.models.COLOR\_MATCH\_RGB\_COLOURSPACE

```
colour.models.COLOR_MATCH_RGB_COLOURSPACE = RGB_Colourspace(ColorMatch RGB, [[ 0.63 , 0.34 ], [ 0.295, 0.605,
```

*ColorMatch RGB colourspace.*

## References

- [\[Lin14\]](#)

COLOR\_MATCH\_RGB\_COLOURSPACE : RGB\_Colourspace

### colour.models.DCI\_P3\_COLOURSPACE

```
colour.models.DCI_P3_COLOURSPACE = RGB_Colourspace(DCI-P3, [[ 0.68 , 0.32 ], [ 0.265, 0.69 ], [ 0.15 , 0.06 ]], "DCI-P3 colourspace.
```

## References

- [\[DigitalCInitiatives07\]](#)
- [\[HewlettPDCompany09\]](#)

DCI\_P3\_COLOURSPACE : RGB\_Colourspace

### colour.models.DCI\_P3\_P\_COLOURSPACE

```
colour.models.DCI_P3_P_COLOURSPACE = RGB_Colourspace(DCI-P3+, [[ 0.74, 0.27], [ 0.22, 0.78], [ 0.09, -0.09]], "DCI-P3+ colourspace.
```

## References

- [\[Can14\]](#)

DCI\_P3\_P\_COLOURSPACE : RGB\_Colourspace

### colour.models.DON\_RGB\_4\_COLOURSPACE

```
colour.models.DON_RGB_4_COLOURSPACE = RGB_Colourspace(Don RGB 4, [[ 0.69612069, 0.29956897], [ 0.21468298, 0.09956897], [ 0.09956897, 0.69612069]], "Don RGB 4 colourspace.
```

## References

- [\[Hutb\]](#)

DON\_RGB\_4\_COLOURSPACE : RGB\_Colourspace

### colour.models.ECI\_RGB\_V2\_COLOURSPACE

```
colour.models.ECI_RGB_V2_COLOURSPACE = RGB_Colourspace(ECI RGB v2, [[ 0.67010309, 0.32989691], [ 0.20990566, 0.09956897], [ 0.09956897, 0.67010309]], "ECI RGB v2 colourspace.
```

## References

- [EuropeanCInitiative02]

ECI\_RGB\_V2\_COLOURSPACE : RGB\_Colourspace

## colour.models.EKTA\_SPACE\_PS\_5\_COLOURSPACE

`colour.models.EKTA_SPACE_PS_5_COLOURSPACE = RGB_Colourspace(Ekta Space PS 5, [[ 0.69473684, 0.30526316], [ 0.299, 0.587, 0.114], [ 0.174, 0.342, 0.522], [ 0.0, 0.0, 0.0], [ 1.0, 1.0, 1.0]], "Ekta Space PS 5 colourspace.`

## References

- [Hol]

EKTA\_SPACE\_PS\_5\_COLOURSPACE : RGB\_Colourspace

## colour.models.PROTUNE\_NATIVE\_COLOURSPACE

colour.models.PROTUNE\_NATIVE\_COLOURSPACE = RGB\_Colourspace(Protune Native, [[ 0.69848046, 0.19302645], [ 0.31740157, 0.92125008], [ 0.19302645, 0.69848046]], "Protune Native colourspace.

## References

- [GDM16]
  - [Man15]

PROTUNE\_NATIVE\_COLOURSPACE : RGB\_Colourspace

## colour.models.MAX\_RGB\_COLOURSPACE

## References

- [Hutc]

MAX\_RGB\_COLOURSPACE : RGB\_Colourspace

colour.models.NTSC COLOURSPACE

colour.models.NTSC\_COLOURSPACE = RGB\_Colourspace(NTSC, [[ 0.67, 0.33], [ 0.21, 0.71], [ 0.14, 0.08]], [ 0.31, 0.31])

## References

- [\[InternationalTUnion98\]](#)

NTSC\_COLOURSPACE : RGB\_Colourspace

### colour.models.PAL\_SECAM\_COLOURSPACE

colour.models.PAL\_SECAM\_COLOURSPACE = RGB\_Colourspace(Pal/Secam, [[ 0.64, 0.33], [ 0.29, 0.6 ], [ 0.15, 0.06 ]],  
Pal/Secam colourspace.

## References

- [\[InternationalTUnion98\]](#)

PAL\_SECAM\_COLOURSPACE : RGB\_Colourspace

### colour.models.RED\_COLOR\_COLOURSPACE

colour.models.RED\_COLOR\_COLOURSPACE = RGB\_Colourspace(REDcolor, [[ 0.699747 , 0.32904693], [ 0.30426404, 0.62403166 ]],  
REDcolor colourspace.

## References

- [\[Man15\]](#)
- [\[SonyImageworks12\]](#)

RED\_COLOR\_COLOURSPACE : RGB\_Colourspace

### colour.models.RED\_COLOR\_2\_COLOURSPACE

colour.models.RED\_COLOR\_2\_COLOURSPACE = RGB\_Colourspace(REDcolor2, [[ 0.87868251, 0.32496401], [ 0.30088871, 0.62403166 ]],  
REDcolor2 colourspace.

## References

- [\[Man15\]](#)
- [\[SonyImageworks12\]](#)

RED\_COLOR\_2\_COLOURSPACE : RGB\_Colourspace

### colour.models.RED\_COLOR\_3\_COLOURSPACE

colour.models.RED\_COLOR\_3\_COLOURSPACE = RGB\_Colourspace(REDcolor3, [[ 0.70118104, 0.32901416], [ 0.3006003 , 0.62403166 ]],  
REDcolor3 colourspace.

## References

- [\[Man15\]](#)
- [\[SonyImageworks12\]](#)

RED\_COLOR\_3\_COLOURSPACE : RGB\_Colourspace

### colour.models.RED\_COLOR\_4\_COLOURSPACE

colour.models.RED\_COLOR\_4\_COLOURSPACE = RGB\_Colourspace(REDcolor4, [[ 0.70118059, 0.3290137 ], [ 0.3006004 ,  
REDcolor4 colourspace.

## References

- [\[Man15\]](#)
- [\[SonyImageworks12\]](#)

RED\_COLOR\_4\_COLOURSPACE : RGB\_Colourspace

### colour.models.RED\_WIDE\_GAMUT\_RGB\_COLOURSPACE

colour.models.RED\_WIDE\_GAMUT\_RGB\_COLOURSPACE = RGB\_Colourspace(REDWideGamutRGB, [[ 0.780308, 0.304253], [ 0.1000000 ,  
REDWideGamutRGB colourspace.

## References

- [\[Man15\]](#)
- [\[Nat16\]](#)
- [\[SonyImageworks12\]](#)

RED\_WIDE\_GAMUT\_RGB\_COLOURSPACE : RGB\_Colourspace

### colour.models.DRAGON\_COLOR\_COLOURSPACE

colour.models.DRAGON\_COLOR\_COLOURSPACE = RGB\_Colourspace(DRAGONcolor, [[ 0.75304422, 0.32783058], [ 0.2995702 ,  
DRAGONcolor colourspace.

## References

- [\[Man15\]](#)
- [\[SonyImageworks12\]](#)

DRAGON\_COLOR\_COLOURSPACE : RGB\_Colourspace

## colour.models.DRAGON\_COLOR\_2\_COLOURSPACE

```
colour.models.DRAGON_COLOR_2_COLOURSPACE = RGB_Colourspace(DRAGONcolor2, [[ 0.75304449, 0.32783103], [ 0.29951441, 0.14441111], [ 0.14441111, 0.06666667]], "DRAGONcolor2 colourspace.
```

### References

- [\[Man15\]](#)
- [\[SonyImageworks12\]](#)

DRAGON\_COLOR\_2\_COLOURSPACE : RGB\_Colourspace

## colour.models.ROMM\_RGB\_COLOURSPACE

```
colour.models.ROMM_RGB_COLOURSPACE = RGB_Colourspace(ROMM RGB, [[ 7.34700000e-01, 2.65300000e-01], [ 1.59600000e-01, 4.19200000e-01], [ 1.59600000e-01, 4.19200000e-01]], "ROMM RGB colourspace.
```

### References

- [\[ANS03\]](#)
- [\[SWG00\]](#)

ROMM\_RGB\_COLOURSPACE : RGB\_Colourspace

## colour.models.RIMM\_RGB\_COLOURSPACE

```
colour.models.RIMM_RGB_COLOURSPACE = RGB_Colourspace(RIMM RGB, [[ 7.34700000e-01, 2.65300000e-01], [ 1.59600000e-01, 4.19200000e-01], [ 1.59600000e-01, 4.19200000e-01]], "RIMM RGB colourspace. In cases in which it is necessary to identify a specific precision level, the notation RIMM8 RGB, RIMM12 RGB and RIMM16 RGB is used.
```

### References

- [\[SWG00\]](#)

RIMM\_RGB\_COLOURSPACE : RGB\_Colourspace

## colour.models.ERIMM\_RGB\_COLOURSPACE

```
colour.models.ERIMM_RGB_COLOURSPACE = RGB_Colourspace(ERIMM RGB, [[ 7.34700000e-01, 2.65300000e-01], [ 1.59600000e-01, 4.19200000e-01], [ 1.59600000e-01, 4.19200000e-01]], "ERIMM RGB colourspace.
```

### References

- [\[SWG00\]](#)

ERIMM\_RGB\_COLOURSPACE : RGB\_Colourspace

## colour.models.PROPHOTO\_RGB\_COLOURSPACE

colour.models.**PROPHOTO\_RGB\_COLOURSPACE** = RGB\_Colourspace(ProPhoto RGB, [[ 7.3470000e-01, 2.6530000e-01], [ 0.1745, 0.1255], [ 0.0945, 0.0745], [ 0.0445, 0.0345], [ 0.0145, 0.0045], [ 0.0045, 0.0045]]).  
ProPhoto RGB colourspace, an alias colourspace for ROMM RGB.

### References

- [\[ANS03\]](#)
- [\[SWG00\]](#)

PROPHOTO\_RGB\_COLOURSPACE : RGB\_Colourspace

## colour.models.RUSSELL\_RGB\_COLOURSPACE

colour.models.**RUSSELL\_RGB\_COLOURSPACE** = RGB\_Colourspace(Russell RGB, [[ 0.69, 0.31], [ 0.18, 0.77], [ 0.1, 0.1], [ 0.01, 0.01], [ 0.001, 0.001]]).  
Russell RGB colourspace.

### References

- [\[Cot\]](#)

RUSSELL\_RGB\_COLOURSPACE : RGB\_Colourspace

## colour.models.SMPTE\_240M\_COLOURSPACE

colour.models.**SMPTE\_240M\_COLOURSPACE** = RGB\_Colourspace(SMPTE 240M, [[ 0.63, 0.34], [ 0.31, 0.595], [ 0.155, 0.155], [ 0.01, 0.01], [ 0.001, 0.001]]).  
SMPTE 240M colourspace.

### References

- [\[SocietyoMPaTEngineers99\]](#)
- [\[SocietyoMPaTEngineers04\]](#)

SMPTE\_240M\_COLOURSPACE : RGB\_Colourspace

## colour.models.S\_GAMUT\_COLOURSPACE

colour.models.**S\_GAMUT\_COLOURSPACE** = RGB\_Colourspace(S-Gamut, [[ 0.73, 0.28], [ 0.14, 0.855], [ 0.1, -0.055], [ 0.01, 0.01], [ 0.001, 0.001]]).  
S-Gamut colourspace.

### References

- [\[GDY+\]](#)
- [\[SonyCorporationb\]](#)

S\_GAMUT\_COLOURSPACE : RGB\_Colourspace

#### colour.models.S\_GAMUT3\_COLOURSPACE

colour.models.S\_GAMUT3\_COLOURSPACE = RGB\_Colourspace(S-Gamut3, [[ 0.73 , 0.28 ], [ 0.14 , 0.855], [ 0.1 , -0.05 ]], S-Gamut3 colourspace.

#### References

- [\[SonyCorporationc\]](#)

S\_GAMUT3\_COLOURSPACE : RGB\_Colourspace

#### colour.models.S\_GAMUT3\_CINE\_COLOURSPACE

colour.models.S\_GAMUT3\_CINE\_COLOURSPACE = RGB\_Colourspace(S-Gamut3.Cine, [[ 0.766, 0.275], [ 0.225, 0.8 ], [ 0.15, -0.05 ]], S-Gamut3.Cine colourspace.

#### References

- [\[SonyCorporationa\]](#)

S\_GAMUT3\_CINE\_COLOURSPACE : RGB\_Colourspace

#### colour.models.sRGB\_COLOURSPACE

colour.models.sRGB\_COLOURSPACE = RGB\_Colourspace(sRGB, [[ 0.64, 0.33], [ 0.3 , 0.6 ], [ 0.15, 0.06]], [ 0.3125, 0.3125, 0.3125 ], sRGB colourspace.

#### References

- [\[InternationalECommission99\]](#)
- [\[InternationalTUnion15b\]](#)

sRGB\_COLOURSPACE : RGB\_Colourspace

#### colour.models.V\_GAMUT\_COLOURSPACE

colour.models.V\_GAMUT\_COLOURSPACE = RGB\_Colourspace(V-Gamut, [[ 0.73 , 0.28 ], [ 0.165, 0.84 ], [ 0.1 , -0.05 ]], V-Gamut colourspace.

#### References

- [\[Pan14\]](#)

V\_GAMUT\_COLOURSPACE : RGB\_Colourspace

## colour.models.XTREME\_RGB\_COLOURSPACE

```
colour.models.XTREME_RGB_COLOURSPACE = RGB_Colourspace(Xtreme RGB, [[ 1., 0.], [ 0., 1.], [ 0., 0.]], [ 0.345, 0.345, 0.345])
```

Xtreme RGB colourspace.

### References

- [\[Hutd\]](#)

XTREME\_RGB\_COLOURSPACE : RGB\_Colourspace

### Ancillary Objects

colour.models

spectral_to_aces_relative_exposure_values(spd)	Converts given spectral power distribution to ACES2065-1 colourspace relative exposure values.
ACES_RICD	Implements support for the <i>CIE RGB</i> colour matching functions.

## colour.models.spectral\_to\_aces\_relative\_exposure\_values

```
colour.models.spectral_to_aces_relative_exposure_values(spd, illuminant=SpectralPowerDistribution(name='D60', ...))
```

Converts given spectral power distribution to ACES2065-1 colourspace relative exposure values.

### Parameters

- **spd** (`SpectralPowerDistribution`) – Spectral power distribution.
- **illuminant** (`SpectralPowerDistribution`, optional) – Illuminant spectral power distribution.

**Returns** ACES2065-1 colourspace relative exposure values array.

**Return type** ndarray, (3,)

### Notes

- Output ACES2065-1 colourspace relative exposure values array is in range [0, 1].

### References

- [\[The AoMPAa Sciences Science T Council Academy CESACESP Subcommittee 14c\]](#)
- [\[The AoMPAa Sciences Science T Council Academy CESACESP Subcommittee 14d\]](#)
- [\[The AoMPAa Sciences Science T Council Academy CESACESP Subcommittee\]](#)

## Examples

```
>>> from colour import COLOURCHECKERS_SPDS
>>> spd = COLOURCHECKERS_SPDS['ColorChecker N Ohta']['dark skin']
>>> spectral_to_aces_relative_exposure_values(spd)
array([ 0.1187697...,  0.0870866...,  0.0589442...])
```

## colour.models.ACES\_RICD

`colour.models.ACES_RICD = RGB_ColourMatchingFunctions(name='ACES RICD', ...)`

Implements support for the *CIE RGB* colour matching functions.

### Parameters

- **data** (Series or Dataframe or Signal or MultiSignal or MultiSpectralPowerDistribution or array\_like or dict\_like, optional) – Data to be stored in the multi-spectral power distribution.
- **domain** (array\_like, optional) – Values to initialise the multiple `colour.SpectralPowerDistribution` class instances `colour.continuous.Signal.wavelengths` attribute with. If both data and domain arguments are defined, the latter will be used to initialise the `colour.continuous.Signal.wavelengths` attribute.
- **labels** (array\_like, optional) – Names to use for the `colour.SpectralPowerDistribution` class instances.

### Other Parameters

- **name** (unicode, optional) – Multi-spectral power distribution name.
- **interpolator** (object, optional) – Interpolator class type to use as interpolating function for the `colour.SpectralPowerDistribution` class instances.
- **interpolator\_args** (dict\_like, optional) – Arguments to use when instantiating the interpolating function of the `colour.SpectralPowerDistribution` class instances.
- **extrapolator** (object, optional) – Extrapolator class type to use as extrapolating function for the `colour.SpectralPowerDistribution` class instances.
- **extrapolator\_args** (dict\_like, optional) – Arguments to use when instantiating the extrapolating function of the `colour.SpectralPowerDistribution` class instances.
- **strict\_labels** (array\_like, optional) – Multi-spectral power distribution labels for figures, default to `colour.colorimetry.RGB_ColourMatchingFunctions.labels` attribute value.

## Opto-Electronic Transfer Functions

`colour`

<code>oetf(value[, function])</code>	Encodes estimated tristimulus values in a scene to $R'G'B'$ video component signal value using given opto-electronic transfer function (OETF / OECF).
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Continued on next page

Table 3.143 – continued from previous page

OETFS	Supported opto-electrical transfer functions (OETFs / OECFs).
<code>oetf_reverse(value[, function])</code>	Decodes $R'G'B'$ video component signal value to tristimulus values at the display using given reverse opto-electronic transfer function (OETF / OECF).
OETFS_REVERSE	Supported reverse opto-electrical transfer functions (OETFs / OECFs).

## colour.oetf

`colour.oetf(value, function='sRGB', **kwargs)`

Encodes estimated tristimulus values in a scene to  $R'G'B'$  video component signal value using given opto-electronic transfer function (OETF / OECF).

### Parameters

- `value` (numeric or array\_like) – Value.
- `function` (unicode, optional) – {‘sRGB’, ‘ARIB STD-B67’, ‘DCI-P3’, ‘DICOM GSDF’, ‘ITU-R BT.2020’, ‘ITU-R BT.2100 HLG’, ‘ITU-R BT.2100 PQ’, ‘ITU-R BT.601’, ‘ITU-R BT.709’, ‘ProPhoto RGB’, ‘RIMM RGB’, ‘ROMM RGB’, ‘SMPTE 240M’, ‘ST 2084’}, Opto-electronic transfer function (OETF / OECF).

### Other Parameters

- `E_clip` (numeric, optional) – {`colour.models.oetf_RIMMRGB()`}, Maximum exposure level.
- `I_max` (numeric, optional) – {`colour.models.oetf_ROMMRGB()`, `colour.models.oetf_RIMMRGB()`}, Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- `L_p` (numeric, optional) – {`colour.models.oetf_ST2084()`}, Display peak luminance  $\text{cd}/\text{m}^2$ .
- `is_12_bits_system` (bool) – {`colour.models.oetf_BT2020()`}, ITU-R BT.2020 alpha and beta constants are used if system is not 12-bit.
- `r` (numeric, optional) – {`colour.models.oetf_ARIBSTDB67()`}, Video level corresponding to reference white level.

**Returns**  $R'G'B'$  video component signal value.

**Return type** numeric or ndarray

## Examples

```
>>> oetf(0.18)
0.4613561...
>>> oetf(0.18, function='ITU-R BT.2020')
0.4090077...
>>> oetf(0.18, function='ST 2084', L_p=1000)
...
0.1820115...
```

## colour.OETFS

```
colour.OETFS = CaseInsensitiveMapping({'ITU-R BT.2020': ..., 'ITU-R BT.2100 PQ': ..., 'SMPTE 240M': ..., 'D...'}
```

Supported opto-electrical transfer functions (OETFs / OECFs).

OETFS [CaseInsensitiveMapping] {'sRGB', 'ARIB STD-B67', 'DCI-P3', 'DICOM GSDF', 'ITU-R BT.2020', 'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ', 'ITU-R BT.601', 'ITU-R BT.709', 'ProPhoto RGB', 'RIMM RGB', 'ROMM RGB', 'SMPTE 240M', 'ST 2084'}

## colour.oetf\_reverse

```
colour.oetf_reverse(value, function='sRGB', **kwargs)
```

Decodes  $R'G'B'$  video component signal value to tristimulus values at the display using given reverse opto-electronic transfer function (OETF / OECF).

### Parameters

- **value** (numeric or array\_like) – Value.
- **function** (unicode, optional) – {'sRGB', 'ARIB STD-B67', 'ITU-R BT.2100 HLD', 'ITU-R BT.2100 PQ', 'ITU-R BT.601', 'ITU-R BT.709'}, Reverse opto-electronic transfer function (OETF / OECF).

**Other Parameters r** (numeric, optional) – {`colour.models.oetf_ARIBSTDB67()`}, Video level corresponding to reference white level.

**Returns** Tristimulus values at the display.

**Return type** numeric or ndarray

## Examples

```
>>> oetf_reverse(0.461356129500442)
0.1...
>>> oetf_reverse(
...     0.409007728864150, function='ITU-R BT.601')
0.1...
```

## colour.OETFS\_REVERSE

```
colour.OETFS_REVERSE = CaseInsensitiveMapping({'ITU-R BT.2100 PQ': ..., 'ITU-R BT.2100 HLD': ..., 'sRGB': ...})
```

Supported reverse opto-electrical transfer functions (OETFs / OECFs).

OETFS\_REVERSE [CaseInsensitiveMapping] {'sRGB', 'ARIB STD-B67', 'ITU-R BT.2100 HLD', 'ITU-R BT.2100 PQ', 'ITU-R BT.601', 'ITU-R BT.709'}

colour.models

<code>oetf_ARIBSTDB67(E[, r])</code>	Defines ARIB STD-B67 (Hybrid Log-Gamma) opto-electrical transfer function (OETF / OECF).
<code>oetf_reverse_ARIBSTDB67(E_p[, r])</code>	Defines ARIB STD-B67 (Hybrid Log-Gamma) reverse opto-electrical transfer function (OETF / OECF).

Continued on next page

Table 3.144 – continued from previous page

<code>oetf_DCIP3(XYZ)</code>	Defines the <i>DCI-P3</i> colourspace opto-electronic transfer function (OETF / OECF).
<code>oetf_DICOMGSDF(L)</code>	Defines the <i>DICOM - Grayscale Standard Display Function</i> opto-electronic transfer function (OETF / OECF).
<code>oetf_BT2020(E[, is_12_bits_system])</code>	Defines <i>Recommendation ITU-R BT.2020</i> opto-electrical transfer function (OETF / OECF).
<code>oetf_BT2100_HLG(E)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference HLG</i> opto-electrical transfer function (OETF / OECF).
<code>oetf_reverse_BT2100_HLG(E)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference HLG</i> reverse opto-electrical transfer function (OETF / OECF).
<code>oetf_BT2100_PQ(E)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference PQ</i> opto-electrical transfer function (OETF / OECF).
<code>oetf_reverse_BT2100_PQ(E_p)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference PQ</i> reverse opto-electrical transfer function (OETF / OECF).
<code>oetf_BT601(L)</code>	Defines <i>Recommendation ITU-R BT.601-7</i> opto-electronic transfer function (OETF / OECF).
<code>oetf_reverse_BT601(E)</code>	Defines <i>Recommendation ITU-R BT.601-7</i> reverse opto-electronic transfer function (OETF / OECF).
<code>oetf_BT709(L)</code>	Defines <i>Recommendation ITU-R BT.709-6</i> opto-electronic transfer function (OETF / OECF).
<code>oetf_reverse_BT709(V)</code>	Defines <i>Recommendation ITU-R BT.709-6</i> reverse opto-electronic transfer function (OETF / OECF).
<code>oetf_ProPhotoRGB(X[, I_max])</code>	Defines the <i>ROMM RGB</i> encoding opto-electronic transfer function (OETF / OECF).
<code>oetf_RIMMRGB(X[, I_max, E_clip])</code>	Defines the <i>RIMM RGB</i> encoding opto-electronic transfer function (OETF / OECF).
<code>oetf_ROMMRGB(X[, I_max])</code>	Defines the <i>ROMM RGB</i> encoding opto-electronic transfer function (OETF / OECF).
<code>oetf_SMPTE240M(L_c)</code>	Defines <i>SMPTE 240M</i> opto-electrical transfer function (OETF / OECF).
<code>oetf_ST2084(C[, L_p])</code>	Defines <i>SMPTE ST 2084:2014</i> optimised perceptual opto-electronic transfer function (OETF / OECF).
<code>oetf_sRGB(L)</code>	Defines the <i>sRGB</i> colourspace opto-electronic transfer function (OETF / OECF).
<code>oetf_reverse_sRGB(V)</code>	Defines the <i>sRGB</i> colourspace reverse opto-electronic transfer function (OETF / OECF).

## colour.models.oetf\_ARIBSTDB67

`colour.models.oetf_ARIBSTDB67(E, r=0.5)`

Defines *ARIB STD-B67 (Hybrid Log-Gamma)* opto-electrical transfer function (OETF / OECF).

### Parameters

- `E` (numeric or array\_like) – Voltage normalized by the reference white level and proportional to the implicit light intensity that would be detected with a reference camera color channel R, G, B.
- `r` (numeric, optional) – Video level corresponding to reference white level.

`Returns` Resulting non-linear signal  $E'$ .

**Return type** numeric or ndarray

## References

- [AssociationoRIaBusinesses15]

## Examples

```
>>> oetf_ARIBSTDB67(0.18)
0.2121320...
```

### colour.models.oetf\_reverse\_ARIBSTDB67

colour.models.**oetf\_reverse\_ARIBSTDB67**(*E\_p*, *r*=0.5)

Defines ARIB STD-B67 (*Hybrid Log-Gamma*) reverse opto-electrical transfer function (OETF / OECF).

#### Parameters

- *E\_p* (numeric or array\_like) – Non-linear signal  $E'$ .
- *r* (numeric, optional) – Video level corresponding to reference white level.

**Returns** Voltage  $E$  normalized by the reference white level and proportional to the implicit light intensity that would be detected with a reference camera color channel R, G, B.

**Return type** numeric or ndarray

## References

- [AssociationoRIaBusinesses15]

## Examples

```
>>> oetf_reverse_ARIBSTDB67(0.212132034355964)
0.1799999...
```

### colour.models.oetf\_DCIP3

colour.models.**oetf\_DCIP3**(*XYZ*)

Defines the DCI-P3 colourspace opto-electronic transfer function (OETF / OECF).

**Parameters** *XYZ* (numeric or array\_like) – CIE XYZ tristimulus values.

**Returns** Non-linear CIE XYZ' tristimulus values.

**Return type** numeric or ndarray

## References

- [DigitalCInitiatives07]

## Examples

```
>>> oetf_DCIP3(0.18)
461.992059...
```

## colour.models.oetf\_DICOMGSDF

colour.models.oetf\_DICOMGSDF( $L$ )

Defines the *DICOM - Grayscale Standard Display Function* opto-electronic transfer function (OETF / OECF).

**Parameters**  $L$  (numeric or array\_like) – Luminance  $L$ .

**Returns** Just-Noticeable Difference (JND) Index,  $j$  in domain 1 to 1023.

**Return type** numeric or ndarray

## References

- [NationalEMAssociation04]

## Examples

```
>>> oetf_DICOMGSDF(130.065284012159790)
511.9964806...
```

## colour.models.oetf\_BT2020

colour.models.oetf\_BT2020( $E$ ,  $is\_12\_bits\_system=False$ )

Defines *Recommendation ITU-R BT.2020* opto-electrical transfer function (OETF / OECF).

**Parameters**

- $E$  (numeric or array\_like) – Voltage  $E$  normalized by the reference white level and proportional to the implicit light intensity that would be detected with a reference camera colour channel R, G, B.
- $is\_12\_bits\_system$  (bool) – *BT.709 alpha* and *beta* constants are used if system is not 12-bit.

**Returns** Resulting non-linear signal  $E'$ .

**Return type** numeric or ndarray

## References

- [InternationalTUnion15a]

## Examples

```
>>> oetf_BT2020(0.18)
0.4090077...
```

### colour.models.oetf\_BT2100\_HLG

`colour.models.oetf_BT2100_HLG(E)`

Defines *Recommendation ITU-R BT.2100 Reference HLG* opto-electrical transfer function (OETF / OECF).

The OETF maps relative scene linear light into the non-linear *HLG* signal value.

**Parameters** `E` (numeric or array\_like) –  $E$  is the signal for each colour component  $R_S, G_S, B_S$  proportional to scene linear light and scaled by camera exposure, normalized to the range [0, 1].

**Returns**  $E$  is the resulting non-linear signal  $R', G', B'$  in the range [0, 1].

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> oetf_BT2100_HLG(0.18 / 12)
0.2121320...
```

### colour.models.oetf\_reverse\_BT2100\_HLG

`colour.models.oetf_reverse_BT2100_HLG(E)`

Defines *Recommendation ITU-R BT.2100 Reference HLG* reverse opto-electrical transfer function (OETF / OECF).

**Parameters** `E_p` (numeric or array\_like) –  $E$  is the resulting non-linear signal  $R', G', B'$  in the range [0, 1].

**Returns**  $E$  is the signal for each colour component  $R_S, G_S, B_S$  proportional to scene linear light and scaled by camera exposure, normalized to the range [0, 1].

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> oetf_reverse_BT2100_HLG(0.212132034355964)
0.0149999...
```

### colour.models.oetf\_BT2100\_PQ

colour.models.**oetf\_BT2100\_PQ**(*E*)

Defines *Recommendation ITU-R BT.2100 Reference PQ* opto-electrical transfer function (OETF / OECF).

The OETF maps relative scene linear light into the non-linear PQ signal value.

**Parameters** *E* (numeric or array\_like) –  $E = R_S, G_S, B_S; Y_S$ ; or  $I_S$  is the signal determined by scene light and scaled by camera exposure. The values  $E, R_S, G_S, B_S, Y_S, I_S$  are in the range [0, 1].

**Returns** *E* is the resulting non-linear signal ( $R', G', B'$ ) in the range [0, 1].

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> oetf_BT2100_PQ(0.1)
0.7247698...
```

### colour.models.oetf\_reverse\_BT2100\_PQ

colour.models.**oetf\_reverse\_BT2100\_PQ**(*E\_p*)

Defines *Recommendation ITU-R BT.2100 Reference PQ* reverse opto-electrical transfer function (OETF / OECF).

**Parameters** *E\_p* (numeric or array\_like) – *E* is the resulting non-linear signal ( $R', G', B'$ ) in the range [0, 1].

**Returns**  $E = R_S, G_S, B_S; Y_S$ ; or  $I_S$  is the signal determined by scene light and scaled by camera exposure. The values  $E, R_S, G_S, B_S, Y_S, I_S$  are in the range [0, 1].

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> oetf_reverse_BT2100_PQ(0.724769816665726)
0.0999999...
```

### colour.models.oetf\_BT601

`colour.models.oetf_BT601(L)`

Defines *Recommendation ITU-R BT.601-7* opto-electronic transfer function (OETF / OECF).

**Parameters** `L` (numeric or array\_like) – Luminance  $L$  of the image.

**Returns** Corresponding electrical signal  $E_s$ .

**Return type** numeric or ndarray

## References

- [InternationalTUnion11b]

## Examples

```
>>> oetf_BT601(0.18)
0.4090077...
```

### colour.models.oetf\_reverse\_BT601

`colour.models.oetf_reverse_BT601(E)`

Defines *Recommendation ITU-R BT.601-7* reverse opto-electronic transfer function (OETF / OECF).

**Parameters** `E` (numeric or array\_like) – Electrical signal  $E$ .

**Returns** Corresponding luminance  $L$  of the image.

**Return type** numeric or ndarray

## References

- [InternationalTUnion11b]

## Examples

```
>>> oetf_reverse_BT601(0.409007728864150)
0.1...
```

## colour.models.oetf\_BT709

colour.models.**oetf\_BT709**(*L*)

Defines *Recommendation ITU-R BT.709-6* opto-electronic transfer function (OETF / OECF).

**Parameters** *L* (numeric or array\_like) – Luminance *L* of the image.

**Returns** Corresponding electrical signal *V*.

**Return type** numeric or ndarray

### References

- [InternationalTUnion15b]

### Examples

```
>>> oetf_BT709(0.18)
0.4090077...
```

## colour.models.oetf\_reverse\_BT709

colour.models.**oetf\_reverse\_BT709**(*V*)

Defines *Recommendation ITU-R BT.709-6* reverse opto-electronic transfer function (OETF / OECF).

**Parameters** *V* (numeric or array\_like) – Electrical signal *V*.

**Returns** Corresponding luminance *L* of the image.

**Return type** numeric or ndarray

### References

- [InternationalTUnion15b]

### Examples

```
>>> oetf_reverse_BT709(0.409007728864150)
0.1...
```

## colour.models.oetf\_ProPhotoRGB

colour.models.**oetf\_ProPhotoRGB**(*X*, *I\_max*=255)

Defines the *ROMM RGB* encoding opto-electronic transfer function (OETF / OECF).

### Parameters

- *X* (numeric or array\_like) – Linear data  $X_{ROMM}$ .
- *I\_max* (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.

**Returns** Non-linear data  $X'_{ROMM}$ .

**Return type** numeric or ndarray

## References

- [\[ANS03\]](#)
- [\[SWG00\]](#)

## Examples

```
>>> oetf_ROMMRGB(0.18)
98.3564133...
```

## colour.models.oetf\_RIMMRGB

colour.models.oetf\_RIMMRGB( $X, I_{max}=255, E_{clip}=2.0$ )

Defines the RIMM RGB encoding opto-electronic transfer function (OETF / OECF).

RIMM RGB encoding non-linearity is based on that specified by *Recommendation ITU-R BT.709-6*.

### Parameters

- **X** (numeric or array\_like) – Linear data  $X_{RIMM}$ .
- **I\_max** (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- **E\_clip** (numeric, optional) – Maximum exposure level.

**Returns** Non-linear data  $X'_{RIMM}$ .

**Return type** numeric or ndarray

## References

- [\[SWG00\]](#)

## Examples

```
>>> oetf_RIMMRGB(0.18)
74.3768017...
```

## colour.models.oetf\_ROMMRGB

colour.models.oetf\_ROMMRGB( $X, I_{max}=255$ )

Defines the ROMM RGB encoding opto-electronic transfer function (OETF / OECF).

### Parameters

- **X** (numeric or array\_like) – Linear data  $X_{ROMM}$ .

- **I\_max** (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.

**Returns** Non-linear data  $X'_{ROMM}$ .

**Return type** numeric or ndarray

## References

- [\[ANS03\]](#)
- [\[SWG00\]](#)

## Examples

```
>>> oetf_ROMMRGB(0.18)
98.3564133...
```

## colour.models.oetf\_SMPTE240M

colour.models.**oetf\_SMPTE240M**( $L_c$ )

Defines SMPTE 240M opto-electrical transfer function (OETF / OECF).

**Parameters**  $L_c$  (numeric or array\_like) – Light input  $L_c$  to the reference camera normalized to the system reference white.

**Returns** Video signal output  $V_c$  of the reference camera normalized to the system reference white.

**Return type** numeric or ndarray

## References

- [\[SocietyoMPaTEngineers99\]](#)

## Examples

```
>>> oetf_SMPTE240M(0.18)
0.4022857...
```

## colour.models.oetf\_ST2084

colour.models.**oetf\_ST2084**( $C$ ,  $L_p=10000$ )

Defines SMPTE ST 2084:2014 optimised perceptual opto-electronic transfer function (OETF / OECF).

**Parameters**

- $C$  (numeric or array\_like) – Target optical output  $C$  in  $cd/m^2$  of the ideal reference display.
- $L_p$  (numeric, optional) – Display peak luminance  $cd/m^2$ .

**Returns** Color value abbreviated as  $N$ , normalized to the range [0, 1], that is directly proportional to the encoded signal representation, and which is not directly proportional to the optical output of a display device.

**Return type** numeric or ndarray

## References

- [\[MDolbyLaboratories14\]](#)
- [\[SocietyoMPaTEngineers14\]](#)

## Examples

```
>>> otf_ST2084(10.0, 1000)
0.5080784...
```

## colour.models.oef\_sRGB

colour.models.oef\_sRGB( $L$ )

Defines the sRGB colourspace opto-electronic transfer function (OETF / OECF).

**Parameters**  $L$  (numeric or array\_like) – Luminance  $L$  of the image.

**Returns** Corresponding electrical signal  $V$ .

**Return type** numeric or ndarray

## References

- [\[InternationalECommission99\]](#)
- [\[InternationalTUnion15b\]](#)

## Examples

```
>>> oef_sRGB(0.18)
0.4613561...
```

## colour.models.oef\_reverse\_sRGB

colour.models.oef\_reverse\_sRGB( $V$ )

Defines the sRGB colourspace reverse opto-electronic transfer function (OETF / OECF).

**Parameters**  $V$  (numeric or array\_like) – Electrical signal  $V$ .

**Returns** Corresponding luminance  $L$  of the image.

**Return type** numeric or ndarray

## References

- [\[International ECommission99\]](#)
- [\[International TUnion15b\]](#)

## Examples

```
>>> otf_reverse_sRGB(0.461356129500442)
0.1...
```

## Ancillary Objects

colour

<code>function_gamma(a[, exponent, ...])</code>	Defines a typical gamma encoding / decoding function.
<code>function_linear(a)</code>	Defines a typical linear encoding / decoding function, essentially a pass-through function.

### colour.function\_gamma

`colour.function_gamma(a, exponent=1, negative_number_handling=u'Indeterminate')`

Defines a typical gamma encoding / decoding function.

#### Parameters

- `a` (numeric or array\_like) – Array to encode / decode.
- `exponent` (numeric or array\_like, optional) – Encoding / decoding exponent.
- `negative_number_handling` (unicode, optional) – {‘Indeterminate’, ‘Mirror’, ‘Preserve’, ‘Clamp’}, Defines the behaviour for a negative numbers and / or the definition return value:
  - *Indeterminate*: The behaviour will be indeterminate and definition return value might contain *nans*.
  - *Mirror*: The definition return value will be mirrored around abscissa and ordinate axis, i.e. Blackmagic Design: Davinci Resolve behaviour.
  - *Preserve*: The definition will preserve any negative number in `a`, i.e. The Foundry Nuke behaviour.
  - *Clamp*: The definition will clamp any negative number in `a` to 0.

**Returns** Encoded / decoded array.

**Return type** numeric or ndarray

**Raises** `ValueError` – If the negative number handling method is not defined.

## Examples

```
>>> function_gamma(0.18, 2.2)
0.0229932...
>>> function_gamma(-0.18, 2.0)
0.0323999...
>>> function_gamma(-0.18, 2.2)
nan
>>> function_gamma(-0.18, 2.2, 'Mirror')
-0.0229932...
>>> function_gamma(-0.18, 2.2, 'Preserve')
-0.1...
>>> function_gamma(-0.18, 2.2, 'Clamp')
0.0
```

## colour.function\_linear

`colour.function_linear(a)`

Defines a typical linear encoding / decoding function, essentially a pass-through function.

**Parameters** `a` (numeric or array\_like) – Array to encode / decode.

**Returns** Encoded / decoded array.

**Return type** numeric or ndarray

### Examples

```
>>> function_linear(0.18)
0.18
```

## Electro-Optical Transfer Functions

`colour`

<code>eotf(value[, function])</code>	Decodes $R'G'B'$ video component signal value to tristimulus values at the display using given electro-optical transfer function (EOTF / EOCF).
<code>EOTFS</code>	Supported electro-optical transfer functions (EOTFs / EOCFs).
<code>eotf_reverse(value[, function])</code>	Encodes estimated tristimulus values in a scene to $R'G'B'$ video component signal value using given reverse electro-optical transfer function (EOTF / EOCF).
<code>EOTFS_REVERSE</code>	Supported reverse electro-optical transfer functions (EOTFs / EOCFs).

## colour.eotf

`colour.eotf(value, function='ITU-R BT.1886', **kwargs)`

Decodes  $R'G'B'$  video component signal value to tristimulus values at the display using given electro-optical transfer function (EOTF / EOCF).

**Parameters**

- **value** (numeric or array\_like) – Value.
- **function** (unicode, optional) – {'ITU-R BT.1886', 'DCI-P3', 'DICOM GSDF', 'ITU-R BT.2020', 'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ', 'ProPhoto RGB', 'RIMM RGB', 'ROMM RGB', 'SMPTE 240M', 'ST 2084'}, Electro-optical transfer function (EOTF / EOCF).

### Other Parameters

- **E\_clip** (numeric, optional) – {colour.models.eotf\_RIMMRGB()}, Maximum exposure level.
- **I\_max** (numeric, optional) – {colour.models.eotf\_RIMMRGB(), colour.models.eotf\_RIMMRGB()}, Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- **L\_B** (numeric, optional) – {colour.models.eotf\_BT1886(), colour.models.eotf\_BT2100\_HLG()}, Screen luminance for black.
- **L\_W** (numeric, optional) – {colour.models.eotf\_BT1886(), colour.models.eotf\_BT2100\_HLG()}, Screen luminance for white.
- **L\_p** (numeric, optional) – {colour.models.eotf\_ST2084()}, Display peak luminance  $cd/m^2$ .
- **gamma** (numeric, optional) – {colour.models.eotf\_BT2100\_HLG()}, System gamma value, 1.2 at the nominal display peak luminance of  $1000cd/m^2$ .
- **is\_12\_bits\_system** (bool) – {colour.models.eotf\_BT2020()}, ITU-R BT.2020 alpha and beta constants are used if system is not 12-bit.

**Returns** Tristimulus values at the display.

**Return type** numeric or ndarray

### Examples

```
>>> eotf(0.461356129500442)
0.1...
>>> eotf(0.409007728864150, function='ITU-R BT.2020')
...
0.1...
>>> eotf(0.182011532850008, function='ST 2084', L_p=1000)
...
0.1...
```

## colour.EOTFS

```
colour.EOTFS = CaseInsensitiveMapping({'ITU-R BT.2020': ..., 'ST 2084': ..., 'DCI-P3': ..., 'ITU-R BT.2100 HLG': ...})
Supported electro-optical transfer functions (EOTFs / EOCFs).

EOTFS [CaseInsensitiveMapping] {'DCI-P3', 'DICOM GSDF', 'ITU-R BT.1886', 'ITU-R BT.2020',
'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ', 'ProPhoto RGB', 'RIMM RGB', 'ROMM RGB',
'SMPTE 240M', 'ST 2084'}
```

## colour.eotf\_reverse

```
colour.eotf_reverse(value, function='ITU-R BT.1886', **kwargs)
```

Encodes estimated tristimulus values in a scene to  $R'G'B'$  video component signal value using given reverse electro-optical transfer function (EOTF / EOCF).

### Parameters

- **value** (numeric or array\_like) – Value.
- **function** (unicode, optional) – {'ITU-R BT.1886', 'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ'}, Reverse electro-optical transfer function (EOTF / EOCF).

### Other Parameters

- **L\_B** (numeric, optional) – {colour.models.eotf\_reverse\_BT1886(), colour.models.eotf\_reverse\_BT2100\_HLG()}, Screen luminance for black.
- **L\_W** (numeric, optional) – {colour.models.eotf\_reverse\_BT1886(), colour.models.eotf\_reverse\_BT2100\_HLG()}, Screen luminance for white.
- **gamma** (numeric, optional) – {colour.models.eotf\_BT2100\_HLG()}, System gamma value, 1.2 at the nominal display peak luminance of  $1000\text{cd/m}^2$ .

**Returns**  $R'G'B'$  video component signal value.

**Return type** numeric or ndarray

## Examples

```
>>> eotf_reverse(0.11699185725296059)
0.4090077...
>>> eotf_reverse(
...     0.11699185725296059, function='ITU-R BT.1886')
0.4090077...
```

## colour.EOTFS\_REVERSE

```
colour.EOTFS_REVERSE = CaseInsensitiveMapping({'ITU-R BT.1886': ..., 'ITU-R BT.2100 PQ': ..., 'ITU-R BT.2100 HLG': ...})
```

Supported reverse electro-optical transfer functions (EOTFs / EOCFs).

**EOTFS\_REVERSE** [CaseInsensitiveMapping] {'ITU-R BT.1886', 'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ'}

colour.models

eotf_DCIP3(XYZ_p)	Defines the <i>DCI-P3</i> colourspace electro-optical transfer function (EOTF / EOCF).
eotf_DICOMSDF(J)	Defines the <i>DICOM - Grayscale Standard Display Function</i> electro-optical transfer function (EOTF / EOCF).
eotf_BT1886(V[, L_B, L_W])	Defines <i>Recommendation ITU-R BT.1886</i> electro-optical transfer function (EOTF / EOCF).
eotf_reverse_BT1886(L[, L_B, L_W])	Defines <i>Recommendation ITU-R BT.1886</i> reverse electro-optical transfer function (EOTF / EOCF).

Continued on next page

Table 3.147 – continued from previous page

<code>eotf_BT2020(E_p[, is_12_bits_system])</code>	Defines <i>Recommendation ITU-R BT.2020</i> electro-optical transfer function (EOTF / EOCF).
<code>eotf_BT2100_HLG(E_p[, L_B, L_W, gamma])</code>	Defines <i>Recommendation ITU-R BT.2100 Reference HLG</i> electro-optical transfer function (EOTF / EOCF).
<code>eotf_reverse_BT2100_HLG(F_D[, L_B, L_W, gamma])</code>	Defines <i>Recommendation ITU-R BT.2100 Reference HLG</i> reverse electro-optical transfer function (EOTF / EOCF).
<code>eotf_BT2100_PQ(E_p)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference PQ</i> electro-optical transfer function (EOTF / EOCF).
<code>eotf_reverse_BT2100_PQ(F_D)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference PQ</i> reverse electro-optical transfer function (EOTF / EOCF).
<code>eotf_ProPhotoRGB(X_p[, I_max])</code>	Defines the <i>ROMM RGB</i> encoding electro-optical transfer function (EOTF / EOCF).
<code>eotf_RIMMRGB(X_p[, I_max, E_clip])</code>	Defines the <i>RIMM RGB</i> encoding electro-optical transfer function (EOTF / EOCF).
<code>eotf_ROMMRGB(X_p[, I_max])</code>	Defines the <i>ROMM RGB</i> encoding electro-optical transfer function (EOTF / EOCF).
<code>eotf SMPTE240M(V_r)</code>	Defines <i>SMPTE 240M</i> electro-optical transfer function (EOTF / EOCF).
<code>eotf_ST2084(N[, L_p])</code>	Defines <i>SMPTE ST 2084:2014</i> optimised perceptual electro-optical transfer function (EOTF / EOCF).

### colour.models.eotf\_DCIP3

`colour.models.eotf_DCIP3(XYZ_p)`

Defines the *DCI-P3* colourspace electro-optical transfer function (EOTF / EOCF).

**Parameters** `XYZ_p` (numeric or array\_like) – Non-linear *CIE XYZ'* tristimulus values.

**Returns** *CIE XYZ* tristimulus values.

**Return type** numeric or ndarray

### References

- [\[DigitalCInitiatives07\]](#)

### Examples

```
>>> eotf_DCIP3(461.99220597484737)
0.18...
```

### colour.models.eotf\_DICOMGSDF

`colour.models.eotf_DICOMGSDF(J)`

Defines the *DICOM - Grayscale Standard Display Function* electro-optical transfer function (EOTF / EOCF).

**Parameters** `J` (numeric or array\_like) – Just-Noticeable Difference (JND) Index,  $j$  in range 1 to 1023.

**Returns** Corresponding *luminance L*.

**Return type** numeric or ndarray

## References

- [NationalEMAssociation04]

## Examples

```
>>> eotf_DICOMGSDF(512)
130.0652840...
```

## colour.models.eotf\_BT1886

`colour.models.eotf_BT1886(V, L_B=0, L_W=1)`

Defines *Recommendation ITU-R BT.1886* electro-optical transfer function (EOTF / EOCF).

### Parameters

- `V` (numeric or array\_like) – Input video signal level (normalized, black at  $V = 0$ , to white at  $V = 1$ . For content mastered per *Recommendation ITU-R BT.709*, 10-bit digital code values  $D$  map into values of  $V$  per the following equation:  $V = (D - 64)/876$
- `L_B` (numeric, optional) – Screen luminance for black.
- `L_W` (numeric, optional) – Screen luminance for white.

**Returns** Screen luminance in  $cd/m^2$ .

**Return type** numeric or ndarray

## References

- [InternationalTUnion11a]

## Examples

```
>>> eotf_BT1886(0.409007728864150)
0.1169918...
```

## colour.models.eotf\_reverse\_BT1886

`colour.models.eotf_reverse_BT1886(L, L_B=0, L_W=1)`

Defines *Recommendation ITU-R BT.1886* reverse electro-optical transfer function (EOTF / EOCF).

### Parameters

- `L` (numeric or array\_like) – Screen luminance in  $cd/m^2$ .
- `L_B` (numeric, optional) – Screen luminance for black.
- `L_W` (numeric, optional) – Screen luminance for white.

**Returns** Input video signal level (normalized, black at  $V = 0$ , to white at  $V = 1$ ).

**Return type** numeric or ndarray

## References

- [InternationalTUnion11a]

## Examples

```
>>> eotf_reverse_BT1886(0.11699185725296059)
0.4090077...
```

## colour.models.eotf\_BT2020

colour.models.eotf\_BT2020(`E_p`, `is_12_bits_system=False`)

Defines Recommendation ITU-R BT.2020 electro-optical transfer function (EOTF / EOCF).

### Parameters

- `E_p` (numeric or array\_like) – Non-linear signal  $E'$ .
- `is_12_bits_system` (bool) – BT.709 alpha and beta constants are used if system is not 12-bit.

**Returns** Resulting voltage  $E$ .

**Return type** numeric or ndarray

## References

- [InternationalTUnion15a]

## Examples

```
>>> eotf_BT2020(0.705515089922121)
0.4999999...
```

## colour.models.eotf\_BT2100\_HLG

colour.models.eotf\_BT2100\_HLG(`E_p`, `L_B=0`, `L_W=1000`, `gamma=None`)

Defines Recommendation ITU-R BT.2100 Reference HLG electro-optical transfer function (EOTF / EOCF).

The EOTF maps the non-linear HLG signal into display light.

### Parameters

- **E\_p** (numeric or array\_like) –  $E'$  denotes a non-linear colour value  $R', G', B'$  or  $L', M', S'$  in HLG space in range  $[0, 1]$ .
- **L\_B** (numeric, optional) –  $L_B$  is the display luminance for black in  $cd/m^2$ .
- **L\_W** (numeric, optional) –  $L_W$  is nominal peak luminance of the display in  $cd/m^2$  for achromatic pixels.
- **gamma** (numeric, optional) – System gamma value, 1.2 at the nominal display peak luminance of  $1000cd/m^2$ .

**Returns** Luminance  $F_D$  of a displayed linear component  $R_D, G_D, B_D$  or  $Y_D$  or  $I_D$ , in  $cd/m^2$ .

**Return type** numeric or ndarray

## References

- [Bor17]
- [InternationalTUnion16]

## Examples

```
>>> eotf_BT2100_HLG(0.212132034355964)
6.4760398...
```

## colour.models.eotf\_reverse\_BT2100\_HLG

colour.models.**eotf\_reverse\_BT2100\_HLG**( $F_D, L_B=0, L_W=1000, \text{gamma}=\text{None}$ )

Defines Recommendation ITU-R BT.2100 Reference HLG reverse electro-optical transfer function (EOTF / EOCF).

### Parameters

- **F\_D** (numeric or array\_like) – Luminance  $F_D$  of a displayed linear component  $R_D, G_D, B_D$  or  $Y_D$  or  $I_D$ , in  $cd/m^2$ .
- **L\_B** (numeric, optional) –  $L_B$  is the display luminance for black in  $cd/m^2$ .
- **L\_W** (numeric, optional) –  $L_W$  is nominal peak luminance of the display in  $cd/m^2$  for achromatic pixels.
- **gamma** (numeric, optional) – System gamma value, 1.2 at the nominal display peak luminance of  $1000cd/m^2$ .

**Returns**  $E'$  denotes a non-linear colour value  $R', G', B'$  or  $L', M', S'$  in HLG space in range  $[0, 1]$ .

**Return type** numeric or ndarray

## References

- [Bor17]
- [InternationalTUnion16]

## Examples

```
>>> eotf_reverse_BT2100_HLG(6.476039825649814)
0.2121320...
```

### colour.models.eotf\_BT2100\_PQ

```
colour.models.eotf_BT2100_PQ(E_p)
```

Defines *Recommendation ITU-R BT.2100 Reference PQ* electro-optical transfer function (EOTF / EOCF).

The EOTF maps the non-linear *PQ* signal into display light.

**Parameters** `E_p` (numeric or array\_like) –  $E'$  denotes a non-linear colour value  $R', G', B'$  or  $L', M', S'$  in *PQ* space [0, 1].

**Returns**  $F_D$  is the luminance of a displayed linear component  $R_D, G_D, B_D$  or  $Y_D$  or  $I_D$ , in  $cd/m^2$ .

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> eotf_BT2100_PQ(0.724769816665726)
779.9883608...
```

### colour.models.eotf\_reverse\_BT2100\_PQ

```
colour.models.eotf_reverse_BT2100_PQ(F_D)
```

Defines *Recommendation ITU-R BT.2100 Reference PQ* reverse electro-optical transfer function (EOTF / EOCF).

**Parameters** `F_D` (numeric or array\_like) –  $F_D$  is the luminance of a displayed linear component  $R_D, G_D, B_D$  or  $Y_D$  or  $I_D$ , in  $cd/m^2$ .

**Returns**  $E'$  denotes a non-linear colour value  $R', G', B'$  or  $L', M', S'$  in *PQ* space [0, 1].

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> eotf_reverse_BT2100_PQ(779.988360834085370)
0.7247698...
```

## colour.models.eotf\_ProPhotoRGB

`colour.models.eotf_ProPhotoRGB(X_p, I_max=255)`

Defines the *ROMM RGB* encoding electro-optical transfer function (EOTF / EOCF).

### Parameters

- `X_p` (numeric or array\_like) – Non-linear data  $X'_{ROMM}$ .
- `I_max` (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.

**Returns** Linear data  $X_{ROMM}$ .

**Return type** numeric or ndarray

## References

- [\[ANS03\]](#)
- [\[SWG00\]](#)

## Examples

```
>>> eotf_ROMMRGB(98.356413311540095)
0.1...
```

## colour.models.eotf\_RIMMRGB

`colour.models.eotf_RIMMRGB(X_p, I_max=255, E_clip=2.0)`

Defines the *RIMM RGB* encoding electro-optical transfer function (EOTF / EOCF).

### Parameters

- `X_p` (numeric or array\_like) – Non-linear data  $X'_{RIMM}$ .
- `I_max` (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- `E_clip` (numeric, optional) – Maximum exposure level.

**Returns** Linear data  $X_{RIMM}$ .

**Return type** numeric or ndarray

## References

- [\[SWG00\]](#)

## Examples

```
>>> eotf_RIMMRGB(74.37680178131521)
0.1...
```

## colour.models.eotf\_ROMMRGB

colour.models.**eotf\_ROMMRGB**( $X_p$ ,  $I_{max}=255$ )

Defines the *ROMM RGB* encoding electro-optical transfer function (EOTF / EOCF).

### Parameters

- $X_p$  (numeric or array\_like) – Non-linear data  $X'_{ROMM}$ .
- $I_{max}$  (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.

**Returns** Linear data  $X_{ROMM}$ .

**Return type** numeric or ndarray

## References

- [\[ANS03\]](#)
- [\[SWG00\]](#)

## Examples

```
>>> eotf_ROMMRGB(98.356413311540095)
0.1...
```

## colour.models.eotf\_SMPTE240M

colour.models.**eotf\_SMPTE240M**( $V_r$ )

Defines *SMPTE 240M* electro-optical transfer function (EOTF / EOCF).

**Parameters**  $V_r$  (numeric or array\_like) – Video signal level  $V_r$  driving the reference reproducer normalized to the system reference white.

**Returns** Light output  $L_r$  from the reference reproducer normalized to the system reference white.

**Return type** numeric or ndarray

## References

- [\[SocietyoMPaTEngineers99\]](#)

## Examples

```
>>> eotf_SMPTE240M(0.402285796753870)
0.1...
```

### colour.models.eotf\_ST2084

`colour.models.eotf_ST2084(N, L_p=10000)`

Defines *SMPTE ST 2084:2014* optimised perceptual electro-optical transfer function (EOTF / EOCF).

This perceptual quantizer (PQ) has been modeled by Dolby Laboratories using *Barten (1999)* contrast sensitivity function.

#### Parameters

- `N` (numeric or array\_like) – Color value abbreviated as  $N$ , normalized to the range [0, 1], that is directly proportional to the encoded signal representation, and which is not directly proportional to the optical output of a display device.
- `L_p` (numeric, optional) – Display peak luminance  $cd/m^2$ .

**Returns** Target optical output  $C$  in  $cd/m^2$  of the ideal reference display.

**Return type** numeric or ndarray

## References

- [\[MDolbyLaboratories14\]](#)
- [\[SocietyoMPaTEngineers14\]](#)

## Examples

```
>>> eotf_ST2084(0.508078421517399)
100.000000...
```

## Opto-Optical Transfer Functions

colour

<code>ootf(value[, function])</code>	Maps relative scene linear light to display linear light using given opto-optical transfer function (OOTF / OOCF).
<code>OOTFS</code>	Supported opto-optical transfer functions (OOTFs / OOCFs).
<code>ootf_reverse(value[, function])</code>	Maps relative display linear light to scene linear light using given reverse opto-optical transfer function (OOTF / OOCF).
<code>OOTFS_REVERSE</code>	Supported reverse opto-optical transfer functions (OOTFs / OOCFs).

## colour.ootf

```
colour.ootf(value, function='ITU-R BT.2100 PQ', **kwargs)
    Maps relative scene linear light to display linear light using given opto-optical transfer function (OOTF / OOCF).
```

### Parameters

- **value** (numeric or array\_like) – Value.
- **function** (unicode, optional) – {'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ'} Opto-optical transfer function (OOTF / OOCF).

**Returns** Luminance of a displayed linear component.

**Return type** numeric or ndarray

## Examples

```
>>> ootf(0.1)
779.9883608...
>>> ootf(0.1, function='ITU-R BT.2100 HLG')
63.0957344...
```

## colour.OOTFS

```
colour.OOTFS = CaseInsensitiveMapping({'ITU-R BT.2100 PQ': ..., 'ITU-R BT.2100 HLG': ...})
    Supported opto-optical transfer functions (OOTFs / OOCFs).
```

**OOTFS** [CaseInsensitiveMapping] {'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ'}

## colour.ootf\_reverse

```
colour.ootf_reverse(value, function='ITU-R BT.2100 PQ', **kwargs)
    Maps relative display linear light to scene linear light using given reverse opto-optical transfer function (OOTF / OOCF).
```

### Parameters

- **value** (numeric or array\_like) – Value.
- **function** (unicode, optional) – {'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ'} Reverse opto-optical transfer function (OOTF / OOCF).

### Other Parameters

- **L\_B** (numeric, optional) – {colour.models.ootf\_reverse\_BT2100\_HLG()},  $L_B$  is the display luminance for black in  $cd/m^2$ .
- **L\_W** (numeric, optional) – {colour.models.ootf\_reverse\_BT2100\_HLG()},  $L_W$  is nominal peak luminance of the display in  $cd/m^2$  for achromatic pixels.
- **gamma** (numeric, optional) – {colour.models.ootf\_reverse\_BT2100\_HLG()}, System gamma value, 1.2 at the nominal display peak luminance of  $1000cd/m^2$ .

**Returns** Luminance of scene linear light.

**Return type** numeric or ndarray

## Examples

```
>>> ootf_reverse(779.988360834115840)
0.100000...
>>> ootf_reverse(
...     63.095734448019336, function='ITU-R BT.2100 HLG')
0.100000...
```

## colour.OOTFS\_REVERSE

`colour.OOTFS_REVERSE = CaseInsensitiveMapping({'ITU-R BT.2100 PQ': ..., 'ITU-R BT.2100 HLG': ...})`  
Supported reverse opto-optical transfer functions (OOTFs / OOCFs).

**OOTFS\_REVERSE** [CaseInsensitiveMapping] {'ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ'}

`colour.models`

<code>ootf_BT2100_HLG(E[, L_B, L_W, gamma])</code>	Defines <i>Recommendation ITU-R BT.2100 Reference HLG</i> opto-optical transfer function (OOTF / OOCF).
<code>ootf_reverse_BT2100_HLG(F_D[, L_B, L_W, gamma])</code>	Defines <i>Recommendation ITU-R BT.2100 Reference HLG</i> reverse opto-optical transfer function (OOTF / OOCF).
<code>ootf_BT2100_PQ(E)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference PQ</i> opto-optical transfer function (OOTF / OOCF).
<code>ootf_reverse_BT2100_PQ(F_D)</code>	Defines <i>Recommendation ITU-R BT.2100 Reference PQ</i> reverse opto-optical transfer function (OOTF / OOCF).

## colour.models.ootf\_BT2100\_HLG

`colour.models.ootf_BT2100_HLG(E, L_B=0, L_W=1000, gamma=None)`

Defines *Recommendation ITU-R BT.2100 Reference HLG* opto-optical transfer function (OOTF / OOCF).

The OOTF maps relative scene linear light to display linear light.

### Parameters

- `E` (numeric or array\_like) –  $E$  is the signal for each colour component  $R_S, G_S, B_S$  proportional to scene linear light and scaled by camera exposure, normalized to the range  $[0, 1]$ .
- `L_B` (numeric, optional) –  $L_B$  is the display luminance for black in  $cd/m^2$ .
- `L_W` (numeric, optional) –  $L_W$  is nominal peak luminance of the display in  $cd/m^2$  for achromatic pixels.
- `gamma` (numeric, optional) – System gamma value, 1.2 at the nominal display peak luminance of  $1000cd/m^2$ .

**Returns**  $F_D$  is the luminance of a displayed linear component  $R_D, G_D, or B_D$ , in  $cd/m^2$ .

**Return type** numeric or ndarray

## References

- [Bor17]

- [InternationalTUnion16]

## Examples

```
>>> ootf_BT2100_HLG(0.1)
63.0957344...
```

## colour.models.ootf\_reverse\_BT2100\_HLG

colour.models.ootf\_reverse\_BT2100\_HLG( $F_D, L_B=0, L_W=1000, \text{gamma}=\text{None}$ )

Defines *Recommendation ITU-R BT.2100 Reference HLG* reverse opto-optical transfer function (OOTF / OOCF).

### Parameters

- $F_D$  (numeric or array\_like) –  $F_D$  is the luminance of a displayed linear component  $R_D, G_D, \text{or } B_D$ , in  $\text{cd}/\text{m}^2$ .
- $L_B$  (numeric, optional) –  $L_B$  is the display luminance for black in  $\text{cd}/\text{m}^2$ .
- $L_W$  (numeric, optional) –  $L_W$  is nominal peak luminance of the display in  $\text{cd}/\text{m}^2$  for achromatic pixels.
- $\text{gamma}$  (numeric, optional) – System gamma value, 1.2 at the nominal display peak luminance of  $1000\text{cd}/\text{m}^2$ .

**Returns**  $E$  is the signal for each colour component  $R_S, G_S, B_S$  proportional to scene linear light and scaled by camera exposure, normalized to the range [0, 1].

**Return type** numeric or ndarray

## References

- [Bor17]
- [InternationalTUnion16]

## Examples

```
>>> ootf_reverse_BT2100_HLG(63.095734448019336)
0.1000000...
```

## colour.models.ootf\_BT2100\_PQ

colour.models.ootf\_BT2100\_PQ( $E$ )

Defines *Recommendation ITU-R BT.2100 Reference PQ* opto-optical transfer function (OOTF / OOCF).

The OOTF maps relative scene linear light to display linear light.

**Parameters**  $E$  (numeric or array\_like) –  $E = R_S, G_S, B_S; Y_S; \text{or } I_S$  is the signal determined by scene light and scaled by camera exposure. The values  $E, R_S, G_S, B_S, Y_S, I_S$  are in the range [0, 1].

**Returns**  $F_D$  is the luminance of a displayed linear component ( $R_D, G_D, B_D; Y_D$ ; or  $I_D$ ).

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> ootf_BT2100_PQ(0.1)
779.9883608...
```

## colour.models.ootf\_reverse\_BT2100\_PQ

colour.models.ootf\_reverse\_BT2100\_PQ( $F_D$ )

Defines *Recommendation ITU-R BT.2100 Reference PQ reverse opto-optical transfer function (OOTF / OOCF)*.

**Parameters**  $F_D$  (numeric or array\_like) –  $F_D$  is the luminance of a displayed linear component ( $R_D, G_D, B_D; Y_D$ ; or  $I_D$ ).

**Returns**  $E = R_S, G_S, B_S; Y_S; or I_S$  is the signal determined by scene light and scaled by camera exposure. The values  $E, R_S, G_S, B_S, Y_S, I_S$  are in the range [0, 1].

**Return type** numeric or ndarray

## References

- [\[Bor17\]](#)
- [\[InternationalTUnion16\]](#)

## Examples

```
>>> ootf_reverse_BT2100_PQ(779.988360834115840)
0.1000000...
```

## Log Encoding and Decoding Curves

colour

---

log\_encoding\_curve(value[, curve])

Encodes linear-light values to  $R'G'B'$  video component signal value using given *log* curve.

---

LOG\_ENCODING\_CURVES

Supported *log* encoding curves.

Continued on next page

Table 3.150 – continued from previous page

<code>log_decoding_curve(value[, curve])</code>	Decodes $R'G'B'$ video component signal value to linear-light values using given log curve.
<code>LOG_DECODING_CURVES</code>	Supported log decoding curves.

`colour.log_encoding_curve``colour.log_encoding_curve(value, curve='Cineon', **kwargs)`Encodes linear-light values to  $R'G'B'$  video component signal value using given log curve.**Parameters**

- `value` (numeric or array\_like) – Value.
- `curve` (unicode, optional) – {‘ACEScc’, ‘ACEScct’, ‘ACESproxy’, ‘ALEXA Log C’, ‘Canon Log 2’, ‘Canon Log 3’, ‘Canon Log’, ‘Cineon’, ‘ERIMM RGB’, ‘Log3G10’, ‘Log3G12’, ‘Panalog’, ‘PLog’, ‘Protune’, ‘REDLog’, ‘REDLogFilm’, ‘S-Log’, ‘S-Log2’, ‘S-Log3’, ‘V-Log’, ‘ViperLog’}, Computation curve.

**Other Parameters**

- `EI` (int, optional) – {`colour.models.log_encoding_ALEXALogC()`}, Ei.
- `E_clip` (numeric, optional) – {`colour.models.log_encoding_ERIMMRGB()`}, Maximum exposure limit.
- `E_min` (numeric, optional) – {`colour.models.log_encoding_ERIMMRGB()`}, Minimum exposure limit.
- `I_max` (numeric, optional) – {`colour.models.log_encoding_ERIMMRGB()`}, Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- `bit_depth` (unicode, optional) – {`colour.models.log_encoding_ACESproxy()`, `colour.models.log_encoding_SLog()`, `colour.models.log_encoding_SLog2()`}, {8, 10, 12}, Bit depth used for conversion, ACESproxy uses {10, 12}.
- `black_offset` (numeric or array\_like) – {`colour.models.log_encoding_Cineon()`, `colour.models.log_encoding_Panalog()`, `colour.models.log_encoding_REDLog()`, `colour.models.log_encoding_REDLogFilm()`}, Black offset.
- `density_per_code_value` (numeric or array\_like) – {`colour.models.log_encoding_PivotedLog()`}, Density per code value.
- `firmware` (unicode, optional) – {`colour.models.log_encoding_ALEXALogC()`}, {‘SUP 3.x’, ‘SUP 2.x’}, Alexa firmware version.
- `in_reflection` (bool, optional) – {`colour.models.log_encoding_SLog()`, `colour.models.log_encoding_SLog2()`}, Whether the light level  $x$  to a camera is reflection.
- `linear_reference` (numeric or array\_like) – {`colour.models.log_encoding_PivotedLog()`}, Linear reference.
- `log_reference` (numeric or array\_like) – {`colour.models.log_encoding_PivotedLog()`}, Log reference.
- `out_legal` (bool, optional) – {`colour.models.log_encoding_SLog()`, `colour.models.log_encoding_SLog2()`, `colour.models.log_encoding_SLog3()`}, Whether the non-linear Sony S-Log, Sony S-Log2 or Sony S-Log3 data  $y$  is encoded in legal range.

- **negative\_gamma** (*numeric or array\_like*) – {`colour.models.log_encoding_PivotedLog()`}, Negative gamma.
- **method** (*unicode, optional*) – {`colour.models.log_encoding_ALEXALogC()`}, {'Linear Scene Exposure Factor', 'Normalised Sensor Signal'}, Conversion method.

**Returns** Log value.

**Return type** numeric or ndarray

## Examples

```
>>> log_encoding_curve(0.18)
0.4573196...
>>> log_encoding_curve(0.18, curve='ACEScc')
0.4135884...
>>> log_encoding_curve(0.18, curve='PLog', log_reference=400)
...
0.3910068...
>>> log_encoding_curve(0.18, curve='S-Log')
0.3849708...
```

## colour.LOG\_ENCODING\_CURVES

`colour.LOG_ENCODING_CURVES = CaseInsensitiveMapping({ 'ACEScc': ... , 'ACEScct': ... , 'Log3G12': ... , 'Log3G10': ... })`  
Supported log encoding curves.

`LOG_ENCODING_CURVES [CaseInsensitiveMapping] {'ACEScc', 'ACEScct', 'ACESproxy', 'ALEXA Log C', 'Canon Log 2', 'Canon Log 3', 'Canon Log', 'Cineon', 'ERIMM RGB', 'Log3G10', 'Log3G12', 'Panalog', 'PLog', 'Protune', 'REDLog', 'REDLogFilm', 'S-Log', 'S-Log2', 'S-Log3', 'V-Log', 'ViperLog'}`

## colour.log\_decoding\_curve

`colour.log_decoding_curve(value, curve='Cineon', **kwargs)`

Decodes  $R'G'B'$  video component signal value to linear-light values using given log curve.

### Parameters

- **value** (*numeric or array\_like*) – Value.
- **curve** (*unicode, optional*) – {'ACEScc', 'ACEScct', 'ACESproxy', 'ALEXA Log C', 'Canon Log 2', 'Canon Log 3', 'Canon Log', 'Cineon', 'ERIMM RGB', 'Log3G10', 'Log3G12', 'Panalog', 'PLog', 'Protune', 'REDLog', 'REDLogFilm', 'S-Log', 'S-Log2', 'S-Log3', 'V-Log', 'ViperLog'}, Computation curve.

### Other Parameters

- **Ei** (*int, optional*) – {`colour.models.log_decoding_ALEXALogC()`}, Ei.
- **E\_clip** (*numeric, optional*) – {`colour.models.log_decoding_ERIMMRGB()`}, Maximum exposure limit.
- **E\_min** (*numeric, optional*) – {`colour.models.log_decoding_ERIMMRGB()`}, Minimum exposure limit.

- **I\_max** (*numeric, optional*) – {`colour.models.log_decoding_ERIMMRGB()`}, Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- **bit\_depth** (*int, optional*) – {`colour.models.log_decoding_ACESproxy()`, `colour.models.log_decoding_SLog()`, `colour.models.log_decoding_SLog2()`}, {8, 10, 12}, Bit depth used for conversion, *ACESproxy* uses {10, 12}.
- **black\_offset** (*numeric or array\_like*) – {`colour.models.log_decoding_Cineon()`, `colour.models.log_decoding_Panalog()`, `colour.models.log_decoding_REDLog()`, `colour.models.log_decoding_REDLogFilm()`}, Black offset.
- **density\_per\_code\_value** (*numeric or array\_like*) – {`colour.models.log_decoding_PivotedLog()`}, Density per code value.
- **firmware** (*unicode, optional*) – {`colour.models.log_decoding_ALEXALogC()`}, {'SUP 3.x', 'SUP 2.x'}, Alexa firmware version.
- **in\_legal** (*bool, optional*) – {`colour.models.log_decoding_SLog()`, `colour.models.log_decoding_SLog2()`, `colour.models.log_decoding_SLog3()`}, Whether the non-linear *Sony S-Log*, *Sony S-Log2* or *Sony S-Log3* data  $y$  is encoded in legal range.
- **linear\_reference** (*numeric or array\_like*) – {`colour.models.log_decoding_PivotedLog()`}, Linear reference.
- **log\_reference** (*numeric or array\_like*) – {`colour.models.log_decoding_PivotedLog()`}, Log reference.
- **negative\_gamma** (*numeric or array\_like*) – {`colour.models.log_decoding_PivotedLog()`}, Negative gamma.
- **out\_reflection** (*bool, optional*) – {`colour.models.log_decoding_SLog()`, `colour.models.log_decoding_SLog2()`}, Whether the light level  $x$  to a camera is reflection.
- **method** (*unicode, optional*) – {`colour.models.log_decoding_ALEXALogC()`}, {'Linear Scene Exposure Factor', 'Normalised Sensor Signal'}, Conversion method.

**Returns** Log value.

**Return type** numeric or ndarray

## Examples

```
>>> log_decoding_curve(0.457319613085418)
0.1...
>>> log_decoding_curve(0.413588402492442, curve='ACEScc')
...
0.1...
>>> log_decoding_curve(0.391006842619746, curve='PLog', log_reference=400)
...
0.1...
>>> log_decoding_curve(0.376512722254600, curve='S-Log')
...
0.1...
```

## colour.LOG\_DECODING\_CURVES

```
colour.LOG_DECODING_CURVES = CaseInsensitiveMapping({'ACEScc': ..., 'ACEScct': ..., 'Log3G12': ..., 'Log3G10': ...})
    Supported log decoding curves.
```

```
LOG_DECODING_CURVES [CaseInsensitiveMapping] {'ACEScc', 'ACEScct', 'ACESproxy', 'ALEXA Log C', 'Canon Log 2', 'Canon Log 3', 'Canon Log', 'Cineon', 'ERIMM RGB', 'Log3G10', 'Log3G12', 'Panalog', 'PLog', 'Protune', 'REDLog', 'REDLogFilm', 'S-Log', 'S-Log2', 'S-Log3', 'V-Log', 'ViperLog'}
```

colour.models

<code>log_encoding_ACEScc(lin_AP1)</code>	Defines the ACEScc colourspace log encoding / opto-electronic transfer function.
<code>log_decoding_ACEScc(ACEScc)</code>	Defines the ACEScc colourspace log decoding / electro-optical transfer function.
<code>log_encoding_ACEScct(lin_AP1)</code>	Defines the ACEScct colourspace log encoding / opto-electronic transfer function.
<code>log_decoding_ACEScct(ACEScct)</code>	Defines the ACEScct colourspace log decoding / electro-optical transfer function.
<code>log_encoding_ACESproxy(lin_AP1[, bit_depth])</code>	Defines the ACESproxy colourspace log encoding curve / opto-electronic transfer function.
<code>log_decoding_ACESproxy(ACESproxy[, bit_depth])</code>	Defines the ACESproxy colourspace log decoding curve / electro-optical transfer function.
<code>log_encoding_ALEXALogC(x[, firmware, method, EI])</code>	Defines the ALEXA Log C log encoding curve / opto-electronic transfer function.
<code>log_decoding_ALEXALogC(t[, firmware, method, EI])</code>	Defines the ALEXA Log C log decoding curve / electro-optical transfer function.
<code>log_encoding_CanonLog2(x[, bit_depth, ...])</code>	Defines the Canon Log 2 log encoding curve / opto-electronic transfer function.
<code>log_decoding_CanonLog2(clog2[, bit_depth, ...])</code>	Defines the Canon Log 2 log decoding curve / electro-optical transfer function.
<code>log_encoding_CanonLog3(x[, bit_depth, ...])</code>	Defines the Canon Log 3 log encoding curve / opto-electronic transfer function.
<code>log_decoding_CanonLog3(clog3[, bit_depth, ...])</code>	Defines the Canon Log 3 log decoding curve / electro-optical transfer function.
<code>log_encoding_CanonLog(x[, bit_depth, ...])</code>	Defines the Canon Log log encoding curve / opto-electronic transfer function.
<code>log_decoding_CanonLog(clog[, bit_depth, ...])</code>	Defines the Canon Log log decoding curve / electro-optical transfer function.
<code>log_encoding_Cineon(x[, black_offset])</code>	Defines the Cineon log encoding curve / opto-electronic transfer function.
<code>log_decoding_Cineon(y[, black_offset])</code>	Defines the Cineon log decoding curve / electro-optical transfer function.
<code>log_encoding_ERIMMRGB(X[, I_max, E_min, E_clip])</code>	Defines the ERIMM RGB log encoding curve / opto-electronic transfer function (OETF / OECF).
<code>log_decoding_ERIMMRGB(X_p[, I_max, E_min, ...])</code>	Defines the ERIMM RGB log decoding curve / electro-optical transfer function (EOTF / EOCF).
<code>log_encoding_Log3G10(x[, legacy_curve])</code>	Defines the Log3G10 log encoding curve / opto-electronic transfer function.
<code>log_decoding_Log3G10(y[, legacy_curve])</code>	Defines the Log3G10 log decoding curve / electro-optical transfer function.

Continued on next page

Table 3.151 – continued from previous page

<code>log_encoding_Log3G12(x)</code>	Defines the <i>Log3G12</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_Log3G12(y)</code>	Defines the <i>Log3G12</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_Panalog(x[, black_offset])</code>	Defines the <i>Panalog</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_Panalog(y[, black_offset])</code>	Defines the <i>Panalog</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_PivotedLog(x[, log_reference, ...])</code>	Defines the <i>Josh Pines</i> style <i>Pivoted Log</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_PivotedLog(y[, log_reference, ...])</code>	Defines the <i>Josh Pines</i> style <i>Pivoted Log</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_Proture(x)</code>	Defines the <i>Proture</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_Proture(y)</code>	Defines the <i>Proture</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_REDLog(x[, black_offset])</code>	Defines the <i>REDLog</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_REDLog(y[, black_offset])</code>	Defines the <i>REDLog</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_REDLogFilm(x[, black_offset])</code>	Defines the <i>REDLogFilm</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_REDLogFilm(y[, black_offset])</code>	Defines the <i>REDLogFilm</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_SLog(x[, bit_depth, out_legal, ...])</code>	Defines the <i>Sony S-Log</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_SLog(y[, bit_depth, in_legal, ...])</code>	Defines the <i>Sony S-Log</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_SLog2(x[, bit_depth, ...])</code>	Defines the <i>Sony S-Log2</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_SLog2(y[, bit_depth, in_legal, ...])</code>	Defines the <i>Sony S-Log2</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_SLog3(x[, bit_depth, ...])</code>	Defines the <i>Sony S-Log3</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_SLog3(y[, bit_depth, in_legal, ...])</code>	Defines the <i>Sony S-Log3</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_VLog(L_in[, bit_depth, ...])</code>	Defines the <i>Panasonic V-Log</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_VLog(V_out[, bit_depth, ...])</code>	Defines the <i>Panasonic V-Log</i> log decoding curve / electro-optical transfer function.
<code>log_encoding_ViperLog(x)</code>	Defines the <i>Viper Log</i> log encoding curve / opto-electronic transfer function.
<code>log_decoding_ViperLog(y)</code>	Defines the <i>Viper Log</i> log decoding curve / electro-optical transfer function.

`colour.models.log_encoding_ACEScc``colour.models.log_encoding_ACEScc(lin_AP1)`

Defines the ACEScc colourspace log encoding / opto-electronic transfer function.

**Parameters** `lin_AP1` (numeric or array\_like) – `lin_AP1` value.

**Returns** ACEScc non-linear value.

**Return type** numeric or ndarray

## References

- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14c]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14d]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14b]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee]

## Examples

```
>>> log_encoding_ACEScc(0.18)
0.4135884...
```

## colour.models.log\_decoding\_ACEScc

colour.models.log\_decoding\_ACEScc(ACEScc)

Defines the ACEScc colourspace log decoding / electro-optical transfer function.

**Parameters** ACEScc (numeric or array\_like) – ACEScc non-linear value.

**Returns** lin\_AP1 value.

**Return type** numeric or ndarray

## References

- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14c]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14d]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14b]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee]

## Examples

```
>>> log_decoding_ACEScc(0.413588402492442)
0.1799999...
```

## colour.models.log\_encoding\_ACEScct

colour.models.log\_encoding\_ACEScct(lin\_AP1)

Defines the ACEScct colourspace log encoding / opto-electronic transfer function.

**Parameters** lin\_AP1 (numeric or array\_like) – lin\_AP1 value.

**Returns** ACEScct non-linear value.

**Return type** numeric or ndarray

## References

- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14c]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14d]
- [The AoMPAa Sciences Science aTC Council Academy CESACES Project 16]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee]

## Examples

```
>>> log_encoding_ACEScct(0.18)
0.4135884...
```

## colour.models.log\_decoding\_ACEScct

colour.models.**log\_decoding\_ACEScct**(ACEScct)

Defines the ACEScct colourspace log decoding / electro-optical transfer function.

**Parameters** ACEScct (numeric or array\_like) – ACEScct non-linear value.

**Returns** lin\_AP1 value.

**Return type** numeric or ndarray

## References

- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14c]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14d]
- [The AoMPAa Sciences Science aTC Council Academy CESACES Project 16]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee]

## Examples

```
>>> log_decoding_ACEScct(0.413588402492442)
0.1799999...
```

## colour.models.log\_encoding\_ACESproxy

colour.models.**log\_encoding\_ACESproxy**(lin\_AP1, bit\_depth=10)

Defines the ACESproxy colourspace log encoding curve / opto-electronic transfer function.

**Parameters**

- lin\_AP1 (numeric or array\_like) – lin\_AP1 value.
- bit\_depth (int, optional) – {10, 12}, ACESproxy bit depth.

**Returns** ACESproxy non-linear value.

**Return type** numeric or ndarray

## References

- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14c]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14d]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14a]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee]

## Examples

```
>>> log_encoding_ACESproxy(0.18)
426
```

## colour.models.log\_decoding\_ACESproxy

colour.models.log\_decoding\_ACESproxy(ACESproxy, bit\_depth=10)

Defines the ACESproxy colourspace log decoding curve / electro-optical transfer function.

### Parameters

- **ACESproxy** (numeric or array\_like) – ACESproxy non-linear value.
- **bit\_depth** (int, optional) – {10, 12}, ACESproxy bit depth.

**Returns** lin\_AP1 value.

**Return type** numeric or ndarray

## References

- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14c]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14d]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee 14a]
- [The AoMPAa Sciences Science aTC Council Academy CESACESP Subcommittee]

## Examples

```
>>> log_decoding_ACESproxy(426)
0.1792444...
```

## colour.models.log\_encoding\_ALEXALogC

colour.models.log\_encoding\_ALEXALogC(*x*, *firmware*=*‘SUP 3.x’*, *method*=*‘Linear Scene Exposure Factor’*, *EI*=800)

Defines the ALEXA Log C log encoding curve / opto-electronic transfer function.

### Parameters

- *x* (numeric or array\_like) – Linear data *x*.
- **firmware** (unicode, optional) – {‘SUP 3.x’, ‘SUP 2.x’}, Alexa firmware version.
- **method** (unicode, optional) – {‘Linear Scene Exposure Factor’, ‘Normalised Sensor Signal’}, Conversion method.
- **EI** (int, optional) – *Ei*.

**Returns** ALEXA Log C encoded data *t*.

**Return type** numeric or ndarray

### References

- [ARR12]

### Examples

```
>>> log_encoding_ALEXALogC(0.18)
0.3910068...
```

## colour.models.log\_decoding\_ALEXALogC

colour.models.log\_decoding\_ALEXALogC(*t*, *firmware*=*‘SUP 3.x’*, *method*=*‘Linear Scene Exposure Factor’*, *EI*=800)

Defines the ALEXA Log C log decoding curve / electro-optical transfer function.

### Parameters

- *t* (numeric or array\_like) – ALEXA Log C encoded data *t*.
- **firmware** (unicode, optional) – {‘SUP 3.x’, ‘SUP 2.x’}, Alexa firmware version.
- **method** (unicode, optional) – {‘Linear Scene Exposure Factor’, ‘Normalised Sensor Signal’}, Conversion method.
- **EI** (int, optional) – *Ei*.

**Returns** Linear data *x*.

**Return type** numeric or ndarray

### References

- [ARR12]

## Examples

```
>>> log_decoding_ALEXALogC(0.391006832034084)
0.18...
```

## colour.models.log\_encoding\_CanonLog2

`colour.models.log_encoding_CanonLog2(x, bit_depth=10, out_legal=True, in_reflection=True)`  
Defines the *Canon Log 2* log encoding curve / opto-electronic transfer function.

### Parameters

- `x` (numeric or array\_like) – Linear data  $x$ .
- `bit_depth` (int, optional) – Bit depth used for conversion.
- `out_legal` (bool, optional) – Whether the *Canon Log 2* non-linear data is encoded in legal range.
- `in_reflection` (bool, optional) – Whether the light level  $x$  to a camera is reflection.

**Returns** *Canon Log 2* non-linear data.

**Return type** numeric or ndarray

## References

- [Can]

## Examples

```
>>> log_encoding_CanonLog2(0.18) * 100
39.8254694...
```

## colour.models.log\_decoding\_CanonLog2

`colour.models.log_decoding_CanonLog2(clog2, bit_depth=10, in_legal=True, out_reflection=True)`  
Defines the *Canon Log 2* log decoding curve / electro-optical transfer function.

### Parameters

- `clog2` (numeric or array\_like) – *Canon Log 2* non-linear data.
- `bit_depth` (int, optional) – Bit depth used for conversion.
- `in_legal` (bool, optional) – Whether the *Canon Log 2* non-linear data is encoded in legal range.
- `out_reflection` (bool, optional) – Whether the light level  $x$  to a camera is reflection.

**Returns** Linear data  $x$ .

**Return type** numeric or ndarray

## References

- [Can]

## Examples

```
>>> log_decoding_CanonLog2(39.825469498316735 / 100)
0.1799999...
```

## colour.models.log\_encoding\_CanonLog3

colour.models.log\_encoding\_CanonLog3(*x*, *bit\_depth*=10, *out\_legal*=True, *in\_reflection*=True)

Defines the *Canon Log 3* log encoding curve / opto-electronic transfer function.

### Parameters

- *x* (numeric or array\_like) – Linear data *x*.
- **bit\_depth** (int, optional) – Bit depth used for conversion.
- **out\_legal** (bool, optional) – Whether the *Canon Log 3* non-linear data is encoded in legal range.
- **in\_reflection** (bool, optional) – Whether the light level *x* to a camera is reflection.

**Returns** *Canon Log 3* non-linear data.

**Return type** numeric or ndarray

## Notes

- Introspection of the grafting points by Shaw, N. (2018) shows that the *Canon Log 3* IDT was likely derived from its encoding curve as the later is grafted at +/-0.014:

```
>>> clog3 = 0.04076162
>>> (clog3 - 0.073059361) / 2.3069815
-0.01400000000000002
>>> clog3 = 0.105357102
>>> (clog3 - 0.073059361) / 2.3069815
0.01399999999999997
```

## References

- [Can]

## Examples

```
>>> log_encoding_CanonLog3(0.18) * 100
34.3389369...
```

## colour.models.log\_decoding\_CanonLog3

```
colour.models.log_decoding_CanonLog3(clog3, bit_depth=10, in_legal=True, out_reflection=True)
```

Defines the *Canon Log 3* log decoding curve / electro-optical transfer function.

### Parameters

- `clog3` (numeric or array\_like) – *Canon Log 3* non-linear data.
- `bit_depth` (int, optional) – Bit depth used for conversion.
- `in_legal` (bool, optional) – Whether the *Canon Log 3* non-linear data is encoded in legal range.
- `out_reflection` (bool, optional) – Whether the light level  $x$  to a camera is reflection.

**Returns** Linear data  $x$ .

**Return type** numeric or ndarray

### References

- [Can]

### Examples

```
>>> log_decoding_CanonLog3(34.338936938868677 / 100)
0.1800000...
```

## colour.models.log\_encoding\_CanonLog

```
colour.models.log_encoding_CanonLog(x, bit_depth=10, out_legal=True, in_reflection=True)
```

Defines the *Canon Log* log encoding curve / opto-electronic transfer function.

### Parameters

- `x` (numeric or array\_like) – Linear data  $x$ .
- `bit_depth` (int, optional) – Bit depth used for conversion.
- `out_legal` (bool, optional) – Whether the *Canon Log* non-linear data is encoded in legal range.
- `in_reflection` (bool, optional) – Whether the light level  $x$  to a camera is reflection.

**Returns** *Canon Log* non-linear data.

**Return type** numeric or ndarray

### References

- [Tho12]

## Examples

```
>>> log_encoding_CanonLog(0.18) * 100
34.3389651...
```

### colour.models.log\_decoding\_CanonLog

colour.models.log\_decoding\_CanonLog(*clog*, *bit\_depth*=10, *in\_legal*=True, *out\_reflection*=True)  
Defines the *Canon Log* log decoding curve / electro-optical transfer function.

#### Parameters

- ***clog*** (numeric or array\_like) – *Canon Log* non-linear data.
- ***bit\_depth*** (int, optional) – Bit depth used for conversion.
- ***in\_legal*** (bool, optional) – Whether the *Canon Log* non-linear data is encoded in legal range.
- ***out\_reflection*** (bool, optional) – Whether the light level *x* to a camera is reflection.

**Returns** Linear data *x*.

**Return type** numeric or ndarray

## References

- [Tho12]

## Examples

```
>>> log_decoding_CanonLog(34.338965172606912 / 100)
0.17999999...
```

### colour.models.log\_encoding\_Cineon

colour.models.log\_encoding\_Cineon(*x*, *black\_offset*=0.0107977516232771)  
Defines the *Cineon* log encoding curve / opto-electronic transfer function.

#### Parameters

- ***x*** (numeric or array\_like) – Linear data *x*.
- ***black\_offset*** (numeric or array\_like) – Black offset.

**Returns** Non-linear data *y*.

**Return type** numeric or ndarray

## References

- [SonyImageworks12]

## Examples

```
>>> log_encoding_Cineon(0.18)
0.4573196...
```

### colour.models.log\_decoding\_Cineon

`colour.models.log_decoding_Cineon(y, black_offset=0.0107977516232771)`  
Defines the *Cineon* log decoding curve / electro-optical transfer function.

#### Parameters

- `y` (numeric or array\_like) – Non-linear data  $y$ .
- `black_offset` (numeric or array\_like) – Black offset.

**Returns** Linear data  $x$ .

**Return type** numeric or ndarray

## References

- [SonyImageworks12]

## Examples

```
>>> log_decoding_Cineon(0.457319613085418)
0.1799999...
```

### colour.models.log\_encoding\_ERIMMRGB

`colour.models.log_encoding_ERIMMRGB(X, I_max=255, E_min=0.001, E_clip=316.2)`  
Defines the *ERIMM RGB* log encoding curve / opto-electronic transfer function (OETF / OECF).

#### Parameters

- `X` (numeric or array\_like) – Linear data  $X_{ERIMM}$ .
- `I_max` (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- `E_min` (numeric, optional) – Minimum exposure limit.
- `E_clip` (numeric, optional) – Maximum exposure limit.

**Returns** Non-linear data  $X'_{ERIMM}$ .

**Return type** numeric or ndarray

## References

- [SWG00]

## Examples

```
>>> log_encoding_ERIMMRGB(0.18)
104.5633593...
```

### colour.models.log\_decoding\_ERIMMRGB

colour.models.log\_decoding\_ERIMMRGB(*X\_p*, *I\_max*=255, *E\_min*=0.001, *E\_clip*=316.2)

Defines the *ERIMM RGB* log decoding curve / electro-optical transfer function (EOTF / EOCF).

#### Parameters

- ***X\_p*** (numeric or array\_like) – Non-linear data  $X'_{ERIMM}$ .
- ***I\_max*** (numeric, optional) – Maximum code value: 255, 4095 and 650535 for respectively 8-bit, 12-bit and 16-bit per channel.
- ***E\_min*** (numeric, optional) – Minimum exposure limit.
- ***E\_clip*** (numeric, optional) – Maximum exposure limit.

**Returns** Linear data  $X_{ERIMM}$ .

**Return type** numeric or ndarray

## References

- [SWG00]

## Examples

```
>>> log_decoding_ERIMMRGB(104.56335932049294)
0.1...
```

### colour.models.log\_encoding\_Log3G10

colour.models.log\_encoding\_Log3G10(*x*, *legacy\_curve*=False)

Defines the *Log3G10* log encoding curve / opto-electronic transfer function.

#### Parameters

- ***x*** (numeric or array\_like) – Linear data *x*.
- ***legacy\_curve*** (bool, optional) – Whether to use the v1 *Log3G10* log encoding curve. Default is *False*.

**Returns** Non-linear data *y*.

**Return type** numeric or ndarray

## Notes

- The v1 *Log3G10* log encoding curve is the one used in *REDCINE-X beta 42*. *Resolve 12.5.2* also uses the v1 curve. *RED* is planning to use v2 *Log3G10* log encoding curve in the release version of the *RED SDK*. Use the *legacy\_curve=True* argument to switch to the v1 curve for compatibility with the current (as of September 21, 2016) *RED SDK*.
- The intent of the v1 *Log3G10* log encoding curve is that zero maps to zero, 0.18 maps to 1/3, and 10 stops above 0.18 maps to 1.0. The name indicates this in a similar way to the naming conventions of *Sony HyperGamma* curves.

The constants used in the functions do not in fact quite hit these values, but rather than use corrected constants, the functions here use the official *RED* values, in order to match the output of the *RED SDK*.

For those interested, solving for constants which exactly hit 1/3 and 1.0 yields the following values:

```
B = 25 * (np.sqrt(4093.0) - 3) / 9
A = 1 / np.log10(B * 184.32 + 1)
```

where the function takes the form:

```
Log3G10(x) = A * np.log10(B * x + 1)
```

Similarly for *Log3G12*, the values which hit exactly 1/3 and 1.0 are:

```
B = 25 * (np.sqrt(16381.0) - 3) / 9
A = 1 / np.log10(B * 737.28 + 1)
```

## References

- [Nat16]

## Examples

```
>>> log_encoding_Log3G10(0.18, legacy_curve=True)
0.3333336...
>>> log_encoding_Log3G10(0.0)
0.0915514...
```

### colour.models.log\_decoding\_Log3G10

`colour.models.log_decoding_Log3G10(y, legacy_curve=False)`

Defines the *Log3G10* log decoding curve / electro-optical transfer function.

#### Parameters

- `y` (numeric or array\_like) – Non-linear data *y*.
- `legacy_curve` (bool, optional) – Whether to use the v1 *Log3G10* log encoding curve. Default is *False*.

`Returns` Linear data *x*.

**Return type** numeric or ndarray

## References

- [Nat16]

## Examples

```
>>> log_decoding_Log3G10(1.0 / 3, legacy_curve=True)
0.1799994...
>>> log_decoding_Log3G10(1.0)
184.3223476...
```

## colour.models.log\_encoding\_Log3G12

colour.models.log\_encoding\_Log3G12(*x*)

Defines the *Log3G12* log encoding curve / opto-electronic transfer function.

**Parameters** *x* (numeric or array\_like) – Linear data *x*.

**Returns** Non-linear data *y*.

**Return type** numeric or ndarray

## References

- [Nat16]

## Examples

```
>>> log_encoding_Log3G12(0.18)
0.3333326...
```

## colour.models.log\_decoding\_Log3G12

colour.models.log\_decoding\_Log3G12(*y*)

Defines the *Log3G12* log decoding curve / electro-optical transfer function.

**Parameters** *y* (numeric or array\_like) – Non-linear data *y*.

**Returns** Linear data *x*.

**Return type** numeric or ndarray

## References

- [Nat16]

## Examples

```
>>> log_decoding_Log3G12(1.0 / 3)
0.1800015...
```

### colour.models.log\_encoding\_Panalog

`colour.models.log_encoding_Panalog(x, black_offset=0.04077184461038074)`  
Defines the *Panalog* log encoding curve / opto-electronic transfer function.

#### Parameters

- `x` (numeric or array\_like) – Linear data  $x$ .
- `black_offset` (numeric or array\_like) – Black offset.

**Returns** Non-linear data  $y$ .

**Return type** numeric or ndarray

**Warning:** These are estimations known to be close enough, the actual log encoding curves are not published.

## References

- [\[SonyImageworks12\]](#)

## Examples

```
>>> log_encoding_Panalog(0.18)
0.3745767...
```

### colour.models.log\_decoding\_Panalog

`colour.models.log_decoding_Panalog(y, black_offset=0.04077184461038074)`  
Defines the *Panalog* log decoding curve / electro-optical transfer function.

#### Parameters

- `y` (numeric or array\_like) – Non-linear data  $y$ .
- `black_offset` (numeric or array\_like) – Black offset.

**Returns** Linear data  $x$ .

**Return type** numeric or ndarray

**Warning:** These are estimations known to be close enough, the actual log encoding curves are not published.

## References

- [\[SonyImageworks12\]](#)

## Examples

```
>>> log_decoding_Panalog(0.374576791382298)
0.1...
```

### colour.models.log\_encoding\_PivotedLog

```
colour.models.log_encoding_PivotedLog(x,    log_reference=445,    linear_reference=0.18,    nega-
                                         tive_gamma=0.6, density_per_code_value=0.002)
```

Defines the *Josh Pines* style *Pivoted Log* log encoding curve / opto-electronic transfer function.

#### Parameters

- **x** (numeric or array\_like) – Linear data *x*.
- **log\_reference** (numeric or array\_like) – Log reference.
- **linear\_reference** (numeric or array\_like) – Linear reference.
- **negative\_gamma** (numeric or array\_like) – Negative gamma.
- **density\_per\_code\_value** (numeric or array\_like) – Density per code value.

**Returns** Non-linear data *y*.

**Return type** numeric or ndarray

## References

- [\[SonyImageworks12\]](#)

## Examples

```
>>> log_encoding_PivotedLog(0.18)
0.4349951...
```

### colour.models.log\_decoding\_PivotedLog

```
colour.models.log_decoding_PivotedLog(y,    log_reference=445,    linear_reference=0.18,    nega-
                                         tive_gamma=0.6, density_per_code_value=0.002)
```

Defines the *Josh Pines* style *Pivoted Log* log decoding curve / electro-optical transfer function.

#### Parameters

- **y** (numeric or array\_like) – Non-linear data *y*.
- **log\_reference** (numeric or array\_like) – Log reference.
- **linear\_reference** (numeric or array\_like) – Linear reference.

- **negative\_gamma** (numeric or array\_like) – Negative gamma.
- **density\_per\_code\_value** (numeric or array\_like) – Density per code value.

**Returns** Linear data  $x$ .

**Return type** numeric or ndarray

## References

- [SonyImageworks12]

## Examples

```
>>> log_decoding_PivotedLog(0.434995112414467)
0.1...
```

## colour.models.log\_encoding\_Proture

colour.models.log\_encoding\_Proture( $x$ )

Defines the *Proture* log encoding curve / opto-electronic transfer function.

**Parameters**  $x$  (numeric or array\_like) – Linear data  $x$ .

**Returns** Non-linear data  $y$ .

**Return type** numeric or ndarray

## References

- [GDM16]

## Examples

```
>>> log_encoding_Proture(0.18)
0.6456234...
```

## colour.models.log\_decoding\_Proture

colour.models.log\_decoding\_Proture( $y$ )

Defines the *Proture* log decoding curve / electro-optical transfer function.

**Parameters**  $y$  (numeric or array\_like) – Non-linear data  $y$ .

**Returns** Linear data  $x$ .

**Return type** numeric or ndarray

## References

- [GDM16]

## Examples

```
>>> log_decoding_Proture(0.645623486803636)
0.1...
```

## colour.models.log\_encoding\_REDLog

colour.models.log\_encoding\_REDLog(*x*, *black\_offset*=0.009955040995908344)

Defines the *REDLog* log encoding curve / opto-electronic transfer function.

### Parameters

- *x* (numeric or array\_like) – Linear data *x*.
- **black\_offset** (numeric or array\_like) – Black offset.

**Returns** Non-linear data *y*.

**Return type** numeric or ndarray

## References

- [SonyImageworks12]

## Examples

```
>>> log_encoding_REDLog(0.18)
0.6376218...
```

## colour.models.log\_decoding\_REDLog

colour.models.log\_decoding\_REDLog(*y*, *black\_offset*=0.009955040995908344)

Defines the *REDLog* log decoding curve / electro-optical transfer function.

### Parameters

- *y* (numeric or array\_like) – Non-linear data *y*.
- **black\_offset** (numeric or array\_like) – Black offset.

**Returns** Linear data *x*.

**Return type** numeric or ndarray

## References

- [SonyImageworks12]

## Examples

```
>>> log_decoding_REDLog(0.637621845988175)
0.1...
```

## colour.models.log\_encoding\_REDLogFilm

colour.models.log\_encoding\_REDLogFilm(*x*, *black\_offset*=0.0107977516232771)  
Defines the *REDLogFilm* log encoding curve / opto-electronic transfer function.

### Parameters

- *x* (numeric or array\_like) – Linear data *x*.
- *black\_offset* (numeric or array\_like) – Black offset.

**Returns** Non-linear data *y*.

**Return type** numeric or ndarray

## References

- [SonyImageworks12]

## Examples

```
>>> log_encoding_REDLogFilm(0.18)
0.4573196...
```

## colour.models.log\_decoding\_REDLogFilm

colour.models.log\_decoding\_REDLogFilm(*y*, *black\_offset*=0.0107977516232771)  
Defines the *REDLogFilm* log decoding curve / electro-optical transfer function.

### Parameters

- *y* (numeric or array\_like) – Non-linear data *y*.
- *black\_offset* (numeric or array\_like) – Black offset.

**Returns** Linear data *x*.

**Return type** numeric or ndarray

## References

- [SonyImageworks12]

## Examples

```
>>> log_decoding_REDLogFilm(0.457319613085418)
0.1799999...
```

## colour.models.log\_encoding\_SLog

colour.models.log\_encoding\_SLog(*x*, *bit\_depth*=10, *out\_legal*=True, *in\_reflection*=True)

Defines the *Sony S-Log* log encoding curve / opto-electronic transfer function.

### Parameters

- *x* (numeric or array\_like) – Reflection or *IRE*/100 input light level *x* to a camera.
- **bit\_depth** (int, optional) – Bit depth used for conversion.
- **out\_legal** (bool, optional) – Whether the non-linear *Sony S-Log* data *y* is encoded in legal range.
- **in\_reflection** (bool, optional) – Whether the light level *x* to a camera is reflection.

**Returns** Non-linear *Sony S-Log* data *y*.

**Return type** numeric or ndarray

## References

- [SonyCorporation12]

## Examples

```
>>> log_encoding_SLog(0.18)
0.3849708...
>>> log_encoding_SLog(0.18, out_legal=False)
0.3765127...
>>> log_encoding_SLog(0.18, in_reflection=False)
0.3708204...
```

## colour.models.log\_decoding\_SLog

colour.models.log\_decoding\_SLog(*y*, *bit\_depth*=10, *in\_legal*=True, *out\_reflection*=True)

Defines the *Sony S-Log* log decoding curve / electro-optical transfer function.

### Parameters

- *y* (numeric or array\_like) – Non-linear *Sony S-Log* data *y*.
- **bit\_depth** (int, optional) – Bit depth used for conversion.
- **in\_legal** (bool, optional) – Whether the non-linear *Sony S-Log* data *y* is encoded in legal range.
- **out\_reflection** (bool, optional) – Whether the light level *x* to a camera is reflection.

**Returns** Reflection or  $IRE/100$  input light level  $x$  to a camera.

**Return type** numeric or ndarray

## References

- [SonyCorporation12]

## Examples

```
>>> log_decoding_SLog(0.384970815928670)
0.1...
>>> log_decoding_SLog(0.376512722254600, in_legal=False)
...
0.1...
>>> log_decoding_SLog(0.370820482371268, out_reflection=False)
...
0.1...
```

## colour.models.log\_encoding\_SLog2

colour.models.**log\_encoding\_SLog2**( $x$ ,  $bit\_depth=10$ ,  $out\_legal=True$ ,  $in\_reflection=True$ )

Defines the Sony S-Log2 log encoding curve / opto-electronic transfer function.

### Parameters

- **x** (numeric or array\_like) – Reflection or  $IRE/100$  input light level  $x$  to a camera.
- **bit\_depth** (int, optional) – Bit depth used for conversion.
- **out\_legal** (bool, optional) – Whether the non-linear Sony S-Log2 data  $y$  is encoded in legal range.
- **in\_reflection** (bool, optional) – Whether the light level  $x$  to a camera is reflection.

**Returns** Non-linear Sony S-Log2 data  $y$ .

**Return type** numeric or ndarray

## References

- [SonyCorporation12]

## Examples

```
>>> log_encoding_SLog2(0.18)
0.3395325...
>>> log_encoding_SLog2(0.18, out_legal=False)
0.3234495...
>>> log_encoding_SLog2(0.18, in_reflection=False)
0.3262865...
```

## colour.models.log\_decoding\_SLog2

colour.models.log\_decoding\_SLog2(*y*, *bit\_depth*=10, *in\_legal*=True, *out\_reflection*=True)

Defines the Sony S-Log2 log decoding curve / electro-optical transfer function.

### Parameters

- *y* (numeric or array\_like) – Non-linear Sony S-Log2 data *y*.
- **bit\_depth** (int, optional) – Bit depth used for conversion.
- **in\_legal** (bool, optional) – Whether the non-linear Sony S-Log2 data *y* is encoded in legal range.
- **out\_reflection** (bool, optional) – Whether the light level *x* to a camera is reflection.

**Returns** Reflection or IRE/100 input light level *x* to a camera.

**Return type** numeric or ndarray

### References

- [SonyCorporation12]

### Examples

```
>>> log_decoding_SLog2(0.339532524633774)
0.1...
>>> log_decoding_SLog2(0.323449512215013, in_legal=False)
...
0.1...
>>> log_decoding_SLog2(0.326286538946799, out_reflection=False)
...
0.1...
```

## colour.models.log\_encoding\_SLog3

colour.models.log\_encoding\_SLog3(*x*, *bit\_depth*=10, *out\_legal*=True, *in\_reflection*=True)

Defines the Sony S-Log3 log encoding curve / opto-electronic transfer function.

### Parameters

- *x* (numeric or array\_like) – Reflection or IRE/100 input light level *x* to a camera.
- **bit\_depth** (int, optional) – Bit depth used for conversion.
- **out\_legal** (bool, optional) – Whether the non-linear Sony S-Log3 data *y* is encoded in legal range.
- **in\_reflection** (bool, optional) – Whether the light level *x* to a camera is reflection.

**Returns** Non-linear Sony S-Log3 data *y*.

**Return type** numeric or ndarray

## References

- [SonyCorporationc]

## Examples

```
>>> log_encoding_SLog3(0.18)
0.4105571...
>>> log_encoding_SLog3(0.18, out_legal=False)
0.4063926...
>>> log_encoding_SLog3(0.18, in_reflection=False)
0.3995079...
```

## colour.models.log\_decoding\_SLog3

`colour.models.log_decoding_SLog3(y, bit_depth=10, in_legal=True, out_reflection=True)`

Defines the Sony *S-Log3* log decoding curve / electro-optical transfer function.

### Parameters

- `y` (numeric or array\_like) – Non-linear Sony *S-Log3* data  $y$ .
- `bit_depth` (int, optional) – Bit depth used for conversion.
- `in_legal` (bool, optional) – Whether the non-linear Sony *S-Log3* data  $y$  is encoded in legal range.
- `out_reflection` (bool, optional) – Whether the light level  $x$  to a camera is reflection.

**Returns** Reflection or  $IRE/100$  input light level  $x$  to a camera.

**Return type** numeric or ndarray

## References

- [SonyCorporationc]

## Examples

```
>>> log_decoding_SLog3(0.410557184750733)
0.1...
>>> log_decoding_SLog3(0.406392694063927, in_legal=False)
...
0.1...
>>> log_decoding_SLog3(0.399507939606216, out_reflection=False)
...
0.1...
```

## colour.models.log\_encoding\_VLog

`colour.models.log_encoding_VLog(L_in, bit_depth=10, out_legal=True, in_reflection=True)`

Defines the *Panasonic V-Log* log encoding curve / opto-electronic transfer function.

### Parameters

- `L_in` (numeric or array\_like) – Linear reflection data :math`L\_{in}`.
- `bit_depth` (int, optional) – Bit depth used for conversion.
- `out_legal` (bool, optional) – Whether the non-linear *Panasonic V-Log* data  $V_{out}$  is encoded in legal range.
- `in_reflection` (bool, optional) – Whether the light level :math`L\_{in}` to a camera is reflection.

**Returns** Non-linear data  $V_{out}$ .

**Return type** numeric or ndarray

### References

- [Pan14]

### Examples

```
>>> log_encoding_VLog(0.18)
0.4233114...
```

## colour.models.log\_decoding\_VLog

`colour.models.log_decoding_VLog(V_out, bit_depth=10, in_legal=True, out_reflection=True)`

Defines the *Panasonic V-Log* log decoding curve / electro-optical transfer function.

### Parameters

- `V_out` (numeric or array\_like) – Non-linear data  $V_{out}$ .
- `bit_depth` (int, optional) – Bit depth used for conversion.
- `in_legal` (bool, optional) – Whether the non-linear *Panasonic V-Log* data  $V_{out}$  is encoded in legal range.
- `out_reflection` (bool, optional) – Whether the light level :math`L\_{in}` to a camera is reflection.

**Returns** Linear reflection data :math`L\_{in}`.

**Return type** numeric or ndarray

### References

- [Pan14]

## Examples

```
>>> log_decoding_VLog(0.423311448760136)
0.1799999...
```

### colour.models.log\_encoding\_ViperLog

```
colour.models.log_encoding_ViperLog(x)
```

Defines the *Viper Log* log encoding curve / opto-electronic transfer function.

**Parameters** `x` (numeric or array\_like) – Linear data  $x$ .

**Returns** Non-linear data  $y$ .

**Return type** numeric or ndarray

## References

- [\[SonyImageworks12\]](#)

## Examples

```
>>> log_encoding_ViperLog(0.18)
0.6360080...
```

### colour.models.log\_decoding\_ViperLog

```
colour.models.log_decoding_ViperLog(y)
```

Defines the *Viper Log* log decoding curve / electro-optical transfer function.

**Parameters** `y` (numeric or array\_like) – Non-linear data  $y$ .

**Returns** Linear data  $x$ .

**Return type** numeric or ndarray

## References

- [\[SonyImageworks12\]](#)

## Examples

```
>>> log_decoding_ViperLog(0.636008067010413)
0.1799999...
```

## Colour Encodings

### Y'CbCr Colour Encoding

colour

RGB_to_YCbCr(RGB[, K, in_bits, in_legal, ...])	Converts an array of $R'G'B'$ values to the corresponding $Y'CbCr$ colour encoding values array.
YCbCr_to_RGB(YCbCr[, K, in_bits, in_legal, ...])	Converts an array of $Y'CbCr$ colour encoding values to the corresponding $R'G'B'$ values array.
YCBCR_WEIGHTS	Implements a case-insensitive mutable mapping / dict object.
RGB_to_YcCbcCrc(RGB[, out_bits, out_legal, ...])	Converts an array of $RGB$ linear values to the corresponding $Yc'Cbc'Crc$ colour encoding values array.
YcCbcCrc_to_RGB(YcCbcCrc[, in_bits, ...])	Converts an array of $Yc'Cbc'Crc$ colour encoding values to the corresponding $RGB$ array of linear values.

#### colour.RGB\_to\_YCbCr

```
colour.RGB_to_YCbCr(RGB, K=array([ 0.2126, 0.0722]), in_bits=10, in_legal=False, in_int=False,
                     out_bits=8, out_legal=True, out_int=False, **kwargs)
```

Converts an array of  $R'G'B'$  values to the corresponding  $Y'CbCr$  colour encoding values array.

#### Parameters

- **RGB** (`array_like`) – Input  $R'G'B'$  array of floats or integer values.
- **K** (`array_like`, optional) – Luma weighting coefficients of red and blue. See `colour.YCBCR_WEIGHTS` for presets. Default is  $(0.2126, 0.0722)$ , the weightings for *ITU-R BT.709*.
- **in\_bits** (`int`, optional) – Bit depth for integer input, or used in the calculation of the denominator for legal range float values, i.e. 8-bit means the float value for legal white is  $235 / 255$ . Default is *10*.
- **in\_legal** (`bool`, optional) – Whether to treat the input values as legal range. Default is *False*.
- **in\_int** (`bool`, optional) – Whether to treat the input values as `in_bits` integer code values. Default is *False*.
- **out\_bits** (`int`, optional) – Bit depth for integer output, or used in the calculation of the denominator for legal range float values, i.e. 8-bit means the float value for legal white is  $235 / 255$ . Ignored if `out_legal` and `out_int` are both *False*. Default is *8*.
- **out\_legal** (`bool`, optional) – Whether to return legal range values. Default is *True*.
- **out\_int** (`bool`, optional) – Whether to return values as `out_bits` integer code values. Default is *False*.

#### Other Parameters

- **in\_range** (`array_like, optional`) – Array overriding the computed range such as `in_range = (RGB_min, RGB_max)`. If `in_range` is undefined, `RGB_min` and `RGB_max` will be computed using `colour.CV_range()` definition.

- **out\_range** (*array\_like, optional*) – Array overriding the computed range such as  $out\_range = (Y_{min}, Y_{max}, C_{min}, C_{max})$ . If “*out\_range*“ is undefined, \* $Y_{min}$ ,  $Y_{max}$ ,  $C_{min}$  and  $C_{max}$  will be computed using `colour.models.rgb.ycbcr.YCbCr_ranges()` definition.

**Returns**  $Y'CbCr$  colour encoding array of integer or float values.

**Return type** ndarray

**Warning:** For Recommendation ITU-R BT.2020, `colour.RGB_to_YCbCr()` definition is only applicable to the non-constant luminance implementation. `colour.RGB_to_YcCbcCrc()` definition should be used for the constant luminance case as per [\[InternationalTUnion15a\]](#).

## Notes

- The default arguments, `**{'in_bits': 10, 'in_legal': False, 'in_int': False, 'out_bits': 8, 'out_legal': True, 'out_int': False}` transform a float  $R'G'B'$  input array in range [0, 1] (*in\_bits* is ignored) to a float  $Y'CbCr$  output array where  $Y'$  is in range [16 / 255, 235 / 255] and  $Cb$  and  $Cr$  are in range [16 / 255, 240./255]. The float values are calculated based on an [0, 255] integer range, but no 8-bit quantisation or clamping are performed.

## References

- [\[InternationalTUnion11c\]](#)
- [\[InternationalTUnion15b\]](#)
- [\[SocietyoMPaTEngineers99\]](#)
- [\[Wik |\]](#)

## Examples

```
>>> RGB = np.array([1.0, 1.0, 1.0])
>>> RGB_to_YCbCr(RGB)
array([ 0.9215686...,  0.5019607...,  0.5019607...])
```

Matching float output of The Foundry Nuke’s Colorspace node set to YCbCr:

```
>>> RGB_to_YCbCr(RGB,
...                      out_range=(16 / 255, 235 / 255, 15.5 / 255, 239.5 / 255))
...
array([ 0.9215686...,  0.5          ,  0.5          ])
```

Matching float output of The Foundry Nuke’s Colorspace node set to YPbPr:

```
>>> RGB_to_YCbCr(RGB, out_legal=False, out_int=False)
...
array([ 1.,  0.,  0.])
```

Creating integer code values as per standard 10-bit SDI:

```
>>> RGB_to_YCbCr(RGB, out_legal=True, out_bits=10, out_int=True)
array([940, 512, 512])
```

For JFIF JPEG conversion as per ITU-T T.871 [*InternationalTUnion11c*]:

```
>>> RGB = np.array([102, 0, 51])
>>> RGB_to_YCbCr(RGB, K=YCBCR_WEIGHTS['ITU-R BT.601'], in_range=(0, 255),
...                 out_range=(0, 255, 0, 256), out_int=True)
array([-36, 136, 175])
```

Note the use of 256 for the max  $Cb / Cr$  value, which is required so that the  $Cb$  and  $Cr$  output is centered about 128. Using 255 centres it about 127.5, meaning that there is no integer code value to represent achromatic colours. This does however create the possibility of output integer codes with value of 256, which cannot be stored in 8-bit integer representation. Recommendation ITU-T T.871 specifies these should be clamped to 255.

These JFIF JPEG ranges are also obtained as follows:

```
>>> RGB_to_YCbCr(RGB, K=YCBCR_WEIGHTS['ITU-R BT.601'], in_bits=8,
...                   in_int=True, out_legal=False, out_int=True)
array([-36, 136, 175])
```

## colour.YCbCr\_to\_RGB

colour.YCbCr\_to\_RGB(YCbCr, K=array([ 0.2126, 0.0722]), in\_bits=8, in\_legal=True, in\_int=False,  
out\_bits=10, out\_legal=False, out\_int=False, \*\*kwargs)

Converts an array of  $Y'CbCr$  colour encoding values to the corresponding  $R'G'B'$  values array.

### Parameters

- **YCbCr** (`array_like`) – Input  $Y'CbCr$  colour encoding array of integer or float values.
- **K** (`array_like`, optional) – Luma weighting coefficients of red and blue. See `colour.YCBCR_WEIGHTS` for presets. Default is  $(0.2126, 0.0722)$ , the weightings for *ITU-R BT.709*.
- **in\_bits** (`int`, optional) – Bit depth for integer input, or used in the calculation of the denominator for legal range float values, i.e. 8-bit means the float value for legal white is  $235 / 255$ . Default is `10`.
- **in\_legal** (`bool`, optional) – Whether to treat the input values as legal range. Default is `False`.
- **in\_int** (`bool`, optional) – Whether to treat the input values as `in_bits` integer code values. Default is `False`.
- **out\_bits** (`int`, optional) – Bit depth for integer output, or used in the calculation of the denominator for legal range float values, i.e. 8-bit means the float value for legal white is  $235 / 255$ . Ignored if `out_legal` and `out_int` are both `False`. Default is `8`.
- **out\_legal** (`bool`, optional) – Whether to return legal range values. Default is `True`.
- **out\_int** (`bool`, optional) – Whether to return values as `out_bits` integer code values. Default is `False`.

### Other Parameters

- **in\_range** (*array\_like, optional*) – Array overriding the computed range such as  $in\_range = (Y\_min, Y\_max, C\_min, C\_max)$ . If *in\_range* is undefined,  $Y\_min$ ,  $Y\_max$ ,  $C\_min$  and  $C\_max$  will be computed using `colour.models.rgb.ycbcr.YCbCr_ranges()` definition.
- **out\_range** (*array\_like, optional*) – Array overriding the computed range such as  $out\_range = (RGB\_min, RGB\_max)$ . If *out\_range* is undefined,  $RGB\_min$  and  $RGB\_max$  will be computed using `colour.CV_range()` definition.

Returns  $R'G'B'$  array of integer or float values.

Return type ndarray

**Warning:** For *Recommendation ITU-R BT.2020*, `colour.YCbCr_to_RGB()` definition is only applicable to the non-constant luminance implementation. `colour.YcCbcCrc_to_RGB()` definition should be used for the constant luminance case as per [\[InternationalTUnion15a\]](#).

## References

- [\[InternationalTUnion11c\]](#)
- [\[InternationalTUnion15b\]](#)
- [\[SocietyoMPaTEngineers99\]](#)
- [\[Wik | \]](#)

## Examples

```
>>> YCbCr = np.array([502, 512, 512])
>>> YCbCr_to_RGB(YCbCr, in_bits=10, in_legal=True, in_int=True)
array([ 0.5,  0.5,  0.5])
```

## colour.YCBCR\_WEIGHTS

`colour.YCBCR_WEIGHTS` = `CaseInsensitiveMapping({u'ITU-R BT.2020': ... , u'ITU-R BT.601': ... , u'ITU-R BT.709': ...})`  
Implements a case-insensitive mutable mapping / `dict` object.

Allows values retrieving from keys while ignoring the key case. The keys are expected to be unicode or string-like objects supporting the `str.lower()` method.

Parameters `data` (`dict`) – `dict` of data to store into the mapping at initialisation.

Other Parameters `**kwargs` (`dict, optional`) – Key / Value pairs to store into the mapping at initialisation.

```
colour.__setitem__()
colour.__getitem__()
colour.__delitem__()
colour.__contains__()
colour.__iter__()
```

```
colour.__len__()
colour.__eq__()
colour.__ne__()
colour.__repr__()
colour.copy()
colour.lower_items()
```

**Warning:** The keys are expected to be unicode or string-like objects.

## References

- [\[Rei\]](#)

## Examples

```
>>> methods = CaseInsensitiveMapping({'McCamy': 1, 'Hernandez': 2})
>>> methods['mccamy']
1
```

## colour.RGB\_to\_YcCbcCrc

colour.RGB\_to\_YcCbcCrc(*RGB*, *out\_bits*=10, *out\_legal*=True, *out\_int*=False, *is\_12\_bits\_system*=False,  
                          \*\**kwargs*)

Converts an array of *RGB* linear values to the corresponding *Yc'Cbc'Crc* colour encoding values array.

### Parameters

- **RGB** (*array\_like*) – Input *RGB* array of linear float values.
- **out\_bits** (*int*, optional) – Bit depth for integer output, or used in the calculation of the denominator for legal range float values, i.e. 8-bit means the float value for legal white is 235 / 255. Ignored if *out\_legal* and *out\_int* are both *False*. Default is 10.
- **out\_legal** (*bool*, optional) – Whether to return legal range values. Default is *True*.
- **out\_int** (*bool*, optional) – Whether to return values as *out\_bits* integer code values. Default is *False*.
- **is\_12\_bits\_system** (*bool*, optional) – *Recommendation ITU-R BT.2020* OETF (OECF) adopts different parameters for 10 and 12 bit systems. Default is *False*.

**Other Parameters** **out\_range** (*array\_like*, optional) – Array overriding the computed range such as *out\_range* = (*Y\_min*, *Y\_max*, *C\_min*, *C\_max*). If *out\_range* is undefined, *Y\_min*, *Y\_max*, *C\_min* and *C\_max* will be computed using `colour.models.rgb.ycbcr.YCbCr_ranges()` definition.

**Returns** *Yc'Cbc'Crc* colour encoding array of integer or float values.

**Return type** ndarray

**Warning:** This definition is specifically for usage with *Recommendation ITU-R BT.2020* when adopting the constant luminance implementation.

## References

- [InternationalTUnion15a]
- [Wik | ]

## Examples

```
>>> RGB = np.array([0.18, 0.18, 0.18])
>>> RGB_to_YcCbcCrc(RGB, out_legal=True, out_bits=10, out_int=True,
...                      is_12_bits_system=False)
array([422, 512, 512])
```

### colour.YcCbcCrc\_to\_RGB

colour.YcCbcCrc\_to\_RGB(YcCbcCrc, in\_bits=10, in\_legal=True, in\_int=False, is\_12\_bits\_system=False,  
\*\*kwargs)

Converts an array of Yc’Cbc’Crc’ colour encoding values to the corresponding RGB array of linear values.

#### Parameters

- **YcCbcCrc** (array\_like) – Input Yc’Cbc’Crc’ colour encoding array of linear float values.
- **in\_bits** (int, optional) – Bit depth for integer input, or used in the calculation of the denominator for legal range float values, i.e. 8-bit means the float value for legal white is  $235 / 255$ . Default is 10.
- **in\_legal** (bool, optional) – Whether to treat the input values as legal range. Default is False.
- **in\_int** (bool, optional) – Whether to treat the input values as in\_bits integer code values. Default is False.
- **is\_12\_bits\_system** (bool, optional) – *Recommendation ITU-R BT.2020* EOTF (EOCF) adopts different parameters for 10 and 12 bit systems. Default is False.

**Other Parameters in\_range (array\_like, optional)** – Array overriding the computed range such as  $in\_range = (Y_{min}, Y_{max}, C_{min}, C_{max})$ . If  $in\_range$  is undefined,  $Y_{min}$ ,  $Y_{max}$ ,  $C_{min}$  and  $C_{max}$  will be computed using colour.models.rgb.ycbcr.YCbCr\_ranges() definition.

**Returns** RGB array of linear float values.

**Return type** ndarray

**Warning:** This definition is specifically for usage with *Recommendation ITU-R BT.2020* when adopting the constant luminance implementation.

## References

- [\[InternationalTUnion15a\]](#)
- [\[Wik | \]](#)

## Examples

```
>>> YcCbcCrc = np.array([1689, 2048, 2048])
>>> YcCbcCrc_to_RGB(YcCbcCrc, in_legal=True, in_bits=12, in_int=True,
...                     is_12_bits_system=True)
...
array([ 0.1800903...,  0.1800903...,  0.1800903...])
```

## Ancillary Objects

colour

<code>full_to_legal(CV[, bit_depth, in_int, out_int])</code>	Converts given code value $CV$ or float equivalent of a code value at a given bit depth from full range (full swing) to legal range (studio swing).
<code>legal_to_full(CV[, bit_depth, in_int, out_int])</code>	Converts given code value $CV$ or float equivalent of a code value at a given bit depth from legal range (studio swing) to full range (full swing).
<code>CV_range([bit_depth, is_legal, is_int])</code>	Returns the code value $CV$ range for given bit depth, range legality and representation.

### colour.full\_to\_legal

`colour.full_to_legal(CV, bit_depth=10, in_int=False, out_int=False)`

Converts given code value  $CV$  or float equivalent of a code value at a given bit depth from full range (full swing) to legal range (studio swing).

#### Parameters

- `CV` (`array_like`) – Full range code value  $CV$  or float equivalent of a code value at a given bit depth.
- `bit_depth` (`int`, optional) – Bit depth used for conversion.
- `in_int` (`bool`, optional) – Whether to treat the input value as integer code value or float equivalent of a code value at a given bit depth.
- `out_int` (`bool`, optional) – Whether to return value as integer code value or float equivalent of a code value at a given bit depth.

**Returns** Legal range code value  $CV$  or float equivalent of a code value at a given bit depth.

**Return type** `ndarray`

## Examples

```
>>> full_to_legal(0.0)
0.0625610...
```

```
>>> full_to_legal(1.0)
0.9188660...
>>> full_to_legal(0.0, out_int=True)
64
>>> full_to_legal(1.0, out_int=True)
940
>>> full_to_legal(0, in_int=True)
0.0625610...
>>> full_to_legal(1023, in_int=True)
0.9188660...
>>> full_to_legal(0, in_int=True, out_int=True)
64
>>> full_to_legal(1023, in_int=True, out_int=True)
940
```

## colour.legal\_to\_full

`colour.legal_to_full(CV, bit_depth=10, in_int=False, out_int=False)`

Converts given code value  $CV$  or float equivalent of a code value at a given bit depth from legal range (studio swing) to full range (full swing).

### Parameters

- `CV` (`array_like`) – Legal range code value  $CV$  or float equivalent of a code value at a given bit depth.
- `bit_depth` (`int`, optional) – Bit depth used for conversion.
- `in_int` (`bool`, optional) – Whether to treat the input value as integer code value or float equivalent of a code value at a given bit depth.
- `out_int` (`bool`, optional) – Whether to return value as integer code value or float equivalent of a code value at a given bit depth.

**Returns** Full range code value  $CV$  or float equivalent of a code value at a given bit depth.

**Return type** ndarray

## Examples

```
>>> legal_to_full(64 / 1023)
0.0
>>> legal_to_full(940 / 1023)
1.0
>>> legal_to_full(64 / 1023, out_int=True)
0
>>> legal_to_full(940 / 1023, out_int=True)
1023
>>> legal_to_full(64, in_int=True)
0.0
>>> legal_to_full(940, in_int=True)
1.0
>>> legal_to_full(64, in_int=True, out_int=True)
0
>>> legal_to_full(940, in_int=True, out_int=True)
1023
```

## colour.CV\_range

`colour.CV_range(bit_depth=10, is_legal=False, is_int=False)`

Returns the code value  $CV$  range for given bit depth, range legality and representation.

### Parameters

- `bit_depth` (`int`, optional) – Bit depth of the code value  $CV$  range.
- `is_legal` (`bool`, optional) – Whether the code value  $CV$  range is legal.
- `is_int` (`bool`, optional) – Whether the code value  $CV$  range represents integer code values.

**Returns** Code value  $CV$  range.

**Return type** ndarray

### Examples

```
>>> CV_range(8, True, True)
array([ 16, 235])
>>> CV_range(8, True, False)
array([ 0.0627451..., 0.9215686...])
>>> CV_range(10, False, False)
array([ 0., 1.])
```

## $IC_T C_P$ Colour Encoding

### colour

---

`RGB_to_ICTCP(RGB[, L_p])`

Converts from ITU-R BT.2020 colourspace to  $IC_T C_P$  colour encoding.

---

`ICTCP_to_RGB(ICTCP[, L_p])`

Converts from  $IC_T C_P$  colour encoding to ITU-R BT.2020 colourspace.

---

### colour.RGB\_to\_ICTCP

`colour.RGB_to_ICTCP(RGB, L_p=10000)`

Converts from ITU-R BT.2020 colourspace to  $IC_T C_P$  colour encoding.

### Parameters

- `RGB` (`array_like`) – ITU-R BT.2020 colourspace array.
- `L_p` (`numeric`, optional) – Display peak luminance  $cd/m^2$  for SMPTE ST 2084:2014 non-linear encoding.

**Returns**  $IC_T C_P$  colour encoding array.

**Return type** ndarray

## References

- [\[Dol16\]](#)
- [\[LPY+16\]](#)

## Examples

```
>>> RGB = np.array([0.35181454, 0.26934757, 0.21288023])
>>> RGB_to_ICTCP(RGB)
array([ 0.0955407..., -0.0089063...,  0.0138928...])
```

## colour.ICTCP\_to\_RGB

`colour.ICTCP_to_RGB(CTCP, L_p=10000)`  
Converts from  $IC_{TCP}$  colour encoding to *ITU-R BT.2020* colourspace.

### Parameters

- `ICTCP` (array\_like) –  $IC_{TCP}$  colour encoding array.
- `L_p` (numeric, optional) – Display peak luminance  $cd/m^2$  for *SMPTE ST 2084:2014* non-linear encoding.

**Returns** *ITU-R BT.2020* colourspace array.

**Return type** ndarray

## References

- [\[Dol16\]](#)
- [\[LPY+16\]](#)

## Examples

```
>>> ICTCP = np.array([0.09554079, -0.00890639, 0.01389286])
>>> ICTCP_to_RGB(ICTCP)
array([ 0.3518145...,  0.2693475...,  0.2128802...])
```

## RGB Representations

### Prismatic Colourspace

colour

<code>RGB_to_Prismatic(RGB)</code>	Converts from <i>RGB</i> colourspace to <i>Prismatic</i> $L\rho\gamma\beta$ colourspace array.
<code>Prismatic_to_RGB(Lrgb)</code>	Converts from <i>Prismatic</i> $L\rho\gamma\beta$ colourspace array to <i>RGB</i> colourspace.

## colour.RGB\_to\_Prismatic

`colour.RGB_to_Prismatic(RGB)`

Converts from *RGB* colourspace to *Prismatic L $\rho\gamma\beta$*  colourspace array.

**Parameters** `RGB` (`array_like`) – *RGB* colourspace array.

**Returns** *Prismatic L $\rho\gamma\beta$*  colourspace array.

**Return type** `ndarray`

### References

- [SH15]

### Examples

```
>>> RGB = np.array([0.25, 0.50, 0.75])
>>> RGB_to_Prismatic(RGB)
array([ 0.75...,  0.166666...,  0.333333...,  0.5...  ])
```

Adjusting saturation of given *RGB* colourspace array: `>>> saturation = 0.5 >>> Lrgb = RGB_to_Prismatic(RGB) >>> Lrgb[..., 1:] = 1 / 3 + saturation * (Lrgb[..., 1:] - 1 / 3) >>> Prismatic_to_RGB(Lrgb) # doctest: +ELLIPSIS array([ 0.45..., 0.6..., 0.75... ])`

## colour.Prismatic\_to\_RGB

`colour.Prismatic_to_RGB(Lrgb)`

Converts from *Prismatic L $\rho\gamma\beta$*  colourspace array to *RGB* colourspace.

**Parameters** `Lrgb` (`array_like`) – *Prismatic L $\rho\gamma\beta$*  colourspace array.

**Returns** *RGB* colourspace array.

**Return type** `ndarray`

### References

- [SH15]

### Examples

```
>>> Lrgb = np.array([0.75000000, 0.16666667, 0.33333333, 0.50000000])
>>> Prismatic_to_RGB(Lrgb)
array([ 0.25...,  0.499999...,  0.75...  ])
```

## HSV Colourspace

`colour`

---

<code>RGB_to_HSV(RGB)</code>	Converts from <i>RGB</i> colourspace to <i>HSV</i> colourspace.
<code>HSV_to_RGB(HSV)</code>	Converts from <i>HSV</i> colourspace to <i>RGB</i> colourspace.

---

## colour.RGB\_to\_HSV

`colour.RGB_to_HSV(RGB)`  
Converts from *RGB* colourspace to *HSV* colourspace.

**Parameters** `RGB` (`array_like`) – *RGB* colourspace array.

**Returns** *HSV* array.

**Return type** `ndarray`

### Notes

- Input *RGB* colourspace array is in domain [0, 1].
- Output *HSV* colourspace array is in range [0, 1].

### References

- [\[Eash\]](#)
- [\[Smi78\]](#)
- [\[Wikj\]](#)

### Examples

```
>>> RGB = np.array([0.49019608, 0.98039216, 0.25098039])
>>> RGB_to_HSV(RGB)
array([ 0.2786738...,  0.744      ,  0.98039216])
```

## colour.HSV\_to\_RGB

`colour.HSV_to_RGB(HSV)`  
Converts from *HSV* colourspace to *RGB* colourspace.

**Parameters** `HSV` (`array_like`) – *HSV* colourspace array.

**Returns** *RGB* colourspace array.

**Return type** `ndarray`

### Notes

- Input *HSV* colourspace array is in domain [0, 1].
- Output *RGB* colourspace array is in range [0, 1].

## References

- [\[Ease\]](#)
- [\[Smi78\]](#)
- [\[Wikj\]](#)

## Examples

```
>>> HSV = np.array([0.27867384, 0.74400000, 0.98039216])
>>> HSV_to_RGB(HSV)
array([ 0.4901960...,  0.9803921...,  0.2509803...])
```

## HSL Colourspace

### colour

<code>RGB_to_HSL(RGB)</code>	Converts from <i>RGB</i> colourspace to <i>HSL</i> colourspace.
<code>HSL_to_RGB(HSL)</code>	Converts from <i>HSL</i> colourspace to <i>RGB</i> colourspace.

### colour.RGB\_to\_HSL

`colour.RGB_to_HSL(RGB)`  
Converts from *RGB* colourspace to *HSL* colourspace.

**Parameters** `RGB` (`array_like`) – *RGB* colourspace array.

**Returns** *HSL* array.

**Return type** ndarray

## Notes

- Input *RGB* colourspace array is in domain [0, 1].
- Output *HSL* colourspace array is in range [0, 1].

## References

- [\[Easg\]](#)
- [\[Smi78\]](#)
- [\[Wikj\]](#)

## Examples

```
>>> RGB = np.array([0.49019608, 0.98039216, 0.25098039])
>>> RGB_to_HSL(RGB)
array([ 0.2786738..., 0.9489796..., 0.6156862...])
```

## colour.HSL\_to\_RGB

`colour.HSL_to_RGB(HSL)`

Converts from *HSL* colourspace to *RGB* colourspace.

**Parameters** `HSL` (`array_like`) – *HSL* colourspace array.

**Returns** *RGB* colourspace array.

**Return type** `ndarray`

### Notes

- Input *HSL* colourspace array is in domain [0, 1].
- Output *RGB* colourspace array is in range [0, 1].

### References

- [\[Easd\]](#)
- [\[Smi78\]](#)
- [\[Wikj\]](#)

### Examples

```
>>> HSL = np.array([0.27867384, 0.94897959, 0.61568627])
>>> HSL_to_RGB(HSL)
array([ 0.4901960..., 0.9803921..., 0.2509803...])
```

## CMY Colourspace

`colour`

<code>RGB_to_CMY(RGB)</code>	Converts from <i>RGB</i> colourspace to <i>CMY</i> colourspace.
<code>CMY_to_RGB(CMY)</code>	Converts from <i>CMY</i> colourspace to <i>CMY</i> colourspace.
<code>CMY_to_CMYK(CMY)</code>	Converts from <i>CMY</i> colourspace to <i>CMYK</i> colourspace.
<code>CMYK_to_CMY(CMYK)</code>	Converts from <i>CMYK</i> colourspace to <i>CMY</i> colourspace.

## colour.RGB\_to\_CMY

`colour.RGB_to_CMY(RGB)`

Converts from *RGB* colourspace to *CMY* colourspace.

**Parameters** `RGB` (`array_like`) – *RGB* colourspace array.  
**Returns** *CMY* array.  
**Return type** `ndarray`

## Notes

- Input *RGB* colourspace array is in domain [0, 1].
- Output *CMY* colourspace array is in range [0, 1].

## References

- [\[Easf\]](#)

## Examples

```
>>> RGB = np.array([0.49019608, 0.98039216, 0.25098039])
>>> RGB_to_CMY(RGB)
array([ 0.5098039...,  0.0196078...,  0.7490196...])
```

## colour.CMY\_to\_RGB

`colour.CMY_to_RGB(CMY)`  
Converts from *CMY* colourspace to *CMY* colourspace.  
**Parameters** `CMY` (`array_like`) – *CMY* colourspace array.  
**Returns** *RGB* colourspace array.  
**Return type** `ndarray`

## Notes

- Input *CMY* colourspace array is in domain [0, 1].
- Output *RGB* colourspace array is in range [0, 1].

## References

- [\[Easb\]](#)

## Examples

```
>>> CMY = np.array([0.50980392, 0.01960784, 0.74901961])
>>> CMY_to_RGB(CMY)
array([ 0.4901960...,  0.9803921...,  0.2509803...])
```

## colour.CMY\_to\_CMYK

`colour.CMY_to_CMYK(CMY)`

Converts from CMY colourspace to CMYK colourspace.

**Parameters** `CMY` (`array_like`) – CMY colourspace array.

**Returns** CMYK array.

**Return type** ndarray

### Notes

- Input CMY colourspace array is in domain [0, 1].
- Output\*CMYK\* colourspace array is in range [0, 1].

### References

- [\[Easa\]](#)

### Examples

```
>>> CMY = np.array([0.50980392, 0.01960784, 0.74901961])
>>> CMY_to_CMYK(CMY)
array([ 0.5         ,  0.          ,  0.744        ,  0.0196078...])
```

## colour.CMYK\_to\_CMY

`colour.CMYK_to_CMY(CMYK)`

Converts from CMYK colourspace to CMY colourspace.

**Parameters** `CMYK` (`array_like`) – CMYK colourspace array.

**Returns** CMY array.

**Return type** ndarray

### Notes

- Input CMYK colourspace array is in domain [0, 1].
- Output CMY colourspace array is in range [0, 1].

### References

- [\[Easc\]](#)

## Examples

```
>>> CMYK = np.array([0.50000000, 0.00000000, 0.74400000, 0.01960784])
>>> CMYK_to_CMY(CMYK)
array([ 0.5098039...,  0.0196078...,  0.7490196...])
```

# Pointer's Gamut

colour

POINTER\_GAMUT\_BOUNDARIES  
POINTER\_GAMUT\_DATA  
POINTER\_GAMUT\_ILLUMINANT

## colour.POINTER\_GAMUT\_BOUNDARIES

```
colour.POINTER_GAMUT_BOUNDARIES = array([[ 0.659,  0.316], [ 0.634,  0.351], [ 0.594,  0.391], [ 0.557,  0.427],
```

## colour.POINTER\_GAMUT\_DATA

#### **colour.POINTER\_GAMUT\_ILLUMINANT**

```
colour.POINTER_GAMUT_ILLUMINANT = array([ 0.31005673, 0.3161457 ])
```

# Colour Notation Systems

- Munsell Renotation System
  - Munsell Value
    - Priest, Gibson and MacNicholas (1920)
    - Munsell, Sloan and Godlove (1933)
    - Moon and Spencer (1943)
    - Saunderson and Milner (1944)
    - Ladd and Pinney (1955)
    - McCamy (1987)
    - ASTM D1535-08e1
  - Hexadecimal Triplet Notation

## Munsell Renotation System

colour

<code>mundsell_colour_to_xyY(mundsell_colour)</code>	Converts given <i>Munsell</i> colour to <i>CIE xyY</i> colourspace.
<code>xyY_to_mundsell_colour(xyY[, hue_decimals, ...])</code>	Converts from <i>CIE xyY</i> colourspace to <i>Munsell</i> colour.

### colour.mundsell\_colour\_to\_xyY

`colour.mundsell_colour_to_xyY(mundsell_colour)`

Converts given *Munsell* colour to *CIE xyY* colourspace.

**Parameters** `mundsell_colour` (`unicode`) – *Munsell* colour.

**Returns** *CIE xyY* colourspace array.

**Return type** `ndarray`, (3,)

### Notes

- Output *CIE xyY* colourspace array is in range [0, 1].

### References

- [\[Cen\]](#)
- [\[Cen12\]](#)

### Examples

```
>>> mundsell_colour_to_xyY('4.2YR 8.1/5.3')
array([ 0.3873694...,  0.3575165...,  0.59362    ])
>>> mundsell_colour_to_xyY('N8.9')
array([ 0.31006   ,  0.31616   ,  0.746134...])
```

### colour.xyY\_to\_mundsell\_colour

`colour.xyY_to_mundsell_colour(xyY, hue_decimals=1, value_decimals=1, chroma_decimals=1)`

Converts from *CIE xyY* colourspace to *Munsell* colour.

**Parameters**

- `xyY` (`array_like`, (3,)) – *CIE xyY* colourspace array.
- `hue_decimals` (`int`) – Hue formatting decimals.
- `value_decimals` (`int`) – Value formatting decimals.
- `chroma_decimals` (`int`) – Chroma formatting decimals.

**Returns** *Munsell* colour.

**Return type** `unicode`

## Notes

- Input *CIE xyY* colourspace array is in domain [0, 1].

## References

- [\[Cen\]](#)
- [\[Cen12\]](#)

## Examples

```
>>> xyY = np.array([0.38736945, 0.35751656, 0.59362000])
>>> # Doctests skip for Python 2.x compatibility.
>>> xyY_to_munsell_colour(xyY)
'4.2YR 8.1/5.3'
```

## Dataset

colour

---

MUNSELL\_COLOURS

Aggregated Munsell colours.

---

## colour.MUNSELL\_COLOURS

colour.MUNSELL\_COLOURS = CaseInsensitiveMapping({'real': ..., 'all': ..., 'Munsell Colours Real': ..., '1929': ...})  
Aggregated Munsell colours.

MUNSELL\_COLOURS : CaseInsensitiveMapping

Aliases:

- ‘all’: ‘Munsell Colours All’
- ‘1929’: ‘Munsell Colours 1929’
- ‘real’: ‘Munsell Colours Real’

## Munsell Value

colour

---

munsell\_value(Y[, method])

Returns the Munsell value *V* of given *luminance Y* using given method.

---

MUNSELL\_VALUE\_METHODS

Supported Munsell value computations methods.

---

## colour.munsell\_value

colour.munsell\_value(*Y*, *method=u'ASTM D1535-08'*)

Returns the Munsell value *V* of given *luminance Y* using given method.

## Parameters

- `Y` (numeric or array\_like) – luminance  $Y$ .
- `method` (unicode, optional) – {‘ASTM D1535-08’, ‘Priest 1920’, ‘Munsell 1933’, ‘Moon 1943’, ‘Saunderson 1944’, ‘Ladd 1955’, ‘McCamy 1987’}, Computation method.

**Returns** Munsell value  $V$ .

**Return type** numeric or ndarray

## Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

## References

- [\[ASTMInternational89\]](#)
- [\[Wikn\]](#)

## Examples

```
>>> munsell_value(10.08)
3.7344764...
>>> munsell_value(10.08, method='Priest 1920')
3.1749015...
>>> munsell_value(10.08, method='Munsell 1933')
3.7918355...
>>> munsell_value(10.08, method='Moon 1943')
3.7462971...
>>> munsell_value(10.08, method='Saunderson 1944')
3.6865080...
>>> munsell_value(10.08, method='Ladd 1955')
3.6952862...
>>> munsell_value(10.08, method='McCamy 1987')
array(3.7347235...)
```

## colour.MUNSELL\_VALUE\_METHODS

`colour.MUNSELL_VALUE_METHODS = CaseInsensitiveMapping({u'Saunderson 1944': ... , u'Mccamy 1987': ... , u'Ladd 1955': ...})`  
Supported Munsell value computations methods.

## References

- [\[ASTMInternational89\]](#)
- [\[Wikn\]](#)

**MUNSELL\_VALUE\_METHODS** [CaseInsensitiveMapping] {'Priest 1920', 'Munsell 1933', 'Moon 1943', 'Saunderson 1944', 'Ladd 1955', 'McCamy 1987', 'ASTM D1535-08'}

Aliases:

- 'astm2008': 'ASTM D1535-08'

### Priest, Gibson and MacNicholas (1920)

colour.notation

---

`mundsell_value_Priest1920(Y)`

Returns the *Munsell* value  $V$  of given luminance  $Y$  using *Priest et alii (1920)* method.

---

### colour.notation.mundsell\_value\_Priest1920

colour.notation.**mundsell\_value\_Priest1920**( $Y$ )

Returns the *Munsell* value  $V$  of given luminance  $Y$  using *Priest et alii (1920)* method.

**Parameters**  $Y$  (numeric or array\_like) – luminance  $Y$ .

**Returns** *Munsell* value  $V$ .

**Return type** numeric or ndarray

### Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

### References

- [\[Wikn\]](#)

### Examples

```
>>> mundsell_value_Priest1920(10.08)
3.1749015...
```

### Munsell, Sloan and Godlove (1933)

colour.notation

---

`mundsell_value_Munsell1933(Y)`

Returns the *Munsell* value  $V$  of given luminance  $Y$  using *Munsell et alii (1933)* method.

---

## colour.notation.munsell\_value\_Munsell1933

`colour.notation.munsell_value_Munsell1933(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using *Munsell et alii (1933)* method.

**Parameters**  $Y$  (numeric or array\_like) – *luminance*  $Y$ .

**Returns** *Munsell* value  $V$ .

**Return type** numeric or ndarray

### Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

### References

- [\[Wikn\]](#)

### Examples

```
>>> munsell_value_Munsell1933(10.08)
3.7918355...
```

## Moon and Spencer (1943)

`colour.notation`

`munsell_value_Moon1943(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using *Moon and Spencer (1943)* method.

## colour.notation.munsell\_value\_Moon1943

`colour.notation.munsell_value_Moon1943(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using *Moon and Spencer (1943)* method.

**Parameters**  $Y$  (numeric or array\_like) – *luminance*  $Y$ .

**Returns** *Munsell* value  $V$ .

**Return type** numeric or ndarray

### Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

## References

- [\[Wikn\]](#)

## Examples

```
>>> munsell_value_Moon1943(10.08)
3.7462971...
```

## Saunderson and Milner (1944)

colour.notation

`munsell_value_Saunderson1944(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using *Saunderson and Milner (1944)* method.

## colour.notation.munsell\_value\_Saunderson1944

`colour.notation.munsell_value_Saunderson1944(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using *Saunderson and Milner (1944)* method.

**Parameters**  $Y$  (numeric) – *luminance*  $Y$ .

**Returns** *Munsell* value  $V$ .

**Return type** numeric

## Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

## References

- [\[Wikn\]](#)

## Examples

```
>>> munsell_value_Saunderson1944(10.08)
3.6865080...
```

## Ladd and Pinney (1955)

colour.notation

---

<code>munsell_value_Ladd1955(Y)</code>	Returns the <i>Munsell</i> value $V$ of given <i>luminance</i> $Y$ using <i>Ladd and Pinney (1955)</i> method.
--	--

---

## colour.notation.munsell\_value\_Ladd1955

`colour.notation.munsell_value_Ladd1955(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using *Ladd and Pinney (1955)* method.

**Parameters**  $Y$  (numeric or array\_like) – *luminance*  $Y$ .

**Returns** *Munsell* value  $V$ .

**Return type** numeric or ndarray

### Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

### References

- [\[Wikn\]](#)

### Examples

```
>>> munsell_value_Ladd1955(10.08)
3.6952862...
```

## McCamy (1987)

`colour.notation`

---

<code>munsell_value_McCamy1987(Y)</code>	Returns the <i>Munsell</i> value $V$ of given <i>luminance</i> $Y$ using <i>McCamy (1987)</i> method.
--	---

---

## colour.notation.munsell\_value\_McCamy1987

`colour.notation.munsell_value_McCamy1987(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using *McCamy (1987)* method.

**Parameters**  $Y$  (numeric or array\_like) – *luminance*  $Y$ .

**Returns** *Munsell* value  $V$ .

**Return type** numeric or ndarray

## Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

## References

- [\[ASTMInternational89\]](#)

## Examples

```
>>> munsell_value_McCamy1987(10.08)
array(3.7347235...)
```

## ASTM D1535-08e1

colour.notation

---

`munsell_value_ASTMD153508(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using a reverse lookup table from *ASTM D1535-08e1* method.

---

## colour.notation.munsell\_value\_ASTMD153508

`colour.notation.munsell_value_ASTMD153508(Y)`

Returns the *Munsell* value  $V$  of given *luminance*  $Y$  using a reverse lookup table from *ASTM D1535-08e1* method.

**Parameters** `Y` (numeric or array\_like) – *luminance*  $Y$

**Returns** *Munsell* value  $V$ .

**Return type** numeric or ndarray

## Notes

- Input  $Y$  is in domain [0, 100].
- Output  $V$  is in range [0, 10].

## References

- [\[ASTMInternational89\]](#)

## Examples

```
>>> munsell_value_ASTMD153508(10.1488096782)
3.7462971...
```

## Hexadecimal Triplet Notation

### colour.notation

<code>RGB_to_HEX(RGB)</code>	Converts from <i>RGB</i> colourspace to hexadecimal triplet representation.
<code>HEX_to_RGB(HEX)</code>	Converts from hexadecimal triplet representation to <i>RGB</i> colourspace.

### colour.notation.RGB\_to\_HEX

#### colour.notation.RGB\_to\_HEX(*RGB*)

Converts from *RGB* colourspace to hexadecimal triplet representation.

**Parameters** `RGB` (`array_like`) – *RGB* colourspace array.

**Returns** Hexadecimal triplet representation.

**Return type** unicode

## Notes

- Input *RGB* colourspace array is in domain [0, 1].

## Examples

```
>>> RGB = np.array([0.66666667, 0.86666667, 1.00000000])
>>> # Doctests skip for Python 2.x compatibility.
>>> RGB_to_HEX(RGB)
'#aaddff'
```

### colour.notation.HEX\_to\_RGB

#### colour.notation.HEX\_to\_RGB(*HEX*)

Converts from hexadecimal triplet representation to *RGB* colourspace.

**Parameters** `HEX` (unicode or `array_like`) – Hexadecimal triplet representation.

**Returns** *RGB* colourspace array.

**Return type** ndarray

## Notes

- Output *RGB* colourspace array is in range [0, 1].

## Examples

```
>>> HEX = '#aaddff'  
>>> HEX_to_RGB(HEX)  
array([ 0.666666...,  0.866666...,  1.        ])
```

## Optical Phenomena

- *Rayleigh Scattering*

### Rayleigh Scattering

#### colour

<code>rayleigh_scattering(wavelength[, ...])</code>	Returns the <i>Rayleigh</i> optical depth $T_r(\lambda)$ as function of wavelength $\lambda$ in centimeters (cm).
<code>rayleigh_scattering_spd([shape, ...])</code>	Returns the <i>Rayleigh</i> spectral power distribution for given spectral shape.
<code>scattering_cross_section(wavelength[, ...])</code>	Returns the scattering cross section per molecule $\sigma$ of dry air as function of wavelength $\lambda$ in centimeters (cm) using given $CO_2$ concentration in parts per million (ppm) and temperature $T[K]$ in kelvin degrees following <i>Van de Hulst (1957)</i> method.

#### colour.rayleigh\_scattering

```
colour.rayleigh_scattering(wavelength,      CO2_concentration=300,      temperature=288.15,  
                           pressure=101325,      latitude=0,      altitude=0,      avo-  
                           gadro_constant=6.02214179e+23,      n_s=<function  
                           air_refraction_index_Bodhaine1999>,      F_air=<function  
                           F_air_Bodhaine1999>)
```

Returns the *Rayleigh* optical depth  $T_r(\lambda)$  as function of wavelength  $\lambda$  in centimeters (cm).

#### Parameters

- **wavelength** (numeric or array\_like) – Wavelength  $\lambda$  in centimeters (cm).
- **CO2\_concentration** (numeric or array\_like, optional) –  $CO_2$  concentration in parts per million (ppm).
- **temperature** (numeric or array\_like, optional) – Air temperature  $T[K]$  in kelvin degrees.
- **pressure** (numeric or array\_like) – Surface pressure  $P$  of the measurement site.

- **latitude** (numeric or array\_like, optional) – Latitude of the site in degrees.
- **altitude** (numeric or array\_like, optional) – Altitude of the site in meters.
- **avogadro\_constant** (numeric or array\_like, optional) – Avogadro's number (molecules  $mol^{-1}$ ).
- **n\_s** (object) – Air refraction index  $n_s$  computation method.
- **F\_air** (object) –  $(6 + 3_p)/(6 - 7_p)$ , the depolarisation term  $F(air)$  or King Factor computation method.

**Returns** Rayleigh optical depth  $T_r(\lambda)$ .

**Return type** numeric or ndarray

**Warning:** Unlike most objects of colour.phenomena.rayleigh module, `colour.phenomena.rayleigh_optical_depth()` expects wavelength  $\lambda$  to be expressed in centimeters (cm).

## References

- [\[BWDS99\]](#)
- [\[Wiku\]](#)

## Examples

```
>>> rayleigh_optical_depth(555 * 10e-8)
0.1004070...
```

## colour.rayleigh\_scattering\_spd

```
colour.rayleigh_scattering_spd(shape=SpectralShape(360, 780, 1), CO2_concentration=300,
                                temperature=288.15, pressure=101325, latitude=0, altitude=0,
                                avogadro_constant=6.02214179e+23, n_s=<function air_refraction_index_Bodhaine1999>, F_air=<function F_air_Bodhaine1999>)
```

Returns the Rayleigh spectral power distribution for given spectral shape.

### Parameters

- **shape** (SpectralShape, optional) – Spectral shape used to create the Rayleigh scattering spectral power distribution.
- **CO2\_concentration** (numeric or array\_like, optional) –  $CO_2$  concentration in parts per million (ppm).
- **temperature** (numeric or array\_like, optional) – Air temperature  $T[K]$  in kelvin degrees.
- **pressure** (numeric or array\_like) – Surface pressure  $P$  of the measurement site.
- **latitude** (numeric or array\_like, optional) – Latitude of the site in degrees.
- **altitude** (numeric or array\_like, optional) – Altitude of the site in meters.

- **avogadro\_constant** (numeric or array\_like, optional) – Avogadro's number (molecules  $mol^{-1}$ ).
- **n\_s** (object) – Air refraction index  $n_s$  computation method.
- **F\_air** (object) –  $(6 + 3_p)/(6 - 7_p)$ , the depolarisation term  $F(air)$  or *King Factor* computation method.

**Returns** Rayleigh optical depth spectral power distribution.

**Return type** *SpectralPowerDistribution*

## References

- [\[BWDS99\]](#)
- [\[Wiku\]](#)

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> with numpy_print_options(suppress=True):
...     rayleigh_scattering_spd()
SpectralPowerDistribution([[ 360.        ,  0.5991013...],
                           [ 361.        ,  0.5921706...],
                           [ 362.        ,  0.5853410...],
                           [ 363.        ,  0.5786105...],
                           [ 364.        ,  0.5719774...],
                           [ 365.        ,  0.5654401...],
                           [ 366.        ,  0.5589968...],
                           [ 367.        ,  0.5526460...],
                           [ 368.        ,  0.5463860...],
                           [ 369.        ,  0.5402153...],
                           [ 370.        ,  0.5341322...],
                           [ 371.        ,  0.5281354...],
                           [ 372.        ,  0.5222234...],
                           [ 373.        ,  0.5163946...],
                           [ 374.        ,  0.5106476...],
                           [ 375.        ,  0.5049812...],
                           [ 376.        ,  0.4993939...],
                           [ 377.        ,  0.4938844...],
                           [ 378.        ,  0.4884513...],
                           [ 379.        ,  0.4830934...],
                           [ 380.        ,  0.4778095...],
                           [ 381.        ,  0.4725983...],
                           [ 382.        ,  0.4674585...],
                           [ 383.        ,  0.4623891...],
                           [ 384.        ,  0.4573889...],
                           [ 385.        ,  0.4524566...],
                           [ 386.        ,  0.4475912...],
                           [ 387.        ,  0.4427917...],
                           [ 388.        ,  0.4380568...],
                           [ 389.        ,  0.4333856...],
                           [ 390.        ,  0.4287771...],
                           [ 391.        ,  0.4242302...],
                           [ 392.        ,  0.4197439...],
                           [ 393.        ,  0.4153172...],
```

[ 394.	,	0.4109493...],
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[ 688.	,	0.0419553...],
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[ 691.	,	0.0412226...],
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[ 697.	,	0.0398047...],
[ 698.	,	0.0395744...],
[ 699.	,	0.0393457...],
[ 700.	,	0.0391187...],
[ 701.	,	0.0388933...],
[ 702.	,	0.0386696...],
[ 703.	,	0.0384474...],
[ 704.	,	0.0382269...],
[ 705.	,	0.0380079...],
[ 706.	,	0.0377905...],
[ 707.	,	0.0375747...],
[ 708.	,	0.0373604...],
[ 709.	,	0.0371476...],
[ 710.	,	0.0369364...],
[ 711.	,	0.0367266...],
[ 712.	,	0.0365184...],
[ 713.	,	0.0363116...],
[ 714.	,	0.0361063...],
[ 715.	,	0.0359024...],
[ 716.	,	0.0357000...],
[ 717.	,	0.0354990...],
[ 718.	,	0.0352994...],
[ 719.	,	0.0351012...],
[ 720.	,	0.0349044...],
[ 721.	,	0.0347090...],
[ 722.	,	0.0345150...],
[ 723.	,	0.0343223...],
[ 724.	,	0.0341310...],
[ 725.	,	0.0339410...],
[ 726.	,	0.0337523...],
[ 727.	,	0.033565 ...],
[ 728.	,	0.0333789...],
[ 729.	,	0.0331941...],
[ 730.	,	0.0330106...],
[ 731.	,	0.0328284...],
[ 732.	,	0.0326474...],
[ 733.	,	0.0324677...],
[ 734.	,	0.0322893...],
[ 735.	,	0.0321120...],
[ 736.	,	0.0319360...],
[ 737.	,	0.0317611...],
[ 738.	,	0.0315875...],
[ 739.	,	0.0314151...],
[ 740.	,	0.0312438...],
[ 741.	,	0.0310737...],

```
[ 742. , 0.0309048...],
[ 743. , 0.0307370...],
[ 744. , 0.0305703...],
[ 745. , 0.0304048...],
[ 746. , 0.0302404...],
[ 747. , 0.0300771...],
[ 748. , 0.0299149...],
[ 749. , 0.0297538...],
[ 750. , 0.0295938...],
[ 751. , 0.0294349...],
[ 752. , 0.0292771...],
[ 753. , 0.0291203...],
[ 754. , 0.0289645...],
[ 755. , 0.0288098...],
[ 756. , 0.0286561...],
[ 757. , 0.0285035...],
[ 758. , 0.0283518...],
[ 759. , 0.0282012...],
[ 760. , 0.0280516...],
[ 761. , 0.0279030...],
[ 762. , 0.0277553...],
[ 763. , 0.0276086...],
[ 764. , 0.027463 ...],
[ 765. , 0.0273182...],
[ 766. , 0.0271744...],
[ 767. , 0.0270316...],
[ 768. , 0.0268897...],
[ 769. , 0.0267487...],
[ 770. , 0.0266087...],
[ 771. , 0.0264696...],
[ 772. , 0.0263314...],
[ 773. , 0.0261941...],
[ 774. , 0.0260576...],
[ 775. , 0.0259221...],
[ 776. , 0.0257875...],
[ 777. , 0.0256537...],
[ 778. , 0.0255208...],
[ 779. , 0.0253888...],
[ 780. , 0.0252576...]],  

interpolator=SpragueInterpolator,  

interpolator_args={},  

extrapolator=Extrapolator,  

extrapolator_args={...})
```

## colour.scattering\_cross\_section

```
colour.scattering_cross_section(wavelength, CO2_concentration=300, temperature=288.15,
                                 avogadro_constant=6.02214179e+23, n_s=<function
                                 air_refraction_index_Bodhaine1999>, F_air=<function
                                 F_air_Bodhaine1999>)
```

Returns the scattering cross section per molecule  $\sigma$  of dry air as function of wavelength  $\lambda$  in centimeters (cm) using given  $CO_2$  concentration in parts per million (ppm) and temperature  $T[K]$  in kelvin degrees following *Van de Hulst (1957)* method.

### Parameters

- **wavelength** (numeric or array\_like) – Wavelength  $\lambda$  in centimeters (cm).

- **CO2\_concentration** (numeric or array\_like, optional) –  $CO_2$  concentration in parts per million (ppm).
- **temperature** (numeric or array\_like, optional) – Air temperature  $T[K]$  in kelvin degrees.
- **avogadro\_constant** (numeric or array\_like, optional) – Avogadro's number (molecules  $mol^{-1}$ ).
- **n\_s** (object) – Air refraction index  $n_s$  computation method.
- **F\_air** (object) –  $(6 + 3_p)/(6 - 7_p)$ , the depolarisation term  $F(air)$  or King Factor computation method.

**Returns** Scattering cross section per molecule  $\sigma$  of dry air.

**Return type** numeric or ndarray

**Warning:** Unlike most objects of colour.phenomena.rayleigh module, `colour.scattering_cross_section()` expects wavelength  $\lambda$  to be expressed in centimeters (cm).

## References

- [\[BWDS99\]](#)
- [\[Wiku\]](#)

## Examples

```
>>> scattering_cross_section(555 * 10e-8)
4.6613309...e-27
```

colour.phenomena

---

<code>rayleigh_optical_depth(wavelength[, ...])</code>	Returns the Rayleigh optical depth $T_r(\lambda)$ as function of wavelength $\lambda$ in centimeters (cm).
--	--

---

## colour.phenomena.rayleigh\_optical\_depth

`colour.phenomena.rayleigh_optical_depth(wavelength, CO2_concentration=300, temperature=288.15, pressure=101325, latitude=0, altitude=0, avogadro_constant=6.02214179e+23, n_s=<function air_refraction_index_Bodhaine1999>, F_air=<function F_air_Bodhaine1999>)`

Returns the Rayleigh optical depth  $T_r(\lambda)$  as function of wavelength  $\lambda$  in centimeters (cm).

### Parameters

- **wavelength** (numeric or array\_like) – Wavelength  $\lambda$  in centimeters (cm).
- **CO2\_concentration** (numeric or array\_like, optional) –  $CO_2$  concentration in parts per million (ppm).
- **temperature** (numeric or array\_like, optional) – Air temperature  $T[K]$  in kelvin

degrees.

- **pressure** (numeric or array\_like) – Surface pressure  $P$  of the measurement site.
- **latitude** (numeric or array\_like, optional) – Latitude of the site in degrees.
- **altitude** (numeric or array\_like, optional) – Altitude of the site in meters.
- **avogadro\_constant** (numeric or array\_like, optional) – Avogadro's number (molecules  $mol^{-1}$ ).
- **n\_s** (object) – Air refraction index  $n_s$  computation method.
- **F\_air** (object) –  $(6 + 3_p)/(6 - 7_p)$ , the depolarisation term  $F(air)$  or King Factor computation method.

**Returns** Rayleigh optical depth  $T_r(\lambda)$ .

**Return type** numeric or ndarray

**Warning:** Unlike most objects of colour.phenomena.rayleigh module, `colour.phenomena.rayleigh_optical_depth()` expects wavelength  $\lambda$  to be expressed in centimeters (cm).

## References

- [\[BWDS99\]](#)
- [\[Wiku\]](#)

## Examples

```
>>> rayleigh_optical_depth(555 * 10e-8)
0.1004070...
```

## Plotting

- [Common](#)
- [Colorimetry](#)
- [Colour Characterisation](#)
- [Corresponding Chromaticities](#)
- [CIE Chromaticity Diagrams](#)
- [Colour Models](#)
- [Colour Notation Systems](#)
- [Optical Phenomena](#)
- [Colour Quality](#)
- [Colour Temperature & Correlated Colour Temperature](#)

- [Colour Models Volume](#)
- [Geometry Plotting Utilities](#)

## Common

### colour.plotting

<code>colour_plotting_defaults([parameters])</code>	Enables <i>Colour</i> default plotting parameters.
<code>colour_cycle(**kwargs)</code>	Returns a colour cycle iterator using given colour map.
<code>canvas(**kwargs)</code>	Sets the figure size.
<code>camera(**kwargs)</code>	Sets the camera settings.
<code>decorate(**kwargs)</code>	Sets the figure decorations.
<code>boundaries(**kwargs)</code>	Sets the plot boundaries.
<code>display(**kwargs)</code>	Sets the figure display.
<code>render([with_boundaries, with_decorate])</code>	Convenient wrapper definition combining <code>colour.plotting.decorate()</code> , <code>colour.plotting.boundaries()</code> and <code>colour.plotting.display()</code> definitions.
<code>label_rectangles(rectangles[, rotation, ...])</code>	Add labels above given rectangles.
<code>equal_axes3d(axes)</code>	Sets equal aspect ratio to given 3d axes.
<code>single_colour_swatch_plot(colour_swatch, ...)</code>	Plots given colour swatch.
<code>multi_colour_swatches_plot(colour_swatches)</code>	Plots given colours swatches.
<code>image_plot(image[, label, label_size, ...])</code>	Plots given image.

### colour.plotting.colour\_plotting\_defaults

`colour.plotting.colour_plotting_defaults(parameters=None)`

Enables *Colour* default plotting parameters.

**Parameters** `parameters` (`dict`, optional) – Parameters to use for plotting.

**Returns** Definition success.

**Return type** `bool`

### colour.plotting.colour\_cycle

`colour.plotting.colour_cycle(**kwargs)`

Returns a colour cycle iterator using given colour map.

**Other Parameters**

- `colour_cycle_map` (`unicode`, optional) – Matplotlib colourmap name.
- `colour_cycle_count` (`int`, optional) – Colours count to pick in the colourmap.

**Returns** Colour cycle iterator.

**Return type** `cycle`

**colour.plotting.canvas**

```
colour.plotting.canvas(**kwargs)
```

Sets the figure size.

**Other Parameters** `figure_size` (*array\_like, optional*) – Array defining figure *width* and *height* such as `figure_size = (width, height)`.

**Returns** Current figure.

**Return type** Figure

**colour.plotting.camera**

```
colour.plotting.camera(**kwargs)
```

Sets the camera settings.

**Other Parameters**

- `camera_aspect` (*unicode, optional*) – Matplotlib axes aspect. Default is *equal*.
- `elevation` (*numeric, optional*) – Camera elevation.
- `azimuth` (*numeric, optional*) – Camera azimuth.

**Returns** Current axes.

**Return type** Axes

**colour.plotting.decorate**

```
colour.plotting.decorate(**kwargs)
```

Sets the figure decorations.

**Other Parameters**

- `title` (*unicode, optional*) – Figure title.
- `x_label` (*unicode, optional*) – *X* axis label.
- `y_label` (*unicode, optional*) – *Y* axis label.
- `legend` (*bool, optional*) – Whether to display the legend. Default is *False*.
- `legend_columns` (*int, optional*) – Number of columns in the legend. Default is 1.
- `legend_location` (*unicode, optional*) – Matplotlib legend location. Default is *upper right*.
- `x_ticker` (*bool, optional*) – Whether to display the *X* axis ticker. Default is *True*.
- `y_ticker` (*bool, optional*) – Whether to display the *Y* axis ticker. Default is *True*.
- `x_ticker_locator` (*Locator, optional*) – Locator type for the *X* axis ticker.
- `y_ticker_locator` (*Locator, optional*) – Locator type for the *Y* axis ticker.
- `grid` (*bool, optional*) – Whether to display the grid. Default is *False*.
- `grid_which` (*unicode, optional*) – Controls whether major tick grids, minor tick grids, or both are affected. Default is *both*.

- **grid\_axis** (*unicode, optional*) – Controls which set of grid-lines are drawn. Default is *both*.
- **x\_axis\_line** (*bool, optional*) – Whether to draw the X axis line. Default is *False*.
- **y\_axis\_line** (*bool, optional*) – Whether to draw the Y axis line. Default is *False*.
- **aspect** (*unicode, optional*) – Matplotlib axes aspect.
- **no\_axes** (*bool, optional*) – Whether to turn off the axes. Default is *False*.

**Returns** Current axes.

**Return type** Axes

## colour.plotting.boundaries

colour.plotting.boundaries(\*\*kwargs)

Sets the plot boundaries.

### Other Parameters

- **bounding\_box** (*array\_like, optional*) – Array defining current axes limits such as *bounding\_box* =  $(x \min, x \max, y \min, y \max)$ .
- **x\_tighten** (*bool, optional*) – Whether to tighten the X axis limit. Default is *False*.
- **y\_tighten** (*bool, optional*) – Whether to tighten the Y axis limit. Default is *False*.
- **limits** (*array\_like, optional*) – Array defining current axes limits such as *limits* =  $(x \text{ limit min}, x \text{ limit max}, y \text{ limit min}, y \text{ limit max})$ . *limits* argument values are added to the *margins* argument values to define the final bounding box for the current axes.
- **margins** (*array\_like, optional*) – Array defining current axes margins such as *margins* =  $(x \text{ margin min}, x \text{ margin max}, y \text{ margin min}, y \text{ margin max})$ . *margins* argument values are added to the *limits* argument values to define the final bounding box for the current axes.

**Returns** Current axes.

**Return type** Axes

## colour.plotting.display

colour.plotting.display(\*\*kwargs)

Sets the figure display.

### Other Parameters

- **standalone** (*bool, optional*) – Whether to show the figure.
- **filename** (*unicode, optional*) – Figure will be saved using given *filename* argument.

**Returns** Current figure or None.

**Return type** Figure

## colour.plotting.render

`colour.plotting.render(with_boundaries=True, with_decorate=True, **kwargs)`  
 Convenient wrapper definition combining `colour.plotting.decorate()`, `colour.plotting.boundaries()` and `colour.plotting.display()` definitions.

### Parameters

- `with_boundaries` (`bool`, optional) – Whether to call `colour.plotting.boundaries()` definition.
- `with_decorate` (`bool`, optional) – Whether to call `colour.plotting.decorate()` definition.

**Other Parameters** `**kwargs` (`dict, optional`) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## colour.plotting.label\_rectangles

`colour.plotting.label_rectangles(rectangles, rotation='vertical', text_size=10, offset=None)`  
 Add labels above given rectangles.

### Parameters

- `rectangles` (`object`) – Rectangles to used to set the labels value and position.
- `rotation` (`unicode`, optional) – `{'horizontal', 'vertical'}`, Labels orientation.
- `text_size` (`numeric`, optional) – Labels text size.
- `offset` (`array_like`, optional) – Labels offset as percentages of the largest rectangle dimensions.

**Returns** Definition success.

**Return type** bool

## colour.plotting.equal\_axes3d

`colour.plotting.equal_axes3d(axes)`  
 Sets equal aspect ratio to given 3d axes.

**Parameters** `axes` (`object`) – Axis to set the equal aspect ratio.

**Returns** Definition success.

**Return type** bool

## colour.plotting.single\_colour\_swatch\_plot

`colour.plotting.single_colour_swatch_plot(colour_swatch, **kwargs)`  
 Plots given colour swatch.

**Parameters** `colour_swatch` (`ColourSwatch`) – ColourSwatch.

**Other Parameters**

- **\*\*kwargs** (*dict, optional*) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **width** (*numeric, optional*) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatch width.
- **height** (*numeric, optional*) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatch height.
- **spacing** (*numeric, optional*) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatches spacing.
- **columns** (*int, optional*) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatches columns count.
- **text\_display** (*bool, optional*) – {`colour.plotting.multi_colour_swatches_plot()`}, Display colour text.
- **text\_size** (*numeric, optional*) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour text size.
- **text\_offset** (*numeric, optional*) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour text offset.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> RGB = (0.32315746, 0.32983556, 0.33640183)
>>> single_colour_swatch_plot(ColourSwatch(RGB))
```

## colour.plotting.multi\_colour\_swatches\_plot

```
colour.plotting.multi_colour_swatches_plot(colour_swatches, width=1, height=1, spacing=0,
                                            columns=3, text_display=True, text_size='large',
                                            text_offset=0.075, background_colour=(1.0, 1.0,
                                            1.0), **kwargs)
```

Plots given colours swatches.

### Parameters

- **colour\_swatches** (*list*) – ColourSwatch sequence.
- **width** (*numeric, optional*) – Colour swatch width.
- **height** (*numeric, optional*) – Colour swatch height.
- **spacing** (*numeric, optional*) – Colour swatches spacing.
- **columns** (*int, optional*) – Colour swatches columns count.
- **text\_display** (*bool, optional*) – Display colour text.
- **text\_size** (*numeric, optional*) – Colour text size.
- **text\_offset** (*numeric, optional*) – Colour text offset.
- **background\_colour** (*array\_like or unicode, optional*) – Background colour.

**Other Parameters** `**kwargs` (*dict, optional*) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> cp1 = ColourSwatch(RGB=(0.45293517, 0.31732158, 0.26414773))
>>> cp2 = ColourSwatch(RGB=(0.77875824, 0.57726450, 0.50453169))
>>> multi_colour_swatches_plot([cp1, cp2])
```

## colour.plotting.image\_plot

`colour.plotting.image_plot(image, label=None, label_size=15, label_colour=None, label_alpha=0.85, interpolation='nearest', colour_map=<matplotlib.colors.LinearSegmentedColormap object>, **kwargs)`

Plots given image.

### Parameters

- `image` (`array_like`) – Image to plot.
- `label` (`unicode, optional`) – Image label.
- `label_size` (`int, optional`) – Image label font size.
- `label_colour` (`array_like or unicode, optional`) – Image label colour.
- `label_alpha` (`numeric, optional`) – Image label alpha.
- `interpolation` (`unicode, optional`) – {‘nearest’, ‘None’, ‘none’, ‘bilinear’, ‘bicubic’, ‘spline16’, ‘spline36’, ‘hanning’, ‘hamming’, ‘hermite’, ‘kaiser’, ‘quadric’, ‘catrom’, ‘gaussian’, ‘bessel’, ‘mitchell’, ‘sinc’, ‘lanczos’} Image display interpolation.
- `colour_map` (`unicode, optional`) – Colour map used to display single channel images.

**Other Parameters** `**kwargs` (*dict, optional*) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> import os
>>> from colour import read_image
>>> path = os.path.join('resources',
...                     ('CIE_1931_Chromaticity_Diagram'
...                      '_CIE_1931_2_Degree_Standard_Observer.png'))
>>> image = read_image(path)
>>> image_plot(image)
```

## Colorimetry

### colour.plotting

<code>single_spd_plot(spd[, cmfs, ...])</code>	Plots given spectral power distribution.
<code>multi_spd_plot(spds[, cmfs, ...])</code>	Plots given spectral power distributions.
<code>single_cmfs_plot([cmfs])</code>	Plots given colour matching functions.
<code>multi_cmfs_plot([cmfs])</code>	Plots given colour matching functions.
<code>single_illuminant_relative_spd_plot([...])</code>	Plots given single illuminant relative spectral power distribution.
<code>multi_illuminants_relative_spd_plot([...])</code>	Plots given illuminants relative spectral power distributions.
<code>visible_spectrum_plot([cmfs, ...])</code>	Plots the visible colours spectrum using given standard observer <i>CIE XYZ</i> colour matching functions.
<code>single_lightness_function_plot([function])</code>	Plots given <i>Lightness</i> function.
<code>multi_lightness_function_plot([functions])</code>	Plots given <i>Lightness</i> functions.
<code>blackbody_spectral_radiance_plot([...])</code>	Plots given blackbody spectral radiance.
<code>blackbody_colours_plot([shape, cmfs])</code>	Plots blackbody colours.

### colour.plotting.single\_spd\_plot

```
colour.plotting.single_spd_plot(spd,    cmfs='CIE 1931 2 Degree Standard Observer',
                                 out_of_gamut_clipping=True, **kwargs)
```

Plots given spectral power distribution.

#### Parameters

- `spd` (`SpectralPowerDistribution`) – Spectral power distribution to plot.
- `out_of_gamut_clipping` (`bool`, optional) – Whether to clip out of gamut colours otherwise, the colours will be offset by the absolute minimal colour leading to a rendering on gray background, less saturated and smoother.
- `cmfs` (unicode) – Standard observer colour matching functions used for spectrum creation.

**Other Parameters** `**kwargs` (*dict, optional*) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## References

- [Spi15]

## Examples

```
>>> from colour import SpectralPowerDistribution
>>> data = {
...     500: 0.0651,
...     520: 0.0705,
```

```

...
    540: 0.0772,
...
    560: 0.0870,
...
    580: 0.1128,
...
    600: 0.1360
...
}
>>> spd = SpectralPowerDistribution(data, name='Custom')
>>> single_spd_plot(spd)

```

## colour.plotting.multi\_spd\_plot

`colour.plotting.multi_spd_plot(spds, cmfs='CIE 1931 2 Degree Standard Observer', use_spds_colours=False, normalise_spds_colours=False, **kwargs)`

Plots given spectral power distributions.

### Parameters

- `spds` (`list`) – Spectral power distributions to plot.
- `cmfs` (`unicode`, optional) – Standard observer colour matching functions used for spectrum creation.
- `use_spds_colours` (`bool`, optional) – Whether to use spectral power distributions colours.
- `normalise_spds_colours` (`bool`) – Whether to normalise spectral power distributions colours.

**Other Parameters** `**kwargs` (`dict, optional`) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```

>>> from colour import SpectralPowerDistribution
>>> data_1 = {
...
    500: 0.004900,
...
    510: 0.009300,
...
    520: 0.063270,
...
    530: 0.165500,
...
    540: 0.290400,
...
    550: 0.433450,
...
    560: 0.594500
...
}
>>> data_2 = {
...
    500: 0.323000,
...
    510: 0.503000,
...
    520: 0.710000,
...
    530: 0.862000,
...
    540: 0.954000,
...
    550: 0.994950,
...
    560: 0.995000
...
}
>>> spd1 = SpectralPowerDistribution(data_1, name='Custom 1')

```

```
>>> spd2 = SpectralPowerDistribution(data_2, name='Custom 2')
>>> multi_spd_plot([spd1, spd2])
```

## colour.plotting.single\_cmfs\_plot

colour.plotting.**single\_cmfs\_plot**(*cmfs*=‘CIE 1931 2 Degree Standard Observer’, \*\**kwargs*)  
Plots given colour matching functions.

**Parameters** **cmfs** (unicode, optional) – Colour matching functions to plot.

**Other Parameters** **\*\*kwargs** (dict, optional) – {colour.plotting.render()}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> single_cmfs_plot()
```

## colour.plotting.multi\_cmfs\_plot

colour.plotting.**multi\_cmfs\_plot**(*cmfs*=None, \*\**kwargs*)  
Plots given colour matching functions.

**Parameters** **cmfs** (array\_like, optional) – Colour matching functions to plot.

**Other Parameters** **\*\*kwargs** (dict, optional) – {colour.plotting.render()}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> cmfs = [
... 'CIE 1931 2 Degree Standard Observer',
... 'CIE 1964 10 Degree Standard Observer']
>>> multi_cmfs_plot(cmfs)
```

## colour.plotting.single\_illuminant\_relative\_spd\_plot

colour.plotting.**single\_illuminant\_relative\_spd\_plot**(*illuminant*=‘A’, *cmfs*=‘CIE 1931 2 Degree Standard Observer’, \*\**kwargs*)  
Plots given single illuminant relative spectral power distribution.

**Parameters**

- **illuminant** (unicode, optional) – Factory illuminant to plot.
- **cmfs** (unicode, optional) – Standard observer colour matching functions to plot.

## Other Parameters

- **\*\*kwargs** (*dict, optional*) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **out\_of\_gamut\_clipping** (*bool, optional*) – {`colour.plotting.single_spd_plot()`}, Whether to clip out of gamut colours otherwise, the colours will be offset by the absolute minimal colour leading to a rendering on gray background, less saturated and smoother.

**Returns** Current figure or None.

**Return type** Figure

## References

- [Spi15]

## Examples

```
>>> single_illuminant_relative_spd_plot()
```

### colour.plotting.multi\_illuminants\_relative\_spd\_plot

`colour.plotting.multi_illuminants_relative_spd_plot(illuminants=None, **kwargs)`  
Plots given illuminants relative spectral power distributions.

**Parameters** `illuminants` (`array_like`, optional) – Factory illuminants to plot.

## Other Parameters

- **\*\*kwargs** (*dict, optional*) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **use\_spds\_colours** (*bool, optional*) – {`colour.plotting.multi_spd_plot()`} Whether to use spectral power distributions colours.
- **normalise\_spds\_colours** (*bool*) – {`colour.plotting.multi_spd_plot()`} Whether to normalise spectral power distributions colours.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> multi_illuminants_relative_spd_plot(['A', 'B', 'C'])
```

### colour.plotting.visible\_spectrum\_plot

`colour.plotting.visible_spectrum_plot(cmfs='CIE 1931 2 Degree Standard Observer', out_of_gamut_clipping=True, **kwargs)`  
Plots the visible colours spectrum using given standard observer CIE XYZ colour matching functions.

## Parameters

- **cmfs** (unicode, optional) – Standard observer colour matching functions used for spectrum creation.
- **out\_of\_gamut\_clipping** (bool, optional) – Whether to clip out of gamut colours otherwise, the colours will be offset by the absolute minimal colour leading to a rendering on gray background, less saturated and smoother.

**Other Parameters** `**kwargs` (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## References

- [Spi15]

## Examples

```
>>> visible_spectrum_plot()
```

### `colour.plotting.single_lightness_function_plot`

```
colour.plotting.single_lightness_function_plot(function='CIE 1976', **kwargs)
```

Plots given *Lightness* function.

**Parameters** `function` (unicode, optional) – *Lightness* function to plot.

**Other Parameters** `**kwargs` (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> single_lightness_function_plot()
```

### `colour.plotting.multi_lightness_function_plot`

```
colour.plotting.multi_lightness_function_plot(functions=None, **kwargs)
```

Plots given *Lightness* functions.

**Parameters** `functions` (array\_like, optional) – *Lightness* functions to plot.

**Other Parameters** `**kwargs` (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

**Raises** `KeyError` – If one of the given *Lightness* function is not found in the factory *Lightness* functions.

## Examples

```
>>> fs = ('CIE 1976', 'Wyszecki 1963')
>>> multi_lightness_function_plot(fs)
```

## colour.plotting.blackbody\_spectral\_radiance\_plot

`colour.plotting.blackbody_spectral_radiance_plot(temperature=3500, cmfs='CIE 1931 2 Degree Standard Observer', blackbody='VY Canis Major', **kwargs)`

Plots given blackbody spectral radiance.

### Parameters

- `temperature` (numeric, optional) – Blackbody temperature.
- `cmfs` (unicode, optional) – Standard observer colour matching functions.
- `blackbody` (unicode, optional) – Blackbody name.

**Other Parameters** `**kwargs` (`dict, optional`) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> blackbody_spectral_radiance_plot()
```

## colour.plotting.blackbody\_colours\_plot

`colour.plotting.blackbody_colours_plot(shape=SpectralShape(150, 12500, 50), cmfs='CIE 1931 2 Degree Standard Observer', **kwargs)`

Plots blackbody colours.

### Parameters

- `shape` (`SpectralShape`, optional) – Spectral shape to use as plot boundaries.
- `cmfs` (unicode, optional) – Standard observer colour matching functions.

**Other Parameters** `**kwargs` (`dict, optional`) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> blackbody_colours_plot()
```

## Colour Characterisation

### colour.plotting

```
colour_checker_plot([colour_checker])
```

Plots given colour checker.

### colour.plotting.colour\_checker\_plot

```
colour.plotting.colour_checker_plot(colour_checker='ColorChecker 2005', **kwargs)
```

Plots given colour checker.

**Parameters** `colour_checker` (unicode, optional) – Color checker name.

#### Other Parameters

- `**kwargs` (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- `width` (numeric, optional) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatch width.
- `height` (numeric, optional) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatch height.
- `spacing` (numeric, optional) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatches spacing.
- `columns` (int, optional) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour swatches columns count.
- `text_display` (bool, optional) – {`colour.plotting.multi_colour_swatches_plot()`}, Display colour text.
- `text_size` (numeric, optional) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour text size.
- `text_offset` (numeric, optional) – {`colour.plotting.multi_colour_swatches_plot()`}, Colour text offset.

**Returns** Current figure or None.

**Return type** Figure

**Raises** `KeyError` – If the given colour rendition chart is not found in the factory colour rendition charts.

## Examples

```
>>> colour_checker_plot()
```

## Corresponding Chromaticities

`colour.plotting`

---

```
corresponding_chromaticities_prediction_plot([...])Plots given chromatic adaptation model corresponding chromaticities prediction.
```

---

### `colour.plotting.corresponding_chromaticities_prediction_plot`

`colour.plotting.corresponding_chromaticities_prediction_plot(experiment=1, model='Von Kries', transform='CAT02', **kwargs)`

Plots given chromatic adaptation model corresponding chromaticities prediction.

#### Parameters

- **experiment** (`int`, optional) – Corresponding chromaticities prediction experiment number.
- **model** (`unicode`, optional) – Corresponding chromaticities prediction model name.
- **transform** (`unicode`, optional) – Transformation to use with *Von Kries* chromatic adaptation model.

#### Other Parameters

- **\*\*kwargs** (`dict`, optional) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (`bool`, optional) – `{colour.plotting.chromaticity_diagram_plot_CIE1976UCS()}` Whether to display the chromaticity diagram background colours.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> corresponding_chromaticities_prediction_plot()
```

## CIE Chromaticity Diagrams

`colour.plotting`

---

<code>chromaticity_diagram_plot_CIE1931([cmfs, ...])</code>	Plots the <i>CIE 1931 Chromaticity Diagram</i> .
<code>chromaticity_diagram_plot_CIE1960UCS([cmfs, ...])</code>	Plots the <i>CIE 1960 UCS Chromaticity Diagram</i> .
<code>chromaticity_diagram_plot_CIE1976UCS([cmfs, ...])</code>	Plots the <i>CIE 1976 UCS Chromaticity Diagram</i> .
<code>spds_chromaticity_diagram_plot_CIE1931(spds)</code>	Plots given spectral power distribution chromaticity coordinates into the <i>CIE 1931 Chromaticity Diagram</i> .

Continued on next page

Table 3.177 – continued from previous page

<code>spds_chromaticity_diagram_plot_CIE1960UCS(spds)</code>	Plots given spectral power distribution chromaticity coordinates into the <i>CIE 1960 UCS Chromaticity Diagram</i> .
<code>spds_chromaticity_diagram_plot_CIE1976UCS(spds)</code>	Plots given spectral power distribution chromaticity coordinates into the <i>CIE 1976 UCS Chromaticity Diagram</i> .

## `colour.plotting.chromaticity_diagram_plot_CIE1931`

```
colour.plotting.chromaticity_diagram_plot_CIE1931(cmfs='CIE 1931 2 Degree Standard Observer',
                                                show_diagram_colours=True,
                                                use_cached_diagram_colours=True,
                                                **kwargs)
```

Plots the *CIE 1931 Chromaticity Diagram*.

### Parameters

- `cmfs` (unicode, optional) – Standard observer colour matching functions used for diagram bounds.
- `show_diagram_colours` (bool, optional) – Whether to display the chromaticity diagram background colours.
- `use_cached_diagram_colours` (bool, optional) – Whether to used the cached chromaticity diagram background colours image.

**Other Parameters** `**kwargs` (dict, optional) – {`colour.plotting.diagrams.chromaticity_diagram_colours_CIE1931()`, `colour.plotting.render()`}, Please refer to the documentation of the previously listed definitions.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> chromaticity_diagram_plot_CIE1931()
```

## `colour.plotting.chromaticity_diagram_plot_CIE1960UCS`

```
colour.plotting.chromaticity_diagram_plot_CIE1960UCS(cmfs='CIE 1931 2 Degree Standard Observer',
                                                       show_diagram_colours=True,
                                                       use_cached_diagram_colours=True,
                                                       **kwargs)
```

Plots the *CIE 1960 UCS Chromaticity Diagram*.

### Parameters

- `cmfs` (unicode, optional) – Standard observer colour matching functions used for diagram bounds.
- `show_diagram_colours` (bool, optional) – Whether to display the chromaticity diagram background colours.

- `use_cached_diagram_colours` (`bool`, optional) – Whether to used the cached chromaticity diagram background colours image.

**Other Parameters** `**kwargs` (`dict`, *optional*) – {`colour.plotting.diagrams.chromaticity_diagram_colours_CIE1960UCS()`, `colour.plotting.render()`}, Please refer to the documentation of the previously listed definitions.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> chromaticity_diagram_plot_CIE1960UCS()
```

## `colour.plotting.chromaticity_diagram_plot_CIE1976UCS`

```
colour.plotting.chromaticity_diagram_plot_CIE1976UCS(cmfs='CIE 1931 2 Degree Standard Observer',
                                                    show_diagram_colours=True,
                                                    use_cached_diagram_colours=True,
                                                    **kwargs)
```

Plots the *CIE 1976 UCS Chromaticity Diagram*.

### Parameters

- `cmfs` (`unicode`, optional) – Standard observer colour matching functions used for diagram bounds.
- `show_diagram_colours` (`bool`, optional) – Whether to display the chromaticity diagram background colours.
- `use_cached_diagram_colours` (`bool`, optional) – Whether to used the cached chromaticity diagram background colours image.

**Other Parameters** `**kwargs` (`dict`, *optional*) – {`colour.plotting.diagrams.chromaticity_diagram_colours_CIE1976UCS()`, `colour.plotting.render()`}, Please refer to the documentation of the previously listed definitions.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> chromaticity_diagram_plot_CIE1976UCS()
```

## colour.plotting.spds\_chromaticity\_diagram\_plot\_CIE1931

```
colour.plotting.spds_chromaticity_diagram_plot_CIE1931(spds,      cmfs='CIE 1931 2 Degree Standard Observer',      annotate=True,      chromaticity_diagram_callable_CIE1931=<function chromaticity_diagram_plot_CIE1931>,      **kwargs)
```

Plots given spectral power distribution chromaticity coordinates into the *CIE 1931 Chromaticity Diagram*.

### Parameters

- **spds** (array\_like, optional) – Spectral power distributions to plot.
- **cmfs** (unicode, optional) – Standard observer colour matching functions used for diagram bounds.
- **annotate** (bool) – Should resulting chromaticity coordinates annotated with their respective spectral power distribution names.
- **chromaticity\_diagram\_callable\_CIE1931** (callable, optional) – Callable responsible for drawing the *CIE 1931 Chromaticity Diagram*.

### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1931()`}, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1931()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> from colour import ILLUMINANTS_RELATIVE_SPDS
>>> A = ILLUMINANTS_RELATIVE_SPDS['A']
>>> D65 = ILLUMINANTS_RELATIVE_SPDS['D65']
>>> spds_chromaticity_diagram_plot_CIE1931([A, D65])
```

## colour.plotting.spds\_chromaticity\_diagram\_plot\_CIE1960UCS

```
colour.plotting.spds_chromaticity_diagram_plot_CIE1960UCS(spds, cmfs='CIE 1931 2 Degree Standard Observer', annotate=True, chromaticity_diagram_callable_CIE1960UCS=<function chromaticity_diagram_plot_CIE1960UCS>, **kwargs)
```

Plots given spectral power distribution chromaticity coordinates into the *CIE 1960 UCS Chromaticity Diagram*.

### Parameters

- **spds** (array\_like, optional) – Spectral power distributions to plot.
- **cmfs** (unicode, optional) – Standard observer colour matching functions used for diagram bounds.
- **annotate** (bool) – Should resulting chromaticity coordinates annotated with their respective spectral power distribution names.
- **chromaticity\_diagram\_callable\_CIE1960UCS** (callable, optional) – Callable responsible for drawing the *CIE 1960 UCS Chromaticity Diagram*.

### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> from colour import ILLUMINANTS_RELATIVE_SPDS
>>> A = ILLUMINANTS_RELATIVE_SPDS['A']
>>> D65 = ILLUMINANTS_RELATIVE_SPDS['D65']
>>> spds_chromaticity_diagram_plot_CIE1960UCS([A, D65])
```

## colour.plotting.spds\_chromaticity\_diagram\_plot\_CIE1976UCS

```
colour.plotting.spds_chromaticity_diagram_plot_CIE1976UCS(spds, cmfs='CIE 1931 2 Degree Standard Observer', annotate=True, chromaticity_diagram_callable_CIE1976UCS=<function chromaticity_diagram_plot_CIE1976UCS>, **kwargs)
```

Plots given spectral power distribution chromaticity coordinates into the *CIE 1976 UCS Chromaticity Diagram*.

### Parameters

- **spds** (array\_like, optional) – Spectral power distributions to plot.
- **cmfs** (unicode, optional) – Standard observer colour matching functions used for diagram bounds.
- **annotate** (bool) – Should resulting chromaticity coordinates annotated with their respective spectral power distribution names.
- **chromaticity\_diagram\_callable\_CIE1976UCS** (callable, optional) – Callable responsible for drawing the *CIE 1976 UCS Chromaticity Diagram*.

### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1976UCS()`}, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1976UCS()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> from colour import ILLUMINANTS_RELATIVE_SPDS
>>> A = ILLUMINANTS_RELATIVE_SPDS['A']
>>> D65 = ILLUMINANTS_RELATIVE_SPDS['D65']
>>> spds_chromaticity_diagram_plot_CIE1976UCS([A, D65])
```

## Colour Models

### colour.plotting

---

`RGB_colourspaces_chromaticity_diagram_plot_CIE1931` Plots given RGB colourespaces in *CIE 1931 Chromaticity Diagram*.

---

Continued on next page

Table 3.178 – continued from previous page

<code>RGB_colourspaces_chromaticity_diagram_plot_CIE196</code>	Plots given RGB colourspaces in CIE 1960 UCS Chromaticity Diagram.
<code>RGB_colourspaces_chromaticity_diagram_plot_CIE197</code>	Plots given RGB colourspaces in CIE 1976 UCS Chromaticity Diagram.
<code>RGB_chromaticity_coordinates_chromaticity_diagram</code>	Plots given RGB colourspace array in CIE 1931 Chromaticity Diagram.
<code>RGB_chromaticity_coordinates_chromaticity_diagram</code>	Plots given RGB colourspace array in CIE 1960 UCS Chromaticity Diagram.
<code>RGB_chromaticity_coordinates_chromaticity_diagram</code>	Plots given RGB colourspace array in CIE 1976 UCS Chromaticity Diagram.
<code>single_cctf_plot([colourspace, decoding_cctf])</code>	Plots given colourspace colour component transfer function.
<code>multi_cctf_plot([colourspaces, decoding_cctf])</code>	Plots given colourspaces colour component transfer functions.

`colour.plotting.RGB_colourspaces_chromaticity_diagram_plot_CIE1931`

```
colour.plotting.RGB_colourspaces_chromaticity_diagram_plot_CIE1931(colourspaces=None,
                                                               cmfs='CIE 1931 2
Degree Standard Observer',
                                                               chromaticity_diagram_callable_CIE1931=<function
                                                               chromaticity_diagram_plot_CIE1931>,
                                                               **kwargs)
```

Plots given RGB colourspaces in CIE 1931 Chromaticity Diagram.

#### Parameters

- `colourspaces` (array\_like, optional) – RGB colourspaces to plot.
- `cmfs` (unicode, optional) – Standard observer colour matching functions used for diagram bounds.
- `chromaticity_diagram_callable_CIE1931` (callable, optional) – Callable responsible for drawing the CIE 1931 Chromaticity Diagram.

#### Other Parameters

- `**kwargs` (dict, optional) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.
- `show_diagram_colours` (bool, optional) – `{colour.plotting.chromaticity_diagram_plot_CIE1931()}`, Whether to display the chromaticity diagram background colours.
- `use_cached_diagram_colours` (bool, optional) – `{colour.plotting.chromaticity_diagram_plot_CIE1931()}`, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> c = ['ITU-R Rec. 709', 'ACEScg', 'S-Gamut']
>>> RGB_colourspaces_chromaticity_diagram_plot_CIE1931(c)
...

```

## colour.plotting.RGB\_colourspaces\_chromaticity\_diagram\_plot\_CIE1960UCS

```
colour.plotting.RGB_colourspaces_chromaticity_diagram_plot_CIE1960UCS(colourspace=None,
                                                                     cmfs='CIE 1931 2
Degree Standard
Observer', chromaticity_diagram_callable_CIE1960UCS=<function chromaticity_diagram_plot_CIE1960UCS>,
**kwargs)
```

Plots given *RGB* colourespaces in *CIE 1960 UCS Chromaticity Diagram*.

### Parameters

- **colourspace** (array\_like, optional) – *RGB* colourespaces to plot.
- **cmfs** (unicode, optional) – Standard observer colour matching functions used for diagram bounds.
- **chromaticity\_diagram\_callable\_CIE1960UCS** (callable, optional) – Callable responsible for drawing the *CIE 1960 UCS Chromaticity Diagram*.

### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> c = ['ITU-R Rec. 709', 'ACEScg', 'S-Gamut']
>>> RGB_colourspaces_chromaticity_diagram_plot_CIE1960UCS(c)
...

```

**colour.plotting.RGB\_colourspace\_chromaticity\_diagram\_plot\_CIE1976UCS**

```
colour.plotting.RGB_colourspace_chromaticity_diagram_plot_CIE1976UCS(colourspace=None,
                                                               cmfs='CIE 1931 2
Degree Standard
Observer', chromatic-
ity_diagram_callable_CIE1976UCS=<function
chromatic-
ity_diagram_plot_CIE1976UCS>,
**kwargs)
```

Plots given *RGB* colourspaces in *CIE 1976 UCS Chromaticity Diagram*.

**Parameters**

- **colourspace** (`array_like`, optional) – *RGB* colourspaces to plot.
- **cmfs** (`unicode`, optional) – Standard observer colour matching functions used for diagram bounds.
- **chromaticity\_diagram\_callable\_CIE1976UCS** (`callable`, optional) – Callable responsible for drawing the *CIE 1976 UCS Chromaticity Diagram*.

**Other Parameters**

- **\*\*kwargs** (`dict`, optional) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (`bool`, optional) – `{colour.plotting.chromaticity_diagram_plot_CIE1976UCS()}`, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (`bool`, optional) – `{colour.plotting.chromaticity_diagram_plot_CIE1976UCS()}`, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

**Examples**

```
>>> c = ['ITU-R Rec. 709', 'ACEScg', 'S-Gamut']
>>> RGB_colourspace_chromaticity_diagram_plot_CIE1976UCS(c)
... 
```

**colour.plotting.RGB\_chromaticity\_coordinates\_chromaticity\_diagram\_plot\_CIE1931**

```
colour.plotting.RGB_chromaticity_coordinates_chromaticity_diagram_plot_CIE1931(RGB,
                                                               colourspace='sRGB',
                                                               chromatic-
                                                               ity_diagram_callable_CIE1931=<function
                                                               RGB_colourspace_chromaticity_di-
                                                               map_callable_CIE1931>,
                                                               **kwargs)
```

Plots given *RGB* colourspace array in *CIE 1931 Chromaticity Diagram*.

**Parameters**

- **RGB** (array\_like) – *RGB* colourspace array.
- **colourspace** (optional, unicode) – *RGB* colourspace of the *RGB* array.
- **chromaticity\_diagram\_callable\_CIE1931** (callable, optional) – Callable responsible for drawing the *CIE 1931 Chromaticity Diagram*.

#### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1931()`}, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1931()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

#### Examples

```
>>> RGB = np.random.random((10, 10, 3))
>>> c = 'ITU-R Rec. 709'
>>> RGB_chromaticity_coordinates_chromaticity_diagram_plot_CIE1931(RGB, c)
... 
```

## colour.plotting.RGB\_chromaticity\_coordinates\_chromaticity\_diagram\_plot\_CIE1960UCS

```
colour.plotting.RGB_chromaticity_coordinates_chromaticity_diagram_plot_CIE1960UCS(RGB,
    colourspace='sRGB',
    chro-
    matic-
    ity_diagram_callable_CIE1960-
    RGB_colourspaces_chromaticity
    **kwargs)
```

Plots given *RGB* colourspace array in *CIE 1960 UCS Chromaticity Diagram*.

#### Parameters

- **RGB** (array\_like) – *RGB* colourspace array.
- **colourspace** (optional, unicode) – *RGB* colourspace of the *RGB* array.
- **chromaticity\_diagram\_callable\_CIE1960UCS** (callable, optional) – Callable responsible for drawing the *CIE 1960 UCS Chromaticity Diagram*.

#### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to display the chromaticity diagram background colours.

- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> RGB = np.random.random((10, 10, 3))
>>> c = 'ITU-R BT.709'
>>> RGB_chromaticity_coordinates_chromaticity_diagram_plot_CIE1960UCS(
...     RGB, c)
```

## colour.plotting.RGB\_chromaticity\_coordinates\_chromaticity\_diagram\_plot\_CIE1976UCS

```
colour.plotting.RGB_chromaticity_coordinates_chromaticity_diagram_plot_CIE1976UCS(RGB,
    colourspace='sRGB',
    chro-
    matic-
    ity_diagram_callable_CIE1976UCS(
        RGB_colourspaces_chromaticity
        **kwargs)
```

Plots given *RGB* colourspace array in *CIE 1976 UCS Chromaticity Diagram*.

### Parameters

- **RGB** (array\_like) – *RGB* colourspace array.
- **colourspace** (optional, unicode) – *RGB* colourspace of the *RGB* array.
- **chromaticity\_diagram\_callable\_CIE1976UCS** (callable, optional) – Callable responsible for drawing the *CIE 1976 UCS Chromaticity Diagram*.

### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1976UCS()`}, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1976UCS()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> RGB = np.random.random((10, 10, 3))
>>> c = 'ITU-R BT.709'
>>> RGB_chromaticity_coordinates_chromaticity_diagram_plot_CIE1976UCS(
...     RGB, c)
```

## colour.plotting.single\_cctf\_plot

colour.plotting.**single\_cctf\_plot**(colourspace='ITU-R BT.709', decoding\_cctf=False, \*\*kwargs)  
Plots given colourspace colour component transfer function.

### Parameters

- **colourspace** (unicode, optional) – RGB Colourspace colour component transfer function to plot.
- **decoding\_cctf** (bool) – Plot decoding colour component transfer function instead.

**Other Parameters** **\*\*kwargs** (dict, optional) – {[colour.plotting.render\(\)](#)}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> single_cctf_plot()
```

## colour.plotting.multi\_cctf\_plot

colour.plotting.**multi\_cctf\_plot**(colourspaces=None, decoding\_cctf=False, \*\*kwargs)  
Plots given colourespaces colour component transfer functions.

### Parameters

- **colourspaces** (array\_like, optional) – Colourspaces colour component transfer function to plot.
- **decoding\_cctf** (bool) – Plot decoding colour component transfer function instead.

**Other Parameters** **\*\*kwargs** (dict, optional) – {[colour.plotting.render\(\)](#)}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> multi_cctf_plot(['ITU-R BT.709', 'sRGB'])
```

## Colour Notation Systems

`colour.plotting`

<code>single_munsell_value_function_plot([function])</code>	Plots given <i>Lightness</i> function.
<code>multi_munsell_value_function_plot([functions])</code>	Plots given <i>Munsell</i> value functions.

### `colour.plotting.single_munsell_value_function_plot`

`colour.plotting.single_munsell_value_function_plot(function='ASTM D1535-08', **kwargs)`  
Plots given *Lightness* function.

**Parameters** `function` (`unicode`, optional) – *Munsell* value function to plot.

**Other Parameters** `**kwargs` (`dict`, optional) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

#### Examples

```
>>> f = 'ASTM D1535-08'
>>> single_munsell_value_function_plot(f)
```

### `colour.plotting.multi_munsell_value_function_plot`

`colour.plotting.multi_munsell_value_function_plot(functions=None, **kwargs)`  
Plots given *Munsell* value functions.

**Parameters** `functions` (`array_like`, optional) – *Munsell* value functions to plot.

**Other Parameters** `**kwargs` (`dict`, optional) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

**Raises** `KeyError` – If one of the given *Munsell* value function is not found in the factory *Munsell* value functions.

#### Examples

```
>>> fs = ('ASTM D1535-08', 'McCamy 1987')
>>> multi_munsell_value_function_plot(fs)
```

## Optical Phenomena

colour.plotting

single_rayleigh_scattering_spd_plot([...])	Plots a single Rayleigh scattering spectral power distribution.
the_blue_sky_plot([cmfs])	Plots the blue sky.

colour.plotting.single\_rayleigh\_scattering\_spd\_plot

colour.plotting.**single\_rayleigh\_scattering\_spd\_plot**(CO2\_concentration=300, temperature=288.15, pressure=101325, latitude=0, altitude=0, cmfs='CIE 1931 2 Degree Standard Observer', \*\*kwargs)

Plots a single Rayleigh scattering spectral power distribution.

### Parameters

- **CO2\_concentration** (numeric, optional) –  $CO_2$  concentration in parts per million (ppm).
- **temperature** (numeric, optional) – Air temperature  $T[K]$  in kelvin degrees.
- **pressure** (numeric) – Surface pressure  $P$  of the measurement site.
- **latitude** (numeric, optional) – Latitude of the site in degrees.
- **altitude** (numeric, optional) – Altitude of the site in meters.
- **cmfs** (unicode, optional) – Standard observer colour matching functions.

### Other Parameters

- **\*\*kwargs** (dict, optional) – {colour.plotting.render()}, Please refer to the documentation of the previously listed definition.
- **out\_of\_gamut\_clipping** (bool, optional) – {colour.plotting.single\_spd\_plot()}, Whether to clip out of gamut colours otherwise, the colours will be offset by the absolute minimal colour leading to a rendering on gray background, less saturated and smoother.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> single_rayleigh_scattering_spd_plot()
```

colour.plotting.the\_blue\_sky\_plot

colour.plotting.**the\_blue\_sky\_plot**(cmfs='CIE 1931 2 Degree Standard Observer', \*\*kwargs)

Plots the blue sky.

**Parameters** **cmfs** (unicode, optional) – Standard observer colour matching functions.

**Other Parameters** `**kwargs` (*dict, optional*) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> the_blue_sky_plot()
```

## Colour Quality

### colour.plotting

<code>single_spd_colour_rendering_index_bars_plot(...)</code>	Plots the <i>Colour Rendering Index</i> (CRI) of given illuminant or light source spectral power distribution.
<code>multi_spd_colour_rendering_index_bars_plot(...)</code>	Plots the <i>Colour Rendering Index</i> (CRI) of given illuminants or light sources spectral power distributions.
<code>single_spd_colour_quality_scale_bars_plot(...)</code>	Plots the <i>Colour Quality Scale</i> (CQS) of given illuminant or light source spectral power distribution.
<code>multi_spd_colour_quality_scale_bars_plot(...)</code>	Plots the <i>Colour Quality Scale</i> (CQS) of given illuminants or light sources spectral power distributions.

### colour.plotting.single\_spd\_colour\_rendering\_index\_bars\_plot

#### colour.plotting.`single_spd_colour_rendering_index_bars_plot`(*spd*, `**kwargs`)

Plots the *Colour Rendering Index* (CRI) of given illuminant or light source spectral power distribution.

**Parameters** `spd` (`SpectralPowerDistribution`) – Illuminant or light source spectral power distribution to plot the *Colour Rendering Index* (CRI).

#### Other Parameters

- `**kwargs` (*dict, optional*) – `{colour.plotting.render()}`, Please refer to the documentation of the previously listed definition.
- `labels` (*bool, optional*) – `{colour.plotting.quality.colour_quality_bars_plot()}`, Add labels above bars.
- `hatching` (*bool or None, optional*) – `{colour.plotting.quality.colour_quality_bars_plot()}`, Use hatching for the bars.
- `hatching_repeat` (*int, optional*) – `{colour.plotting.quality.colour_quality_bars_plot()}`, Hatching pattern repeat.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> from colour import ILLUMINANTS_RELATIVE_SPDS
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['F2']
>>> single_spd_colour_rendering_index_bars_plot(illuminant)
...

```

### colour.plotting.multi\_spd\_colour\_rendering\_index\_bars\_plot

colour.plotting.**multi\_spd\_colour\_rendering\_index\_bars\_plot**(*spds*, \*\**kwargs*)

Plots the *Colour Rendering Index* (CRI) of given illuminants or light sources spectral power distributions.

**Parameters** **spds** (array\_like) – Array of illuminants or light sources spectral power distributions to plot the *Colour Rendering Index* (CRI).

#### Other Parameters

- **\*\*kwargs** (dict, optional) – {colour.plotting.render()}, Please refer to the documentation of the previously listed definition.
- **labels** (bool, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Add labels above bars.
- **hatching** (bool or None, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Use hatching for the bars.
- **hatching\_repeat** (int, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Hatching pattern repeat.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> from colour import (ILLUMINANTS_RELATIVE_SPDS,
...                      LIGHT_SOURCES_RELATIVE_SPDS)
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['F2']
>>> light_source = LIGHT_SOURCES_RELATIVE_SPDS['Kinoton 75P']
>>> multi_spd_colour_rendering_index_bars_plot([illuminant, light_source])
...

```

### colour.plotting.single\_spd\_colour\_quality\_scale\_bars\_plot

colour.plotting.**single\_spd\_colour\_quality\_scale\_bars\_plot**(*spd*, \*\**kwargs*)

Plots the *Colour Quality Scale* (CQS) of given illuminant or light source spectral power distribution.

**Parameters** **spd** (SpectralPowerDistribution) – Illuminant or light source spectral power distribution to plot the *Colour Quality Scale* (CQS).

#### Other Parameters

- **\*\*kwargs** (dict, optional) – {colour.plotting.render()}, Please refer to the documentation of the previously listed definition.

- **labels** (bool, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Add labels above bars.
- **hatching** (bool or None, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Use hatching for the bars.
- **hatching\_repeat** (int, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Hatching pattern repeat.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> from colour import ILLUMINANTS_RELATIVE_SPDS
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['F2']
>>> single_spd_colour_quality_scale_bars_plot(illuminant)
... 
```

## colour.plotting.multi\_spd\_colour\_quality\_scale\_bars\_plot

colour.plotting.**multi\_spd\_colour\_quality\_scale\_bars\_plot**(spds, \*\*kwargs)

Plots the *Colour Quality Scale* (CQS) of given illuminants or light sources spectral power distributions.

**Parameters** **spds** (array\_like) – Array of illuminants or light sources spectral power distributions to plot the *Colour Quality Scale* (CQS).

### Other Parameters

- **\*\*kwargs** (dict, optional) – {colour.plotting.render()}, Please refer to the documentation of the previously listed definition.
- **labels** (bool, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Add labels above bars.
- **hatching** (bool or None, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Use hatching for the bars.
- **hatching\_repeat** (int, optional) – {colour.plotting.quality.colour\_quality\_bars\_plot()}, Hatching pattern repeat.

**Returns** Current figure or None.

**Return type** Figure

## Examples

```
>>> from colour import (ILLUMINANTS_RELATIVE_SPDS,
...                      LIGHT_SOURCES_RELATIVE_SPDS)
>>> illuminant = ILLUMINANTS_RELATIVE_SPDS['F2']
>>> light_source = LIGHT_SOURCES_RELATIVE_SPDS['Kinoton 75P']
>>> multi_spd_colour_quality_scale_bars_plot([illuminant, light_source])
... 
```

## Colour Temperature & Correlated Colour Temperature

colour.plotting

`planckian_locus_chromaticity_diagram_plot_CIE1931` [Plots] the planckian locus and given illuminants in *CIE 1931 Chromaticity Diagram*.

`planckian_locus_chromaticity_diagram_plot_CIE1960` [Plots] the planckian locus and given illuminants in *CIE 1960 UCS Chromaticity Diagram*.

---

### colour.plotting.planckian\_locus\_chromaticity\_diagram\_plot\_CIE1931

colour.plotting.`planckian_locus_chromaticity_diagram_plot_CIE1931`(`illuminants=None,`  
`chromatic-`  
`ity_diagram_callable_CIE1931=<function`  
`chromatic-`  
`ity_diagram_plot_CIE1931>,`  
`**kwargs)`

Plots the planckian locus and given illuminants in *CIE 1931 Chromaticity Diagram*.

#### Parameters

- `illuminants` (array\_like, optional) – Factory illuminants to plot.
- `chromaticity_diagram_callable_CIE1931` (callable, optional) – Callable responsible for drawing the *CIE 1931 Chromaticity Diagram*.

#### Other Parameters

- `**kwargs` (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- `show_diagram_colours` (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1931()`}, Whether to display the chromaticity diagram background colours.
- `use_cached_diagram_colours` (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1931()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

**Raises** `KeyError` – If one of the given illuminant is not found in the factory illuminants.

#### Examples

```
>>> planckian_locus_chromaticity_diagram_plot_CIE1931(['A', 'B', 'C'])
...
```

## colour.plotting.planckian\_locus\_chromaticity\_diagram\_plot\_CIE1960UCS

```
colour.plotting.planckian_locus_chromaticity_diagram_plot_CIE1960UCS(illuminants=None,
                                                               chromatic-
                                                               ity_diagram_callable_CIE1960UCS=<function
                                                               chromatic-
                                                               ity_diagram_plot_CIE1960UCS>,
                                                               **kwargs)
```

Plots the planckian locus and given illuminants in *CIE 1960 UCS Chromaticity Diagram*.

### Parameters

- **illuminants** (array\_like, optional) – Factory illuminants to plot.
- **chromaticity\_diagram\_callable\_CIE1960UCS** (callable, optional) – Callable responsible for drawing the *CIE 1960 UCS Chromaticity Diagram*.

### Other Parameters

- **\*\*kwargs** (dict, optional) – {`colour.plotting.render()`}, Please refer to the documentation of the previously listed definition.
- **show\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to display the chromaticity diagram background colours.
- **use\_cached\_diagram\_colours** (bool, optional) – {`colour.plotting.chromaticity_diagram_plot_CIE1960UCS()`}, Whether to used the cached chromaticity diagram background colours image.

**Returns** Current figure or None.

**Return type** Figure

**Raises** `KeyError` – If one of the given illuminant is not found in the factory illuminants.

### Examples

```
>>> planckian_locus_chromaticity_diagram_plot_CIE1960UCS(['A', 'C', 'E'])
...
```

## Colour Models Volume

### colour.plotting

<code>RGB_colourspaces_gamuts_plot([colourspaces, ...])</code>	Plots given <i>RGB</i> colourspaces gamuts in given reference colourspace.
<code>RGB_scatter_plot(RGB, colourspace[, ...])</code>	Plots given <i>RGB</i> colourspace array in a scatter plot.

## colour.plotting.RGB\_colourspaces\_gamuts\_plot

```
colour.plotting.RGB_colourspaces_gamuts_plot(colourspaces=None, reference_colourspace='CIE xyY', segments=8, display_grid=True, grid_segments=10, spectral_locus=False, spectral_locus_colour=None, cmfs='CIE 1931 2 Degree Standard Observer', **kwargs)
```

Plots given *RGB* colourspaces gamuts in given reference colourspace.

### Parameters

- **colourspaces** (`array_like`, optional) – *RGB* colourspaces to plot the gamuts.
- **reference\_colourspace** (`unicode`, optional) – {‘CIE XYZ’, ‘CIE xyY’, ‘CIE Lab’, ‘CIE Luv’, ‘CIE UCS’, ‘CIE UVW’, ‘IPT’, ‘Hunter Lab’, ‘Hunter Rdab’}, Reference colourspace to plot the gamuts into.
- **segments** (`int`, optional) – Edge segments count for each *RGB* colourspace cubes.
- **display\_grid** (`bool`, optional) – Display a grid at the bottom of the *RGB* colourspace cubes.
- **grid\_segments** (`bool`, optional) – Edge segments count for the grid.
- **spectral\_locus** (`bool`, optional) – Is spectral locus line plotted.
- **spectral\_locus\_colour** (`array_like`, optional) – Spectral locus line colour.
- **cmfs** (`unicode`, optional) – Standard observer colour matching functions used for spectral locus.

### Other Parameters

- **\*\*kwargs** (`dict`, optional) – {`colour.plotting.volume.nadir_grid()`}, Please refer to the documentation of the previously listed definition.
- **face\_colours** (`array_like`, optional) – Face colours array such as `face_colours = (None, (0.5, 0.5, 1.0))`.
- **edge\_colours** (`array_like`, optional) – Edge colours array such as `edge_colours = (None, (0.5, 0.5, 1.0))`.
- **face\_alpha** (`numeric`, optional) – Face opacity value such as `face_alpha = (0.5, 1.0)`.
- **edge\_alpha** (`numeric`, optional) – Edge opacity value such as `edge_alpha = (0.0, 1.0)`.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> c = ['ITU-R BT.709', 'ACEScg', 'S-Gamut']
>>> RGB_colourspaces_gamuts_plot(c)
```

## colour.plotting.RGB\_scatter\_plot

```
colour.plotting.RGB_scatter_plot(RGB,      colourspace,      reference_colourspace='CIE xyY',
                                  colourspaces=None,    segments=8,        display_grid=True,
                                  grid_segments=10,     spectral_locus=False,   spec-
                                  tral_locus_colour=None, points_size=12, cmfs='CIE 1931 2
                                  Degree Standard Observer', **kwargs)
```

Plots given *RGB* colourspace array in a scatter plot.

### Parameters

- **RGB** (array\_like) – *RGB* colourspace array.
- **colourspace** (RGB\_Colourspace) – *RGB* colourspace of the *RGB* array.
- **reference\_colourspace** (unicode, optional) – {‘CIE XYZ’, ‘CIE xyY’, ‘CIE Lab’, ‘CIE Luv’, ‘CIE UCS’, ‘CIE UVW’, ‘IPT’, ‘Hunter Lab’, ‘Hunter Rdab’}, Reference colourspace for colour conversion.
- **colourspaces** (array\_like, optional) – *RGB* colourespaces to plot the gamuts.
- **segments** (int, optional) – Edge segments count for each *RGB* colourspace cubes.
- **display\_grid** (bool, optional) – Display a grid at the bottom of the *RGB* colourspace cubes.
- **grid\_segments** (bool, optional) – Edge segments count for the grid.
- **spectral\_locus** (bool, optional) – Is spectral locus line plotted.
- **spectral\_locus\_colour** (array\_like, optional) – Spectral locus line colour.
- **points\_size** (numeric, optional) – Scatter points size.
- **cmfs** (unicode, optional) – Standard observer colour matching functions used for spectral locus.

**Other Parameters** **\*\*kwargs** (dict, optional) – {`colour.plotting.RGB_colourspace_gamuts_plot()`}, Please refer to the documentation of the previously listed definition.

**Returns** Current figure or None.

**Return type** Figure

### Examples

```
>>> c = 'ITU-R BT.709'
>>> RGB_scatter_plot(c)
```

## Geometry Plotting Utilities

### colour.plotting

<code>quad([plane, origin, width, height, depth])</code>	Returns the vertices of a quad geometric element in counter-clockwise order.
--	--

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Table 3.184 – continued from previous page

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<code>cube([plane, origin, width, height, depth, ...])</code>	Returns the vertices of a cube made of grids.
---	---

---

## colour.plotting.quad

`colour.plotting.quad(plane='xy', origin=None, width=1, height=1, depth=0)`

Returns the vertices of a quad geometric element in counter-clockwise order.

### Parameters

- **plane** (array\_like, optional) – {'xy', 'xz', 'yz'}, Construction plane of the quad.
- **origin** (array\_like, optional) – Quad origin on the construction plane.
- **width** (numeric, optional) – Quad width.
- **height** (numeric, optional) – Quad height.
- **depth** (numeric, optional) – Quad depth.

**Returns** Quad vertices.

**Return type** ndarray

### Examples

```
>>> quad()
array([[0, 0, 0],
       [1, 0, 0],
       [1, 1, 0],
       [0, 1, 0]])
```

## colour.plotting.grid

`colour.plotting.grid(plane='xy', origin=None, width=1, height=1, depth=0, width_segments=1, height_segments=1)`

Returns the vertices of a grid made of quads.

### Parameters

- **plane** (array\_like, optional) – {'xy', 'xz', 'yz'}, Construction plane of the grid.
- **origin** (array\_like, optional) – Grid origin on the construction plane.
- **width** (numeric, optional) – Grid width.
- **height** (numeric, optional) – Grid height.
- **depth** (numeric, optional) – Grid depth.
- **width\_segments** (int, optional) – Grid segments, quad counts along the width.
- **height\_segments** (int, optional) – Grid segments, quad counts along the height.

**Returns** Grid vertices.

**Return type** ndarray

## Examples

```
>>> grid(width_segments=2, height_segments=2)
array([[[ 0. ,  0. ,  0. ],
       [ 0.5,  0. ,  0. ],
       [ 0.5,  0.5,  0. ],
       [ 0. ,  0.5,  0. ]],

      [[ 0. ,  0.5,  0. ],
       [ 0.5,  0.5,  0. ],
       [ 0.5,  1. ,  0. ],
       [ 0. ,  1. ,  0. ]],

      [[ 0.5,  0. ,  0. ],
       [ 1. ,  0. ,  0. ],
       [ 1. ,  0.5,  0. ],
       [ 0.5,  0.5,  0. ]],

      [[ 0.5,  0.5,  0. ],
       [ 1. ,  0.5,  0. ],
       [ 1. ,  1. ,  0. ],
       [ 0.5,  1. ,  0. ]]])
```

## colour.plotting.cube

`colour.plotting.cube(plane=None, origin=None, width=1, height=1, depth=1, width_segments=1, height_segments=1, depth_segments=1)`

Returns the vertices of a cube made of grids.

### Parameters

- **plane** (array\_like, optional) – Any combination of {‘+x’, ‘-x’, ‘+y’, ‘-y’, ‘+z’, ‘-z’}, Included grids in the cube construction.
- **origin** (array\_like, optional) – Cube origin.
- **width** (numeric, optional) – Cube width.
- **height** (numeric, optional) – Cube height.
- **depth** (numeric, optional) – Cube depth.
- **width\_segments** (int, optional) – Cube segments, quad counts along the width.
- **height\_segments** (int, optional) – Cube segments, quad counts along the height.
- **depth\_segments** (int, optional) – Cube segments, quad counts along the depth.

**Returns** Cube vertices.

**Return type** ndarray

## Examples

```
>>> cube()
array([[[ 0.,  0.,  0.],
       [ 1.,  0.,  0.],
       [ 1.,  1.,  0.],
```

```
[ 0.,  1.,  0.]],  
[[ 0.,  0.,  1.],  
 [ 1.,  0.,  1.],  
 [ 1.,  1.,  1.],  
 [ 0.,  1.,  1.]],  
  
[[ 0.,  0.,  0.],  
 [ 1.,  0.,  0.],  
 [ 1.,  0.,  1.],  
 [ 0.,  0.,  1.]],  
  
[[ 0.,  1.,  0.],  
 [ 1.,  1.,  0.],  
 [ 1.,  1.,  1.],  
 [ 0.,  1.,  1.]],  
  
[[ 0.,  0.,  0.],  
 [ 0.,  1.,  0.],  
 [ 0.,  1.,  1.],  
 [ 0.,  0.,  1.]],  
  
[[ 1.,  0.,  0.],  
 [ 1.,  1.,  0.],  
 [ 1.,  1.,  1.],  
 [ 1.,  0.,  1.]])
```

## Colour Quality

- *Colour Rendering Index*
- *Colour Quality Scale*

### Colour Rendering Index

colour

`colour.colour_rendering_index(spd_test[, ...])`

Returns the *Colour Rendering Index* (CRI)  $Q_a$  of given spectral power distribution.

### `colour.colour_rendering_index`

`colour.colour_rendering_index(spd_test, additional_data=False)`

Returns the *Colour Rendering Index* (CRI)  $Q_a$  of given spectral power distribution.

#### Parameters

- `spd_test` (`SpectralPowerDistribution`) – Test spectral power distribution.
- `additional_data` (`bool`, optional) – Output additional data.

**Returns** Colour Rendering Index (CRI).

**Return type** numeric or [CRI\\_Specification](#)

## References

- [\[OD08\]](#)

## Examples

```
>>> from colour import ILLUMINANTS_RELATIVE_SPDS
>>> spd = ILLUMINANTS_RELATIVE_SPDS['F2']
>>> colour_rendering_index(spd)
64.1515202...
```

colour.quality

[CRI\\_Specification](#)

Defines the *Colour Rendering Index (CRI)* colour quality specification.

colour.quality.CRI\_Specification

**class** colour.quality.CRI\_Specification

Defines the *Colour Rendering Index (CRI)* colour quality specification.

### Parameters

- **name** (unicode) – Name of the test spectral power distribution.
- **Q\_a** (numeric) – *Colour Rendering Index (CRI)*  $Q_a$ .
- **Q\_as** ([dict](#)) – Individual *colour rendering indexes* data for each sample.
- **colorimetry\_data** ([tuple](#)) – Colorimetry data for the test and reference computations.

## References

- [\[OD08\]](#)

Create new instance of CRI\_Specification(name, Q\_a, Q\_as, colorimetry\_data)

**\_\_init\_\_()**

x.\_\_init\_\_(...) initializes x; see help(type(x)) for signature

## Methods

---

[count\(...\)](#)

[index\(\(value, \[start, ...\]\)](#)

Raises ValueError if the value is not present.

## Colour Quality Scale

colour

---

colour.colour\_quality\_scale(spd\_test[, additional\_data])      Returns the *Colour Quality Scale* (CQS) of given spectral power distribution.

---

### colour.colour\_quality\_scale

colour.colour\_quality\_scale(spd\_test, additional\_data=False)

Returns the *Colour Quality Scale* (CQS) of given spectral power distribution.

#### Parameters

- **spd\_test** (`SpectralPowerDistribution`) – Test spectral power distribution.
- **additional\_data** (`bool`, optional) – Output additional data.

**Returns** Color quality scale.

**Return type** numeric or `CQS_Specification`

## References

- [\[DO10\]](#)
- [\[OD08\]](#)

## Examples

```
>>> from colour import ILLUMINANTS_RELATIVE_SPDS
>>> spd = ILLUMINANTS_RELATIVE_SPDS['F2']
>>> colour.colour_quality_scale(spd)
64.6863391...
```

colour.quality

---

`CQS_Specification`

Defines the *Colour Quality Scale* (CQS) colour quality specification.

---

### colour.quality.CQS\_Specification

**class** colour.quality.CQS\_Specification

Defines the *Colour Quality Scale* (CQS) colour quality specification.

#### Parameters

- **name** (`unicode`) – Name of the test spectral power distribution.
- **Q\_a** (`numeric`) – Colour quality scale  $Q_a$ .
- **Q\_f** (`numeric`) – Colour fidelity scale  $Q_f$  intended to evaluate the fidelity of object colour appearances (compared to the reference illuminant of the same correlated

colour temperature and illuminance).

- `Q_p` (numeric) – Colour preference scale  $Q_p$  similar to colour quality scale  $Q_a$  but placing additional weight on preference of object colour appearance. This metric is based on the notion that increases in chroma are generally preferred and should be rewarded.
- `Q_g` (numeric) – Gamut area scale  $Q_g$  representing the relative gamut formed by the  $(a^*, b^*)$  coordinates of the 15 samples illuminated by the test light source in the *CIE L\*a\*b\** object colourspace.
- `Q_d` (numeric) – Relative gamut area scale  $Q_d$ .
- `Q_as` (dict) – Individual *Colour Quality Scale* (CQS) data for each sample.
- `colorimetry_data` (tuple) – Colorimetry data for the test and reference computations.

## References

- [\[DO10\]](#)
- [\[OD08\]](#)

Create new instance of `CQS_Specification(name, Q_a, Q_f, Q_p, Q_g, Q_d, Q_as, colorimetry_data)`

`__init__()`  
`x.__init__(...)` initializes `x`; see `help(type(x))` for signature

## Methods

---

`count(...)`

`index((value, [start, ...])`

Raises `ValueError` if the value is not present.

---

## Reflectance Recovery

- *CIE XYZ Colourspace to Spectral*
  - *Smits (1999)*
  - *Meng, Simon and Hanika (2015)*

## CIE XYZ Colourspace to Spectral

`colour`

---

`XYZ_to_spectral(XYZ[, method])`

Recovers the spectral power distribution of given *CIE XYZ* tristimulus values using given method.

---

`REFLECTANCE_RECOVERY_METHODS`

Supported reflectance recovery methods.

---

## colour.XYZ\_to\_spectral

`colour.XYZ_to_spectral(XYZ, method='Meng 2015', **kwargs)`

Recovers the spectral power distribution of given CIE XYZ tristimulus values using given method.

### Parameters

- `XYZ` (array\_like) – CIE XYZ tristimulus values to recover the spectral power distribution from.
- `method` (unicode, optional) – {‘Meng 2015’, ‘Smits 1999’}, Computation method.

### Other Parameters

- `cmfs` (XYZ\_ColourMatchingFunctions) – {`colour.recovery.XYZ_to_spectral_Meng2015()`}, Standard observer colour matching functions.
- `interval` (numeric, optional) – {`colour.recovery.XYZ_to_spectral_Meng2015()`}, Wavelength  $\lambda_i$  range interval in nm. The smaller interval is, the longer the computations will be.
- `tolerance` (numeric, optional) – {`colour.recovery.XYZ_to_spectral_Meng2015()`}, Tolerance for termination. The lower tolerance is, the smoother the recovered spectral power distribution will be.
- `maximum_iterations` (int, optional) – {`colour.recovery.XYZ_to_spectral_Meng2015()`}, Maximum number of iterations to perform.

**Returns** Recovered spectral power distribution.

**Return type** `SpectralPowerDistribution`

## Notes

- Smits (1999) method will internally convert given CIE XYZ tristimulus values to RGB colourspace array assuming equal energy illuminant  $E$ .

## References

- [MSHD15]
- [Smi99]

## Examples

Meng (2015) reflectance recovery:

```
>>> from colour.utilities import numpy_print_options
>>> from colour.colorimetry import spectral_to_XYZ_integration
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])
>>> spd = XYZ_to_spectral(XYZ, interval=10)
>>> with numpy_print_options(suppress=True):
...     spd
SpectralPowerDistribution([[ 360.        ,      0.0788075...],
                           [ 370.        ,      0.0788543...],
                           [ 380.        ,      0.0788825...],
                           [ 390.        ,      0.0788714...],
```

```

[ 400.      ,  0.0788911...],
[ 410.      ,  0.07893   ...],
[ 420.      ,  0.0797471...],
[ 430.      ,  0.0813339...],
[ 440.      ,  0.0840145...],
[ 450.      ,  0.0892826...],
[ 460.      ,  0.0965359...],
[ 470.      ,  0.1053176...],
[ 480.      ,  0.1150921...],
[ 490.      ,  0.1244252...],
[ 500.      ,  0.1326083...],
[ 510.      ,  0.1390282...],
[ 520.      ,  0.1423548...],
[ 530.      ,  0.1414636...],
[ 540.      ,  0.1365195...],
[ 550.      ,  0.1277319...],
[ 560.      ,  0.1152622...],
[ 570.      ,  0.1004513...],
[ 580.      ,  0.0844187...],
[ 590.      ,  0.0686863...],
[ 600.      ,  0.0543013...],
[ 610.      ,  0.0423486...],
[ 620.      ,  0.0333861...],
[ 630.      ,  0.0273558...],
[ 640.      ,  0.0233407...],
[ 650.      ,  0.0211208...],
[ 660.      ,  0.0197248...],
[ 670.      ,  0.0187157...],
[ 680.      ,  0.0181510...],
[ 690.      ,  0.0179691...],
[ 700.      ,  0.0179247...],
[ 710.      ,  0.0178665...],
[ 720.      ,  0.0178005...],
[ 730.      ,  0.0177570...],
[ 740.      ,  0.0177090...],
[ 750.      ,  0.0175743...],
[ 760.      ,  0.0175058...],
[ 770.      ,  0.0174492...],
[ 780.      ,  0.0174984...],
[ 790.      ,  0.0175667...],
[ 800.      ,  0.0175657...],
[ 810.      ,  0.0175319...],
[ 820.      ,  0.0175184...],
[ 830.      ,  0.0175390...]],
interpolator=SpragueInterpolator,
interpolator_args={},
extrapolator=Extrapolator,
extrapolator_args={...})
>>> spectral_to_XYZ_integration(spd) / 100
array([ 0.0705100...,  0.1007987...,  0.0956738...])

```

*Smits (1999)* reflectance recovery:

```

>>> spd = XYZ_to_spectral(XYZ, method='Smits 1999')
>>> with numpy_print_options(suppress=True):
...     spd
SpectralPowerDistribution([[ 380.      ,  0.0908046...],
                           [ 417.7778  ,  0.0887761...],

```

```
[ 455.5556 , 0.0939795...],  
[ 493.3333 , 0.1236033...],  
[ 531.1111 , 0.1315788...],  
[ 568.8889 , 0.1293411...],  
[ 606.6667 , 0.0392680...],  
[ 644.4444 , 0.0214496...],  
[ 682.2222 , 0.0214496...],  
[ 720. , 0.0215462...]],  
interpolator=CubicSplineInterpolator,  
interpolator_args={},  
extrapolator=Extrapolator,  
extrapolator_args={...})  
>>> spectral_to_XYZ_integration(spd) / 100  
array([ 0.0753341..., 0.1054586..., 0.0977855...])
```

## colour.REFLECTANCE\_RECOVERY\_METHODS

```
colour.REFLECTANCE_RECOVERY_METHODS = CaseInsensitiveMapping({'Meng 2015': ... , 'Smits 1999': ...})  
Supported reflectance recovery methods.
```

### References

- [MSHD15]
- [Smi99]

**REFLECTANCE\_RECOVERY\_METHODS** [CaseInsensitiveMapping] {'Meng 2015', 'Smits 1999'}

## Smits (1999)

### colour.recovery

RGB_to_spectral_Smits1999(RGB)	Recovers the spectral power distribution of given <i>RGB</i> colourspace array using <i>Smits (1999)</i> method.
SMITS_1999_SPDS	<i>Smits (1999)</i> spectral power distributions.

### colour.recovery.RGB\_to\_spectral\_Smits1999

colour.recovery.RGB\_to\_spectral\_Smits1999(*RGB*)

Recovers the spectral power distribution of given *RGB* colourspace array using *Smits (1999)* method.

**Parameters** *RGB* (array\_like, (3,)) – *RGB* colourspace array to recover the spectral power distribution from.

**Returns** Recovered spectral power distribution.

**Return type** *SpectralPowerDistribution*

## References

- [Smi99]

## Examples

```
>>> from colour.utilities import numpy_print_options
>>> RGB = np.array([0.02144962, 0.13154603, 0.09287601])
>>> with numpy_print_options(suppress=True):
...     RGB_to_spectral_Smits1999(RGB)
SpectralPowerDistribution([[ 380.        ,  0.0908046...],
                           [ 417.7778    ,  0.0887761...],
                           [ 455.5556    ,  0.0939795...],
                           [ 493.3333    ,  0.1236033...],
                           [ 531.1111    ,  0.1315788...],
                           [ 568.8889    ,  0.1293411...],
                           [ 606.6667    ,  0.0392680...],
                           [ 644.4444    ,  0.0214496...],
                           [ 682.2222    ,  0.0214496...],
                           [ 720.        ,  0.0215463...]],
                           interpolator=CubicSplineInterpolator,
                           interpolator_args={},
                           extrapolator=Extrapolator,
                           extrapolator_args={...})
```

## colour.recovery.SMITS\_1999\_SPDS

colour.recovery.SMITS\_1999\_SPDS = CaseInsensitiveMapping({u'blue': ..., u'yellow': ..., u'green': ..., u'cyan': ...},  
*Smits (1999)* spectral power distributions.

## References

- [Smi99]

SMITS\_1999\_SPDS : CaseInsensitiveMapping

## Meng, Simon and Hanika (2015)

colour.recovery

XYZ\_to\_spectral\_Meng2015(XYZ[, cmfs, ...])

Recovers the spectral power distribution of given  
*CIE XYZ* tristimulus values using *Meng et alii (2015)*  
method.

## colour.recovery.XYZ\_to\_spectral\_Meng2015

```
colour.recovery.XYZ_to_spectral_Meng2015(XYZ, cmfs=XYZ_ColourMatchingFunctions(name='CIE  
1931 2 Degree Standard Observer', ...), interval=5,  
tolerance=1e-10, maximum_iterations=2000)
```

Recovers the spectral power distribution of given CIE XYZ tristimulus values using Meng et alii (2015) method.

### Parameters

- **XYZ** (array\_like, (3,)) – CIE XYZ tristimulus values to recover the spectral power distribution from.
- **cmfs** (XYZ\_ColourMatchingFunctions) – Standard observer colour matching functions.
- **interval** (numeric, optional) – Wavelength  $\lambda_i$  range interval in nm. The smaller interval is, the longer the computations will be.
- **tolerance** (numeric, optional) – Tolerance for termination. The lower tolerance is, the smoother the recovered spectral power distribution will be.
- **maximum\_iterations** (int, optional) – Maximum number of iterations to perform.

**Returns** Recovered spectral power distribution.

**Return type** SpectralPowerDistribution

### Notes

- The definition used to convert spectrum to CIE XYZ tristimulus values is colour.colorimetry.spectral\_to\_XYZ\_integration() definition because it processes any measurement interval opposed to colour.colorimetry.spectral\_to\_XYZ\_ASTME30815() definition that handles only measurement interval of 1, 5, 10 or 20nm.

### References

- [MSHD15]

### Examples

```
>>> from colour.utilities import numpy_print_options  
>>> XYZ = np.array([0.07049534, 0.10080000, 0.09558313])  
>>> spd = XYZ_to_spectral_Meng2015(XYZ, interval=10)  
>>> with numpy_print_options(suppress=True):  
...     spd  
SpectralPowerDistribution([[ 360.        ,  0.0788075...],  
                          [ 370.        ,  0.0788543...],  
                          [ 380.        ,  0.0788825...],  
                          [ 390.        ,  0.0788714...],  
                          [ 400.        ,  0.0788911...],  
                          [ 410.        ,  0.07893    ...],  
                          [ 420.        ,  0.0797471...],  
                          [ 430.        ,  0.0813339...],  
                          [ 440.        ,  0.0840145...],
```

```

[ 450.      ,  0.0892826...],
[ 460.      ,  0.0965359...],
[ 470.      ,  0.1053176...],
[ 480.      ,  0.1150921...],
[ 490.      ,  0.1244252...],
[ 500.      ,  0.1326083...],
[ 510.      ,  0.1390282...],
[ 520.      ,  0.1423548...],
[ 530.      ,  0.1414636...],
[ 540.      ,  0.1365195...],
[ 550.      ,  0.1277319...],
[ 560.      ,  0.1152622...],
[ 570.      ,  0.1004513...],
[ 580.      ,  0.0844187...],
[ 590.      ,  0.0686863...],
[ 600.      ,  0.0543013...],
[ 610.      ,  0.0423486...],
[ 620.      ,  0.0333861...],
[ 630.      ,  0.0273558...],
[ 640.      ,  0.0233407...],
[ 650.      ,  0.0211208...],
[ 660.      ,  0.0197248...],
[ 670.      ,  0.0187157...],
[ 680.      ,  0.0181510...],
[ 690.      ,  0.0179691...],
[ 700.      ,  0.0179247...],
[ 710.      ,  0.0178665...],
[ 720.      ,  0.0178005...],
[ 730.      ,  0.0177570...],
[ 740.      ,  0.0177090...],
[ 750.      ,  0.0175743...],
[ 760.      ,  0.0175058...],
[ 770.      ,  0.0174492...],
[ 780.      ,  0.0174984...],
[ 790.      ,  0.0175667...],
[ 800.      ,  0.0175657...],
[ 810.      ,  0.0175319...],
[ 820.      ,  0.0175184...],
[ 830.      ,  0.0175390...]],
interpolator=SpragueInterpolator,
interpolator_args={},
extrapolator=Extrapolator,
extrapolator_args={...})
>>> spectral_to_XYZ_integration(spd) / 100
array([ 0.0705100...,  0.1007987...,  0.0956738...])

```

## Colour Temperature

- *Correlated Colour Temperature*
  - Robertson (1968)
  - Krystek (1985)
  - Ohno (2013)

- *Hernandez-Andres, Lee and Romero (1999)*
- *Kang, Moon, Hong, Lee, Cho and Kim (2002)*
- *CIE Illuminant D Series*

## Correlated Colour Temperature

colour

<code>CCT_to_uv(CCT[, method])</code>	Returns the <i>CIE UCS</i> colourspace <i>uv</i> chromaticity coordinates from given correlated colour temperature $T_{cp}$ using given method.
<code>CCT_TO_UV_METHODS</code>	Supported correlated colour temperature $T_{cp}$ to <i>CIE UCS</i> colourspace <i>uv</i> chromaticity coordinates computation methods.
<code>uv_to_CCT(uv[, method])</code>	Returns the correlated colour temperature $T_{cp}$ and $\Delta_{uv}$ from given <i>CIE UCS</i> colourspace <i>uv</i> chromaticity coordinates using given method.
<code>UV_TO_CCT_METHODS</code>	Supported <i>CIE UCS</i> colourspace <i>uv</i> chromaticity coordinates to correlated colour temperature $T_{cp}$ computation methods.
<code>CCT_to_xy(CCT[, method])</code>	Returns the <i>CIE XYZ</i> tristimulus values <i>xy</i> chromaticity coordinates from given correlated colour temperature $T_{cp}$ using given method.
<code>CCT_TO_XY_METHODS</code>	Supported correlated colour temperature $T_{cp}$ to <i>CIE XYZ</i> tristimulus values <i>xy</i> chromaticity coordinates computation methods.
<code>xy_to_CCT(xy[, method])</code>	Returns the correlated colour temperature $T_{cp}$ from given <i>CIE XYZ</i> tristimulus values <i>xy</i> chromaticity coordinates using given method.
<code>XY_TO_CCT_METHODS</code>	Supported <i>CIE XYZ</i> tristimulus values <i>xy</i> chromaticity coordinates to correlated colour temperature $T_{cp}$ computation methods.

`colour.CCT_to_uv`

`colour.CCT_to_uv(CCT, method=u'Ohno 2013', **kwargs)`

Returns the *CIE UCS* colourspace *uv* chromaticity coordinates from given correlated colour temperature  $T_{cp}$  using given method.

### Parameters

- `CCT` (numeric) – Correlated colour temperature  $T_{cp}$ .
- `method` (unicode, optional) – {‘Ohno 2013’, ‘Robertson 1968’, ‘Krystek 1985’}, Computation method.

### Other Parameters

- `D_uv` (numeric) – {`CCT_to_uv_Ohno2013`, `CCT_to_uv_Robertson1968()`},  $\Delta_{uv}$ .
- `cmfs` (*XYZ\_ColourMatchingFunctions*, optional) – {`colour.temperature.CCT_to_uv_Ohno2013()`}, Standard observer colour matching functions.

**Returns** *CIE UCS* colourspace *uv* chromaticity coordinates.

**Return type** ndarray

## References

- [AdobeSystems13a]
- [AdobeSystems13b]
- [Kry85]
- [Ohn14]
- [WS00c]

## Examples

```
>>> from colour import STANDARD_OBSERVERS_CMFS
>>> cmfs = STANDARD_OBSERVERS_CMFS['CIE 1931 2 Degree Standard Observer']
>>> CCT = 6507.47380460
>>> D_uv = 0.00322335
>>> CCT_to_uv(CCT, D_uv=D_uv, cmfs=cmfs)
array([ 0.1977999...,  0.3121999...])
```

## colour.CCT\_TO\_UV\_METHODS

colour.CCT\_TO\_UV\_METHODS = CaseInsensitiveMapping({u'Ohno 2013': ... , u'Krystek 1985': ... , u'robertson1968': ...})  
Supported correlated colour temperature  $T_{cp}$  to *CIE UCS* colourspace *uv* chromaticity coordinates computation methods.

## References

- [AdobeSystems13a]
- [AdobeSystems13b]
- [Kry85]
- [Ohn14]
- [WS00c]

CCT\_TO\_UV\_METHODS [CaseInsensitiveMapping] {'Ohno 2013', 'Robertson 1968', 'Krystek 1985'}

Aliases:

- ‘ohno2013’: ‘Ohno 2013’
- ‘robertson1968’: ‘Robertson 1968’

## colour.uv\_to\_CCT

`colour.uv_to_CCT(uv, method=u'Ohno 2013', **kwargs)`

Returns the correlated colour temperature  $T_{cp}$  and  $\Delta_{uv}$  from given CIE UCS colourspace  $uv$  chromaticity coordinates using given method.

### Parameters

- `uv` (array\_like) – CIE UCS colourspace  $uv$  chromaticity coordinates.
- `method` (unicode, optional) – {‘Ohno 2013’, ‘Robertson 1968’}, Computation method.

### Other Parameters

- `cmfs` (XYZ\_ColourMatchingFunctions, optional) – {`colour.temperature.uv_to_CCT_Ohno2013()`}, Standard observer colour matching functions.
- `start` (numeric, optional) – {`colour.temperature.uv_to_CCT_Ohno2013()`}, Temperature range start in kelvins.
- `end` (numeric, optional) – {`colour.temperature.uv_to_CCT_Ohno2013()`}, Temperature range end in kelvins.
- `count` (int, optional) – {`colour.temperature.uv_to_CCT_Ohno2013()`}, Temperatures count in the planckian tables.
- `iterations` (int, optional) – {`colour.temperature.uv_to_CCT_Ohno2013()`}, Number of planckian tables to generate.

**Returns** Correlated colour temperature  $T_{cp}$ ,  $\Delta_{uv}$ .

**Return type** ndarray

## References

- [\[AdobeSystems13a\]](#)
- [\[AdobeSystems13b\]](#)
- [\[Ohn14\]](#)
- [\[WS00c\]](#)

## Examples

```
>>> from colour import STANDARD_OBSERVERS_CMFS
>>> cmfs = STANDARD_OBSERVERS_CMFS['CIE 1931 2 Degree Standard Observer']
>>> uv = np.array([0.1978, 0.3122])
>>> uv_to_CCT(uv, cmfs=cmfs)
array([ 6.5074738...e+03,  3.2233461...e-03])
```

## colour.UV\_TO\_CCT\_METHODS

`colour.UV_TO_CCT_METHODS = CaseInsensitiveMapping({u'Ohno 2013': ... , u'robertson1968': ... , u'ohno2013': ...})`

Supported CIE UCS colourspace  $uv$  chromaticity coordinates to correlated colour temperature  $T_{cp}$  computation methods.

## References

- [\[AdobeSystems13a\]](#)
- [\[AdobeSystems13b\]](#)
- [\[Ohn14\]](#)
- [\[WS00c\]](#)

**UV\_TO\_CCT\_METHODS** [CaseInsensitiveMapping] {‘Ohno 2013’, ‘Robertson 1968’}

Aliases:

- ‘ohno2013’: ‘Ohno 2013’
- ‘robertson1968’: ‘Robertson 1968’

## colour.CCT\_to\_xy

`colour.CCT_to_xy(CCT, method=u'Kang 2002')`

Returns the *CIE XYZ* tristimulus values *xy* chromaticity coordinates from given correlated colour temperature  $T_{cp}$  using given method.

### Parameters

- `CCT` (numeric or array\_like) – Correlated colour temperature  $T_{cp}$ .
- `method` (unicode, optional) – {‘Kang 2002’, ‘CIE Illuminant D Series’}, Computation method.

**Returns** *xy* chromaticity coordinates.

**Return type** ndarray

## References

- [\[KMH+02\]](#)
- [\[Wiki\]](#)
- [\[WS00b\]](#)

## colour.CCT\_TO\_XY\_METHODS

`colour.CCT_TO_XY_METHODS = CaseInsensitiveMapping({u'cie_d': ..., u'CIE Illuminant D Series': ..., u'kang2002': ...})`  
Supported correlated colour temperature  $T_{cp}$  to *CIE XYZ* tristimulus values *xy* chromaticity coordinates computation methods.

## References

- [\[KMH+02\]](#)
- [\[Wiki\]](#)
- [\[WS00b\]](#)

**CCT\_TO\_XY\_METHODS** [CaseInsensitiveMapping] {‘Kang 2002’, ‘CIE Illuminant D Series’}

Aliases:

- ‘kang2002’: ‘Kang 2002’
- ‘cie\_d’: ‘Hernandez 1999’

**colour.xy\_to\_CCT**

`colour.xy_to_CCT(xy, method=u'McCamy 1992')`

Returns the correlated colour temperature  $T_{cp}$  from given CIE XYZ tristimulus values xy chromaticity coordinates using given method.

**Parameters**

- `xy` (array\_like) – xy chromaticity coordinates.
- `method` (unicode, optional) – {‘McCamy 1992’, ‘Hernandez 1999’}, Computation method.

**Returns** Correlated colour temperature  $T_{cp}$ .

**Return type** numeric or ndarray

**References**

- [\[HernandezAndresLR99\]](#)
- [\[Wika\]](#)
- [\[Wiki\]](#)

**colour.XY\_TO\_CCT\_METHODS**

`colour.XY_TO_CCT_METHODS = CaseInsensitiveMapping({u'hernandez1999': ..., u'Hernandez 1999': ..., u'McCam`

Supported CIE XYZ tristimulus values xy chromaticity coordinates to correlated colour temperature  $T_{cp}$  computation methods.

**References**

- [\[HernandezAndresLR99\]](#)
- [\[Wika\]](#)
- [\[Wiki\]](#)

**XY\_TO\_CCT\_METHODS** [CaseInsensitiveMapping] {‘McCamy 1992’, ‘Hernandez 1999’}

Aliases:

- ‘mccamy1992’: ‘McCamy 1992’
- ‘hernandez1999’: ‘Hernandez 1999’

## Robertson (1968)

`colour.temperature`

<code>CCT_to_uv_Robertson1968(CCT[, D_uv])</code>	Returns the <i>CIE UCS</i> colourspace <i>uv</i> chromaticity coordinates from given correlated colour temperature $T_{cp}$ and $\Delta_{uv}$ using <i>Robertson (1968)</i> method.
<code>uv_to_CCT_Robertson1968(uv)</code>	Returns the correlated colour temperature $T_{cp}$ and $\Delta_{uv}$ from given <i>CIE UCS</i> colourspace <i>uv</i> chromaticity coordinates using <i>Robertson (1968)</i> method.

### `colour.temperature.CCT_to_uv_Robertson1968`

`colour.temperature.CCT_to_uv_Robertson1968(CCT, D_uv=0)`

Returns the *CIE UCS* colourspace *uv* chromaticity coordinates from given correlated colour temperature  $T_{cp}$  and  $\Delta_{uv}$  using *Robertson (1968)* method.

#### Parameters

- `CCT` (numeric) – Correlated colour temperature  $T_{cp}$ .
- `D_uv` (numeric) –  $\Delta_{uv}$ .

**Returns** *CIE UCS* colourspace *uv* chromaticity coordinates.

**Return type** ndarray

#### References

- [AdobeSystems13b]
- [WS00c]

#### Examples

```
>>> CCT = 6500.0081378199056
>>> D_uv = 0.008333331244225
>>> CCT_to_uv_Robertson1968(CCT, D_uv)
array([ 0.1937413...,  0.3152210...])
```

### `colour.temperature.uv_to_CCT_Robertson1968`

`colour.temperature.uv_to_CCT_Robertson1968(uv)`

Returns the correlated colour temperature  $T_{cp}$  and  $\Delta_{uv}$  from given *CIE UCS* colourspace *uv* chromaticity coordinates using *Robertson (1968)* method.

**Parameters** `uv` (array\_like) – *CIE UCS* colourspace *uv* chromaticity coordinates.

**Returns** Correlated colour temperature  $T_{cp}$ ,  $\Delta_{uv}$ .

**Return type** ndarray

## References

- [\[AdobeSystems13a\]](#)
- [\[WS00c\]](#)

## Examples

```
>>> uv = np.array([0.193741375998230, 0.315221043940594])
>>> uv_to_CCT_Robertson1968(uv)
array([-6.5000162...e+03, 8.3333289...e-03])
```

## Krystek (1985)

colour.temperature

`CCT_to_uv_Krystek1985(CCT)`

Returns the *CIE UCS* colourspace *uv* chromaticity coordinates from given correlated colour temperature  $T_{cp}$  using *Krystek (1985)* method.

`colour.temperature.CCT_to_uv_Krystek1985`

`colour.temperature.CCT_to_uv_Krystek1985(CCT)`

Returns the *CIE UCS* colourspace *uv* chromaticity coordinates from given correlated colour temperature  $T_{cp}$  using *Krystek (1985)* method.

**Parameters** `CCT` (numeric) – Correlated colour temperature  $T_{cp}$ .

**Returns** *CIE UCS* colourspace *uv* chromaticity coordinates.

**Return type** ndarray

## Notes

- *Krystek (1985)* method computations are valid for correlated colour temperature  $T_{cp}$  in domain [1000, 15000].

## References

- [\[Kry85\]](#)

## Examples

```
>>> CCT_to_uv_Krystek1985(6504.38938305)
array([-0.1837669..., 0.3093443...])
```

## Ohno (2013)

`colour.temperature`

<code>CCT_to_uv_Ohno2013(CCT[, D_uv, cmfs])</code>	Returns the <i>CIE UCS</i> colourspace $uv$ chromaticity coordinates from given correlated colour temperature $T_{cp}$ , $\Delta_{uv}$ and colour matching functions using <i>Ohno (2013)</i> method.
<code>uv_to_CCT_Ohno2013(uv[, cmfs, start, end, ...])</code>	Returns the correlated colour temperature $T_{cp}$ and $\Delta_{uv}$ from given <i>CIE UCS</i> colourspace $uv$ chromaticity coordinates, colour matching functions and temperature range using <i>Ohno (2013)</i> method.

### `colour.temperature.CCT_to_uv_Ohno2013`

`colour.temperature.CCT_to_uv_Ohno2013(CCT, D_uv=0, cmfs=XYZ_ColourMatchingFunctions(name='CIE 1931 2 Degree Standard Observer', ...))`

Returns the *CIE UCS* colourspace  $uv$  chromaticity coordinates from given correlated colour temperature  $T_{cp}$ ,  $\Delta_{uv}$  and colour matching functions using *Ohno (2013)* method.

#### Parameters

- `CCT` (numeric) – Correlated colour temperature  $T_{cp}$ .
- `D_uv` (numeric, optional) –  $\Delta_{uv}$ .
- `cmfs` (`XYZ_ColourMatchingFunctions`, optional) – Standard observer colour matching functions.

**Returns** *CIE UCS* colourspace  $uv$  chromaticity coordinates.

**Return type** ndarray

#### References

- [Ohn14]

#### Examples

```
>>> from colour import STANDARD_OBSERVERS_CMFS
>>> cmfs = STANDARD_OBSERVERS_CMFS['CIE 1931 2 Degree Standard Observer']
>>> CCT = 6507.4342201047066
>>> D_uv = 0.003223690901513
>>> CCT_to_uv_Ohno2013(CCT, D_uv, cmfs)
array([ 0.197799...,  0.3122004...])
```

### `colour.temperature.uv_to_CCT_Ohno2013`

`colour.temperature.uv_to_CCT_Ohno2013(uv, cmfs=XYZ_ColourMatchingFunctions(name='CIE 1931 2 Degree Standard Observer', ...), start=1000, end=100000, count=10, iterations=6)`

Returns the correlated colour temperature  $T_{cp}$  and  $\Delta_{uv}$  from given *CIE UCS* colourspace  $uv$  chromaticity

coordinates, colour matching functions and temperature range using *Ohno* (2013) method.

The iterations parameter defines the calculations precision: The higher its value, the more planckian tables will be generated through cascade expansion in order to converge to the exact solution.

#### Parameters

- `uv` (array\_like) – *CIE UCS* colourspace *uv* chromaticity coordinates.
- `cmfs` (`XYZ_ColourMatchingFunctions`, optional) – Standard observer colour matching functions.
- `start` (numeric, optional) – Temperature range start in kelvins.
- `end` (numeric, optional) – Temperature range end in kelvins.
- `count` (int, optional) – Temperatures count in the planckian tables.
- `iterations` (int, optional) – Number of planckian tables to generate.

**Returns** Correlated colour temperature  $T_{cp}$ ,  $\Delta_{uv}$ .

**Return type** ndarray

#### References

- [\[Ohn14\]](#)

#### Examples

```
>>> from colour import STANDARD_OBSERVERS_CMFS
>>> cmfs = STANDARD_OBSERVERS_CMFS['CIE 1931 2 Degree Standard Observer']
>>> uv = np.array([0.1978, 0.3122])
>>> uv_to_CCT_Ohno2013(uv, cmfs)
array([ 6.5074738...e+03,  3.2233461...e-03])
```

## Hernandez-Andres, Lee and Romero (1999)

colour.temperature

---

`xy_to_CCT_Hernandez1999(xy)`

Returns the correlated colour temperature  $T_{cp}$  from given *CIE XYZ* tristimulus values *xy* chromaticity coordinates using *Hernandez-Andres et alii* (1999) method.

---

`colour.temperature.xy_to_CCT_Hernandez1999`

`colour.temperature.xy_to_CCT_Hernandez1999(xy)`

Returns the correlated colour temperature  $T_{cp}$  from given *CIE XYZ* tristimulus values *xy* chromaticity coordinates using *Hernandez-Andres et alii* (1999) method.

**Parameters** `xy` (array\_like) – *xy* chromaticity coordinates.

**Returns** Correlated colour temperature  $T_{cp}$ .

**Return type** numeric

## References

- [HernandezAndresLR99]

## Examples

```
>>> xy = np.array([0.31270, 0.32900])
>>> xy_to_CCT_Hernandez1999(xy)
6500.7420431...
```

Kang, Moon, Hong, Lee, Cho and Kim (2002)

colour.temperature

`CCT_to_xy_Kang2002(CCT)`

Returns the *CIE XYZ* tristimulus values *xy* chromaticity coordinates from given correlated colour temperature  $T_{cp}$  using Kang et alii (2002) method.

`colour.temperature.CCT_to_xy_Kang2002`

`colour.temperature.CCT_to_xy_Kang2002(CCT)`

Returns the *CIE XYZ* tristimulus values *xy* chromaticity coordinates from given correlated colour temperature  $T_{cp}$  using Kang et alii (2002) method.

**Parameters** `CCT` (numeric or array\_like) – Correlated colour temperature  $T_{cp}$ .

**Returns** *xy* chromaticity coordinates.

**Return type** ndarray

**Raises** `ValueError` – If the correlated colour temperature is not in appropriate domain.

## References

- [KMH+02]

## Examples

```
>>> CCT_to_xy_Kang2002(6504.38938305)
array([ 0.313426...,  0.323595...])
```

CIE Illuminant D Series

colour.temperature

<code>CCT_to_xy_CIE_D(CCT)</code>	Converts from the correlated colour temperature $T_{cp}$ of a <i>CIE Illuminant D Series</i> to the chromaticity of that <i>CIE Illuminant D Series</i> illuminant.
-----------------------------------	---

## colour.temperature.CCT\_to\_xy\_CIE\_D

`colour.temperature.CCT_to_xy_CIE_D(CCT)`

Converts from the correlated colour temperature  $T_{cp}$  of a *CIE Illuminant D Series* to the chromaticity of that *CIE Illuminant D Series* illuminant.

**Parameters** `CCT` (numeric or array\_like) – Correlated colour temperature  $T_{cp}$ .

**Returns** `xy` chromaticity coordinates.

**Return type** ndarray

**Raises** `ValueError` – If the correlated colour temperature is not in appropriate domain.

## References

- [WS00b]

## Examples

```
>>> CCT_to_xy_CIE_D(6504.38938305)
array([ 0.3127077...,  0.3291128...])
```

## Utilities

- *Common*
- *Array*
- *Data Structures*
- *Verbose*

## Common

### colour.utilities

<code>handle_numpy_errors(**kwargs)</code>	Decorator for handling Numpy errors.
<code>ignore_numpy_errors(function)</code>	Wrapper for given function.
<code>raise_numpy_errors(function)</code>	Wrapper for given function.
<code>print_numpy_errors(function)</code>	Wrapper for given function.
<code>warn_numpy_errors(function)</code>	Wrapper for given function.
<code>ignore_python_warnings(function)</code>	Decorator for ignoring Python warnings.

Continued on next page

Table 3.201 – continued from previous page

<code>batch(iterable[, k])</code>	Returns a batch generator from given iterable.
<code>is_openimageio_installed([raise_exception])</code>	Returns if <i>OpenImageIO</i> is installed and available.
<code>is_pandas_installed([raise_exception])</code>	Returns if <i>Pandas</i> is installed and available.
<code>is_iterable(a)</code>	Returns if given <i>a</i> variable is iterable.
<code>is_string(a)</code>	Returns if given <i>a</i> variable is a <i>string</i> like variable.
<code>is_numeric(a)</code>	Returns if given <i>a</i> variable is a number.
<code>is_integer(a)</code>	Returns if given <i>a</i> variable is an integer under given threshold.
<code>filter_kwargs(function, **kwargs)</code>	Filters keyword arguments incompatible with the given function signature.
<code>first_item(a)</code>	Return the first item of an iterable.

## colour.utilities.handle\_numpy\_errors

`colour.utilities.handle_numpy_errors(**kwargs)`

Decorator for handling *Numpy* errors.

**Other Parameters** `**kwargs` (*dict, optional*) – Keywords arguments.

**Returns**

**Return type** `object`

## References

- [\[KPK11\]](#)

## Examples

```
>>> import numpy
>>> @handle_numpy_errors(all='ignore')
... def f():
...     1 / numpy.zeros(3)
>>> f()
```

## colour.utilities.ignore\_numpy\_errors

`colour.utilities.ignore_numpy_errors(function)`

Wrapper for given function.

## colour.utilities.raise\_numpy\_errors

`colour.utilities.raise_numpy_errors(function)`

Wrapper for given function.

## colour.utilities.print\_numpy\_errors

colour.utilities.**print\_numpy\_errors**(*function*)  
Wrapper for given function.

## colour.utilities.warn\_numpy\_errors

colour.utilities.**warn\_numpy\_errors**(*function*)  
Wrapper for given function.

## colour.utilities.ignore\_python\_warnings

colour.utilities.**ignore\_python\_warnings**(*function*)  
Decorator for ignoring *Python* warnings.

**Parameters** `function` (`object`) – Function to decorate.

**Returns**

**Return type** `object`

### Examples

```
>>> @ignore_python_warnings
... def f():
...     warnings.warn('This is an ignored warning!')
>>> f()
```

## colour.utilities.batch

colour.utilities.**batch**(*iterable*, *k*=3)  
Returns a batch generator from given iterable.

**Parameters**

- `iterable` (`iterable`) – Iterable to create batches from.
- `k` (`integer`) – Batches size.

**Returns** Is `string_like` variable.

**Return type** `bool`

### Examples

```
>>> batch(tuple(range(10)))
<generator object batch at 0x...>
```

`colour.utilities.is_openimageio_installed``colour.utilities.is_openimageio_installed(raise_exception=False)`

Returns if *OpenImageIO* is installed and available.

**Parameters** `raise_exception` (`bool`) – Raise exception if *OpenImageIO* is unavailable.

**Returns** Is *OpenImageIO* installed.

**Return type** `bool`

**Raises** `ImportError` – If *OpenImageIO* is not installed.

`colour.utilities.is_pandas_installed``colour.utilities.is_pandas_installed(raise_exception=False)`

Returns if *Pandas* is installed and available.

**Parameters** `raise_exception` (`bool`) – Raise exception if *Pandas* is unavailable.

**Returns** Is *Pandas* installed.

**Return type** `bool`

**Raises** `ImportError` – If *Pandas* is not installed.

`colour.utilities.is_iterable``colour.utilities.is_iterable(a)`

Returns if given *a* variable is iterable.

**Parameters** `a` (`object`) – Variable to check the iterability.

**Returns** *a* variable iterability.

**Return type** `bool`

### Examples

```
>>> is_iterable([1, 2, 3])
True
>>> is_iterable(1)
False
```

`colour.utilities.is_string``colour.utilities.is_string(a)`

Returns if given *a* variable is a *string* like variable.

**Parameters** `a` (`object`) – Data to test.

**Returns** Is *a* variable a *string* like variable.

**Return type** `bool`

## Examples

```
>>> is_string("I'm a string!")
True
>>> is_string(["I'm a string!"])
False
```

## colour.utilities.is\_numeric

colour.utilities.**is\_numeric**(*a*)

Returns if given *a* variable is a number.

**Parameters** *a* (`object`) – Variable to check.

**Returns** Is *a* variable a number.

**Return type** `bool`

## Examples

```
>>> is_numeric(1)
True
>>> is_numeric((1,))
False
```

## colour.utilities.is\_integer

colour.utilities.**is\_integer**(*a*)

Returns if given *a* variable is an integer under given threshold.

**Parameters** *a* (`object`) – Variable to check.

**Returns** Is *a* variable an integer.

**Return type** `bool`

## Notes

- The determination threshold is defined by the `colour.algebra.common.INTEGER_THRESHOLD` attribute.

## Examples

```
>>> is_integer(1)
True
>>> is_integer(1.01)
False
```

## colour.utilities.filter\_kwargs

`colour.utilities.filter_kwargs(function, **kwargs)`

Filters keyword arguments incompatible with the given function signature.

**Parameters** `function` (`callable`) – Callable to filter the incompatible keyword arguments.

**Other Parameters** `**kwargs` (`dict, optional`) – Keywords arguments.

**Returns** Filtered keyword arguments.

**Return type** `dict`

### Examples

```
>>> def fn_a(a):
...     return a
>>> def fn_b(a, b=0):
...     return a, b
>>> def fn_c(a, b=0, c=0):
...     return a, b, c
>>> fn_a(1, **filter_kwargs(fn_a, b=2, c=3))
1
>>> fn_b(1, **filter_kwargs(fn_b, b=2, c=3))
(1, 2)
>>> fn_c(1, **filter_kwargs(fn_c, b=2, c=3))
(1, 2, 3)
```

## colour.utilities.first\_item

`colour.utilities.first_item(a)`

Return the first item of an iterable.

**Parameters** `a` (`object`) – Iterable to get the first item from.

**Returns**

**Return type** `object`

**Raises** `StopIteration` – If the iterable is empty.

### Examples

```
>>> a = range(10)
>>> first_item(a)
0
```

## Array

`colour.utilities`

---

`as_numeric(a[, type_])`

Converts given `a` variable to *numeric*.

---

Continued on next page

Table 3.202 – continued from previous page

<code>as_namedtuple(a, named_tuple)</code>	Converts given <i>a</i> variable to given <i>namedtuple</i> class instance.
<code>closest_indexes(a, b)</code>	Returns the <i>a</i> variable closest element indexes to reference <i>b</i> variable elements.
<code>closest(a, b)</code>	Returns the <i>a</i> variable closest elements to reference <i>b</i> variable elements.
<code>normalise_maximum(a[, axis, factor, clip])</code>	Normalises given <i>array_like</i> <i>a</i> variable values by <i>a</i> variable maximum value and optionally clip them between.
<code>interval(distribution[, unique])</code>	Returns the interval size of given distribution.
<code>is_uniform(distribution)</code>	Returns if given distribution is uniform.
<code>in_array(a, b[, tolerance])</code>	Tests whether each element of an array is also present in a second array within given tolerance.
<code>tstack(a)</code>	Stacks arrays in sequence along the last axis (tail).
<code>tsplit(a)</code>	Splits arrays in sequence along the last axis (tail).
<code>row_as_diagonal(a)</code>	Returns the per row diagonal matrices of the given array.
<code>dot_vector(m, v)</code>	Convenient wrapper around <code>np.einsum()</code> with the following subscripts: ' $\dots ij, \dots j > \dots i$ '.
<code>dot_matrix(a, b)</code>	Convenient wrapper around <code>np.einsum()</code> with the following subscripts: ' $\dots ij, \dots jk > \dots ik$ '.
<code>orient(a, orientation)</code>	Orient given array according to given orientation value.
<code>centroid(a)</code>	Computes the centroid indexes of given <i>a</i> array.
<code>linear_conversion(a, old_range, new_range)</code>	Performs a simple linear conversion of given array between the old and new ranges.
<code>fill_nan(a[, method, default])</code>	Fills given array NaNs according to given method.
<code>ndarray_write(*args, **kwds)</code>	A context manager setting given array writeable to perform an operation and then read-only.

## colour.utilities.as\_numeric

`colour.utilities.as_numeric(a, type_=<type 'numpy.float64'>)`

Converts given *a* variable to *numeric*. In the event where *a* cannot be converted, it is passed as is.

### Parameters

- `a (object)` – Variable to convert.
- `type (object)` – Type to use for conversion.

**Returns** *a* variable converted to *numeric*.

**Return type** ndarray

### Examples

```
>>> as_numeric(np.array([1]))
1.0
>>> as_numeric(np.arange(10))
array([ 0.,  1.,  2.,  3.,  4.,  5.,  6.,  7.,  8.,  9.])
```

## colour.utilities.as\_namedtuple

`colour.utilities.as_namedtuple(a, named_tuple)`

Converts given *a* variable to given *namedtuple* class instance.

*a* can be either a *Numpy* structured array, a *namedtuple*, a *mapping*, or an *array\_like* object. The definition will attempt to convert it to given *namedtuple*.

### Parameters

- **a** (`object`) – Variable to convert.
- **named\_tuple** (`namedtuple`) – *namedtuple* class.

**Returns** `math:a` variable converted to *namedtuple*.

**Return type** `namedtuple`

## Examples

```
>>> from collections import namedtuple
>>> a_a = 1
>>> a_b = 2
>>> a_c = 3
>>> NamedTuple = namedtuple('NamedTuple', 'a b c')
>>> as_namedtuple(NamedTuple(a=1, b=2, c=3), NamedTuple)
NamedTuple(a=1, b=2, c=3)
>>> as_namedtuple({'a': a_a, 'b': a_b, 'c': a_c}, NamedTuple)
NamedTuple(a=1, b=2, c=3)
>>> as_namedtuple([a_a, a_b, a_c], NamedTuple)
NamedTuple(a=1, b=2, c=3)
```

## colour.utilities.closest\_indexes

`colour.utilities.closest_indexes(a, b)`

Returns the *a* variable closest element indexes to reference *b* variable elements.

### Parameters

- **a** (`array_like`) – Variable to search for the closest element indexes.
- **b** (`numeric`) – Reference variable.

**Returns** Closest *a* variable element indexes.

**Return type** `numeric`

## Examples

```
>>> a = np.array([24.31357115, 63.62396289, 55.71528816,
...               62.70988028, 46.84480573, 25.40026416])
>>> closest_indexes(a, 63)
array([3])
>>> closest_indexes(a, [63, 25])
array([3, 5])
```

## colour.utilities.closest

`colour.utilities.closest(a, b)`

Returns the *a* variable closest elements to reference *b* variable elements.

### Parameters

- **a** (array\_like) – Variable to search for the closest elements.
- **b** (numeric) – Reference variable.

**Returns** Closest *a* variable elements.

**Return type** numeric

### Examples

```
>>> a = np.array([24.31357115, 63.62396289, 55.71528816,
...                 62.70988028, 46.84480573, 25.40026416])
>>> closest(a, 63)
array([ 62.70988028])
>>> closest(a, [63, 25])
array([ 62.70988028,  25.40026416])
```

## colour.utilities.normalise\_maximum

`colour.utilities.normalise_maximum(a, axis=None, factor=1, clip=True)`

Normalises given *array\_like* *a* variable values by *a* variable maximum value and optionally clip them between.

### Parameters

- **a** (array\_like) – *a* variable to normalise.
- **axis** (numeric, optional) – Normalization axis.
- **factor** (numeric, optional) – Normalization factor.
- **clip** (bool, optional) – Clip values between in domain [0, ‘factor’].

**Returns** Maximum normalised *a* variable.

**Return type** ndarray

### Examples

```
>>> a = np.array([0.48222001, 0.31654775, 0.22070353])
>>> normalise_maximum(a)
array([ 1.          ,  0.6564384...,  0.4576822...])
```

## colour.utilities.interval

`colour.utilities.interval(distribution, unique=True)`

Returns the interval size of given distribution.

**Parameters**

- **distribution** (array\_like) – Distribution to retrieve the interval.
- **unique** (bool, optional) – Whether to return unique intervals if the distribution is non-uniformly spaced or the complete intervals

**Returns** Distribution interval.**Return type** ndarray**Examples**

Uniformly spaced variable:

```
>>> y = np.array([1, 2, 3, 4, 5])
>>> interval(y)
array([1])
>>> interval(y, False)
array([1, 1, 1, 1])
```

Non-uniformly spaced variable:

```
>>> y = np.array([1, 2, 3, 4, 8])
>>> interval(y)
array([1, 4])
>>> interval(y, False)
array([1, 1, 1, 4])
```

**colour.utilities.is\_uniform****colour.utilities.is\_uniform(distribution)**

Returns if given distribution is uniform.

**Parameters** **distribution** (array\_like) – Distribution to check for uniformity.**Returns** Is distribution uniform.**Return type** bool**Examples**

Uniformly spaced variable:

```
>>> a = np.array([1, 2, 3, 4, 5])
>>> is_uniform(a)
True
```

Non-uniformly spaced variable:

```
>>> a = np.array([1, 2, 3.1415, 4, 5])
>>> is_uniform(a)
False
```

## colour.utilities.in\_array

`colour.utilities.in_array(a, b, tolerance=2.2204460492503131e-16)`

Tests whether each element of an array is also present in a second array within given tolerance.

### Parameters

- `a` (array\_like) – Array to test the elements from.
- `b` (array\_like) – The values against which to test each value of array `a`.
- `tolerance` (numeric, optional) – Tolerance value.

**Returns** A boolean array with `a` shape describing whether an element of `a` is present in `b` within given tolerance.

**Return type** ndarray

### References

- [Yor14]

### Examples

```
>>> a = np.array([0.50, 0.60])
>>> b = np.linspace(0, 10, 101)
>>> np.in1d(a, b)
array([ True, False], dtype=bool)
>>> in_array(a, b)
array([ True,  True], dtype=bool)
```

## colour.utilities.vstack

`colour.utilities.vstack(a)`

Stacks arrays in sequence along the last axis (tail).

Rebuilds arrays divided by `colour.utilities.tsplits()`.

**Parameters** `a` (array\_like) – Array to perform the stacking.

### Returns

**Return type** ndarray

### Examples

```
>>> a = []
>>> vstack((a, a, a))
array([[], [], []])
>>> a = np.arange(0, 6)
>>> vstack((a, a, a))
array([[0, 0, 0],
       [1, 1, 1],
       [2, 2, 2],
       [3, 3, 3],
```

```

[4, 4, 4],
[5, 5, 5]])
>>> a = np.reshape(a, (1, 6))
>>> tstack((a, a, a))
array([[[0, 0, 0],
       [1, 1, 1],
       [2, 2, 2],
       [3, 3, 3],
       [4, 4, 4],
       [5, 5, 5]]])
>>> a = np.reshape(a, (1, 1, 6))
>>> tstack((a, a, a))
array([[[[0, 0, 0],
       [1, 1, 1],
       [2, 2, 2],
       [3, 3, 3],
       [4, 4, 4],
       [5, 5, 5]]]])

```

## colour.utilities.tsplit

colour.utilities.**tsplit**(*a*)

Splits arrays in sequence along the last axis (tail).

**Parameters** *a* (array\_like) – Array to perform the splitting.

**Returns**

**Return type** ndarray

## Examples

```

>>> a = np.array([0, 0, 0])
>>> tsplit(a)
array([0, 0, 0])
>>> a = np.array(
...     [[0, 0, 0],
...      [1, 1, 1],
...      [2, 2, 2],
...      [3, 3, 3],
...      [4, 4, 4],
...      [5, 5, 5]])
...
>>> tsplit(a)
array([[0, 1, 2, 3, 4, 5],
       [0, 1, 2, 3, 4, 5],
       [0, 1, 2, 3, 4, 5]])
>>> a = np.array(
...     [[[0, 0, 0],
...       [1, 1, 1],
...       [2, 2, 2],
...       [3, 3, 3],
...       [4, 4, 4],
...       [5, 5, 5]]]
... )

```

```
>>> tsplit(a)
array([[[0, 1, 2, 3, 4, 5]],
       [[0, 1, 2, 3, 4, 5]],
       [[0, 1, 2, 3, 4, 5]])
```

## colour.utilities.row\_as\_diagonal

```
colour.utilities.row_as_diagonal(a)
```

Returns the per row diagonal matrices of the given array.

**Parameters** `a` (array\_like) – Array to perform the diagonal matrices computation.

**Returns**

**Return type** ndarray

## References

- [Cas14]

## Examples

```
>>> a = np.array(
...     [[0.25891593, 0.07299478, 0.36586996],
...      [0.30851087, 0.37131459, 0.16274825],
...      [0.71061831, 0.67718718, 0.09562581],
...      [0.71588836, 0.76772047, 0.15476079],
...      [0.92985142, 0.22263399, 0.88027331]])
...
>>> row_as_diagonal(a)
array([[[ 0.25891593,  0.          ,  0.          ],
        [ 0.          ,  0.07299478,  0.          ],
        [ 0.          ,  0.          ,  0.36586996]],

       [[ 0.30851087,  0.          ,  0.          ],
        [ 0.          ,  0.37131459,  0.          ],
        [ 0.          ,  0.          ,  0.16274825]],

       [[ 0.71061831,  0.          ,  0.          ],
        [ 0.          ,  0.67718718,  0.          ],
        [ 0.          ,  0.          ,  0.09562581]],

       [[ 0.71588836,  0.          ,  0.          ],
        [ 0.          ,  0.76772047,  0.          ],
        [ 0.          ,  0.          ,  0.15476079]],

       [[ 0.92985142,  0.          ,  0.          ],
        [ 0.          ,  0.22263399,  0.          ],
        [ 0.          ,  0.          ,  0.88027331]])
```

**colour.utilities.dot\_vector**

```
colour.utilities.dot_vector(m, v)
```

Convenient wrapper around `np.einsum()` with the following subscripts: ‘ $\dots ij, \dots j->\dots i$ ’.

It performs the dot product of two arrays where `m` parameter is expected to be an array of 3x3 matrices and parameter `v` an array of vectors.

**Parameters**

- `m` (`array_like`) – Array of 3x3 matrices.
- `v` (`array_like`) – Array of vectors.

**Returns**

**Return type** `ndarray`

**Examples**

```
>>> m = np.array(
...     [[0.7328, 0.4296, -0.1624],
...      [-0.7036, 1.6975, 0.0061],
...      [0.0030, 0.0136, 0.9834]])
...
>>> m = np.reshape(np.tile(m, (6, 1)), (6, 3, 3))
>>> v = np.array([0.07049534, 0.10080000, 0.09558313])
>>> v = np.tile(v, (6, 1))
>>> dot_vector(m, v)
array([[ 0.0794399...,  0.1220905...,  0.0955788...],
       [ 0.0794399...,  0.1220905...,  0.0955788...],
       [ 0.0794399...,  0.1220905...,  0.0955788...],
       [ 0.0794399...,  0.1220905...,  0.0955788...],
       [ 0.0794399...,  0.1220905...,  0.0955788...],
       [ 0.0794399...,  0.1220905...,  0.0955788...]])
```

**colour.utilities.dot\_matrix**

```
colour.utilities.dot_matrix(a, b)
```

Convenient wrapper around `np.einsum()` with the following subscripts: ‘ $\dots ij, \dots jk->\dots ik$ ’.

It performs the dot product of two arrays where `a` parameter is expected to be an array of 3x3 matrices and parameter `b` another array of of 3x3 matrices.

**Parameters**

- `a` (`array_like`) – Array of 3x3 matrices.
- `b` (`array_like`) – Array of 3x3 matrices.

**Returns**

**Return type** `ndarray`

## Examples

```
>>> a = np.array(  
...     [[0.7328, 0.4296, -0.1624],  
...      [-0.7036, 1.6975, 0.0061],  
...      [0.0030, 0.0136, 0.9834]]  
... )  
>>> a = np.reshape(np.tile(a, (6, 1)), (6, 3, 3))  
>>> b = a  
>>> dot_matrix(a, b)  
array([[[ 0.2342420..., 1.0418482..., -0.2760903...],  
       [-1.7099407..., 2.5793226..., 0.1306181...],  
       [-0.0044203..., 0.0377490..., 0.9666713...]],  
  
[[ 0.2342420..., 1.0418482..., -0.2760903...],  
[-1.7099407..., 2.5793226..., 0.1306181...],  
[-0.0044203..., 0.0377490..., 0.9666713...]],  
  
[[ 0.2342420..., 1.0418482..., -0.2760903...],  
[-1.7099407..., 2.5793226..., 0.1306181...],  
[-0.0044203..., 0.0377490..., 0.9666713...]],  
  
[[ 0.2342420..., 1.0418482..., -0.2760903...],  
[-1.7099407..., 2.5793226..., 0.1306181...],  
[-0.0044203..., 0.0377490..., 0.9666713...]],  
  
[[ 0.2342420..., 1.0418482..., -0.2760903...],  
[-1.7099407..., 2.5793226..., 0.1306181...],  
[-0.0044203..., 0.0377490..., 0.9666713...]],  
  
[[ 0.2342420..., 1.0418482..., -0.2760903...],  
[-1.7099407..., 2.5793226..., 0.1306181...],  
[-0.0044203..., 0.0377490..., 0.9666713...]]])
```

## colour.utilities.orient

colour.utilities.orient(*a*, *orientation*)

Orient given array according to given orientation value.

### Parameters

- ***a*** (array\_like) – Array to perform the orientation onto.
- ***orientation*** (unicode, optional) – {‘Flip’, ‘Flop’, ‘90 CW’, ‘90 CCW’, ‘180’} Orientation to perform.

**Returns** Oriented array.

**Return type** ndarray

## Examples

```
>>> a = np.tile(np.arange(5), (5, 1))  
>>> a  
array([[0, 1, 2, 3, 4],  
       [0, 1, 2, 3, 4],
```

```
[0, 1, 2, 3, 4],
[0, 1, 2, 3, 4],
[0, 1, 2, 3, 4]])
>>> orient(a, '90 CW')
array([[0, 0, 0, 0, 0],
       [1, 1, 1, 1, 1],
       [2, 2, 2, 2, 2],
       [3, 3, 3, 3, 3],
       [4, 4, 4, 4, 4]])
>>> orient(a, 'Flip')
array([[4, 3, 2, 1, 0],
       [4, 3, 2, 1, 0],
       [4, 3, 2, 1, 0],
       [4, 3, 2, 1, 0],
       [4, 3, 2, 1, 0]])
```

## colour.utilities.centroid

`colour.utilities.centroid(a)`

Computes the centroid indexes of given *a* array.

**Parameters** *a* (array\_like) – *a* array to compute the centroid indexes.

**Returns** *a* array centroid indexes.

**Return type** ndarray

### Examples

```
>>> a = np.tile(np.arange(0, 5), (5, 1))
>>> centroid(a)
array([2, 3])
```

## colour.utilities.linear\_conversion

`colour.utilities.linear_conversion(a, old_range, new_range)`

Performs a simple linear conversion of given array between the old and new ranges.

**Parameters**

- *a* (array\_like) – Array to perform the linear conversion onto.
- **old\_range** (array\_like) – Old range.
- **new\_range** (array\_like) – New range.

**Returns**

**Return type** ndarray

### Examples

```
>>> a = np.linspace(0, 1, 10)
>>> linear_conversion(a, np.array([0, 1]), np.array([1, 10]))
array([ 1.,  2.,  3.,  4.,  5.,  6.,  7.,  8.,  9., 10.])
```

## colour.utilities.fill\_nan

colour.utilities.**fill\_nan**(*a*, *method='Interpolation'*, *default=0*)

Fills given array NaNs according to given method.

### Parameters

- ***a*** (array\_like) – Array to fill the NaNs of.
- ***method*** (unicode) – {‘Interpolation’, ‘Constant’}, *Interpolation* method linearly interpolates through the NaNs, *Constant* method replaces NaNs with *default*.
- ***default*** (numeric) – Value to use with the *Constant* method.

**Returns** NaNs filled array.

**Return type** ndarray

## Examples

```
>>> a = np.array([0.1, 0.2, np.nan, 0.4, 0.5])
>>> fill_nan(a)
array([ 0.1,  0.2,  0.3,  0.4,  0.5])
>>> fill_nan(a, method='Constant')
array([ 0.1,  0.2,  0. ,  0.4,  0.5])
```

## colour.utilities.ndarray\_write

colour.utilities.**ndarray\_write**(\*args, \*\*kwds)

A context manager setting given array writeable to perform an operation and then read-only.

**Parameters** ***a*** (array\_like) – Array to perform an operation.

**Returns** Array.

**Return type** ndarray

## Examples

```
>>> a = np.linspace(0, 1, 10)
>>> a.setflags(write=False)
>>> try:
...     a += 1
... except ValueError:
...     pass
>>> with ndarray_write(a):
...     a +=1
```

## Data Structures

`colour.utilities`

<code>CaseInsensitiveMapping([data])</code>	Implements a case-insensitive mutable mapping / <i>dict</i> object.
<code>Lookup</code>	Extends <i>dict</i> type to provide a lookup by value(s).
<code>Structure(*args, **kwargs)</code>	Defines an object similar to C/C++ structured type.

### `colour.utilities.CaseInsensitiveMapping`

`class colour.utilities.CaseInsensitiveMapping(data=None, **kwargs)`  
Implements a case-insensitive mutable mapping / *dict* object.

Allows values retrieving from keys while ignoring the key case. The keys are expected to be unicode or string-like objects supporting the `str.lower()` method.

**Parameters** `data` (`dict`) – *dict* of data to store into the mapping at initialisation.

**Other Parameters** `**kwargs` (`dict, optional`) – Key / Value pairs to store into the mapping at initialisation.

`__setitem__()`  
`__getitem__()`  
`__delitem__()`  
`__contains__()`  
`__iter__()`  
`__len__()`  
`__eq__()`  
`__ne__()`  
`__repr__()`  
`copy()`  
`lower_items()`

**Warning:** The keys are expected to be unicode or string-like objects.

## References

- [\[Rei\]](#)

## Examples

```
>>> methods = CaseInsensitiveMapping({'McCamy': 1, 'Hernandez': 2})
>>> methods['mccamy']
1
```

`__init__(data=None, **kwargs)`

## Methods

---

<code>__init__([data])</code>	
<code>clear() -&gt; None. Remove all items from D.)</code>	
<code>copy()</code>	Returns a copy of the mapping.
<code>get((k,d]) -&gt; D[k] if k in D, ...)</code>	
<code>items() -&gt; list of D's (key, value) pairs, ...)</code>	
<code>iteritems() -&gt; an iterator over the (key, ...)</code>	
<code>iterkeys() -&gt; an iterator over the keys of D)</code>	
<code>itervalues(...)</code>	
<code>keys() -&gt; list of D's keys)</code>	
<code>lower_items()</code>	Iterates over the lower items names.
<code>pop((k,d]) -&gt; v, ...)</code>	If key is not found, d is returned if given, otherwise KeyError is raised.
<code>popitem() -&gt; (k, v), ...)</code>	as a 2-tuple; but raise KeyError if D is empty.
<code>setdefault((k,d]) -&gt; D.get(k,d), ...)</code>	
<code>update(([E, ...)</code>	If E present and has a .keys() method, does: for k in E: D[k] = E[k]
<code>values() -&gt; list of D's values)</code>	

---

## colour.utilities.Lookup

`class colour.utilities.Lookup`  
Extends `dict` type to provide a lookup by value(s).

`first_key_from_value()`

`keys_from_value()`

## References

- [Mana]

## Examples

```
>>> person = Lookup(first_name='Doe', last_name='John', gender='male')
>>> person.first_key_from_value('Doe')
'first_name'
>>> persons = Lookup(John='Doe', Jane='Doe', Luke='Skywalker')
>>> sorted(persons.keys_from_value('Doe'))
['Jane', 'John']
```

`__init__()`  
`x.__init__(...) initializes x; see help(type(x)) for signature`

## Methods

clear() -> None. Remove all items from D.)	
copy() -> a shallow copy of D)	
first_key_from_value(value)	Gets the first key with given value.
fromkeys(...)	v defaults to None.
get((k,[d]) -> D[k] if k in D, ...)	
has_key((k) -> True if D has a key k, else False)	
items() -> list of D's (key, value) pairs, ...)	
iteritems() -> an iterator over the (key, ...)	
iterkeys() -> an iterator over the keys of D)	
itervalues(...)	
keys() -> list of D's keys)	
keys_from_value(value)	Gets the keys with given value.
pop((k,[d]) -> v, ...)	If key is not found, d is returned if given, otherwise KeyError is raised
popitem() -> (k, v), ...)	2-tuple; but raise KeyError if D is empty.
setdefault((k,[d]) -> D.get(k,d), ...)	
update(([E, ...)	If E present and has a .keys() method, does: for k in E: D[k] = E[k]
values() -> list of D's values)	
viewitems(...)	
viewkeys(...)	
viewvalues(...)	

## colour.utilities.Structure

**class** colour.utilities.Structure(\*args, \*\*kwargs)  
 Defines an object similar to C/C++ structured type.

### Other Parameters

- \*args (*list, optional*) – Arguments.
- \*\*kwargs (*dict, optional*) – Key / Value pairs.

`__getattr__()`  
`__setattr__()`  
`__delattr__()`  
`update()`

### References

- [Manb]

### Examples

```
>>> person = Structure(first_name='Doe', last_name='John', gender='male')
>>> # Doctests skip for Python 2.x compatibility.
>>> person.first_name
'Doe'
>>> sorted(person.keys())
```

```
['first_name', 'gender', 'last_name']
>>> # Doctests skip for Python 2.x compatibility.
>>> person['gender']
'male'
```

```
__init__(*args, **kwargs)
```

## Methods

<code>__init__(*args, **kwargs)</code>	
<code>clear() -&gt; None.</code> Remove all items from D.)	
<code>copy() -&gt; a shallow copy of D)</code>	
<code>fromkeys(...)</code>	v defaults to None.
<code>get((k,d)) -&gt; D[k] if k in D, ...)</code>	
<code>has_key((k) -&gt; True if D has a key k, else False)</code>	
<code>items() -&gt; list of D's (key, value) pairs, ...)</code>	
<code>iteritems() -&gt; an iterator over the (key, ...)</code>	
<code>iterkeys() -&gt; an iterator over the keys of D)</code>	
<code>itervalues(...)</code>	
<code>keys() -&gt; list of D's keys)</code>	
<code>pop((k,d)) -&gt; v, ...)</code>	If key is not found, d is returned if given, otherwise KeyError is raised
<code>popitem() -&gt; (k, v), ...)</code>	2-tuple; but raise KeyError if D is empty.
<code>setdefault((k,d)) -&gt; D.get(k,d), ...)</code>	
<code>update(*args, **kwargs)</code>	Updates both keys and sibling attributes.
<code>values() -&gt; list of D's values)</code>	
<code>viewitems(...)</code>	
<code>viewkeys(...)</code>	
<code>viewvalues(...)</code>	

## Verbose

### colour.utilities

<code>message_box(message[, width, padding])</code>	Prints a message inside a box.
<code>warning(*args, **kwargs)</code>	Issues a warning.
<code>filter_warnings([state, colour_warnings_only])</code>	Filters Colour and also optionally overall Python warnings.
<code>suppress_warnings(*args, **kwds)</code>	A context manager filtering Colour and also optionally overall Python warnings.
<code>numpy_print_options(*args, **kwds)</code>	A context manager implementing context changes to Numpy print behaviour.

### colour.utilities.message\_box

```
colour.utilities.message_box(message, width=79, padding=3)
```

Prints a message inside a box.

#### Parameters

- **message** (`unicode`) – Message to print.
- **width** (`int`, optional) – Message box width.
- **padding** (`unicode`) – Padding on each sides of the message.

**Returns** Definition success.

**Return type** `bool`

## Examples

```
>>> message = ('Lorem ipsum dolor sit amet, consectetur adipiscing elit, '
...           'sed do eiusmod tempor incididunt ut labore et dolore magna '
...           'aliqua.')
>>> message_box(message, width=75)
=====
*                                         *
*   Lorem ipsum dolor sit amet, consectetur adipiscing elit, sed do   *
*   eiusmod tempor incididunt ut labore et dolore magna aliqua.       *
*                                         *
=====
True
>>> message_box(message, width=60)
=====
*                                         *
*   Lorem ipsum dolor sit amet, consectetur adipiscing   *
*   elit, sed do eiusmod tempor incididunt ut labore et   *
*   dolore magna aliqua.                         *
*                                         *
=====
True
>>> message_box(message, width=75, padding=16)
=====
*                                         *
*           Lorem ipsum dolor sit amet, consectetur   *
*           adipiscing elit, sed do eiusmod tempor   *
*           incididunt ut labore et dolore magna   *
*           aliqua.                           *
*                                         *
=====
True
```

## colour.utilities.warning

```
colour.utilities.warning(*args, **kwargs)
```

Issues a warning.

### Other Parameters

- **\*args** (`list`, optional) – Arguments.
- **\*\*kwargs** (`dict`, optional) – Keywords arguments.

**Returns** Definition success.

**Return type** `bool`

## Examples

```
>>> warning('This is a warning!')
/Users/.../colour/utilities/verbose.py:132: UserWarning: This is a warning!
```

## colour.utilities.filter\_warnings

`colour.utilities.filter_warnings(state=True, colour_warnings_only=True)`  
Filters *Colour* and also optionally overall Python warnings.

### Parameters

- `state` (`bool`, optional) – Warnings filter state.
- `colour_warnings_only` (`bool`, optional) – Whether to only filter *Colour* warnings or also overall Python warnings.

**Returns** Definition success.

**Return type** `bool`

## Examples

```
# Filtering Colour only warnings: >>> filter_warnings() True
# Filtering Colour and also Python warnings: >>> filter_warnings(colour_warnings_only=False) True
```

## colour.utilities.suppress\_warnings

`colour.utilities.suppress_warnings(*args, **kwds)`  
A context manager filtering *Colour* and also optionally overall Python warnings.

**Parameters** `colour_warnings_only` (`bool`, optional) – Whether to only filter *Colour* warnings or also overall Python warnings.

## colour.utilities.numpy\_print\_options

`colour.utilities.numpy_print_options(*args, **kwds)`  
A context manager implementing context changes to *Numpy* print behaviour.

### Other Parameters

- `*args` (`list, optional`) – Arguments.
- `**kwargs` (`dict, optional`) – Keywords arguments.

## Examples

```
>>> np.array([np.pi])
array([ 3.1415926...])
>>> with numpy_print_options(formatter={'float': '{:0.1f}'.format}):
...     np.array([np.pi])
array([3.1])
```

## Ancillary Objects

`colour.utilities`

---

### `ColourWarning`

This is the base class of *Colour* warnings.

---

## `colour.utilities.ColourWarning`

**exception** `colour.utilities.ColourWarning`

This is the base class of *Colour* warnings. It is a subclass of `Warning`.

## Colour Volume

- *Optimal Colour Stimuli - MacAdam Limits*
- *Mesh Volume*
- *Pointer's Gamut*
- *RGB Volume*
- *Visible Spectrum*

## Optimal Colour Stimuli - MacAdam Limits

`colour`

---

<code>is_within_macadam_limits(xyY, illuminant[, ...])</code>	Returns if given <i>CIE xyY</i> colourspace array is within MacAdam limits of given illuminant.
<code>ILLUMINANTS_OPTIMAL_COLOUR_STIMULI</code>	Illuminants <i>Optimal Colour Stimuli</i> .

---

## `colour.is_within_macadam_limits`

`colour.is_within_macadam_limits(xyY, illuminant, tolerance=None)`

Returns if given *CIE xyY* colourspace array is within MacAdam limits of given illuminant.

### Parameters

- `xyY` (array\_like) – *CIE xyY* colourspace array.
- `illuminant` (unicode) – Illuminant.
- `tolerance` (numeric, optional) – Tolerance allowed in the inside-triangle check.

**Returns** Is within MacAdam limits.

**Return type** `bool`

### Notes

- Input *CIE xyY* colourspace array is in domain [0, 1].

## Examples

```
>>> is_within_macadam_limits(np.array([0.3205, 0.4131, 0.51]), 'A')
array(True, dtype=bool)
>>> a = np.array([[0.3205, 0.4131, 0.51],
...                 [0.0005, 0.0031, 0.001]])
>>> is_within_macadam_limits(a, 'A')
array([ True, False], dtype=bool)
```

## colour.ILLUMINANTS\_OPTIMAL\_COLOUR\_STIMULI

colour.ILLUMINANTS\_OPTIMAL\_COLOUR\_STIMULI = CaseInsensitiveMapping({u'A': ..., u'D65': ..., u'C': ...})  
Illuminants *Optimal Colour Stimuli*.

## References

- [\[Wikw\]](#)

ILLUMINANTS\_OPTIMAL\_COLOUR\_STIMULI [CaseInsensitiveMapping] {'A', 'C', 'D65'}

## Mesh Volume

colour

---

is_within_mesh_volume(points, mesh[, tolerance])	Returns if given points are within given mesh volume using Delaunay triangulation.
--	--

---

## colour.is\_within\_mesh\_volume

colour.is\_within\_mesh\_volume(points, mesh, tolerance=None)

Returns if given points are within given mesh volume using Delaunay triangulation.

### Parameters

- **points** (array\_like) – Points to check if they are within mesh volume.
- **mesh** (array\_like) – Points of the volume used to generate the Delaunay triangulation.
- **tolerance** (numeric, optional) – Tolerance allowed in the inside-triangle check.

**Returns** Is within mesh volume.

**Return type** bool

## Examples

```
>>> mesh = np.array(
...     [[-1.0, -1.0, 1.0],
...      [1.0, -1.0, 1.0],
...      [1.0, -1.0, -1.0],
```

```

...
[-1.0, -1.0, -1.0],
[0.0, 1.0, 0.0]]
...
)
>>> is_within_mesh_volume(np.array([0.0005, 0.0031, 0.0010]), mesh)
array(True, dtype=bool)
>>> a = np.array([[0.0005, 0.0031, 0.0010],
...                 [0.3205, 0.4131, 0.5100]])
>>> is_within_mesh_volume(a, mesh)
array([ True, False], dtype=bool)

```

## Pointer's Gamut

colour

<code>is_within_pointer_gamut(XYZ[, tolerance])</code>	Returns if given <i>CIE XYZ</i> tristimulus values are within Pointer's Gamut volume.
--	---

`colour.is_within_pointer_gamut`

`colour.is_within_pointer_gamut(XYZ, tolerance=None)`

Returns if given *CIE XYZ* tristimulus values are within Pointer's Gamut volume.

### Parameters

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `tolerance` (`numeric`, `optional`) – Tolerance allowed in the inside-triangle check.

**Returns** Is within Pointer's Gamut.

**Return type** `bool`

### Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].

### Examples

```

>>> import numpy as np
>>> is_within_pointer_gamut(np.array([0.3205, 0.4131, 0.5100]))
array(True, dtype=bool)
>>> a = np.array([[0.3205, 0.4131, 0.5100], [0.0005, 0.0031, 0.0010]])
>>> is_within_pointer_gamut(a)
array([ True, False], dtype=bool)

```

## RGB Volume

colour

RGB_colourspace_limits(colourspace[, illuminant])	Computes given <i>RGB</i> colourspace volume limits in <i>Lab</i> colourspace.
RGB_colourspace_pointer_gamut_coverage_MonteCarlo()	Returns given <i>RGB</i> colourspace percentage coverage of Pointer's Gamut volume using <i>Monte Carlo</i> method.
RGB_colourspace_visible_spectrum_coverage_MonteCarlo()	Returns given <i>RGB</i> colourspace percentage coverage of visible spectrum volume using <i>Monte Carlo</i> method.
RGB_colourspace_volume_MonteCarlo(colourspace)	Performs given <i>RGB</i> colourspace volume computation using <i>Monte Carlo</i> method and multiprocessing.
RGB_colourspace_volume_coverage_MonteCarlo(...)	Returns given <i>RGB</i> colourspace percentage coverage of an arbitrary volume.

## colour.RGB\_colourspace\_limits

colour.RGB\_colourspace\_limits(*colourspace*, *illuminant*=array([ 0.3457, 0.3585]))  
Computes given *RGB* colourspace volume limits in *Lab* colourspace.

### Parameters

- **colourspace** (*RGB\_Colourspace*) – *RGB* colourspace to compute the volume of.
- **illuminant** (*array\_like*, optional) – *Lab* colourspace *illuminant* chromaticity coordinates.

**Returns** *RGB* colourspace volume limits.

**Return type** *ndarray*

### Examples

```
>>> from colour import sRGB_COLOURSPACE as sRGB
>>> RGB_colourspace_limits(sRGB)
array([[ 0..., 100.0000848...],
       [-79.2197012..., 94.6760011...],
       [-114.7814393..., 96.7261797...]])
```

## colour.RGB\_colourspace\_pointer\_gamut\_coverage\_MonteCarlo

colour.RGB\_colourspace\_pointer\_gamut\_coverage\_MonteCarlo(*colourspace*, *samples*=10000000.0,  
*random\_generator*=<*function random\_triplet\_generator*>, *random\_state*=None)

Returns given *RGB* colourspace percentage coverage of Pointer's Gamut volume using *Monte Carlo* method.

### Parameters

- **colourspace** (*RGB\_Colourspace*) – *RGB* colourspace to compute the *Pointer's Gamut* coverage percentage.
- **samples** (*numeric*, optional) – Samples count.
- **random\_generator** (*generator*, optional) – Random triplet generator providing the random samples.

- **random\_state** (RandomState, optional) – Mersenne Twister pseudo-random number generator to use in the random number generator.

**Returns** Percentage coverage of *Pointer's Gamut* volume.

**Return type** float

## Examples

```
>>> from colour import sRGB_COLOURSPACE as sRGB
>>> prng = np.random.RandomState(2)
>>> RGB_colourspace_pointer_gamut_coverage_MonteCarlo(
...     sRGB, 10e3, random_state=prng)
83...
```

## colour.RGB\_colourspace\_visible\_spectrum\_coverage\_MonteCarlo

```
colour.RGB_colourspace_visible_spectrum_coverage_MonteCarlo(colourspace,
    samples=10000000.0, random_generator=<function random_triplet_generator>,
    random_state=None)
```

Returns given *RGB* colourspace percentage coverage of visible spectrum volume using *Monte Carlo* method.

### Parameters

- **colourspace** (RGB\_Colourspace) – *RGB* colourspace to compute the visible spectrum coverage percentage.
- **samples** (numeric, optional) – Samples count.
- **random\_generator** (generator, optional) – Random triplet generator providing the random samples.
- **random\_state** (RandomState, optional) – Mersenne Twister pseudo-random number generator to use in the random number generator.

**Returns** Percentage coverage of visible spectrum volume.

**Return type** float

## Examples

```
>>> from colour import sRGB_COLOURSPACE as sRGB
>>> prng = np.random.RandomState(2)
>>> RGB_colourspace_visible_spectrum_coverage_MonteCarlo(
...     sRGB, 10e3, random_state=prng)
36...
```

## colour.RGB\_colourspace\_volume\_MonteCarlo

```
colour.RGB_colourspace_volume_MonteCarlo(colourspace, samples=10000000.0, limits=array([[0, 100], [-150, 150], [-150, 150]]), illuminant_Lab=array([ 0.3457, 0.3585]), chromatic_adaptation_method=u'CAT02', random_generator=<function random_triplet_generator>, random_state=None, processes=None)
```

Performs given *RGB* colourspace volume computation using *Monte Carlo* method and multiprocessing.

### Parameters

- **colourspace** (*RGB\_Colourspace*) – *RGB* colourspace to compute the volume of.
- **samples** (numeric, optional) – Samples count.
- **limits** (array\_like, optional) – *Lab* colourspace volume.
- **illuminant\_Lab** (array\_like, optional) – *Lab* colourspace *illuminant* chromaticity coordinates.
- **chromatic\_adaptation\_method** (unicode, optional) – {‘CAT02’, ‘XYZ Scaling’, ‘Von Kries’, ‘Bradford’, ‘Sharp’, ‘Fairchild’, ‘CMCCAT97’, ‘CMCCAT2000’, ‘CAT02\_BRILL\_CAT’, ‘Bianco’, ‘Bianco PC’}, *Chromatic adaptation* method.
- **random\_generator** (generator, optional) – Random triplet generator providing the random samples within the *Lab* colourspace volume.
- **random\_state** (RandomState, optional) – Mersenne Twister pseudo-random number generator to use in the random number generator.
- **processes** (integer, optional) – Processes count, default to `multiprocessing.cpu_count()` definition.

**Returns** *RGB* colourspace volume.

**Return type** `float`

### Notes

- The doctest is assuming that `np.random.RandomState()` definition will return the same sequence no matter which *OS* or *Python* version is used. There is however no formal promise about the *prng* sequence reproducibility of either *Python* or *Numpy* implementations: Laurent. (2012). Reproducibility of python pseudo-random numbers across systems and versions? Retrieved January 20, 2015, from <http://stackoverflow.com/questions/8786084/reproducibility-of-python-pseudo-random-numbers-across-systems-and-versions>

### Examples

```
>>> from colour import sRGB_COLOURSPACE as sRGB
>>> prng = np.random.RandomState(2)
>>> processes = 1
>>> RGB_colourspace_volume_MonteCarlo(sRGB, 10e3, random_state=prng,
...                                         processes=processes)
...
858...
```

**colour.RGB\_colourspace\_volume\_coverage\_MonteCarlo**

```
colour.RGB_colourspace_volume_coverage_MonteCarlo(colourspace, coverage_sampler,
                                                    samples=10000000.0, ran-
                                                    dom_generator=<function ran-
                                                    dom_triplet_generator>, ran-
                                                    dom_state=None)
```

Returns given *RGB* colourspace percentage coverage of an arbitrary volume.

**Parameters**

- **colourspace** (*RGB\_Colourspace*) – *RGB* colourspace to compute the volume coverage percentage.
- **coverage\_sampler** (*object*) – Python object responsible for checking the volume coverage.
- **samples** (*numeric, optional*) – Samples count.
- **random\_generator** (*generator, optional*) – Random triplet generator providing the random samples.
- **random\_state** (*RandomState, optional*) – Mersenne Twister pseudo-random number generator to use in the random number generator.

**Returns** Percentage coverage of volume.

**Return type** *float*

**Examples**

```
>>> from colour import sRGB_COLOURSPACE as sRGB
>>> prng = np.random.RandomState(2)
>>> RGB_colourspace_volume_coverage_MonteCarlo(
...     sRGB, is_within_pointer_gamut, 10e3, random_state=prng)
...
83...
```

**Visible Spectrum****colour**


---

**is\_within\_visible\_spectrum**(*XYZ*[, *cmfs*, ...])

Returns if given *CIE XYZ* tristimulus values are within visible spectrum volume / given colour matching functions volume.

---

**colour.is\_within\_spectrum**

```
colour.is_within_spectrum(XYZ, cmfs=XYZ_ColourMatchingFunctions(name='CIE 1931 2
                                                               Degree Standard Observer', ...), tolerance=None)
```

Returns if given *CIE XYZ* tristimulus values are within visible spectrum volume / given colour matching functions volume.

**Parameters**

- `XYZ` (`array_like`) – *CIE XYZ* tristimulus values.
- `cmfs` (`XYZ_ColourMatchingFunctions`) – Standard observer colour matching functions.
- `tolerance` (`numeric`, `optional`) – Tolerance allowed in the inside-triangle check.

**Returns** Is within visible spectrum.

**Return type** `bool`

## Notes

- Input *CIE XYZ* tristimulus values are in domain [0, 1].

## Examples

```
>>> import numpy as np
>>> is_within_visible_spectrum(np.array([0.3205, 0.4131, 0.51]))
array(True, dtype=bool)
>>> a = np.array([[0.3205, 0.4131, 0.51],
...                 [-0.0005, 0.0031, 0.001]])
>>> is_within_visible_spectrum(a)
array([ True, False], dtype=bool)
```

### 3.1.2.2 Indices and tables

- `genindex`
- `search`

### 3.1.3 Bibliography

#### 3.1.3.1 Indirect References

Some extra references used in the codebase but not directly part of the public api:

- [\[Cen14e\]](#)
- [\[Cen14k\]](#)
- [\[Cen14h\]](#)
- [\[Cen14c\]](#)
- [\[Cen14j\]](#)
- [\[Cen14i\]](#)
- [\[Cen14g\]](#)
- [\[Cen14d\]](#)
- [\[Cen14f\]](#)
- [\[Cen14b\]](#)
- [\[Cen14a\]](#)

- [CIET13805d]
- [Hou15]
- [Lau12]
- [Mac35]
- [MunsellCScienceb]
- [Poi80]
- [RenewableRDCenter03]
- [SWD05]
- [WSOOh]
- [WSO0j]
- [WSO0f]

## 3.2 Examples

### Chromatic Adaptation

```
>>> import colour
>>> XYZ = [0.07049534, 0.10080000, 0.09558313]
>>> A = colour.ILLUMINANTS['CIE 1931 2 Degree Standard Observer']['A']
>>> D65 = colour.ILLUMINANTS['CIE 1931 2 Degree Standard Observer']['D65']
>>> colour.chromatic_adaptation(
...     XYZ, colour.xy_to_XYZ(A), colour.xy_to_XYZ(D65))
array([ 0.08398225,  0.11413379,  0.28629643])
>>> sorted(colour.CHROMATIC_ADAPTATION_METHODS.keys())
['CIE 1994', 'CMCCAT2000', 'Fairchild 1990', 'Von Kries']
```

### Algebra

```
>>> import colour
>>> y = [5.9200, 9.3700, 10.8135, 4.5100, 69.5900, 27.8007, 86.0500]
>>> x = range(len(y))
>>> colour.KernelInterpolator(x, y)([0.25, 0.75, 5.50])
array([-6.18062083,  8.08238488,  57.85783403])
>>> colour.SpragueInterpolator(x, y)([0.25, 0.75, 5.50])
array([-6.72951612,  7.81406251,  43.77379185])
```

### Spectral Computations

```
>>> import colour
>>> colour.spectral_to_XYZ(colour.LIGHT_SOURCES_RELATIVE_SPDS['Neodium Incandescent'])
array([-36.94726204,  32.62076174,  13.0143849])
>>> sorted(colour.SPECTRAL_TO_XYZ_METHODS.keys())
[u'ASTM E308-15', u'Integration', u'astm2015']
```

### Blackbody Spectral Radiance Computation

```
>>> import colour
>>> colour.blackbody_spd(5000)
SpectralPowerDistribution([[ 3.60000000e+02,  6.65427827e+12],
                           [-3.61000000e+02,  6.70960528e+12],
```

```
[ 3.62000000e+02,   6.76482512e+12],  
...  
[ 7.78000000e+02,   1.06068004e+13],  
[ 7.79000000e+02,   1.05903327e+13],  
[ 7.80000000e+02,   1.05738520e+13]],  
interpolator=SpragueInterpolator,  
interpolator_args={},  
extrapolator=Extrapolator,  
extrapolator_args={u'right': None, u'method': u'Constant', u'left': None})
```

## Dominant, Complementary Wavelength & Colour Purity Computation

```
>>> import colour  
>>> xy = [0.26415, 0.37770]  
>>> xy_n = [0.31270, 0.32900]  
>>> colour.dominant_wavelength(xy, xy_n)  
(array(504.0),  
 array([ 0.00369694,  0.63895775]),  
 array([ 0.00369694,  0.63895775]))
```

## Lightness Computation

```
>>> import colour  
>>> colour.lightness(10.08)  
24.902290269546651  
>>> sorted(colour.LIGHTNESS_METHODS.keys())  
[u'CIE 1976',  
 u'Fairchild 2010',  
 u'Glasser 1958',  
 u'Lstar1976',  
 u'Wyszecki 1963']
```

## Luminance Computation

```
>>> import colour  
>>> colour.luminance(37.98562910)  
10.080000001314646  
>>> sorted(colour.LUMINANCE_METHODS.keys())  
[u'ASTM D1535-08',  
 u'CIE 1976',  
 u'Fairchild 2010',  
 u'Newhall 1943',  
 u'astm2008',  
 u'cie1976']
```

## Whiteness Computation

```
>>> import colour  
>>> colour.whiteness(xy=[0.3167, 0.3334], Y=100, xy_n=[0.3139, 0.3311])  
array([-93.85, -1.305])  
>>> sorted(colour.WHITENESS_METHODS.keys())  
[u'ASTM E313',  
 u'Berger 1959',  
 u'CIE 2004',  
 u'Ganz 1979',  
 u'Stensby 1968',  
 u'Taube 1960',  
 u'cie2004']
```

## Yellowness Computation

```
>>> import colour
>>> XYZ = [95.00000000, 100.00000000, 105.00000000]
>>> colour.yellowness(XYZ)
11.065000000000003
>>> sorted(colour.YELLOWNESS_METHODS.keys())
[u'ASTM D1925', u'ASTM E313']
```

## Luminous Flux, Efficiency & Efficacy Computation

```
>>> import colour
>>> spd = colour.LIGHT_SOURCES_RELATIVE_SPDS['Neodium Incandescent']
>>> colour.luminous_flux(spd)
3807.655527367202
>>> colour.luminous_efficiency(spd)
0.19943935624521045
>>> colour.luminous_efficiency(spd)
136.21708031547874
```

## Colour Models

```
>>> import colour
>>> XYZ = [0.07049534, 0.10080000, 0.09558313]
>>> colour.XYZ_to_Lab(XYZ)
array([ 37.9856291, -23.62907688, -4.41746615])
>>> colour.XYZ_to_Luv(XYZ)
array([ 37.9856291, -28.80219593, -1.35800706])
>>> colour.XYZ_to_UCS(XYZ)
array([ 0.04699689, 0.1008, 0.1637439])
>>> colour.XYZ_to_UVW(XYZ)
array([ 4.0680797, 0.12787175, -5.36516614])
>>> colour.XYZ_to_xyY(XYZ)
array([ 0.26414772, 0.37770001, 0.1008])
>>> colour.XYZ_to_hdr_CIELab(XYZ)
array([ 24.90206646, -46.83127607, -10.14274843])
>>> colour.XYZ_to_hdr_IPT(XYZ)
array([ 25.18261761, -22.62111297, 3.18511729])
>>> colour.XYZ_to_Hunter_Lab([7.049534, 10.080000, 9.558313])
array([ 31.74901573, -15.11462629, -2.78660758])
>>> colour.XYZ_to_Hunter_Rdab([7.049534, 10.080000, 9.558313])
array([ 10.08, -18.67653764, -3.44329925])
>>> colour.XYZ_to_IPT(XYZ)
array([ 0.36571124, -0.11114798, 0.01594746])

>>> XYZ = np.array([19.01, 20.00, 21.78])
>>> XYZ_w = np.array([95.05, 100.00, 108.88])
>>> L_A = 318.31
>>> Y_b = 20.0
>>> surround = colour.CIECAM02_VIEWING_CONDITIONS['Average']
>>> specification = colour.XYZ_to_CIECAM02(
    XYZ, XYZ_w, L_A, Y_b, surround)
>>> JMh = (specification.J, specification.M, specification.h)
>>> colour.JMh_CIECAM02_to_CAM02UCS(JMh)
array([ 54.90433134, -0.08442362, -0.06848314])
>>> specification = colour.XYZ_to_CAM16(
    XYZ, XYZ_w, L_A, Y_b, surround)
>>> JMh = (specification.J, specification.M, specification.h)
>>> colour.JMh_CAM16_to_CAM16UCS(JMh)
```

```
array([ 54.89102616, -9.42910274, -5.52845976])

>>> XYZ = [0.07049534, 0.10080000, 0.09558313]
>>> illuminant_XYZ = [0.34570, 0.35850]
>>> illuminant_RGB = [0.31270, 0.32900]
>>> chromatic_adaptation_transform = 'Bradford'
>>> XYZ_to_RGB_matrix = [
    [3.24062548, -1.53720797, -0.49862860],
    [-0.96893071, 1.87575606, 0.04151752],
    [0.05571012, -0.20402105, 1.05699594]]
>>> colour.XYZ_to_RGB(
    XYZ,
    illuminant_XYZ,
    illuminant_RGB,
    XYZ_to_RGB_matrix,
    chromatic_adaptation_transform)
array([ 0.01100154, 0.12735048, 0.11632713])

>>> colour.RGB_to_ICTCP([0.35181454, 0.26934757, 0.21288023])
array([ 0.09554079, -0.00890639, 0.01389286])

>>> colour.RGB_to_HSV([0.49019608, 0.98039216, 0.25098039])
array([ 0.27867383, 0.744       , 0.98039216])

>>> p = [0.73470, 0.26530, 0.00000, 1.00000, 0.00010, -0.07700]
>>> w = [0.32168, 0.33767]
>>> colour.normalised_primary_matrix(p, w)
array([[ 9.52552396e-01, 0.00000000e+00, 9.36786317e-05],
       [ 3.43966450e-01, 7.28166097e-01, -7.21325464e-02],
       [ 0.00000000e+00, 0.00000000e+00, 1.00882518e+00]])

>>> colour.RGB_to_Prismatic([0.25, 0.50, 0.75])
array([ 0.75       , 0.16666667, 0.33333333, 0.5       ])

>>> colour.RGB_to_YCbCr([1.0, 1.0, 1.0])
array([ 0.92156863, 0.50196078, 0.50196078])
```

## RGB Colourspaces

```
>>> import colour
>>> sorted(colour.RGB_COLOURSPACES.keys())
[u'ACES2065-1',
 u'ACEScc',
 u'ACEScct',
 u'ACEScg',
 u'ACESproxy',
 u'ALEXA Wide Gamut',
 u'Adobe RGB (1998)',
 u'Adobe Wide Gamut RGB',
 u'Apple RGB',
 u'Best RGB',
 u'Beta RGB',
 u'CIE RGB',
 u'Cinema Gamut',
 u'ColorMatch RGB',
 u'DCI-P3',
 u'DCI-P3+',
 u'DRAGONcolor',
```

```
u'DRAGONcolor2',
u'Don RGB 4',
u'ECI RGB v2',
u'ERIMM RGB',
u'Ekta Space PS 5',
u'ITU-R BT.2020',
u'ITU-R BT.470 - 525',
u'ITU-R BT.470 - 625',
u'ITU-R BT.709',
u'Max RGB',
u'NTSC',
u'Pal/Secam',
u'ProPhoto RGB',
u'Protune Native',
u'REDWideGamutRGB',
u'REDcolor',
u'REDcolor2',
u'REDcolor3',
u'REDcolor4',
u'RIMM RGB',
u'ROMM RGB',
u'Russell RGB',
u'S-Gamut',
u'S-Gamut3',
u'S-Gamut3.Cine',
u'SMPTE 240M',
u'V-Gamut',
u'Xtreme RGB',
'aces',
'adobe1998',
'prophoto',
'u'sRGB']
```

## OETFs

```
>>> import colour
>>> sorted(colour.OETFS.keys())
['ARIB STD-B67',
 'DCI-P3',
 'DICOM GSDF',
 'ITU-R BT.2020',
 'ITU-R BT.2100 HLG',
 'ITU-R BT.2100 PQ',
 'ITU-R BT.601',
 'ITU-R BT.709',
 'ProPhoto RGB',
 'RIMM RGB',
 'ROMM RGB',
 'SMPTE 240M',
 'ST 2084',
 'sRGB']
```

## EOTFs

```
>>> import colour
>>> sorted(colour.EOTFS.keys())
['DCI-P3',
 'DICOM GSDF',
```

```
'ITU-R BT.1886',
'ITU-R BT.2020',
'ITU-R BT.2100 HLG',
'ITU-R BT.2100 PQ',
'ProPhoto RGB',
'RIMM RGB',
'ROMM RGB',
'SMPTE 240M',
'ST 2084']
```

## OOTFs

```
>>> import colour
>>> sorted(colour.OOTFS.keys())
['ITU-R BT.2100 HLG', 'ITU-R BT.2100 PQ']
```

## Log Encoding / Decoding Curves

```
>>> import colour
>>> sorted(colour.LOG_ENCODING_CURVES.keys())
['ACEScc',
 'ACEScct',
 'ACESproxy',
 'ALEXA Log C',
 'Canon Log',
 'Canon Log 2',
 'Canon Log 3',
 'Cineon',
 'ERIMM RGB',
 'Log3G10',
 'Log3G12',
 'PLog',
 'Panalog',
 'Protune',
 'REDLog',
 'REDLogFilm',
 'S-Log',
 'S-Log2',
 'S-Log3',
 'V-Log',
 'ViperLog']
```

## Chromatic Adaptation Models

```
>>> import colour
>>> XYZ = [0.07049534, 0.10080000, 0.09558313]
>>> XYZ_w = [1.09846607, 1.00000000, 0.35582280]
>>> XYZ_wr = [0.95042855, 1.00000000, 1.08890037]
>>> colour.chromatic_adaptation_VonKries(XYZ, XYZ_w, XYZ_wr)
array([ 0.08397461,  0.11413219,  0.28625545])
```

## Colour Appearance Models

```
>>> import colour
>>> XYZ = [19.01, 20.00, 21.78]
>>> XYZ_w = [95.05, 100.00, 108.88]
>>> L_A = 318.31
>>> Y_b = 20.0
```

```
>>> colour.XYZ_to_CIECAM02(XYZ, XYZ_w, L_A, Y_b)
CIECAM02_Specification(J=41.731091132513917, C=0.10470775717103062, h=219.04843265831178, s=2.
↪ 3603053739196032, Q=195.37132596607671, M=0.10884217566914849, H=278.06073585667758, HC=None)
```

## Colour Difference

```
>>> import colour
>>> Lab_1 = [100.0000000, 21.57210357, 272.22819350]
>>> Lab_2 = [100.0000000, 426.67945353, 72.39590835]
>>> colour.delta_E(Lab_1, Lab_2)
94.035649026659485
>>> sorted(colour.DELTA_E_METHODS.keys())
['CAM02-LCD',
 'CAM02-SCD',
 'CAM02-UCS',
 'CAM16-LCD',
 'CAM16-SCD',
 'CAM16-UCS',
 'CIE 1976',
 'CIE 1994',
 'CIE 2000',
 'CMC',
 'cie1976',
 'cie1994',
 'cie2000']
```

## Colour Notation Systems

```
>>> import colour
>>> colour.munsell_value(10.1488096782)
3.7462971142584354
>>> sorted(colour.MUNSELL_VALUE_METHODS.keys())
[u'ASTM D1535-08',
 u'Ladd 1955',
 u'McCamy 1987',
 u'Moon 1943',
 u'Munsell 1933',
 u'Priest 1920',
 u'Saunderson 1944',
 u'astm2008']
>>> colour.xyY_to_munsell_colour([0.38736945, 0.35751656, 0.59362000])
u'4.2YR 8.1/5.3'
>>> colour.munsell_colour_to_xyY('4.2YR 8.1/5.3')
array([ 0.38736945,  0.35751656,  0.59362    ])
```

## Optical Phenomena

```
>>> import colour
>>> colour.rayleigh_scattering_spd()
SpectralPowerDistribution([[ 3.6000000e+02,  5.99101337e-01],
                           [ 3.6100000e+02,  5.92170690e-01],
                           [ 3.6200000e+02,  5.85341006e-01],
                           ...
                           [ 7.7800000e+02,  2.55208377e-02],
                           [ 7.7900000e+02,  2.53887969e-02],
                           [ 7.8000000e+02,  2.52576106e-02]],,
                           interpolator=SpragueInterpolator,
                           interpolator_args={})
```

```
    extrapolator=Extrapolator,
    extrapolator_args={u'right': None, u'method': u'Constant', u'left': None})
```

## Light Quality

```
>>> import colour
>>> colour.colour_quality_scale(colour.ILLUMINANTS_RELATIVE_SPDS['F2'])
64.686416902221609
>>> colour.colour_rendering_index(colour.ILLUMINANTS_RELATIVE_SPDS['F2'])
64.151520202968015
```

## Reflectance Recovery

```
>>> import colour
>>> colour.XYZ_to_spectral([0.07049534, 0.10080000, 0.09558313])
SpectralPowerDistribution([[ 3.6000000e+02,    7.96361498e-04],
                           [ 3.6500000e+02,    7.96489667e-04],
                           [ 3.7000000e+02,    7.96543669e-04],
                           ...
                           [ 8.2000000e+02,    1.71014294e-04],
                           [ 8.2500000e+02,    1.71621924e-04],
                           [ 8.3000000e+02,    1.72026883e-04]],
                           interpolator=SpragueInterpolator,
                           interpolator_args={},
                           extrapolator=Extrapolator,
                           extrapolator_args={u'right': None, u'method': u'Constant', u'left': None})
>>> sorted(colour.REFLECTANCE_RECOVERY_METHODS.keys())
['Meng 2015', 'Smits 1999']
```

## Correlated Colour Temperature Computation Methods

```
>>> import colour
>>> colour.uv_to_CCT([0.1978, 0.3122])
array([ 6.50751282e+03,   3.22335875e-03])
>>> sorted(colour.UV_TO_CCT_METHODS.keys())
[u'Ohno 2013', u'Robertson 1968', u'ohno2013', u'robertson1968']
>>> sorted(colour.UV_TO_CCT_METHODS.keys())
[u'Krystek 1985',
 u'Ohno 2013',
 u'Robertson 1968',
 u'ohno2013',
 u'robertson1968']
>>> sorted(colour.XY_TO_CCT_METHODS.keys())
[u'Hernandez 1999', u'McCamy 1992', u'hernandez1999', u'mccamy1992']
>>> sorted(colour.CCT_TO_XY_METHODS.keys())
[u'CIE Illuminant D Series', u'Kang 2002', u'cie_d', u'kang2002']
```

## Volume

```
>>> import colour
>>> colour.RGB_colourspace_volume_MonteCarlo(colour.sRGB_COLOURSPACE)
857011.5
```

---

CHAPTER  
**FOUR**

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## **CONTRIBUTING**

If you would like to contribute to Colour, please refer to the following [Contributing](#) guide.



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**CHAPTER  
FIVE**

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**CHANGES**

The changes are viewable on the [Releases](#) page.



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**CHAPTER  
SIX**

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## **BIBLIOGRAPHY**

The bibliography is available on the [Bibliography](#) page.

It is also viewable directly from the repository in [BibTeX](#) format.



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CHAPTER  
SEVEN

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SEE ALSO

Here is a list of notable colour science packages sorted by languages:

**Python**

- [ColorPy](#) by Kness, M.
- [Colorsapcious](#) by Smith, N. J., et al.
- [python-colormath](#) by Taylor, G., et al.

**.NET**

- [Colourful](#) by Pažourek, T., et al.

**Julia**

- [Colors.jl](#) by Holy, T., et al.

**Matlab & Octave**

- [COLORLAB](#) by Malo, J., et al.
- [Psychtoolbox](#) by Brainard, D., et al.
- [The Munsell and Kubelka-Munk Toolbox](#) by Centore, P.



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**CHAPTER  
EIGHT**

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**ABOUT**

**Colour** by Colour Developers - 2013-2018

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This software is released under terms of New BSD License: <http://opensource.org/licenses/BSD-3-Clause>

<http://github.com/colour-science/colour>



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- [ARR12] ARRI. ALEXA - Log C Curve - Usage in VFX. 2012. URL: [http://www.arri.com/?eID=registration&file\\_uid=8026](http://www.arri.com/?eID=registration&file_uid=8026).
- [Bab12a] BabelColor. ColorChecker RGB and spectra. 2012. URL: [http://www.babelcolor.com/download/ColorChecker\\_RGB\\_and\\_spectra.xls](http://www.babelcolor.com/download/ColorChecker_RGB_and_spectra.xls).
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# INDEX

## Symbols

`_add_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_call_()` (colour.KernelInterpolator method), 48  
`_call_()` (colour.LinearInterpolator method), 49  
`_call_()` (colour.NullInterpolator method), 50  
`_call_()` (colour.SpragueInterpolator method), 52  
`_class_()` (colour.Extrapolator method), 46  
`_contains_()` (colour.SpectralShape method), 102  
`_contains_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_contains_()` (colour.continuous.MultiSignal method), 187  
`_contains_()` (colour.continuous.Signal method), 184  
`_contains_()` (colour.utilities.CaseInsensitiveMapping method), 439  
`_contains_()` (in module colour), 331  
`_delattr_()` (colour.utilities.Structure method), 441  
`_delitem_()` (colour.utilities.CaseInsensitiveMapping method), 439  
`_delitem_()` (in module colour), 331  
`_div_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_eq_()` (colour.SpectralShape method), 102  
`_eq_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_eq_()` (colour.continuous.MultiSignal method), 187  
`_eq_()` (colour.continuous.Signal method), 184  
`_eq_()` (colour.utilities.CaseInsensitiveMapping method), 439  
`_eq_()` (in module colour), 332  
`_getattr_()` (colour.utilities.Structure method), 441  
`_getitem_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_getitem_()` (colour.continuous.MultiSignal method), 187  
`_getitem_()` (colour.continuous.Signal method), 184  
`_getitem_()` (colour.utilities.CaseInsensitiveMapping method), 439  
`_getitem_()` (in module colour), 331  
`_hash_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_iadd_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_idiv_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_imul_()` (colour.continuous.AbstractContinuousFunction method), 182  
`_init_()` (colour.ATD95\_Specification method), 66  
`_init_()` (colour.CAM16\_Specification method), 74  
`_init_()` (colour.CIECAM02\_Specification method), 70  
`_init_()` (colour.Extrapolator method), 47  
`_init_()` (colour.Hunt\_Specification method), 77  
`_init_()` (colour.IES\_TM2714\_Spd method), 213  
`_init_()` (colour.KernelInterpolator method), 49  
`_init_()` (colour.LLAB\_Specification method), 80  
`_init_()` (colour.LinearInterpolator method), 50  
`_init_()` (colour.MultiSpectralPowerDistribution method), 100  
`_init_()` (colour.Nayatani95\_Specification method), 83  
`_init_()` (colour.NullInterpolator method), 51  
`_init_()` (colour.PchipInterpolator method), 51  
`_init_()` (colour.RGB\_Colourspace method), 256  
`_init_()` (colour.RLAB\_Specification method), 85  
`_init_()` (colour.SpectralPowerDistribution method), 96, 97  
`_init_()` (colour.SpectralShape method), 102  
`_init_()` (colour.SpragueInterpolator method), 52  
`_init_()` (colour.adaptation.CMCCAT2000\_InductionFactors method), 41  
`_init_()` (colour.algebra.LineSegmentsIntersections\_Specification method), 62  
`_init_()` (colour.appearance.CAM16\_InductionFactors method), 75  
`_init_()` (colour.appearance.CIECAM02\_InductionFactors method), 71  
`_init_()` (colour.appearance.LLAB\_InductionFactors

method), 81  
\_init\_(colour.characterisation.RGB\_DisplayPrimaries repr\_0 (colour.continuous.MultiSignal method),  
method), 93  
\_init\_(colour.characterisation.RGB\_SpectralSensitivities repr\_0 (colour.continuous.Signal method), 183  
method), 91  
\_repr\_0 (colour.utilities.CaseInsensitiveMapping  
method), 439  
\_init\_(colour.colorimetry.LMS\_ConeFundamentals  
method), 131  
\_repr\_0 (in module colour), 332  
\_init\_(colour.colorimetry.RGB\_ColourMatchingFunction repr\_0 (colour.utilities.Structure method), 441  
method), 133  
\_repr\_0 (colour.continuous.AbstractContinuousFunction  
method), 182  
\_init\_(colour.colorimetry.XYZ\_ColourMatchingFunctions  
method), 182  
\_setitem\_(colour.continuous.MultiSignal  
method), 187  
\_init\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_setitem\_(colour.continuous.Signal method),  
184  
\_setitem\_(colour.utilities.CaseInsensitiveMapping  
method), 439  
\_init\_(colour.continuous.Signal method), 185  
\_init\_(colour.quality.CQS\_Specification  
method), 405  
\_init\_(colour.quality.CRI\_Specification method),  
403  
\_init\_(colour.utilities.CaseInsensitiveMapping  
method), 439  
\_init\_(colour.utilities.Lookup method), 440  
\_init\_(colour.utilities.Structure method), 442  
\_ipow\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_ipow\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_isub\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_iter\_(colour.SpectralShape method), 102  
\_iter\_(colour.utilities.CaseInsensitiveMapping  
method), 439  
\_iter\_(in module colour), 331  
\_len\_(colour.SpectralShape method), 102  
\_len\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_len\_(colour.utilities.CaseInsensitiveMapping  
method), 439  
\_len\_(in module colour), 331  
\_mul\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_ne\_(colour.SpectralShape method), 102  
\_ne\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_ne\_(colour.continuous.MultiSignal method),  
187  
\_ne\_(colour.continuous.Signal method), 184  
\_ne\_(colour.utilities.CaseInsensitiveMapping  
method), 439  
\_ne\_(in module colour), 332  
\_pow\_(colour.continuous.AbstractContinuousFunction  
method), 182  
\_repr\_(colour.RGB\_Colourspace method), 255  
\_repr\_(colour.SpectralShape method), 102  
\_repr\_(colour.continuous.AbstractContinuousFunction  
method), 182  
method), 182  
\_repr\_0 (colour.continuous.Signal method), 187  
\_repr\_0 (colour.utilities.CaseInsensitiveMapping  
method), 439  
\_repr\_0 (in module colour), 332  
\_repr\_0 (colour.continuous.AbstractContinuousFunction  
method), 182  
\_repr\_0 (colour.continuous.Signal method),  
184  
\_repr\_0 (colour.utilities.CaseInsensitiveMapping  
method), 439  
\_repr\_0 (in module colour), 331  
\_str\_(colour.RGB\_Colourspace method), 255  
\_str\_(colour.SpectralShape method), 101  
\_str\_(colour.continuous.AbstractContinuousFunction  
method), 181  
\_str\_(colour.continuous.MultiSignal method),  
187  
\_str\_(colour.continuous.Signal method), 183  
\_str\_(colour.continuous.AbstractContinuousFunction  
method), 182

**A**

absolute\_tolerance (colour.NullInterpolator attribute), 50  
AbstractContinuousFunction (class) in colour.continuous, 181  
ACES\_2065\_1\_COLOURSPACE (in module colour.models), 258  
ACES\_CC\_COLOURSPACE (in module colour.models), 258  
ACES\_CCT\_COLOURSPACE (in module colour.models), 258  
ACES(CG\_COLOURSPACE (in module colour.models), 259  
ACES\_PROXY\_COLOURSPACE (in module colour.models), 259  
ACES\_RICD (in module colour.models), 271  
adjust\_tristimulus\_weighting\_factors\_ASTME308150 (in module colour.colorimetry), 124  
ADOBE\_RGB\_1998\_COLOURSPACE (in module colour.models), 259  
ADOBE\_WIDE\_GAMUT\_RGB\_COLOURSPACE (in module colour.models), 260  
AIXA\_WIDE\_GAMUT\_COLOURSPACE (in module colour.models), 260  
align() (colour.MultiSpectralPowerDistribution method), 99

align() (colour.SpectralPowerDistribution method), 96	BT470_525_COLOURSPACE (in module colour.models), 261	(in module colour.models), 261
APPLE_RGB_COLOURSPACE (in module colour.models), 260	BT470_625_COLOURSPACE (in module colour.models), 261	(in module colour.models), 261
arithmetical_operation() (colour.continuous.AbstractContinuousFunction method), 182	BT709_COLOURSPACE (in module colour.models), 261	
arithmetical_operation() (colour.continuous.MultiSignal method), 187	C	
arithmetical_operation() (colour.continuous.Signal method), 184	CAM02LCD_to_JMh_CIECAM02() (in module colour), 234	(in module colour), 234
as_namedtuple() (in module colour.utilities), 429	CAM02SCD_to_JMh_CIECAM02() (in module colour), 235	(in module colour), 235
as_numeric() (in module colour.utilities), 428	CAM02UCS_to_JMh_CIECAM02() (in module colour), 236	(in module colour), 236
ASTME30815_PRACTISE_SHAPE (in module colour), 102	CAM16_InductionFactors (class in colour.appearance), 74	
ATD95_Specification (class in colour), 66	CAM16_Specification (class in colour), 73	
AVOGADRO_CONSTANT (in module colour.constants), 179	CAM16_to_XYZ() (in module colour), 72	
B	CAM16_VIEWING_CONDITIONS (in module colour), 74	
bandpass_correction() (in module colour), 129	CAM16LCD_to_JMh_CAM16 (in module colour), 237	
BANDPASS_CORRECTION_METHODS (in module colour), 129	CAM16SCD_to_JMh_CAM16 (in module colour), 238	
bandpass_correction_Stearns1988() (in module colour.colorimetry), 130	CAM16UCS_to_JMh_CAM16 (in module colour), 239	
bandwidth_corrected (colour.IES_TM2714_Spd attribute), 212	camera() (in module colour.plotting), 367	
bandwidth_FWHM (colour.IES_TM2714_Spd attribute), 212	CAMERAS_RGB_SPECTRAL_SENSITIVITIES (in module colour), 92	
batch() (in module colour.utilities), 424	canvas() (in module colour.plotting), 367	
BEST_RGB_COLOURSPACE (in module colour.models), 260	cartesian_to_cylindrical() (in module colour.algebra), 58	
BETA_RGB_COLOURSPACE (in module colour.models), 261	cartesian_to_polar() (in module colour.algebra), 57	
blackbody_colours_plot() (in module colour.plotting), 377	cartesian_to_spherical() (in module colour.algebra), 56	
blackbody_spd() (in module colour), 103	CaseInsensitiveMapping (class in colour.utilities), 439	
blackbody_spectral_radiance() (in module colour.colorimetry), 116	CAT02_BRILL_CAT (in module colour.adaptation), 43	
blackbody_spectral_radiance_plot() (in module colour.plotting), 377	CAT02_CAT (in module colour.adaptation), 44	
BOLTZMANN_CONSTANT (in module colour.constants), 180	CCT_to_uv() (in module colour), 412	
boundaries (colour.SpectralShape attribute), 101	CCT_to_uv_Krystek1985() (in module colour.temperature), 418	
boundaries() (in module colour.plotting), 368	CCT_TO_UV_METHODS (in module colour), 413	
BRADFORD_CAT (in module colour.adaptation), 43	CCT_to_uv_Ohno2013() (in module colour.temperature), 419	
BRENEMAN_EXPERIMENTS (in module colour), 192	CCT_to_uv_Robertson1968() (in module colour.temperature), 417	
BRENEMAN_EXPERIMENTS_PRIMARIES_CHROMATICITIES_xy() (in module colour), 192	CIE_xy() (in module colour), 415	
BS_CAT (in module colour.adaptation), 43	CCT_to_xy_CIE_D() (in module colour.temperature), 422	
BS_PC_CAT (in module colour.adaptation), 43	CCT_to_xy_Kang2002() (in module colour.temperature), 421	
BT2020_COLOURSPACE (in module colour.models), 262	CCT_TO_XY_METHODS (in module colour), 415	

centroid() (in module colour.utilities), 437  
chromatic\_adaptation() (in module colour), 31  
chromatic\_adaptation\_CIE1994() (in module colour.adaptation), 36  
chromatic\_adaptation\_CMCCAT2000() (in module colour.adaptation), 37  
chromatic\_adaptation\_Fairchild1990() (in module colour.adaptation), 35  
chromatic\_adaptation\_forward\_CMCCAT2000() (in module colour.adaptation), 39  
chromatic\_adaptation\_matrix\_VonKries() (in module colour.adaptation), 44  
CHROMATIC\_ADAPTATION\_METHODS (in module colour), 33  
chromatic\_adaptation\_reverse\_CMCCAT2000() (in module colour.adaptation), 40  
CHROMATIC\_ADAPTATION\_TRANSFORMS (in module colour), 34  
CHROMATIC\_ADAPTATION\_TRANSFORMS (in module colour.adaptation), 42  
chromatic\_adaptation\_VonKries() (in module colour.adaptation), 42  
chromatically\_adapted\_primaries() (in module colour), 251  
chromaticity\_diagram\_plot\_CIE1931() (in module colour.plotting), 380  
chromaticity\_diagram\_plot\_CIE1960UCS() (in module colour.plotting), 380  
chromaticity\_diagram\_plot\_CIE1976UCS() (in module colour.plotting), 381  
CIE\_E (in module colour.constants), 178  
CIE\_K (in module colour.constants), 178  
CIE\_RGB\_COLOURSPACE (in module colour.models), 262  
CIE\_standard\_illuminant\_A\_function() (in module colour), 112  
CIECAM02\_InductionFactors (class in colour.appearance), 70  
CIECAM02\_Specification (class in colour), 69  
CIECAM02\_to\_XYZ() (in module colour), 68  
CIECAM02\_VIEWING\_CONDITIONS (in module colour), 70  
CINEMA\_GAMUT\_COLOURSPACE (in module colour.models), 262  
closest() (in module colour.utilities), 430  
closest\_indexes() (in module colour.utilities), 429  
CMCCAT2000\_CAT (in module colour.adaptation), 44  
CMCCAT2000\_InductionFactors (class in colour.adaptation), 41  
CMCCAT2000\_VIEWING\_CONDITIONS (in module colour), 34  
CMCCAT2000\_VIEWING\_CONDITIONS (in module colour.adaptation), 38  
CMCCAT97\_CAT (in module colour.adaptation), 44  
CMFS (in module colour), 135  
CMY\_to\_CMYK() (in module colour), 343  
CMY\_to\_RGB() (in module colour), 342  
CMYK\_to\_CMY() (in module colour), 343  
COLOR\_MATCH\_RGB\_COLOURSPACE (in module colour.models), 262  
colorimetric\_purity() (in module colour), 145  
colour\_checker\_plot() (in module colour.plotting), 378  
colour\_cycle() (in module colour.plotting), 366  
colour\_plotting\_defaults() (in module colour.plotting), 366  
colour\_quality\_scale() (in module colour), 404  
colour\_rendering\_index() (in module colour), 402  
COLOURCHECKERS (in module colour), 89  
COLOURCHECKERS\_SPDS (in module colour), 90  
ColourWarning, 445  
complementary\_wavelength() (in module colour), 143  
constant\_spd() (in module colour), 114  
copy() (colour.continuous.AbstractContinuousFunction method), 182  
copy() (colour.utilities.CaseInsensitiveMapping method), 439  
copy() (in module colour), 332  
corresponding\_chromaticities\_prediction() (in module colour), 191  
corresponding\_chromaticities\_prediction\_CIE1994() (in module colour.corresponding), 194  
corresponding\_chromaticities\_prediction\_CMCCAT2000() (in module colour.corresponding), 195  
corresponding\_chromaticities\_prediction\_Fairchild1990() (in module colour.corresponding), 193  
CORRESPONDING\_CHROMATICITIES\_PREDICTION\_MODELS (in module colour), 192  
corresponding\_chromaticities\_prediction\_plot() (in module colour.plotting), 379  
corresponding\_chromaticities\_prediction\_VonKries() (in module colour.corresponding), 195  
CQS\_Specification (class in colour.quality), 404  
CRI\_Specification (class in colour.quality), 403  
cube() (in module colour.plotting), 401  
CV\_range() (in module colour), 336  
cylindrical\_to\_cartesian() (in module colour.algebra), 59

**D**

D\_illuminant\_relative\_spd() (in module colour), 112  
DCI\_P3\_COLOURSPACE (in module colour.models), 263  
DCI\_P3\_P\_COLOURSPACE (in module colour.models), 263

decoding\_cctf (colour.RGB\_Colourspace attribute), 255  
 decorate() (in module colour.plotting), 367  
 default (colour.NullInterpolator attribute), 50  
 DEFAULT\_FLOAT\_DTYPE (in module colour.constants), 180  
 DEFAULT\_SPECTRAL\_SHAPE (in module colour), 102  
 delta\_E() (in module colour), 197  
 delta\_E\_CAM02LCD() (in module colour.difference), 202  
 delta\_E\_CAM02SCD() (in module colour.difference), 203  
 delta\_E\_CAM02UCS() (in module colour.difference), 203  
 delta\_E\_CAM16LCD() (in module colour.difference), 204  
 delta\_E\_CAM16SCD() (in module colour.difference), 204  
 delta\_E\_CAM16UCS() (in module colour.difference), 205  
 delta\_E\_CIE1976() (in module colour.difference), 198  
 delta\_E\_CIE1994() (in module colour.difference), 199  
 delta\_E\_CIE2000() (in module colour.difference), 200  
 delta\_E\_CMC() (in module colour.difference), 201  
 DELTA\_E\_METHODS (in module colour), 198  
 display() (in module colour.plotting), 368  
 DISPLAYS\_RGB\_PRIMARIES (in module colour), 94  
 domain (colour.continuous.AbstractContinuousFunction attribute), 181  
 domain (colour.continuous.MultiSignal attribute), 187  
 domain (colour.continuous.Signal attribute), 183  
 domain\_distance() (colour.continuous.AbstractContinuousFunction method), 182  
 dominant\_wavelength() (in module colour), 142  
 DON\_RGB\_4\_COLOURSPACE (in module colour.models), 263  
 dot\_matrix() (in module colour.utilities), 435  
 dot\_vector() (in module colour.utilities), 435  
 DRAGON\_COLOR\_2\_COLOURSPACE (in module colour.models), 267  
 DRAGON\_COLOR\_COLOURSPACE (in module colour.models), 266  
 dtype (colour.continuous.MultiSignal attribute), 186  
 dtype (colour.continuous.Signal attribute), 183

**E**

ECI\_RGB\_V2\_COLOURSPACE (in module colour.models), 263

**F**

FAIRCHILD\_CAT (in module colour.adaptation), 44

EKTA\_SPACE\_PS\_5\_COLOURSPACE (in module colour.models), 264  
 encoding\_cctf (colour.RGB\_Colourspace attribute), 255  
 end (colour.SpectralShape attribute), 101  
 eotf() (in module colour), 285  
 eotf\_BT1886() (in module colour.models), 289  
 eotf\_BT2020() (in module colour.models), 290  
 eotf\_BT2100\_HLG() (in module colour.models), 290  
 eotf\_BT2100\_PQ() (in module colour.models), 292  
 eotf\_DCIP3() (in module colour.models), 288  
 eotf\_DICOMGSDF() (in module colour.models), 288  
 eotf\_ProPhotoRGB() (in module colour.models), 293  
 eotf\_reverse() (in module colour), 287  
 eotf\_reverse\_BT1886() (in module colour.models), 289  
 eotf\_reverse\_BT2100\_HLG() (in module colour.models), 291  
 eotf\_reverse\_BT2100\_PQ() (in module colour.models), 292  
 eotf\_RIMMRGB() (in module colour.models), 293  
 eotf\_ROMMRGB() (in module colour.models), 294  
 eotf\_SMPTE240M() (in module colour.models), 294  
 eotf\_ST2084() (in module colour.models), 295  
 EOTFS (in module colour), 286  
 EOTFS\_REVERSE (in module colour), 287  
 EPSILON (in module colour.constants), 180  
 equal\_axes3d() (in module colour.plotting), 369  
 ERIMM\_RGB\_COLOURSPACE (in module colour.models), 267  
 euclidean\_distance() (in module colour.algebra), 60  
 extrapolate() (colour.MultiSpectralPowerDistribution method), 99  
 extrapolator() (colour.SpectralPowerDistribution method), 96  
 Extrapolator (class in colour), 46  
 extrapolator (colour.continuous.AbstractContinuousFunction attribute), 181  
 extrapolator (colour.continuous.MultiSignal attribute), 187  
 extrapolator (colour.continuous.Signal attribute), 183  
 extrapolator\_args (colour.continuous.AbstractContinuousFunction attribute), 181  
 extrapolator\_args (colour.continuous.MultiSignal attribute), 187  
 extrapolator\_args (colour.continuous.Signal attribute), 183

fill\_nan() (colour.continuous.AbstractContinuousFunction method), 182

fill\_nan() (colour.continuous.MultiSignal method), 187

fill\_nan() (colour.continuous.Signal method), 184

fill\_nan() (in module colour.utilities), 438

filter\_kwargs() (in module colour.utilities), 427

filter\_warnings() (in module colour.utilities), 444

first\_item() (in module colour.utilities), 427

first\_key\_from\_value() (colour.utilities.Lookup method), 440

first\_order\_colour\_fit() (in module colour), 88

FLOATING\_POINT\_NUMBER\_PATTERN (in module colour.constants), 180

full\_to\_legal() (in module colour), 334

function (colour.continuous.AbstractContinuousFunction attribute), 181

function (colour.continuous.MultiSignal attribute), 187

function (colour.continuous.Signal attribute), 183

function\_gamma() (in module colour), 284

function\_linear() (in module colour), 285

**G**

grid() (in module colour.plotting), 400

**H**

handle\_numpy\_errors() (in module colour.utilities), 423

HDR\_CIELAB\_METHODS (in module colour), 245

hdr\_CIELab\_to\_XYZ() (in module colour), 244

HDR\_IPT\_METHODS (in module colour), 242

hdr\_IPT\_to\_XYZ() (in module colour), 242

header (colour.IES\_TM2714\_Spd attribute), 212

HEX\_to\_RGB() (in module colour.notation), 353

HSL\_to\_RGB() (in module colour), 341

HSV\_to\_RGB() (in module colour), 339

Hunt\_Specification (class in colour), 77

HUNT\_VIEWING\_CONDITIONS (in module colour), 77

Hunter\_Lab\_to\_XYZ() (in module colour), 231

HUNTERLAB\_ILLUMINANTS (in module colour), 141

**I**

ICTCP\_to\_RGB() (in module colour), 337

IES\_TM2714\_Spd (class in colour), 211

ignore\_numpy\_errors() (in module colour.utilities), 423

ignore\_python\_warnings() (in module colour.utilities), 424

illuminant (colour.RGB\_Colourspace attribute), 255

ILLUMINANTS (in module colour), 140

ILLUMINANTS\_OPTIMAL\_COLOUR\_STIMULI (in module colour), 446

ILLUMINANTS\_RELATIVE\_SPDS (in module colour), 141

image\_plot() (in module colour.plotting), 371

in\_array() (in module colour.utilities), 432

INTEGER\_THRESHOLD (in module colour.constants), 181

interpolate() (colour.MultiSpectralPowerDistribution method), 99

interpolate() (colour.SpectralPowerDistribution method), 96

interpolator (colour.continuous.AbstractContinuousFunction attribute), 181

interpolator (colour.continuous.MultiSignal attribute), 187

interpolator (colour.continuous.Signal attribute), 183

interpolator\_args (colour.continuous.AbstractContinuousFunction attribute), 181

interpolator\_args (colour.continuous.MultiSignal attribute), 187

interpolator\_args (colour.continuous.Signal attribute), 183

intersect\_line\_segments() (in module colour.algebra), 61

interval (colour.SpectralShape attribute), 101

interval() (in module colour.utilities), 430

IPT\_hue\_angle() (in module colour), 240

IPT\_to\_XYZ() (in module colour), 240

is\_identity() (in module colour.algebra), 63

is\_integer() (in module colour.utilities), 426

is\_iterable() (in module colour.utilities), 425

is\_numeric() (in module colour.utilities), 426

is\_openimageio\_installed() (in module colour.utilities), 425

is\_pandas\_installed() (in module colour.utilities), 425

is\_string() (in module colour.utilities), 425

is\_uniform() (colour.continuous.AbstractContinuousFunction method), 182

is\_uniform() (in module colour.utilities), 431

is\_within\_macadam\_limits() (in module colour), 445

is\_within\_mesh\_volume() (in module colour), 446

is\_within\_pointer\_gamut() (in module colour), 447

is\_within\_visible\_spectrum() (in module colour), 451

**J**

JMh\_CAM16\_to\_CAM16LCD (in module colour), 237

JMh\_CAM16\_to\_CAM16SCD (in module colour), 238

JMh_CAM16_to_CAM16UCS (in module colour),	LineSegmentsIntersections_Specification (class in colour.algebra),	62
JMh_CIECAM02_to_CAM02LCD() (in module colour),	LLAB_InductionFactors (class in colour.appearance),	81
JMh_CIECAM02_to_CAM02SCD() (in module colour),	LLAB_Specification (class in colour),	79
JMh_CIECAM02_to_CAM02UCS() (in module colour),	LLAB_VIEWING_CONDITIONS (in module colour),	80
	LMS_10_degree_cmfs_to_XYZ_10_degree_cmfs()	
	(in module colour.colorimetry),	140
	LMS_2_degree_cmfs_to_XYZ_2_degree_cmfs()	(in module colour.colorimetry),
	139	
	LMS_CMFS (in module colour),	136
	LMS_ConeFundamentals (class in colour.colorimetry),	131
	log_decoding_ACEScc() (in module colour.models),	305
	log_decoding_ACEScct() (in module colour.models),	306
	log_decoding_ACESproxy() (in module colour.models),	307
	log_decoding_ALEXALogC() (in module colour.models),	308
	log_decoding_CanonLog() (in module colour.models),	312
	log_decoding_CanonLog2() (in module colour.models),	309
	log_decoding_CanonLog3() (in module colour.models),	311
	log_decoding_Cineon() (in module colour.models),	313
	log_decoding_curve() (in module colour),	301
	LOG_DECODING_CURVES (in module colour),	303
	log_decoding_ERIMMRGB() (in module colour.models),	314
	log_decoding_Log3G10() (in module colour.models),	315
	log_decoding_Log3G12() (in module colour.models),	316
	log_decoding_Panalog() (in module colour.models),	317
	log_decoding_PivotedLog() (in module colour.models),	318
	log_decoding_Proture() (in module colour.models),	319
	log_decoding_REDLog() (in module colour.models),	320
	log_decoding_REDLogFilm() (in module colour.models),	321
	log_decoding_SLog() (in module colour.models),	322
	log_decoding_SLog2() (in module colour.models),	324
	log_decoding_SLog3() (in module colour.models),	325
K		
K_M (in module colour.constants),	179	
kernel (colour.KernelInterpolator attribute),	48	
kernel_args (colour.KernelInterpolator attribute),	48	
kernel_cardinal_spline() (in module colour),	55	
kernel_lanczos() (in module colour),	55	
kernel_linear() (in module colour),	54	
kernel_nearest_neighbour() (in module colour),	53	
kernel_sinc() (in module colour),	54	
KernelInterpolator (class in colour),	47	
keys_from_value() (colour.utilities.Lookup method),	440	
KP_M (in module colour.constants),	179	
L		
Lab_to_LCHab() (in module colour),	221	
Lab_to_XYZ() (in module colour),	220	
label_rectangles() (in module colour.plotting),	369	
labels (colour.continuous.MultiSignal attribute),	187	
lagrange_coefficients() (in module colour),	52	
lagrange_coefficients_ASTME202211() (in module colour.colorimetry),	125	
LCHab_to_Lab() (in module colour),	222	
LCHuv_to_Luv() (in module colour),	224	
LEFS (in module colour),	155	
legal_to_full() (in module colour),	335	
LIGHT_SOURCES (in module colour),	141	
LIGHT_SOURCES_RELATIVE_SPDS (in module colour),	142	
LIGHT_SPEED (in module colour.constants),	180	
lightness() (in module colour),	156	
lightness_CIE1976() (in module colour.colorimetry),	160	
lightness_Fairchild2010() (in module colour.colorimetry),	160	
lightness_Fairchild2011() (in module colour.colorimetry),	161	
lightness_Glasser1958() (in module colour.colorimetry),	158	
LIGHTNESS_METHODS (in module colour),	158	
lightness_Wyszecki1963() (in module colour.colorimetry),	159	
linear_conversion() (in module colour.utilities),	437	
LinearInterpolator (class in colour),	49	

log\_decoding\_ViperLog() (in module colour.models),  
    327  
log\_decoding\_VLog() (in module colour.models),  
    326  
log\_encoding\_ACEScc() (in module colour.models),  
    304  
log\_encoding\_ACEScct() (in module colour.models),  
    305  
log\_encoding\_ACESproxy() (in module colour.models),  
    306  
log\_encoding\_ALEXALogC() (in module colour.models),  
    308  
log\_encoding\_CanonLog() (in module colour.models),  
    311  
log\_encoding\_CanonLog2() (in module colour.models),  
    309  
log\_encoding\_CanonLog3() (in module colour.models),  
    310  
log\_encoding\_Cineon() (in module colour.models),  
    312  
log\_encoding\_curve() (in module colour), 300  
LOG\_ENCODING\_CURVES (in module colour), 301  
log\_encoding\_ERIMMRGB() (in module colour.models),  
    313  
log\_encoding\_Log3G10() (in module colour.models),  
    314  
log\_encoding\_Log3G12() (in module colour.models),  
    316  
log\_encoding\_Panalog() (in module colour.models),  
    317  
log\_encoding\_PivotedLog() (in module colour.models),  
    318  
log\_encoding\_Proture() (in module colour.models),  
    319  
log\_encoding\_REDLog() (in module colour.models),  
    320  
log\_encoding\_REDLogFilm() (in module colour.models),  
    321  
log\_encoding\_SLog() (in module colour.models),  
    322  
log\_encoding\_SLog2() (in module colour.models),  
    323  
log\_encoding\_SLog3() (in module colour.models),  
    324  
log\_encoding\_ViperLog() (in module colour.models),  
    327  
log\_encoding\_VLog() (in module colour.models),  
    326  
Lookup (class in colour.utilities), 440  
lower\_items() (colour.utilities.CaseInsensitiveMapping  
    method), 439  
lower\_items() (in module colour), 332  
luminance() (in module colour), 162  
luminance\_ASTMD153508() (in module colour.colorimetry), 165  
luminance\_CIE1976() (in module colour.colorimetry), 165  
luminance\_Fairchild2010() (in module colour.colorimetry), 166  
luminance\_Fairchild2011() (in module colour.colorimetry), 167  
LUMINANCE\_METHODS (in module colour), 163  
luminance\_Newhall1943() (in module colour.colorimetry), 164  
luminous\_efficiency() (in module colour), 146  
luminous\_efficiency() (in module colour), 146  
luminous\_flux() (in module colour), 147  
Luv\_to\_LCHuv() (in module colour), 224  
Luv\_to\_uv() (in module colour), 225  
Luv\_to\_XYZ() (in module colour), 223  
Luv\_uv\_to\_xy() (in module colour), 226

## M

mapping (colour.IES\_TM2714\_Spd attribute), 212  
MAX\_RGB\_COLOURSPACE (in module colour.models), 264  
mesopic\_luminous\_efficiency\_function() (in module colour), 147  
message\_box() (in module colour.utilities), 442  
multi\_cctf\_plot() (in module colour.plotting), 390  
multi\_cmfs\_plot() (in module colour.plotting), 374  
multi\_colour\_swatches\_plot() (in module colour.plotting), 370  
multi\_illuminants\_relative\_spd\_plot() (in module colour.plotting), 375  
multi\_lightness\_function\_plot() (in module colour.plotting), 376  
multi\_munsell\_value\_function\_plot() (in module colour.plotting), 391  
multi\_signal\_unpack\_data()  
    (colour.continuous.MultiSignal method), 187  
multi\_spd\_colour\_quality\_scale\_bars\_plot() (in module colour.plotting), 395  
multi\_spd\_colour\_rendering\_index\_bars\_plot() (in module colour.plotting), 394  
multi\_spd\_plot() (in module colour.plotting), 373  
MultiSignal (class in colour.continuous), 186  
MultiSpectralPowerDistribution (class in colour), 98  
munsell\_colour\_to\_xyY() (in module colour), 345  
MUNSELL\_COLOURS (in module colour), 346  
munsell\_value() (in module colour), 346  
munsell\_value\_ASTMD153508() (in module colour.notation), 352  
munsell\_value\_Ladd1955() (in module colour.notation), 351  
munsell\_value\_McCamy1987() (in module colour.notation), 351

MUNSELL\_VALUE\_METHODS (in module colour), 347  
 munsell\_value\_Moon1943() (in module colour.notation), 349  
 munsell\_value\_Munsell1933() (in module colour.notation), 349  
 munsell\_value\_Priest1920() (in module colour.notation), 348  
 munsell\_value\_Saunderson1944() (in module colour.notation), 350

**N**

name (colour.continuous.AbstractContinuousFunction attribute), 181  
 name (colour.RGB\_Colourspace attribute), 255  
 Nayatani95\_Specification (class in colour), 82  
 ndarray\_write() (in module colour.utilities), 438  
 normalise() (colour.MultiSpectralPowerDistribution method), 99  
 normalise() (colour.SpectralPowerDistribution method), 97  
 normalise\_maximum() (in module colour.utilities), 430  
 normalise\_vector() (in module colour.algebra), 59  
 normalised\_primary\_matrix() (in module colour), 251  
 NTSC\_COLOURSPACE (in module colour.models), 264  
 NullInterpolator (class in colour), 50  
 numpy\_print\_options() (in module colour.utilities), 444

**O**

oef() (in module colour), 272  
 oef\_ARIBSTDB67() (in module colour.models), 274  
 oef\_BT2020() (in module colour.models), 276  
 oef\_BT2100\_HLG() (in module colour.models), 277  
 oef\_BT2100\_PQ() (in module colour.models), 278  
 oef\_BT601() (in module colour.models), 279  
 oef\_BT709() (in module colour.models), 280  
 oef\_DCIP3() (in module colour.models), 275  
 oef\_DICOMGSDF() (in module colour.models), 276  
 oef\_ProPhotoRGB() (in module colour.models), 280  
 oef\_reverse() (in module colour), 273  
 oef\_reverse\_ARIBSTDB67() (in module colour.models), 275  
 oef\_reverse\_BT2100\_HLG() (in module colour.models), 277  
 oef\_reverse\_BT2100\_PQ() (in module colour.models), 278  
 oef\_reverse\_BT601() (in module colour.models), 279  
 oef\_reverse\_BT709() (in module colour.models), 280

oef\_reverse\_sRGB() (in module colour.models), 283  
 oef\_RIMMRGB() (in module colour.models), 281  
 oef\_ROMMRGB() (in module colour.models), 281  
 oef\_SMPTE240M() (in module colour.models), 282  
 oef\_sRGB() (in module colour.models), 283  
 oef\_ST2084() (in module colour.models), 282  
 OETFS (in module colour), 273  
 OETFS\_REVERSE (in module colour), 273  
 ones\_spd() (in module colour), 114  
 ootf() (in module colour), 296  
 ootf\_BT2100\_HLG() (in module colour.models), 297  
 ootf\_BT2100\_PQ() (in module colour.models), 298  
 ootf\_reverse() (in module colour), 296  
 ootf\_reverse\_BT2100\_HLG() (in module colour.models), 298  
 ootf\_reverse\_BT2100\_PQ() (in module colour.models), 299  
 OOTFS (in module colour), 296  
 OOTFS\_REVERSE (in module colour), 297  
 orient() (in module colour.utilities), 436

**P**

padding\_args (colour.KernelInterpolator attribute), 48  
 PAL\_SECAM\_COLOURSPACE (in module colour.models), 265  
 path (colour.IES\_TM2714\_Spd attribute), 212  
 PchipInterpolator (class in colour), 51  
 PHOTOPIC\_LEFS (in module colour), 156  
 PLANCK\_CONSTANT (in module colour.constants), 180  
 planck\_law() (in module colour.colorimetry), 116  
 planckian\_locus\_chromaticity\_diagram\_plot\_CIE1931() (in module colour.plotting), 396  
 planckian\_locus\_chromaticity\_diagram\_plot\_CIE1960UCS() (in module colour.plotting), 397  
 POINTER\_GAMUT\_BOUNDARIES (in module colour), 344  
 POINTER\_GAMUT\_DATA (in module colour), 344  
 POINTER\_GAMUT\_ILLUMINANT (in module colour), 344  
 polar\_to\_cartesian() (in module colour.algebra), 58  
 primaries (colour.RGB\_Colourspace attribute), 255  
 primaries\_whitepoint() (in module colour), 252  
 print\_numpy\_errors() (in module colour.utilities), 424  
 Prismatic\_to\_RGB() (in module colour), 338  
 PROPHOTO\_RGB\_COLOURSPACE (in module colour.models), 268  
 PROTUNE\_NATIVE\_COLOURSPACE (in module colour.models), 264

**Q**

quad() (in module colour.plotting), 400

## R

**raise\_numpy\_errors()** (in module colour.utilities), 423  
**random\_triplet\_generator()** (in module colour.algebra), 63  
**range** (colour.continuous.AbstractContinuousFunction attribute), 181  
**range** (colour.continuous.MultiSignal attribute), 187  
**range** (colour.continuous.Signal attribute), 183  
**range()** (colour.SpectralShape method), 102  
**rayleigh\_optical\_depth()** (in module colour.phenomena), 364  
**rayleigh\_scattering()** (in module colour), 354  
**rayleigh\_scattering\_spd()** (in module colour), 355  
**reaction\_rate\_MichealisMenten()** (in module colour.biochemistry), 86  
**read()** (colour.IES\_TM2714\_Spd method), 212  
**read\_image()** (in module colour), 206  
**read\_spds\_from\_csv\_file()** (in module colour), 207  
**read\_spds\_from\_xrite\_file()** (in module colour), 214  
**read\_spectral\_data\_from\_csv\_file()** (in module colour), 209  
**RED\_COLOR\_2\_COLOURSPACE** (in module colour.models), 265  
**RED\_COLOR\_3\_COLOURSPACE** (in module colour.models), 265  
**RED\_COLOR\_4\_COLOURSPACE** (in module colour.models), 266  
**RED\_COLOR\_COLOURSPACE** (in module colour.models), 265  
**RED\_WIDE\_GAMUT\_RGB\_COLOURSPACE** (in module colour.models), 266  
**REFLECTANCE\_RECOVERY\_METHODS** (in module colour), 408  
**reflection\_geometry** (colour.IES\_TM2714\_Spd attribute), 212  
**relative\_tolerance** (colour.NullInterpolator attribute), 50  
**render()** (in module colour.plotting), 369  
**RGB\_10\_degree\_cmfs\_to\_LMS\_10\_degree\_cmfs()** (in module colour.colorimetry), 138  
**RGB\_10\_degree\_cmfs\_to\_XYZ\_10\_degree\_cmfs()** (in module colour.colorimetry), 138  
**RGB\_2\_degree\_cmfs\_to\_XYZ\_2\_degree\_cmfs()** (in module colour.colorimetry), 137  
**RGB\_chromaticity\_coordinates\_chromaticity\_diagram\_plot\_CIE1931()** (in module colour.plotting), 387  
**RGB\_chromaticity\_coordinates\_chromaticity\_diagram\_plot\_CIE1976()** (in module colour.plotting), 388  
**RGB\_chromaticity\_coordinates\_chromaticity\_diagram\_plot\_CIE1960UCS()** (in module colour.plotting), 389  
**RGB\_CMFS** (in module colour), 136  
**RGB\_ColourMatchingFunctions** (class in colour.colorimetry), 132  
**RGB\_Colourspace** (class in colour), 254  
**RGB\_colourspace\_limits()** (in module colour), 448  
**RGB\_colourspace\_pointer\_gamut\_coverage\_MonteCarlo()** (in module colour), 448  
**RGB\_colourspace\_visible\_spectrum\_coverage\_MonteCarlo()** (in module colour), 449  
**RGB\_colourspace\_volume\_coverage\_MonteCarlo()** (in module colour), 451  
**RGB\_colourspace\_volume\_MonteCarlo()** (in module colour), 450  
**RGB\_COLOURSPACES** (in module colour), 256  
**RGB\_colourspaces\_chromaticity\_diagram\_plot\_CIE1931()** (in module colour.plotting), 385  
**RGB\_colourspaces\_chromaticity\_diagram\_plot\_CIE1960UCS()** (in module colour.plotting), 386  
**RGB\_colourspaces\_chromaticity\_diagram\_plot\_CIE1976UCS()** (in module colour.plotting), 387  
**RGB\_colourspaces\_gamuts\_plot()** (in module colour.plotting), 398  
**RGB\_DisplayPrimaries** (class in colour.characterisation), 92  
**RGB\_luminance()** (in module colour), 252  
**RGB\_luminance\_equation()** (in module colour), 253  
**RGB\_scatter\_plot()** (in module colour.plotting), 399  
**RGB\_SpectralSensitivities** (class in colour.characterisation), 90  
**RGB\_to\_CMY()** (in module colour), 341  
**RGB\_to\_HEX()** (in module colour.notation), 353  
**RGB\_to\_HSL()** (in module colour), 340  
**RGB\_to\_HSV()** (in module colour), 339  
**RGB\_to\_ICTCP()** (in module colour), 336  
**RGB\_to\_Prismatic()** (in module colour), 338  
**RGB\_to\_RGB()** (in module colour), 247  
**RGB\_to\_RGB\_matrix()** (in module colour), 248  
**RGB\_to\_spectral\_Smits1999()** (in module colour.recovery), 408  
**RGB\_to\_XYZ()** (in module colour), 246  
**RGB\_to\_XYZ\_matrix** (colour.RGB\_Colourspace attribute), 255  
**RGB\_to\_YCbCr()** (in module colour), 328  
**RGB\_to\_YcCbcCrc()** (in module colour), 332  
**RIMM\_RGB\_COLOURSPACE** (in module colour.models), 267  
**RLAB\_D\_FACTOR** (in module colour), 85  
**RLAB\_Specification** (class in colour), 85  
**RLAB\_VIEWING\_CONDITIONS** (in module colour), 86  
**RIMM\_RGB\_COLOURSPACE** (in module colour.models), 267  
**RLAB\_SPECIFICATION** (in module colour), 85  
**RLAB\_VIEWING\_CONDITIONS** (in module colour), 86  
**ROMM\_RGB\_COLOURSPACE** (in module colour.models), 267  
**ROW\_AS\_DIAGNOD** (in module colour.utilities), 434  
**RUSSELL\_RGB\_COLOURSPACE** (in module colour.models), 268

**S**

S\_GAMUT3\_CINE\_COLOURSPACE (in module colour.models), 269  
 S\_GAMUT3\_COLOURSPACE (in module colour.models), 269  
 S\_GAMUT\_COLOURSPACE (in module colour.models), 268  
 scattering\_cross\_section() (in module colour), 363  
 SCOTOPIC\_LEFS (in module colour), 156  
 shape (colour.MultiSpectralPowerDistribution attribute), 99  
 shape (colour.SpectralPowerDistribution attribute), 96  
 SHARP\_CAT (in module colour.adaptation), 44  
 Signal (class in colour.continuous), 183  
 signal\_unpack\_data() (colour.continuous.Signal method), 184  
 signals (colour.continuous.MultiSignal attribute), 187  
 single\_ccf\_plot() (in module colour.plotting), 390  
 single\_cmfs\_plot() (in module colour.plotting), 374  
 single\_colour\_swatch\_plot() (in module colour.plotting), 369  
 single\_illuminant\_relative\_spd\_plot() (in module colour.plotting), 374  
 single\_lightness\_function\_plot() (in module colour.plotting), 376  
 single\_munsell\_value\_function\_plot() (in module colour.plotting), 391  
 single\_rayleigh\_scattering\_spd\_plot() (in module colour.plotting), 392  
 single\_spd\_colour\_quality\_scale\_bars\_plot() (in module colour.plotting), 394  
 single\_spd\_colour\_rendering\_index\_bars\_plot() (in module colour.plotting), 393  
 single\_spd\_plot() (in module colour.plotting), 372  
 SMITS\_1999\_SPDS (in module colour.recovery), 409  
 SMPTE\_240M\_COLOURSPACE (in module colour.models), 268  
 spds\_chromaticity\_diagram\_plot\_CIE1931() (in module colour.plotting), 382  
 spds\_chromaticity\_diagram\_plot\_CIE1960UCS() (in module colour.plotting), 383  
 spds\_chromaticity\_diagram\_plot\_CIE1976UCS() (in module colour.plotting), 384  
 spectral\_quantity (colour.IES\_TM2714\_Spd attribute), 212  
 spectral\_to\_aces\_relative\_exposure\_values() (in module colour.models), 270  
 spectral\_to\_XYZ() (in module colour), 118  
 spectral\_to\_XYZ\_ASTME30815() (in module colour.colorimetry), 121  
 spectral\_to\_XYZ\_integration() (in module colour.colorimetry), 128

SPECTRAL\_TO\_XYZ\_METHODS (in module colour), 119  
 spectral\_to\_XYZ\_tristimulus\_weighting\_factors\_ASTME30815() (in module colour.colorimetry), 123  
 SpectralPowerDistribution (class in colour), 96  
 SpectralShape (class in colour), 101  
 spherical\_to\_cartesian() (in module colour.algebra), 57  
 SpragueInterpolator (class in colour), 51  
 sRGB\_COLOURSPACE (in module colour.models), 269  
 sRGB\_to\_XYZ() (in module colour), 250  
 STANDARD\_OBSERVERS\_CMFS (in module colour), 136  
 start (colour.SpectralShape attribute), 101  
 strict\_labels (colour.MultiSpectralPowerDistribution attribute), 99  
 strict\_name (colour.MultiSpectralPowerDistribution attribute), 99  
 strict\_name (colour.SpectralPowerDistribution attribute), 96  
 Structure (class in colour.utilities), 441  
 substrate\_concentration\_MichealisMenten() (in module colour.biochemistry), 87  
 suppress\_warnings() (in module colour.utilities), 444

**T**

the\_blue\_sky\_plot() (in module colour.plotting), 392  
 to\_dataframe() (colour.continuous.MultiSignal method), 187  
 to\_series() (colour.continuous.Signal method), 184  
 transmission\_geometry (colour.IES\_TM2714\_Spd attribute), 212  
 trim() (colour.MultiSpectralPowerDistribution method), 99  
 trim() (colour.SpectralPowerDistribution method), 96  
 tristimulus\_weighting\_factors\_ASTME202211() (in module colour.colorimetry), 126  
 tsplit() (in module colour.utilities), 433  
 tstack() (in module colour.utilities), 432

**U**

UCS\_to\_uv() (in module colour), 228  
 UCS\_to\_XYZ() (in module colour), 227  
 UCS\_uv\_to\_xy() (in module colour), 228  
 update() (colour.utilities.Structure method), 441  
 use\_derived\_RGB\_to\_XYZ\_matrix (colour.RGB\_Colourspace attribute), 255  
 use\_derived\_transformation\_matrices() (colour.RGB\_Colourspace method), 255  
 use\_derived\_XYZ\_to\_RGB\_matrix (colour.RGB\_Colourspace attribute), 255

uv\_to\_CCT() (in module colour), 414  
UV\_TO\_CCT\_METHODS (in module colour), 414  
uv\_to\_CCT\_Ohno2013() (in module colour.temperature), 419  
uv\_to\_CCT\_Robertson1968() (in module colour.temperature), 417

**V**

V\_GAMUT\_COLOURSPACE (in module colour.models), 269  
values (colour.MultiSpectralPowerDistribution attribute), 99  
values (colour.SpectralPowerDistribution attribute), 96  
visible\_spectrum\_plot() (in module colour.plotting), 375  
VON\_KRIES\_CAT (in module colour.adaptation), 44

**W**

warn\_numpy\_errors() (in module colour.utilities), 424  
warning() (in module colour.utilities), 443  
wavelength\_to\_XYZ() (in module colour), 120  
wavelengths (colour.MultiSpectralPowerDistribution attribute), 99  
wavelengths (colour.SpectralPowerDistribution attribute), 96  
whiteness() (in module colour), 168  
whiteness\_ASTME3130 (in module colour.colorimetry), 172  
whiteness\_Berger1959() (in module colour.colorimetry), 169  
whiteness\_CIE2004() (in module colour.colorimetry), 174  
whiteness\_Ganz1979() (in module colour.colorimetry), 173  
WHITENESS\_METHODS (in module colour), 169  
whiteness\_Stensby1968() (in module colour.colorimetry), 171  
whiteness\_Taube1960() (in module colour.colorimetry), 170  
whitepoint (colour.RGB\_Colourspace attribute), 255  
window (colour.KernelInterpolator attribute), 48  
write() (colour.IES\_TM2714\_Spd method), 212  
write\_image() (in module colour), 206  
write\_spds\_to\_csv\_file() (in module colour), 210

**X**

x (colour.KernelInterpolator attribute), 48  
x (colour.LinearInterpolator attribute), 49  
x (colour.NullInterpolator attribute), 50  
x (colour.SpragueInterpolator attribute), 52  
XTREME\_RGB\_COLOURSPACE (in module colour.models), 270

xy\_to\_CCT() (in module colour), 416  
xy\_to\_CCT\_Hernandez1999() (in module colour.temperature), 420  
XY\_TO\_CCT\_METHODS (in module colour), 416  
xy\_to\_xyY() (in module colour), 219  
xy\_to\_XYZ() (in module colour), 217  
xyY\_to\_munsell\_colour() (in module colour), 345  
xyY\_to\_xy() (in module colour), 218  
xyY\_to\_XYZ() (in module colour), 216  
XYZ\_ColourMatchingFunctions (class in colour.colorimetry), 134  
XYZ\_SCALING\_CAT (in module colour.adaptation), 44  
XYZ\_to\_ATD95() (in module colour), 65  
XYZ\_to\_CAM16() (in module colour), 71  
XYZ\_to\_CIECAM02() (in module colour), 67  
XYZ\_to\_hdr\_CIELab() (in module colour), 243  
XYZ\_to\_hdr\_IPT() (in module colour), 241  
XYZ\_to\_Hunt() (in module colour), 75  
XYZ\_to\_Hunter\_Lab() (in module colour), 230  
XYZ\_to\_Hunter\_Rdab() (in module colour), 232  
XYZ\_to\_IPT() (in module colour), 239  
XYZ\_to\_K\_ab\_HunterLab1966() (in module colour), 231  
XYZ\_to\_Lab() (in module colour), 220  
XYZ\_to\_LLAB() (in module colour), 78  
XYZ\_to\_Luv() (in module colour), 222  
XYZ\_to\_Nayatani95() (in module colour), 81  
XYZ\_to\_RGB() (in module colour), 245  
XYZ\_to\_RGB\_matrix (colour.RGB\_Colourspace attribute), 255  
XYZ\_to\_RLAB() (in module colour), 84  
XYZ\_to\_spectral() (in module colour), 406  
XYZ\_to\_spectral\_Meng2015() (in module colour.recovery), 410  
XYZ\_to\_sRGB() (in module colour), 249  
XYZ\_to\_UCS() (in module colour), 226  
XYZ\_to\_UVW() (in module colour), 229  
XYZ\_to\_xy() (in module colour), 217  
XYZ\_to\_xyY() (in module colour), 215

**Y**

y (colour.KernelInterpolator attribute), 48  
y (colour.LinearInterpolator attribute), 49  
y (colour.NullInterpolator attribute), 50  
y (colour.PchipInterpolator attribute), 51  
y (colour.SpragueInterpolator attribute), 52  
YCbCr\_to\_RGB() (in module colour), 330  
YCBCR\_WEIGHTS (in module colour), 331  
YcCbcCrc\_to\_RGB() (in module colour), 333  
yellowness() (in module colour), 175  
yellowness\_ASTMD1925() (in module colour.colorimetry), 176

yellowness\_ASTME313() (in module colour.colorimetry), [177](#)  
YELLOWNESS\_METHODS (in module colour), [175](#)

## Z

zeros\_spd() (in module colour), [115](#)